

1 **Aerosol Lidar Observations of Atmospheric Mixing in**
2 **Los Angeles: Climatology and Implications for**
3 **Greenhouse Gas Observations**

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D R A F T

Key Points.

- Aerosol lidar maps LA mixing depth in space (pilot mobile study) and time (2 years data)
- Automatic mixing depth retrieval system finds daily variability far exceeds seasonal difference
- PBL heights in models used for GHG monitoring show biases that will carry over to flux estimates

4 **Abstract.** Atmospheric observations of greenhouse gases provide essen-
5 tial information on sources and sinks of these key atmospheric constituents.
6 To quantify fluxes from atmospheric observations, representation of trans-
7 port – especially vertical mixing – is a necessity and often a source of error.
8 We report on remotely sensed profiles of vertical aerosol distribution taken
9 over a two-year period in Pasadena, California. Using an automated anal-
10 ysis system, we estimate daytime mixing layer depth, achieving high confi-
11 dence in the afternoon maximum on 51% of days with profiles from a Sigma
12 Space Mini Micropulse LiDAR (MiniMPL) and on 36% of days with a Vaisala
13 CL51 ceilometer. We note that considering ceilometer data on a logarithmic
14 scale, a standard method, introduces an offset in mixing height retrievals.
15 The mean afternoon maximum mixing height is 770 m AGL in summer and
16 670 m in winter, with significant day-to-day variance (within-season $\sigma =$
17 220 m \approx 30%). Taking advantage of the MiniMPL's portability, we demon-
18 strate the feasibility of measuring the detailed horizontal structure of the mix-
19 ing layer by automobile. We compare our observations to PBL heights from
20 sonde launches, NARR reanalysis, and a custom WRF model developed for

21 GHG monitoring in Los Angeles. NARR and WRF PBL heights at Pasadena
22 are both systematically higher than measured, NARR by 2.5 times; these
23 biases will cause proportional errors in GHG flux estimates using modeled
24 transport. We discuss how sustained lidar observations can be used to re-
25 duce flux inversion error by selecting suitable analysis periods, calibrating
26 models, or characterizing bias for correction in post-processing.

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1. Introduction

27 Improved understanding of sources, sinks, and controlling processes of CO₂ and other
28 greenhouse gases (GHGs) will require robust methods for estimating surface fluxes. Ob-
29 servations of GHG concentrations capture the influence of known and unknown sources
30 and sinks alike, making these observations an important complement to models and in-
31 ventories. Top-down GHG inversions have been used for some time to estimate fluxes
32 on global [*Tans et al.*, 1990], continental [*Bousquet et al.*, 2000], and regional [*Lauvaux*
33 *et al.*, 2013; *Peters et al.*, 2007; *Schuh et al.*, 2010] scales, and there is increasing focus
34 on bringing a similar approach to individual cities [*McKain et al.*, 2012; *Lauvaux et al.*,
35 2013; *Breon et al.*, 2014; *Turnbull et al.*, 2015]. However, relating observed concentrations
36 to surface fluxes requires a representation of atmospheric transport. On the regional and
37 urban scales, the extent and variability of vertical mixing is a dominant source of uncer-
38 tainty [*McKain et al.*, 2012] that can easily overwhelm the effects of instrument error. It
39 is therefore critical to represent vertical mixing accurately.

40 The spatiotemporal structure of vertical mixing and diffusion can be complex. However,
41 it can be useful to approximate gases recently emitted from the surface as being confined
42 to and uniformly distributed throughout a near-surface layer. A cluster of related concepts
43 – atmospheric or planetary boundary layer (PBL), convective boundary layer, mixed layer
44 – are commonly used to describe the part of the atmosphere which “responds to surface
45 forcings with a timescale of about an hour or less.” [*Stull*, 1988] Various specific definitions
46 of these layers are in use [*Seibert et al.*, 2000], some referring to thermodynamic variables

47 and others directly to mixing or turbulence. Layers identified by different definitions can
48 be conceptually distinct and therefore need to be considered differently.

49 The layer relevant to the dilution of GHGs is that within which substantial vertical
50 mixing takes place. The time scale of mixing under turbulent conditions has been es-
51 timated to be tens of minutes [*Stull*, 1988; *van Stratum et al.*, 2012; *Janssen and Pozzer*,
52 2015]. Nonetheless, a fully well-mixed equilibrium may not exist; we therefore follow *Seib-*
53 *ert et al.* [2000] in referring to the *mixing layer*. When we refer to the *mixing height* or
54 *mixing depth*, we mean the altitude of the top of the mixing layer. In addition to GHGs,
55 the mixing depth also controls the dilution of aerosols and of other trace gases produced
56 primarily within the mixing layer, including those that contribute to poor air quality. It
57 is well known that shallow mixing contributes to air quality exceedances as these species
58 are trapped near the surface, and observations such as those presented here can help in
59 defining the presence of these conditions.

60 While it is difficult to measure the vertical distribution of GHGs directly, especially
61 on an ongoing basis, we can measure the mixing height by observing the distribution
62 of aerosol. Lidar systems measure the backscatter of a laser from particulate matter in
63 the atmosphere, providing a vertical (or skew) profile of the concentration of scattering
64 particles. We make use of such an instrument, the Sigma Space Mini-Micropulse LiDAR
65 (MiniMPL), as well as a Vaisala CL51 ceilometer. These and other remote sensing in-
66 struments benefit from continuous operation, making observations at a rate of once per
67 minute or more. New models like the MiniMPL are smaller and more portable than earlier
68 research lidars and have better signal-to-noise performance than ceilometers.

69 Note that the mixing layer may not always coincide with the boundary layer commonly
70 diagnosed by applying thermodynamic criteria to data from radiosondes (e.g. using the
71 parcel method). In Pasadena, comparison to results from a series of sonde launches
72 suggests that the mixing depth is related to, though not identical with, the depth of
73 the boundary layer as defined using thermodynamic criteria. This finding is consistent
74 with past results. Working in Indiana and the Amazon basin, respectively, *Coulter* [1979]
75 and *Martin et al.* [1988] found that mixing depths determined using lidar observations
76 were similar to and well-correlated with, though generally somewhat higher than, those
77 determined from temperature profiles. *Marsik et al.* [1995] found that mixing depths from
78 lidar in Atlanta were slightly *lower* than those measured using sondes. We discuss the
79 comparison to sonde data in greater detail in section 3.3.

80 Given the high frequency of observations, operational use of lidar to measure the mixing
81 height benefits from an at least partially automated method of analysis. A variety of
82 schemes have been used. The simplest, the gradient method [*Endlich et al.*, 1979], searches
83 for the maximum (most negative) vertical gradient of the backscatter signal, indicating a
84 sudden decrease in density of scatterers. Related is the inflection point method [*Menut*
85 *et al.*, 1999], which searches for zeros of the second spatial derivative of the backscatter.
86 The wavelet method [*Ehret et al.*, 1996; *Davis et al.*, 1997, 2000; *Baars et al.*, 2008], which
87 we use, is a refinement of the gradient method that takes into account the typical spatial
88 scale of the boundary region at the top of the mixing layer. The variance method [*Hooper*
89 *and Eloranta*, 1986; *Menut et al.*, 1999] identifies the entrainment zone at the top of the
90 mixing layer by detecting a maximum in the temporal variance of backscatter, indicating
91 the presence of turbulent vertical mixing. The idealized-profile method [*Steyn et al.*, 1999;

92 *Eresmaa et al.*, 2006; *Münkel et al.*, 2006] attempts to fit the vertical backscatter profile
93 to an ideal representation of aerosol density in and above the mixing layer – typically
94 an error function. Some studies have applied a combination of methods: for example,
95 using the gradient to refine a spatially [*Lammert and Bösenberg*, 2006] or temporally
96 [*Hennemuth and Lammert*, 2006] coarse estimate generated by the variance method, or
97 using gradient methods to select a number of candidate heights, then selecting between
98 them by minimizing disagreement with a physical model [*Di Giuseppe et al.*, 2012].

99 In any method, the most serious challenge in automated mixing layer detection is to
100 distinguish between the mixing layer top and other similar boundaries in the atmosphere,
101 such as fog, low clouds, or residual layers of scatterers remaining aloft from previous days
102 [*Haeffelin et al.*, 2012; *Lewis et al.*, 2013]. One approach to this challenge is to use the
103 automated system only to generate a set of candidate heights and then rely on a human
104 expert to distinguish between them. A person with some knowledge of atmospheric physics
105 can often, though not always, identify the top of the mixing layer by visual inspection of a
106 whole day's backscatter data. We take a different approach, aiming to automate the entire
107 process in order to allow for long-term continuous operation. Following recent work [*Gan*
108 *et al.*, 2011; *Lewis et al.*, 2013], we apply criteria that constrain the detected boundary to
109 behavior that is physically reasonable, and we automatically detect and exclude conditions
110 in which the instrument beam is blocked by fog or clouds. Finally, modifying a method
111 introduced by *Lewis et al.* [2013], we implement a voting scheme, processing the day's
112 data in several different ways and interpreting the degree of concurrence as a measure of
113 confidence that our algorithm has selected the correct boundary.

114 In section 2, we describe the backscatter data, the instrument used to obtain them, and
115 our automated method for extracting the mixing height. Section 3 presents our findings
116 as to the climatological mixing state in the LA area and its temporal and spatial variation.
117 We compare the results obtained with the MiniMPL to mixing depth estimates from a
118 ceilometer to a one-day sonde intensive, and to PBL heights from models and reanalysis.
119 Finally, in section 4, we discuss the implications of our work for GHG flux estimation and
120 suggest possible future applications.

2. Method

2.1. Instrumentation

121 We collected aerosol backscatter data using a Sigma Space Mini-Micropulse LiDAR
122 (MiniMPL) operating at the Caltech campus in Pasadena, California. The MiniMPL is a
123 compact version of the standard MPL, also manufactured by Sigma Space, that populates
124 the NASA MPLNET lidar network. The MiniMPL inherits many of the design features
125 of the MPL, such as a fiber coupled detector and robust optical train. Compared to the
126 MPL, the MiniMPL reduces the power-aperture product to minimize cost, size, weight,
127 and power requirements. As a result, detection range is limited to the troposphere while
128 the MPL measures into the stratosphere. For tropospheric applications such as GHG flux
129 estimation and air quality monitoring, however, the MiniMPL is designed to match the
130 data quality of a standard MPL.

131 The MiniMPL transceiver shown in Fig. 1 weighs 13 kg and measures 380 x 305 x 480
132 mm in width, depth and height. The system consists of a laptop and the lidar transceiver,
133 which are connected by a USB cable and consume 100W during normal operation. The
134 whole system fits in a storm case with a telescopic handle and wheels that can be checked

135 in as regular luggage during a domestic or international flight. The system's portability
136 allows for applications that would not be possible with the standard MPL. In section 3.6,
137 we demonstrate the feasibility of operating the MiniMPL out of a moving car, enabling
138 us to observe the spatial structure of the mixing layer without the use of aircraft.

139 The MiniMPL's Nd:YAG laser emits polarized 532 nm light at a 4 KHz repetition
140 rate and 3.5 uJ nominal pulse energy. The laser beam is expanded to the size of the
141 telescope aperture (80 mm) to satisfy the eye safe requirements in ANSI Z136.1.2000 and
142 IEC 60825 standards. Laser light is scattered back toward the instrument by particles and
143 molecules in the atmosphere and collected by an 80mm diameter receiver. Distance to the
144 scattering event is calculated from the time of flight. The instrument reports the number
145 of scattering events recorded during a user-defined accumulation time (in our case, 30 s)
146 originating in each vertical bin. We use a vertical range resolution of 30 m. Although
147 this study does not make use of it, the MiniMPL also measures the depolarization [*Flynn*
148 *et al.*, 2007] of the scattered light with a contrast ratio greater than 100:1.

149 The receiver uses a pair of narrowband filters with bandwidth less than 180 pm to reject
150 the majority of solar background noise. The filtered light is then collected by a 100 um
151 multimode fiber and fed into a Silicon Avalanche Photodetector (Si APD) operating in
152 photon-counting mode (Geiger mode). Photon-counting detection enables the MiniMPL
153 design to be lightweight and compact with high signal-to-noise ratio (SNR) throughout
154 the troposphere.

155 To further maximize the SNR, MiniMPL uses a coaxial design; the transmitter and
156 receiver Field of View (FOV) overlap with each other from range zero. This design
157 eliminates the need for a wide FOV in order to minimize the overlap distance as in some

158 biaxial lidar systems [Kuze *et al.*, 1998]. A wide FOV can result in measuring multiple
159 scattering from aerosol [Spinhirne, 1982] and can distort depolarization measurements
160 [Tatarov *et al.*, 2000]. On the other hand, a very narrow receiver FOV could make the
161 lidar system sensitive to external factors like shock, vibration and temperature, making
162 the system unsuitable for field deployment. The design of MiniMPL balances the above
163 requirements and constraints, with an FOV of 240 urad.

164 Additional technical specifications of the MiniMPL, along with those of the standard
165 MPL for comparison, are given in table 1.

2.2. Calibration

166 The raw event count reported by the MiniMPL must be calibrated and normalized in
167 order to arrive at the quantity of interest, Normalized Relative Backscatter (NRB), which
168 is approximately proportional to the concentration of scatterers at a given distance above
169 the instrument. First, the event count is corrected for the deadtime of the detector, a pe-
170 riod after each photon incidence during which no additional photons can be detected. The
171 likely number of missed incidences can be extrapolated from the rate of detected photons.
172 After the deadtime correction, the background (no laser light) value is subtracted. The
173 event rate is then scaled by the laser pulse energy, which prevents changes in pulse energy
174 from appearing as variation of the measured backscatter. Next, a correction is applied
175 to account for laser light, called afterpulse, that strikes the inside of the instrument and
176 returns to the detector without interacting with the atmosphere.

177 Finally, two corrections account for the fraction of scattered photons that are intercepted
178 by the detector. The solid angle subtended by the collecting lens is inversely proportional
179 to the square of the distance to the scattering event, so the event rate is multiplied by r^2 .

180 Since the MiniMPL laser beam overlaps with the receiver field of view from range zero,
 181 there is no need for an overlap correction in the sense required by a biaxial instrument.
 182 However, because not all of the light incident on the collecting lens is focused onto the
 183 photon counter, a geometric factor calibration is still required. For historical reasons, this
 184 factor is also referred to as an overlap correction.

These steps are summarized in the following calibration equation:

$$B_{nr} = \left[\frac{x(z)C(x(z)) - bC(b)}{E} - \frac{x_{ap}(z)C(x_{ap}(z)) - b_{ap}C(b_{ap})}{E_{ap}} \right] \frac{z^2}{O(z)} \quad (1)$$

185 where $x(z)$ is the raw event rate signal at distance z from the instrument, $C(x)$ is the
 186 deadtime correction factor for event rate x , b is the background, E is the laser pulse energy,
 187 $x_{ap}(z)$ is the afterpulse signal at the time corresponding to distance z , b_{ap} is the background
 188 of the afterpulse signal, E_{ap} is the energy of the afterpulse, $O(z)$ is the overlap correction
 189 factor at distance z , and B_{nr} is the Normalized Relative Backscatter (NRB). An example
 190 vertical profile of NRB can be seen in Figure 3. In order to reduce the impact of short
 191 time scale fluctuations on our mixing depth retrieval, we apply a additional two-minute
 192 sliding average to the NRB values already aggregated to a thirty-second accumulation
 193 time by the instrument.

2.3. Observations

194 The MiniMPL collected backscatter data at Caltech on 530 days between August 1,
 195 2012 and October 23, 2014, operating between dawn and dusk. Of those, 54 included
 196 data gaps of longer than one hour, including late starts to data collection, persistent
 197 midday rain or fog, or obstruction of the beam by obstacles. We exclude those days from
 198 the analysis. The remaining 476 days are distributed across all months other than July.

199 Table 2 shows the number of days of data by month as well as the concurrence scores of
 200 the mixing depth estimates (see section 2.4).

201 We analyze backscatter data in daily increments. Over the course of the day, the
 202 changing backscatter profile gives a picture of the distribution of scatterers in the lower
 203 atmosphere (Figure 2a). In the Los Angeles area, the scattering signal is typically quite
 204 strong due to the high levels of anthropogenic aerosols. Since they are produced primarily
 205 within the mixing layer, aerosols are concentrated near the surface. During the day, solar
 206 heating of the surface drives vertical mixing, causing the mixing layer to deepen and
 207 carrying aerosols to higher altitudes. As surface heating decreases in the late afternoon,
 208 the region of active vertical mixing shrinks, but the aerosols may remain aloft for some
 209 time. Frequently, aerosols carried aloft by one day’s mixing can still be observed the
 210 following day in a residual layer disconnected from the surface. In the coastal mountain
 211 environment of Los Angeles, aerosols can also be carried above the mixing layer by the
 212 dominant circulation pattern, resulting in a sometimes complex stratification structure
 213 with thin, lofted aerosol layers [Lu and Turco, 1994, 1995].

2.4. Analysis

214 We use a Haar wavelet covariance method to identify boundaries between layers with
 215 high and low aerosol density. At a given height z , the wavelet covariance w is given by
 216 integrating the product of the backscatter profile with a Haar wavelet H centered at z :

$$217 \quad w(z) = \frac{1}{d} \int dz' B_{nr}(z') H(z, z', d) \quad (2)$$

$$218 \quad H(z, z', d) = \begin{cases} 1 & z - d/2 < z' < z \\ -1 & z < z' < z + d/2 \\ 0 & |z' - z| > d/2 \end{cases} \quad (3)$$

219 The width d , or dilation, of the wavelet is chosen to correspond to the typical size of
220 the transition zone at the top of the mixing layer, 200 m. As illustrated in Figure 3,
221 the covariance is highest where the backscatter decreases rapidly with height. Because
222 aerosols are concentrated within the mixing layer, such a rapid decrease in backscatter
223 occurs at the top of the mixing layer. We therefore use high wavelet covariance values to
224 identify the mixing layer top.

225 In order to increase the likelihood of detecting the mixing layer top rather than some
226 other boundary – for example, a structure within the mixing layer or a residual layer of
227 aerosols further aloft – each day’s data is considered as a whole. Call the set of times
228 during a single day at which backscatter data is available t_1, \dots, t_{max} . First, designate a
229 single time t_k and compute the altitude $z_k(t_k)$ at which the Haar covariance is maximum.
230 Any later timepoints are then considered in order, beginning with t_{k+1} . The altitude z_k of
231 the detected boundary is constrained to vary at a rate no faster than v . For the MiniMPL,
232 we set $v = 100$ m/min, a conservative upper bound on typical rates of change of the mixing
233 layer height [Stull, 1988]. This is equivalent to setting the Haar covariance to zero outside
234 the range $(z_k(t_{i-1}) - v(t_i - t_{i-1}), z_k(t_{i-1}) + v(t_i - t_{i-1}))$. In addition, a multiplicative bias
235 factor is applied to suppress the Haar covariance for unlikely but possible rates of change,
236 decreasing linearly from one at $(2/3)v$ to zero at v . Similarly, any timepoints earlier than
237 t_k are considered in reverse order, beginning with t_{k-1} .

238 For $t_k = t_1$ an additional physical constraint is applied: the mixing layer top must
239 begin each day within 500 m of the ground. This aids in selecting a boundary that is
240 continuous with the top of the nocturnal boundary layer, as the mixing layer should be.

241 This process is repeated for five values of t_k distributed evenly throughout the day,
242 including the earliest time t_1 and the latest time t_{max} . The result is a set of estimates
243 $z_{k_1}(t), \dots, z_{k_n}(t)$ of the mixing height as a function of time.

244 A voting procedure is then used to select one estimate from the set. First, estimates
245 are checked for pairwise agreement according to one of several criteria. In this study,
246 we consider two estimates to be in agreement if they differ by no more than one unit
247 of instrument vertical resolution (30 m for the MiniMPL as we operate it) as to the
248 maximum depth of the mixing layer during the midday period. This criterion is optimized
249 for determining that maximum; other criteria, such as agreement to within a tolerance
250 over a specified fraction of the data period, might be better suited for other purposes.
251 Next, this pairwise agreement is used to calculate a concurrence score for each estimate.
252 An estimate E has a concurrence score equal to the fraction of all estimates that agree
253 with E according to the selected criterion – see Figure 2 for an example. A 3/10 penalty is
254 applied to the concurrence score of any estimate that violates the start-of-day condition,
255 i.e. that puts the mixing layer top above 500 m at the start of the day. This was already
256 forbidden during processing for the estimate beginning at t_1 , but it may occur in other
257 cases, and it generally indicates that the estimate has been fooled by a residual layer. After
258 applying the penalty, the estimate with the highest score is selected for reporting, and the
259 concurrence score can be used as a measure of confidence. Concurrence ties are broken by
260 selecting the estimate with the earliest start time t_k ; note that for concurrence scores of
261 better than one-half, the tied estimates necessarily agree as to the chosen criterion. We
262 recommend excluding estimates with scores less than one-half.

263 Because fog or clouds can completely extinguish the instrument beam, preventing any
264 information from being returned from higher altitudes, it is important that our processing
265 algorithm be able to detect this circumstance. Under foggy conditions, the altitude of
266 highest Haar wavelet covariance does not represent the top of the mixing layer – in fact,
267 there likely is no mixing layer – but only the maximum altitude to which the beam was
268 able to pierce the fog before being extinguished. This situation is common in Pasadena
269 in the early morning. We detect fog by checking directly for beam extinction, i.e. a layer
270 of very high backscatter values with close to zero signal from above, and do not report
271 any mixing height while fog is present.

272 Although the altitude of maximum Haar covariance on a foggy morning does not repre-
273 sent the mixing layer top, it remains useful, since that altitude transitions smoothly into
274 the mixing layer top as the fog burns off. Fog, clouds, or rain that occur in the middle of
275 the day are more problematic, since they often produce discontinuous changes in signal.
276 We treat such occurrences as data gaps, and we exclude days on which gaps, including fog
277 or rain, persist for too much of the total data period. In any case, we report for each day
278 the maximum length of any gap in data, including instrument malfunction, a late start
279 to data gathering, or beam extinction. It is important to check the maximum gap length
280 before making use of the data, and to establish a standard for maximum allowable gap
281 length, since long gaps can produce nonsensical results.

3. Results

3.1. Climatology and Variation

282 On the basis of our estimates, we emphasize the very large daily variability in the mixing
283 height in the LA basin. The maximum depth of the mixing layer in afternoon may differ

284 by a factor of two from one day to the next. On average, the greater degree of insolation
285 does produce deeper afternoon mixing layers in summer than in winter. Using backscatter
286 data from the MiniMPL on days with concurrence scores of 4/5 or higher and without
287 gaps longer than 1 hr, we find the mean afternoon maximum mixing depth to be 770 m
288 AGL in summer (June and August) and 670 m AGL in winter (December-February).

289 However, as illustrated in Figure 4, this seasonal difference is overwhelmed by the
290 very large day- to week-scale variability. Within-season standard deviations in afternoon
291 maximum mixing height are about 220 m in both summer and winter, representing 29%
292 and 32% of the means, respectively. Similarly, a given day's mixing height cannot reliably
293 be extrapolated from measurements made on previous days. Across 105 cases across all
294 seasons in which we achieve concurrence scores of 4/5 or higher on both of two consecutive
295 days, the root-mean-square difference in afternoon maximum mixing depth at Caltech is
296 230 m.

297 The high variability reinforces that applications of climatological mixing depth values
298 are subject to large uncertainties; sustained observations like those we present here can
299 quantify those uncertainties. Such observations can also be used to calibrate models or
300 to choose between parameterization schemes in meteorological models, as we discuss in
301 sections 3.4 and 3.5. Comprehensive comparisons to a model and/or to other meteorological
302 observations over a long period could also provide a more granular understanding
303 of the mixing dynamics. A robust explanation is needed for the variation we observe,
304 which takes place too consistently and on too short a time scale to be attributed solely
305 to unusual events such as forest fires or the LA basin's periodic Santa Ana winds.

3.2. Ceilometer

306 Alongside the MiniMPL, we also operated a Vaisala CL51 ceilometer (a successor in-
307 strument to the CL31, for details see e.g. *Münkel and Rasanen* [2004]; *McKendry et al.*
308 [2009]; *Münkel et al.* [2011]) at the same site. The measurement principle of the ceilometer
309 is similar to that of the lidar, but the overlap correction and other calibration steps are
310 performed by proprietary software not visible to or modifiable by the user [*Wiegner et al.*,
311 2014]. The resulting quantity is referred to simply as the backscatter profile. The CL51
312 operates at 910 nm, in the near-infrared; it uses a 16 s temporal bin and a 10 m vertical
313 range resolution.

314 We apply a version of the same algorithm to estimate mixing depths based on ceilometer
315 backscatter data as we use with the MiniMPL. An example is shown in Figure 2b for
316 comparison to the MiniMPL results on the same day. As is visible in the figure, especially
317 in regions of low backscatter signal, the ceilometer’s signal-to-noise performance is not as
318 good as that of the MiniMPL. As a result, some adjustments are necessary. First, the
319 maximum allowed rate of change v in the mixing layer height must be relaxed; for the
320 ceilometer we set it to 150 m/min. This change is necessary because noise can temporarily
321 disguise a change in the boundary location; the algorithm must be able to “snap back”
322 to the true location of the boundary even after it has moved some distance away.

323 Second, the ceilometer tends to show an unrealistically large signal in the near field.
324 *He et al.* [2006] note a similar artifact, which they attribute to an imperfectly corrected
325 overlap error [see also *Wiegner et al.*, 2014]. Such errors are caused by differences in the
326 optical geometry of the outgoing beam aperture and the detector that collects scattered
327 photons. Because the erroneous backscatter signal associated with the artifact decays
328 very rapidly, it has a high wavelet covariance. The algorithm therefore tends to detect the

329 artifact in place of the real boundary, estimating the mixing depth at the lowest possible
330 altitude.

331 To solve this problem, and to dampen noise in general, it is standard to take the
332 logarithm of the ceilometer backscatter data prior to processing. Although physically
333 unmotivated, this preprocessing step flattens out large signals, decreasing the influence of
334 the low-altitude artifact. Figure 5 shows an example of a case in which the log transform
335 allows the algorithm to detect the correct boundary. However, the log transform also
336 introduces an offset. It suppresses the magnitude of the gradient of the backscatter more
337 where backscatter values are higher:

$$338 \quad \frac{d}{dz} \log(b(z)) = \frac{1}{b(z)} \frac{db(z)}{dz} \quad (4)$$

339 Backscatter decreases with height in the transition from the mixing layer to the free tropo-
340 sphere above, so the strongest gradient in $\log(b(z))$ generally occurs at a higher altitude
341 than the strongest gradient in $b(z)$. This effect carries over to the wavelet covariance
342 method, causing a positive offset of about 50 m. The offset is due to a methodological
343 choice to identify the altitude of greatest relative change in scattering, not a difference in
344 physical reality. It should therefore be noted and compensated for in comparisons with
345 estimates that identify the altitude of greatest absolute change, i.e. those that do not
346 employ a log transform.

347 The effect of the log transform on the whole dataset is shown in Figure 5. There are two
348 distinct populations. On some days, the low-altitude artifact traps the maximum mixing
349 depth at the bottom of the instrument range. Applying the log transform removes the
350 effect of the artifact, allowing the true mixing depth (which is variable) to be detected.
351 On days on which the algorithm is not fooled by the artifact, the offset introduced by the

352 log transform is visible: applying the transform results in an average increase of about 50
353 m in the estimated mixing depth.

354 Even with adjustments, our confidence in mixing height estimates derived from the
355 ceilometer is not as high as in those derived from the MiniMPL. One proxy for confi-
356 dence in a given day's results is the degree of concurrence among estimates in the voting
357 procedure (see section 2). As can be seen in Figure 6, the MiniMPL achieves unanimity
358 (concurrence score of 5/5) or near-unanimity (score of 4/5) on 51% days for which data
359 is available. By contrast, the ceilometer achieves a score of 4/5 or better on only 36% of
360 days. It is for this reason that we focus our results on estimates derived from MiniMPL
361 observations.

3.3. Sonde Comparison

362 In September 2012, a one-day intensive campaign of sonde launches was conducted for
363 comparison to mixing layer information from the MiniMPL. Sondes were launched every
364 three hours between 7:00am and 7:00pm local time. The results are displayed in Figure
365 7. In each case, the PBL height is extracted from the sonde using the method of *Heffter*
366 [1980]. At 7:00am, morning fog is still present and the mixing layer has not yet developed.
367 At 10:00am, 1:00pm, and 4:00pm, the mixing height identified using the backscatter data
368 coincides with the sonde-derived PBL height to within 150 m. Since the top of the mixing
369 layer is in fact a transitional zone of 100 to 200 m thickness, it should not be considered to
370 have a well-defined exact location. Some discrepancy should therefore be expected even
371 between methods that detect substantially the same layer. In this one-day comparison,
372 the backscatter method displays no identifiable systematic bias with respect to the sonde
373 method; of course, the comparison presented is too limited to conclude that no bias exists.

374 We emphasize that, while the sonde comparison provides some confidence that, at least
375 during the day, the layer in which elevated aerosol levels are present does correspond to
376 the thermodynamic boundary layer, it is in any case the former that is of most interest
377 for interpreting atmospheric concentrations of trace gases. For the purpose of linking
378 atmospheric measurements to emissions rates, the important question is what part of the
379 atmosphere should be considered in contact with the surface. In other words, through
380 what volume are species emitted from the surface dispersed?

381 By 7:00pm, the mixing layer has begun to collapse and the structure is becoming more
382 complicated. Two distinct boundaries are visible in both the potential temperature profile
383 and the backscatter distribution, and both methods select the higher of these. Indeed,
384 the day's aerosol emissions are distributed up to the higher boundary at 920 m. However,
385 with the decrease in solar heating to drive vertical motion, the upper part of the identi-
386 fied layer (above about 500 m) is probably no longer interacting with the surface. Our
387 method has therefore failed to detect a region of substantial, active vertical mixing. This
388 case serves as a reminder that the mixing layer concept is not always straightforwardly
389 applicable, particularly in the evening as vertical mixing tapers off. Care should be taken
390 in interpreting and applying our or any other mixing depth estimates around sundown,
391 even on days – like this one – with otherwise robust retrievals.

3.4. North American Regional Reanalysis (NARR) Comparison

392 GHG flux inversion studies typically make use of PBL heights derived from meteorolo-
393 gical models or reanalysis products. We compare afternoon maximum mixing depth
394 estimates based on MiniMPL data to PBL height estimates from the Weather Research
395 and Forecasting model (WRF) and the North American Regional Reanalysis (NARR).

396 NARR is a reanalysis product providing a variety of atmospheric and surface variables
397 over North America at 32 km spatial resolution and at 3 hour intervals [*Mesinger et al.*,
398 2005]. We find a large and persistent difference between afternoon maxima of MiniMPL-
399 derived mixing depths at Caltech and PBL height estimates at the nearest NARR grid
400 location. Figure 8 shows the distributions of these quantities over 227 days on which
401 the MiniMPL estimate achieves a concurrence score of at least 4/5 and without data
402 gaps longer than one hour. The maximum NARR PBL height exceeds the maximum
403 MiniMPL-derived mixing depth on all but one day, differing by a factor of two or more
404 on 63% of days. Summary statistics are in Table 3.

405 Interestingly, although maximum NARR PBL heights are an average of 2.5 times Min-
406 iMPL derived mixing depths, the two quantities are similarly distributed. Both show
407 substantial variability, with standard deviations about 32% of the respective means, and
408 both are skewed toward high values, with skewness 0.84 (NARR) and 0.88 (MiniMPL).
409 However, NARR does not reproduce the detailed timing of this variability. Even after
410 scaling maximum NARR PBL heights down by a factor of 2.5 to account for the mean
411 difference, a root-mean-square difference of 360 m remains between scaled NARR esti-
412 mates and MiniMPL estimates on the same days. This is almost as large as the RMS
413 difference of 370 m in a sample of 10^6 random pairs of MiniMPL estimates and scaled
414 NARR estimates.

415 We can attribute NARR's failure to accurately represent the boundary layer in Pasadena
416 at least in part to its coarse spatial grid. The meteorology of the Los Angeles basin is
417 strongly influenced by the coastal mountain topography (see Figure 9), resulting in a
418 complex pattern of circulation [*Lu and Turco*, 1994, 1995]. It comes as no surprise that

419 a product unable to resolve the rapid changes in elevation will struggle to predict PBL
420 heights in this environment. If NARR is used to drive a transport model for GHG flux
421 estimation in Los Angeles or in other areas with meteorology strongly influenced by the
422 detailed topography, careful evaluation and correction of mixing depth biases will be
423 critical for avoiding large errors. Since a biased mixing depth results in a proportional
424 bias in flux estimates (see the general argument in section 4), we would expect a 250%
425 bias in an LA flux inversion using NARR.

3.5. Weather Research and Forecasting (WRF) Comparison

426 Given the difficulty posed by the rapidly-varying topography of the LA basin, one might
427 expect a high-resolution model to better represent the mixing dynamics. We compare
428 mixing depth estimates from MiniMPL data taken during a deployment of the instrument
429 in October-November 2015 to PBL heights from such a high-resolution model, a WRF
430 setup developed specifically for the Los Angeles environment by *Feng et al.* [2016] to
431 simulate CO₂ concentrations. The model is initialized with NARR and with sea surface
432 temperatures from NCEP and uses three nested domains, with the innermost domain
433 covering the LA basin at a resolution of 1.3 km. Using observations from the intensive
434 Calnex campaign in 2010, including aircraft and ceilometer PBL measurements, *Feng et al.*
435 [2016] tested a variety of WRF configurations. We employ only the MYNN_UCM_d03
436 configuration, which they found to minimize errors.

437 We redeployed the MiniMPL to Caltech for the three-week period of October 21 to
438 November 9, 2015. Of these twenty days of observations, the mixing depth estimation
439 algorithm achieves a concurrence score of 4/5 or better on six days and a score of 3/5
440 on another nine days. Although this comparison period is too short to allow robust sta-

441 tistical conclusions, we make some preliminary observations. Like NARR, WRF PBL
442 heights show variability that is similar in relative terms to that of Mini-MPL derived mix-
443 ing heights. Over the three-week comparison period, the standard deviation of maximum
444 afternoon WRF PBL heights is 540 m, about 37% of the mean. However, WRF estimates
445 PBL heights that are greater than MiniMPL-derived mixing depths on all but one after-
446 noon. On average, afternoon maximum WRF PBL height exceeds afternoon maximum
447 MiniMPL mixing depth by 730 m. Considering only days with high concurrence scores
448 reduces the discrepancy considerably. The mean difference on days with scores of 4/5 or
449 better is 380 m, suggesting that the concurrence voting scheme effectively identifies days
450 that are easier to analyze.

451 The discrepancy we find between modeled PBL height and MiniMPL-derived mixing
452 depth is surprising given the excellent agreement reported by *Feng et al.* [2016]. Dur-
453 ing the 2010 Calnex campaign period, they report a mean WRF-derived daytime PBL
454 height (using the same MYNN_UCM_d03 configuration we use here) of 828.8 m, in good
455 agreement with a mean mixing depth of 835.7 m obtained from ceilometer measurements
456 using the gradient method. They also report substantially less variability in modeled PBL
457 height than in measured mixing depth. Further work, including a model-data comparison
458 covering a longer period, is clearly needed to resolve this perplexing difference. While
459 such a comparison is beyond the scope of this study, we do note that NARR PBL height
460 estimates for May-June 2010 are generally similar to those from our comparison period
461 in October-November 2015, with a mean daily maximum of 2.1 km.

462 Our analysis here cannot distinguish between differences due to errors in mixing depth
463 estimation, errors in modeled PBL depth, or conditions under which the mixing layer

464 fails to correspond to the thermodynamic PBL. Still, it is prudent to expect that the
465 same complex stratification which can cause the mixing depth estimation algorithm to
466 fail might also indicate challenging conditions for the model. By selecting days with high
467 concurrence scores, MiniMPL observations can be used to choose “golden days” for model
468 analysis. Alternatively, if a model is run over a long period, days with good agreement
469 between the model and lidar estimates can be selected for flux estimation. For example,
470 Figure 10 shows a pair of days which would not be readily distinguished on the basis of
471 model results alone. The additional information provided by the lidar estimates lets us
472 assign greater confidence to modeling on the day with good agreement (panel a) than that
473 with poor agreement (panel b).

474 Sustained lidar can also inform the choice of model configurations or parameters, as *Feng*
475 *et al.* [2016] and others [e.g., *Nehrkorn et al.*, 2013] have done with PBL observations from
476 limited campaigns. In addition to increasing confidence in that choice simply by virtue
477 of a larger volume of data, long-term observations can provide more detailed information
478 about how model errors depend on season or on other meteorological conditions. For
479 example, *Lewis et al.* [2013] found that PBL height as estimated by the general circulation
480 model GEOS-5 differs most from that measured by the lidar network MPLNET in winter.
481 Unlike sondes, lidar data can validate not only the depth of the mixing layer but also the
482 timing of its development and collapse. That timing can be critical; for example, in an
483 urban setting, a difference of one hour may determine whether the mixing layer begins
484 to develop before, during, or after the emissions peak associated with the morning rush
485 hour.

486 An alternate method for integrating mixing depth observations into flux estimation is
487 to characterize a known model bias and correct for it after the modeling stage. *Zhao et al.*
488 [2009] use three months of wind profiler measurements to derive an empirical relationship
489 between observed and modeled PBL heights. They apply that relationship to scale down
490 modeled PBL height before computing fluxes, reducing the residual error by a factor of
491 1.5. Among the advantages of postprocessing corrections of this kind are that they are
492 simple to apply, allowing accuracy to be improved even in less detailed inversions, and
493 that they can be combined with the strategies discussed above to further control any
494 errors remaining after tuning model parameterization and/or selecting out “golden days.”

3.6. Spatial Variation

495 Taking advantage of the MiniMPL’s portability, we also conducted a one-time pilot
496 mobile study in which backscatter data was collected over a period of about twenty min-
497 utes as the MiniMPL was transported due west toward the Pacific coast in the back of a
498 passenger car. This observing strategy, which could not have been implemented with a
499 full-size research lidar, is made possible by the compact size and low power requirements
500 of the MiniMPL. Now that we have demonstrated its feasibility, we hope that this new
501 approach will allow for both more regular mapping of the spatial structure of the mixing
502 layer and more nimble mobile deployment of lidar in response to irregular events like fires
503 and gas leaks.

504 The spatial profile of aerosol backscatter near the Pacific coast is shown in Figure 11.
505 The transition between the shallow marine layer, which extends two to three kilometers
506 onto land, and the convective regime that dominates further inland is clearly visible. The
507 vertical structure in this case is simple, with a well-defined mixing layer of high backscatter

508 adjacent to the ground and a sharp decrease in backscatter at the top of that layer. The
509 mixing depth as estimated by the minimum backscatter gradient is indicated in the figure
510 by the black circles (our retrieval algorithm is not suitable, since it relies on the temporal
511 evolution of the boundary at a fixed location).

512 Figure 11 also shows the PBL height as predicted by WRF. The WRF prediction agrees
513 well with the MiniMPL-derived mixing depth near the coast, but does not increase as
514 sharply further inland. Unlike at Caltech, in this case the WRF PBL height is lower than
515 the observed mixing depth. Repeated measurements of this kind could reveal whether the
516 difference is consistent with time and at locations elsewhere along the coast, both in the
517 immediate Los Angeles area and elsewhere, which could contribute to model development
518 in the challenging coastal environment. Further work characterizing the coastal transition
519 could also aid in understanding the fate of GHG emissions from sources like ports and
520 marine industry.

4. Conclusions

521 Researchers have recognized that the representation of mixing dynamics is both critical
522 for the interpretation of top-down emissions estimates and also a major source of uncer-
523 tainty [e.g., *Newman et al.*, 2013; *Zimnoch et al.*, 2010]. *McKain et al.* [2012] advocate
524 the use of column-integrated concentration measurements in urban studies, among other
525 reasons in order to avoid the impact of mixing height errors. A common strategy [*Breon*
526 *et al.*, 2014] is to rely only on observations made during midafternoon, when the mixing
527 layer is at or near its maximum depth and the detailed timing of its dynamics are less
528 important. But we observe even afternoon maximum mixing height in Los Angeles to
529 vary substantially from day to day, typically differing from the seasonal mean by 30%.

530 A simple dimensional argument demonstrates the impact of such variations. Suppose
531 that an instrument measures the in situ concentration of a trace gas at a particular
532 location. This concentration is expressed as a molar fraction, or, equivalently given the
533 local density of dry air, as a volume concentration C in moles of gas per unit volume, i.e.
534 in n/L^3 . The goal is to use the measurement to infer a surface flux F , expressed in moles
535 of gas emitted or absorbed per unit area per unit time, i.e. as $n/(L^2T)$. On dimensional
536 grounds, any method for relating the concentration to the flux must incorporate some
537 temporal information, such as the time τ during which the sampled air mass was exposed
538 to the flux, and also some vertical length scale.

539 The relevant vertical length scale is the mixing height h , which controls the height of
540 the space into which the emitted gas is diluted. We therefore expect

$$541 \quad F \propto hC/\tau \quad (5)$$

542 in which case an error in the mixing height h will result in a proportional error in the
543 flux estimate. In detailed models, this picture is complicated to some degree by higher-
544 order effects, e.g. the coupling between vertical motion and horizontal wind shear, but
545 the essential proportionality remains. Applying sustained observations to control mixing
546 depth errors, whether by validating models, choosing suitable periods for analysis, or
547 characterizing and correcting for errors in postprocessing, is critical for accurate GHG
548 flux estimation.

549 We have focused above on determining the depth of the mixing layer, especially at
550 its afternoon maximum. But the mixing layer concept is not always applicable. Even
551 when the mixing height is applicable, it does not fully describe the complex structure
552 of the lower troposphere. The potential exists to extract much more information about

553 that structure from lidar backscatter data. Among other applications, a more complete
554 picture of the mixing state could contribute to our understanding of the transport of
555 species emitted from the surface. Here we suggest one direction in particular for future
556 work.

557 The mixing layer itself may exhibit internal structure. For example, in Los Angeles, the
558 sea breeze circulation pushes near-surface air inland during the day. As a result, the air
559 mass within the mixing layer over Pasadena in the afternoon has traveled over downtown in
560 the preceding hours. The time scale of this horizontal motion, and the varying emissions
561 rates and compositions from the traversed areas, may create a stratification, in which
562 fresh emissions from Pasadena are concentrated in the lowest part of the mixing layer
563 while those from downtown are more thoroughly mixed throughout. If we were able to
564 observe and understand within-layer dynamics of this kind, we could much more precisely
565 link trace gases observed in the atmosphere to their points of emission, allowing us to
566 answer more specific questions about the sources and composition of emissions in the
567 urban environment.

568 Since the lidar is primarily sensitive to aerosols and not to trace gases, the distribution
569 of aerosol would need to be used as a proxy for the distribution of co-emitted trace gases,
570 assuming that the two are transported within the mixing layer in a similar way, at least
571 on short time scales and over small distances. That assumption would need to be tested
572 before it could form the basis of any future work. Challenges notwithstanding, this is an
573 exciting possibility for future applications, including more detailed validation of transport
574 models and finer-scale attribution of emissions sources within complex urban environments
575 like that of Los Angeles.

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585 To obtain a copy of the analysis system used to generate the estimates, please contact the
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Figure 1. A complete MiniMPL lidar system consists of an optical transceiver (shown) and a laptop running data acquisition and post-processing software.

Figure 2. A sample day of backscatter data (heatmap) from the MiniMPL (panel a) and ceilometer (panel b, see section 3.2) with mixing heights as estimated by our algorithm (black symbols: majority opinion; green symbols: estimates initialized at other times of day). Prior to 8am, both instrument beams are completely extinguished near the surface; the algorithm recognizes the presence of fog and does not attempt to make an estimate. In the late afternoon – and in the morning in the case of the ceilometer – the various estimates disagree as to the mixing height, identifying two different boundaries. We report the majority opinion together with the degree of concurrence (4/5 for the MiniMPL, 2/5 for the ceilometer). Note that MiniMPL NRB values and ceilometer backscatter values do not use comparable scales.

Figure 3. An illustration of the wavelet method. The instrument returns a vertical profile of normalized relative backscatter (NRB, left). To compute the wavelet covariance at a given altitude z , the backscatter profile is integrated against a Haar wavelet centered at z (middle). In this example, the covariance is given by the difference in area between the orange (upper) shaded region and the blue (lower) region, which indicates the decrease in backscatter over the scale of the wavelet. The resulting Haar wavelet covariance is shown at right.

Figure 4. Solid curves: average diurnal cycles of mixing height in June-August (orange) and December-February (blue). Shaded regions: one standard deviation of between-days variability. Estimates according to the MiniMPL, retaining only days on which the concurrence score was at least 4/5.

Figure 5. (a) Representative backscatter profiles from the ceilometer, with (orange, triangular symbols) and without (blue, round symbols) applying a log transform. The horizontal lines show the corresponding mixing height as estimated by the algorithms: solid blue line, without transform; dashed orange line, with transform. Note the very high backscatter values at low altitudes in the untransformed data, which fool the algorithm into selecting an unrealistically low mixing height.

(b) Maximum afternoon mixing depths as estimated using ceilometer data with (vertical axis) or without (horizontal axis) applying the log transform. Days on which the untransformed data is affected by the low-altitude artifact are indicated by the dashed green ellipse; taking the log transform removes the effect of the artifact. On other days (indicated by the solid pink ellipse), the bias introduced by the transform is visible. The solid black line is the 1-1 line. Only days with concurrence scores of at least 3/5 are shown.

Figure 6. Degree of concurrence achieved by the algorithm using backscatter data from the MiniMPL (blue, solid) or from the ceilometer (orange, dashed), shown as a fraction of days on which both instruments were operating.

Figure 7. Orange: potential temperature profiles from sonde launches, with the corresponding PBL height as calculated using Heffter’s method (horizontal dashed line). Blue with triangles: contemporaneous MiniMPL backscatter profiles, with the mean (center horizontal line) and range (shaded area) of the algorithmically-estimated mixing height over the 30-minute period surrounding the sonde launch.

Table 1. Technical specifications for the MiniMPL (used in this study) and the standard MPL.

Figure 8. Gaussian kernel density (smoothed relative frequency) of maximum afternoon mixing depth according to the MiniMPL (pink, left peak) and according to NARR (green, right peak) over 227 days with MiniMPL concurrence score at least 4/5. Solid vertical lines: median; dashed vertical lines: quartiles.

Figure 9. (a) Elevation map of the Los Angeles Basin [*U.S. Geological Survey, 2015*]. The labeled diamonds indicates the location of the measurement site at Caltech (in Pasadena). The solid line shows the route taken in the mobile study; the dashed line corresponds to the cross section in panel (b).

(b) Elevation cross section along the dashed line in panel (a); the longitude scale is the same for both panels.

Figure 10. Examples of days with good (panel a) and poor (panel b) agreement between MiniMPL derived mixing depths (small circles) and PBL heights as estimated by WRF (large diamonds). NARR PBL heights (large triangles) show large discrepancies in both cases.

Figure 11. Heatmap: MiniMPL backscatter intensity near the Pacific coast (located at longitude 118.41). Small black circles: mixing depth as estimated by the gradient method using MiniMPL data. Large black diamonds: PBL height as estimated by WRF. Black curve at bottom: topography (same vertical scale).

Table 2. Number of days of MiniMPL data collection without gaps longer than one hour, by month and by concurrence score of the mixing depth estimation algorithm.

Month	Total	1/5	2/5	3/5	4/5	5/5
January	52	3	4	11	10	24
February	26	4	7	4	5	6
March	29	1	5	8	7	8
April	31	1	3	12	10	5
May	26	0	7	5	8	6
June	14	1	2	1	3	7
July	0	0	0	0	0	0
August	49	1	5	15	14	14
September	85	1	18	24	16	26
October	64	5	13	16	13	17
November	48	4	13	15	11	5
December	52	4	14	16	7	11
All Months	476	25	91	127	104	129

Table 3. Mean, median, 1st and 3rd quartiles, and standard deviation of afternoon maximum PBL height (NARR) or mixing depth (MiniMPL), in km AGL, over 227 days with concurrence score at least 4/5 and without data gaps longer than one hour.

Method	Q_1	Median	Mean	Q_3	σ	σ/Mean
MiniMPL	0.63	0.75	0.84	0.98	0.27	32%
NARR	1.46	1.84	1.92	2.20	0.62	32%

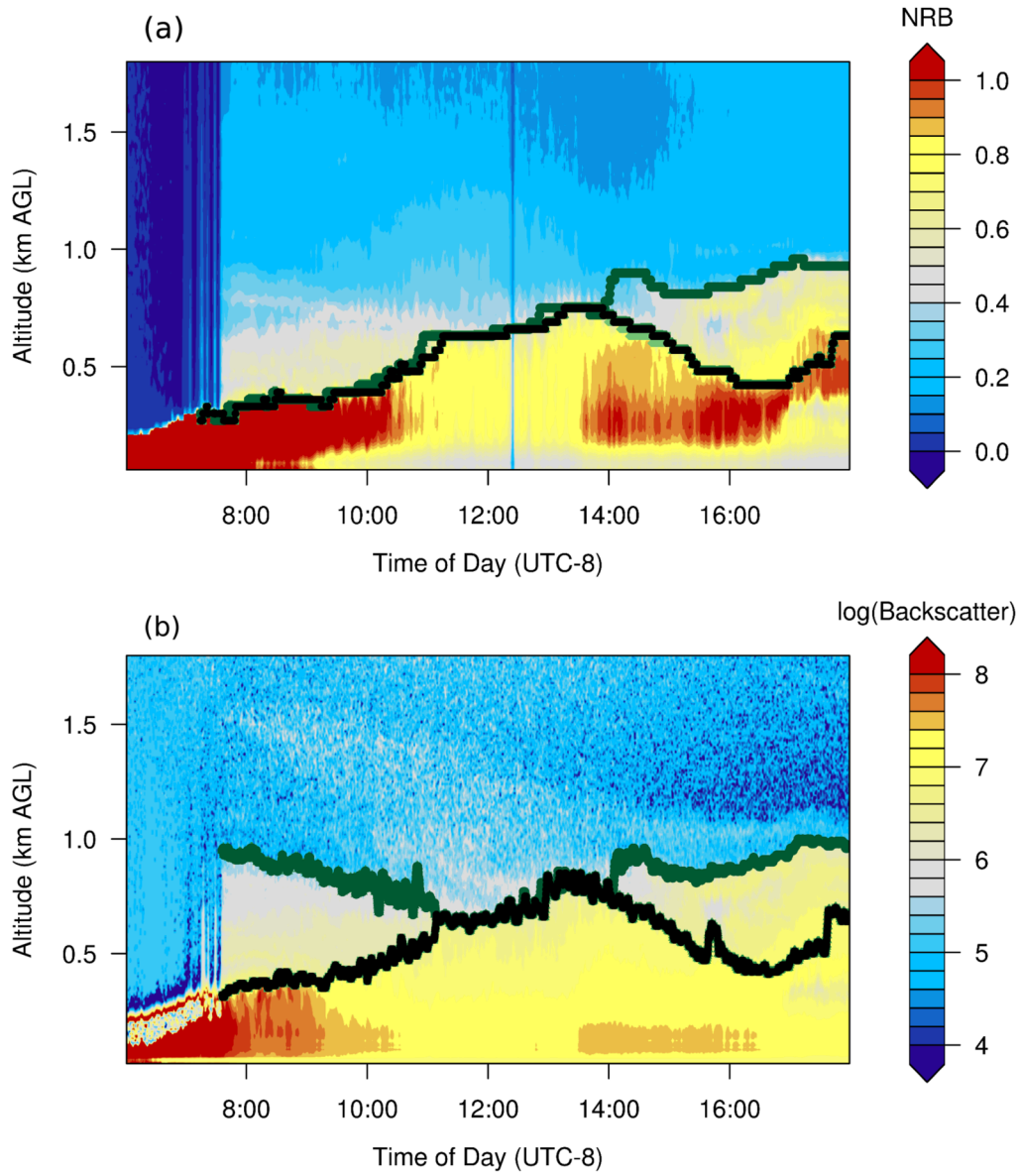
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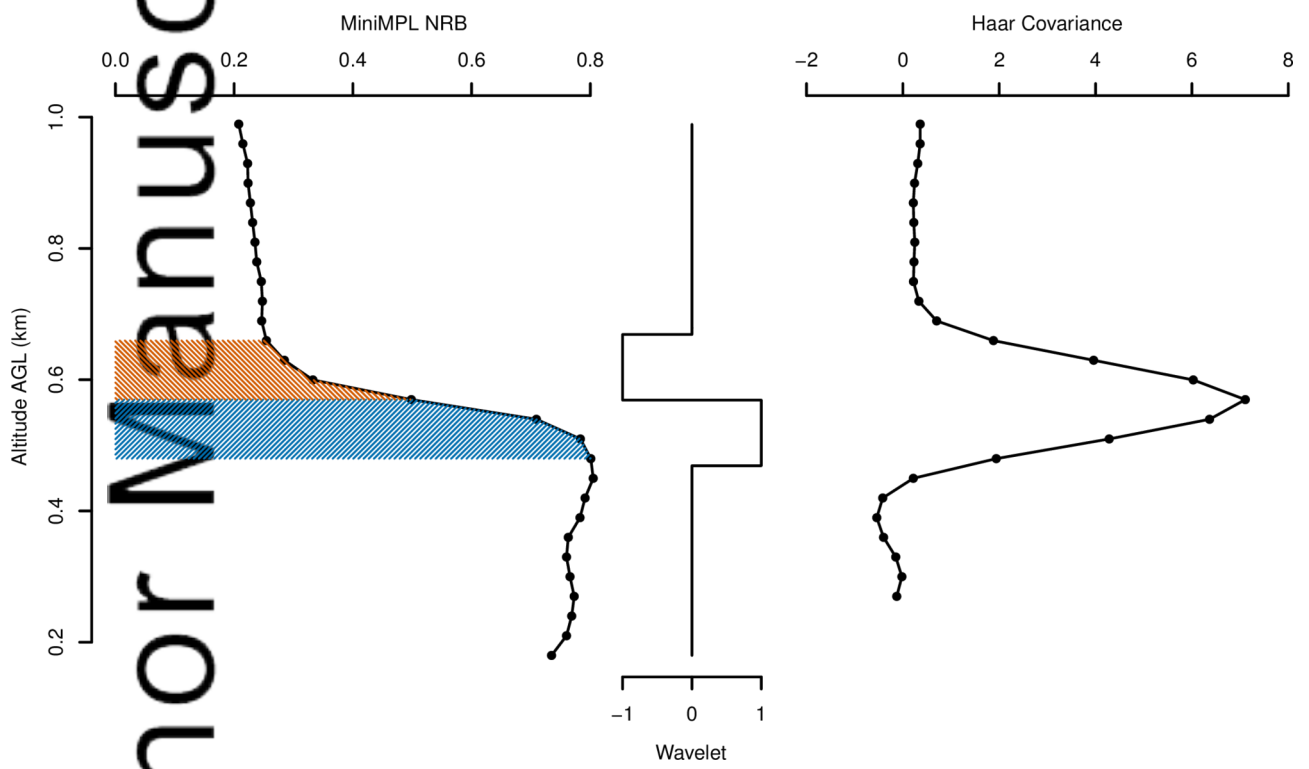
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		MPL	MiniMPL
PERFORMANCE	range resolution	5/15/30/75 m (software programmable)	5/15/30/75 m (software programmable)
	minimum range	150 m	150 m
	accumulation time	1 sec - 15 min	1 sec - 15 min
	detection range	up to 25 km	up to 25 km
	polarization	standard	standard
	scanning	optional	optional
OPTICS	laser wavelength	532 nm	532 nm
	laser pulse energy	6 - 10 μ J @ 2500 Hz	3 - 4 μ J @ 2500 Hz
	eye-safety	ANSI Z136.1 2000 IEC 60825	ANSI Z136.1 2000 IEC 60825
	receiver diameter	178 mm	80 mm
	pump laser diode	fiber coupled user replaceable	
	detector	fiber coupled user replaceable	fiber coupled user replaceable
DIMENSIONS	size	300 x 350 x 850 mm	240 x 305 x 480 mm
	weight (portability)	25 + 2 kg	13 kg
DATA	operating system	Windows 7/10	Windows 7/10
	computer interface	USB	USB
	data transfer	LAN ethernet	LAN ethernet
ENVIRONMENT	temperature	NEMA-4 enclosure	NEMA-4 enclosure
	humidity	enclosure	enclosure
POWER	supply	100/240 V AC 50 - 60 Hz	100/240 V AC 50 - 60 Hz
	consumption	500 W	100 W

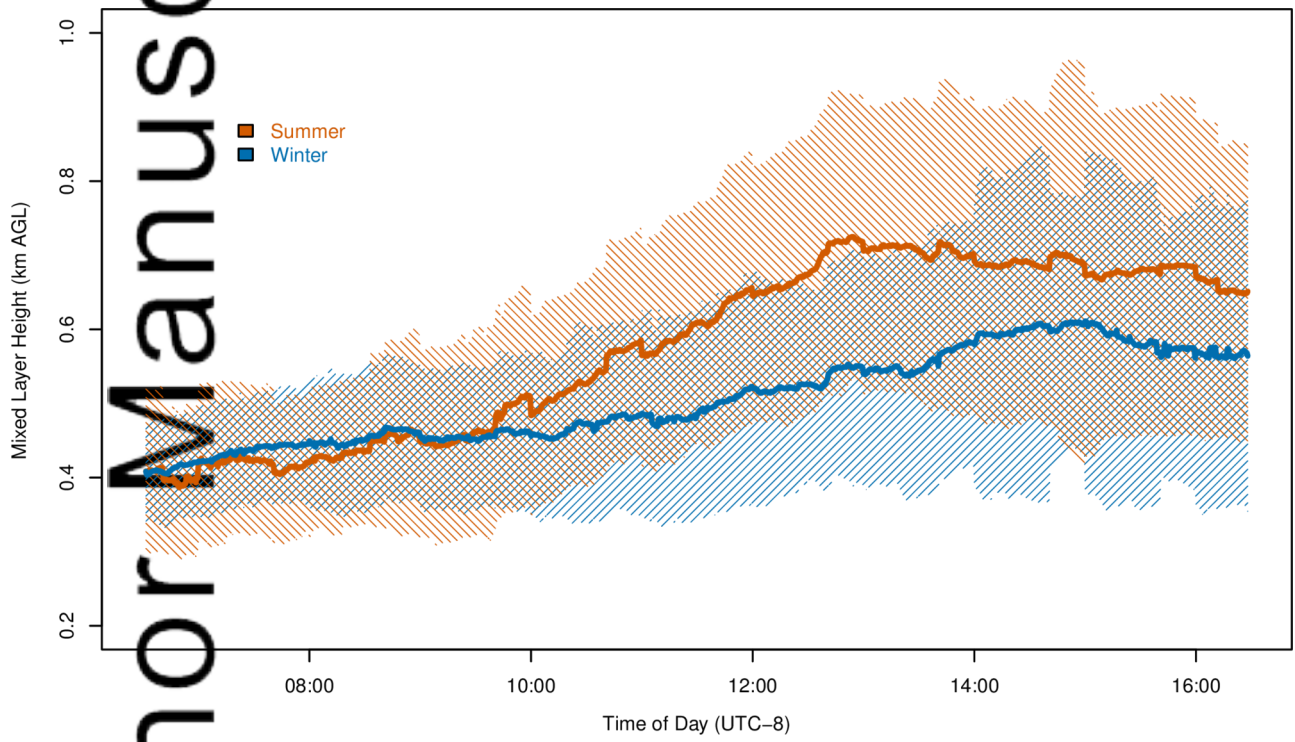
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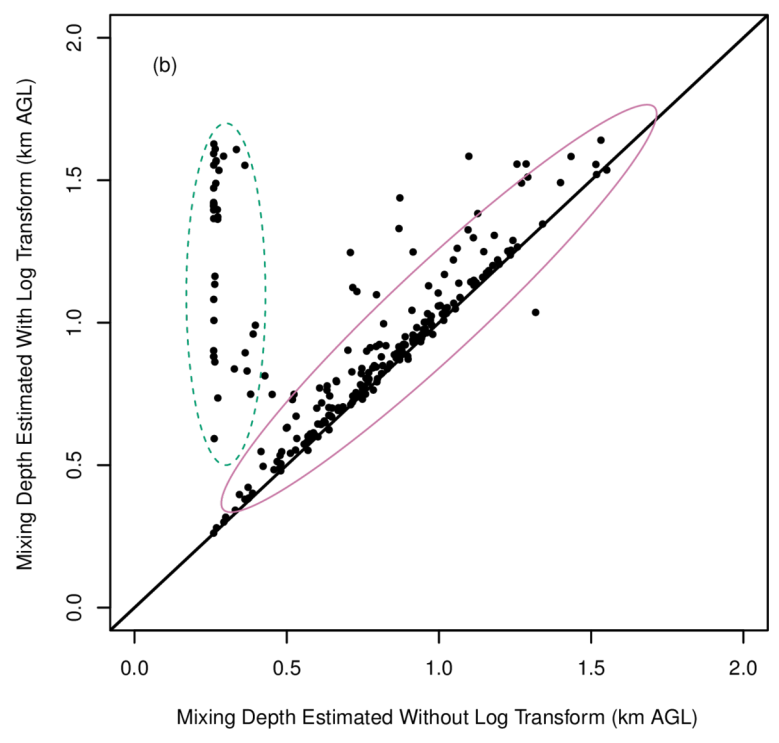
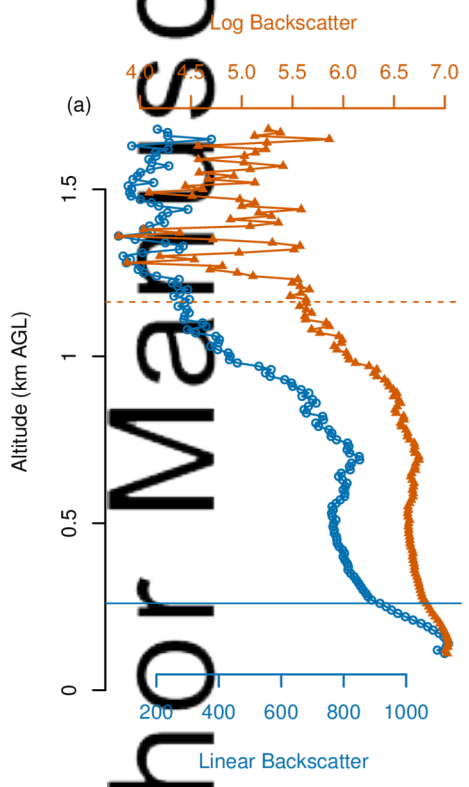
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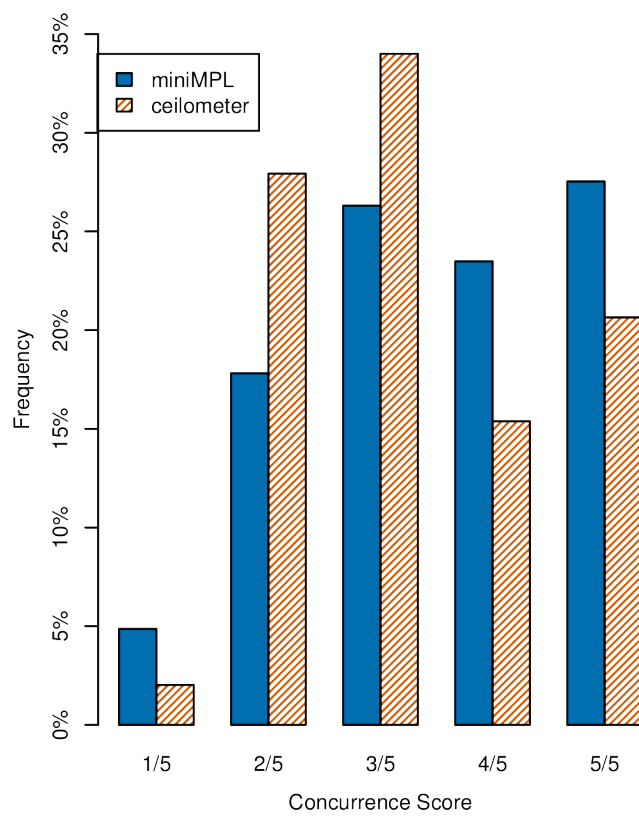


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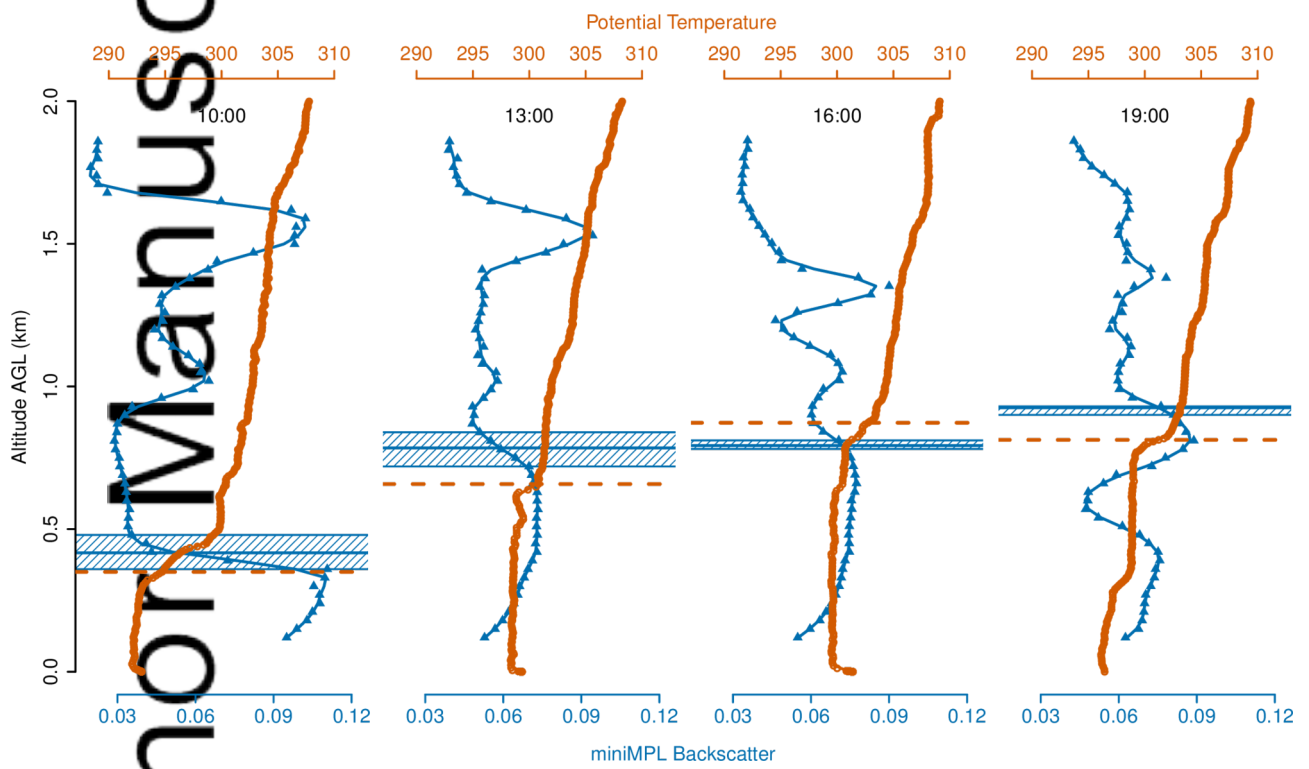


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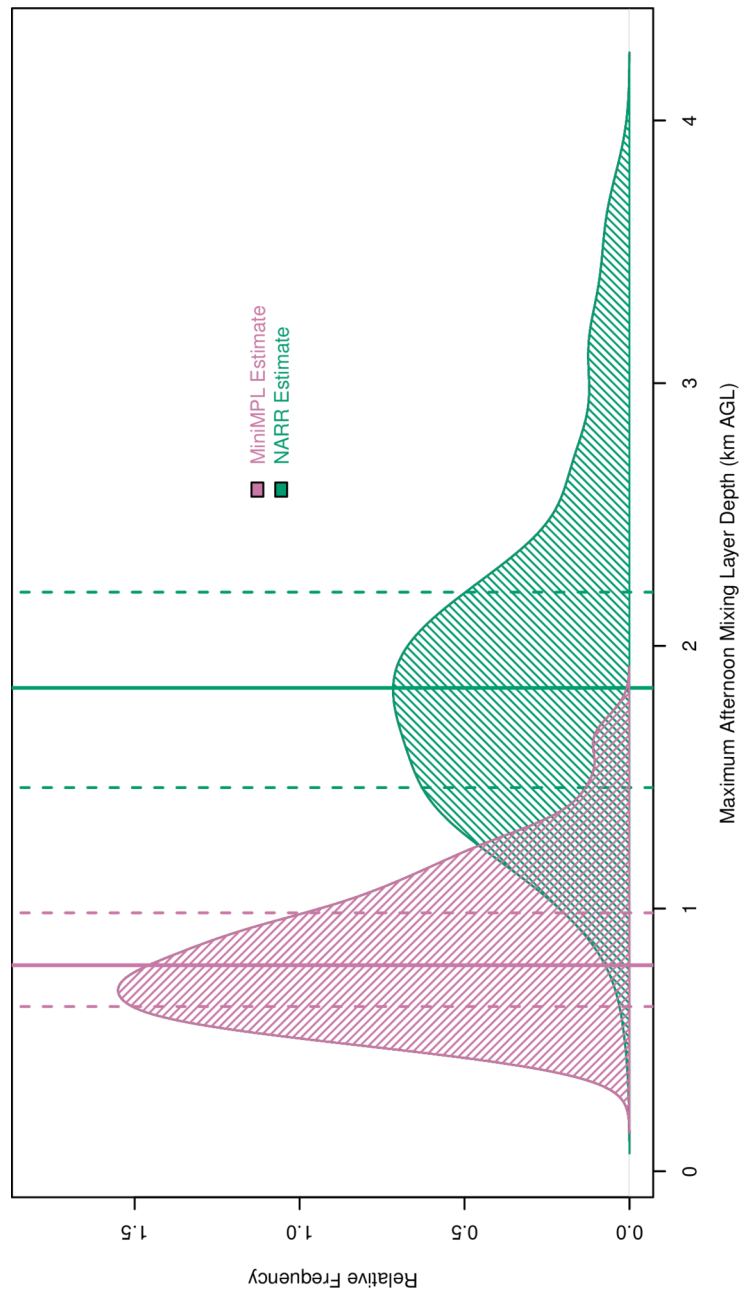


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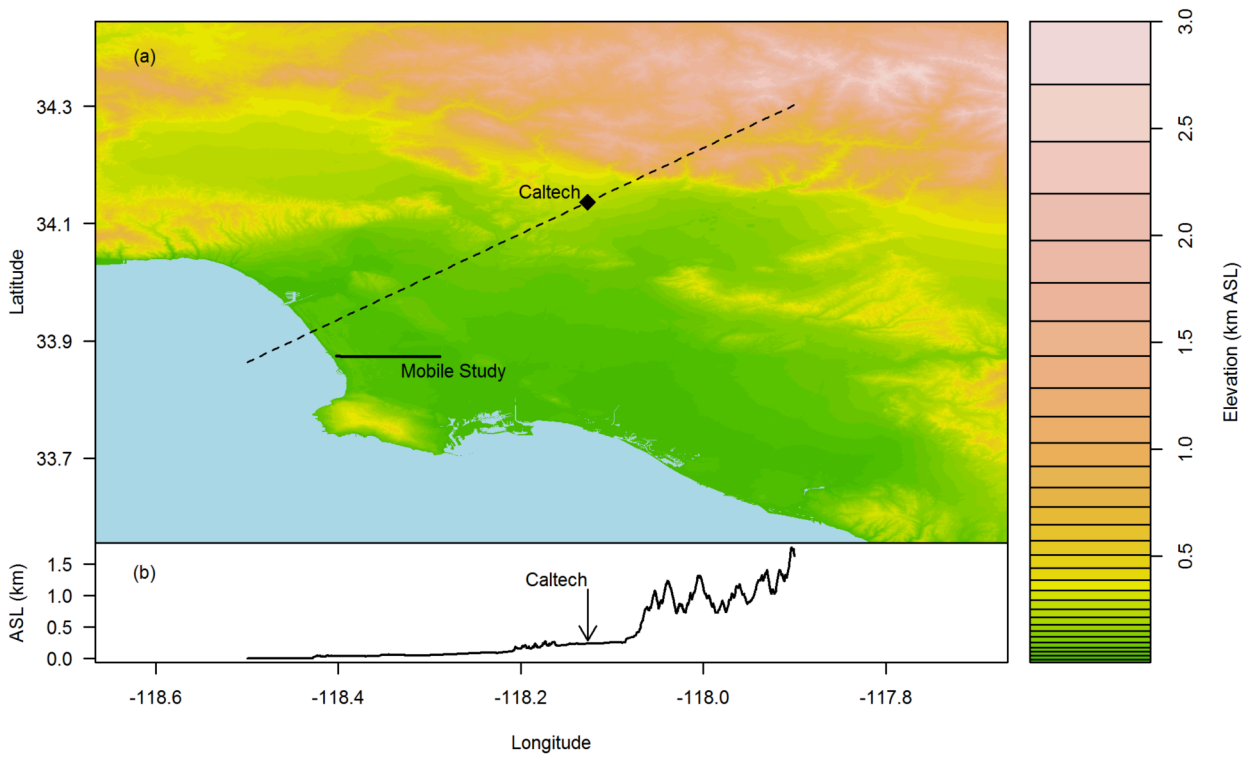
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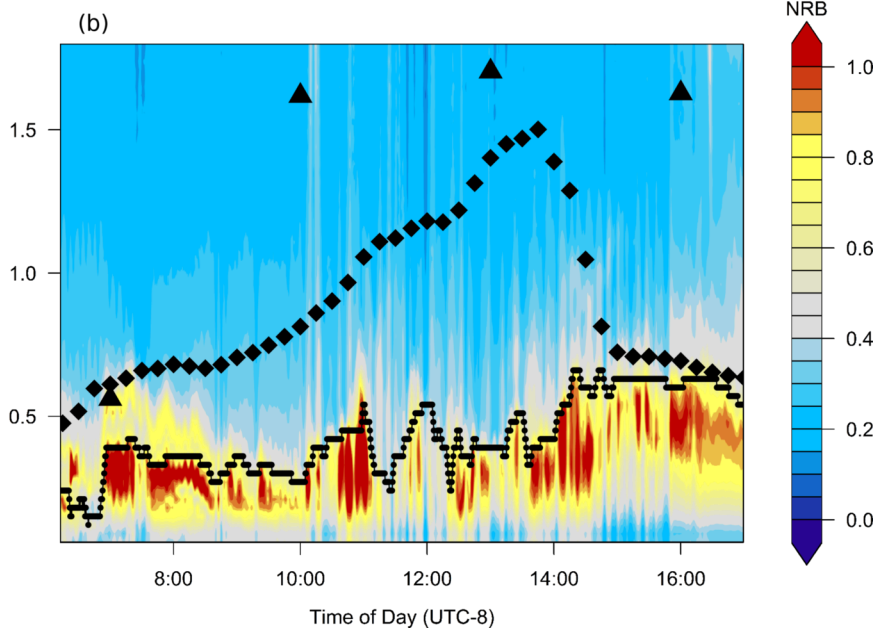
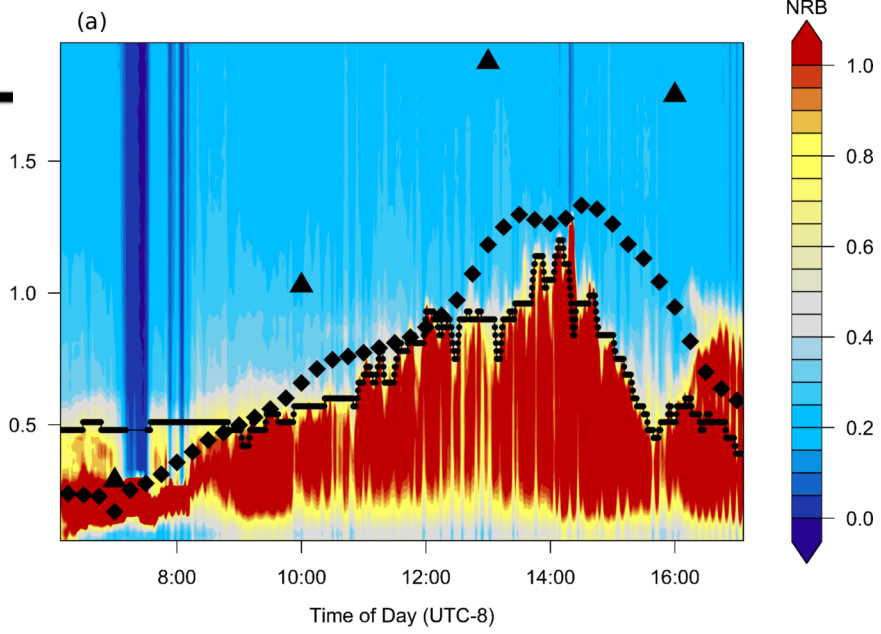
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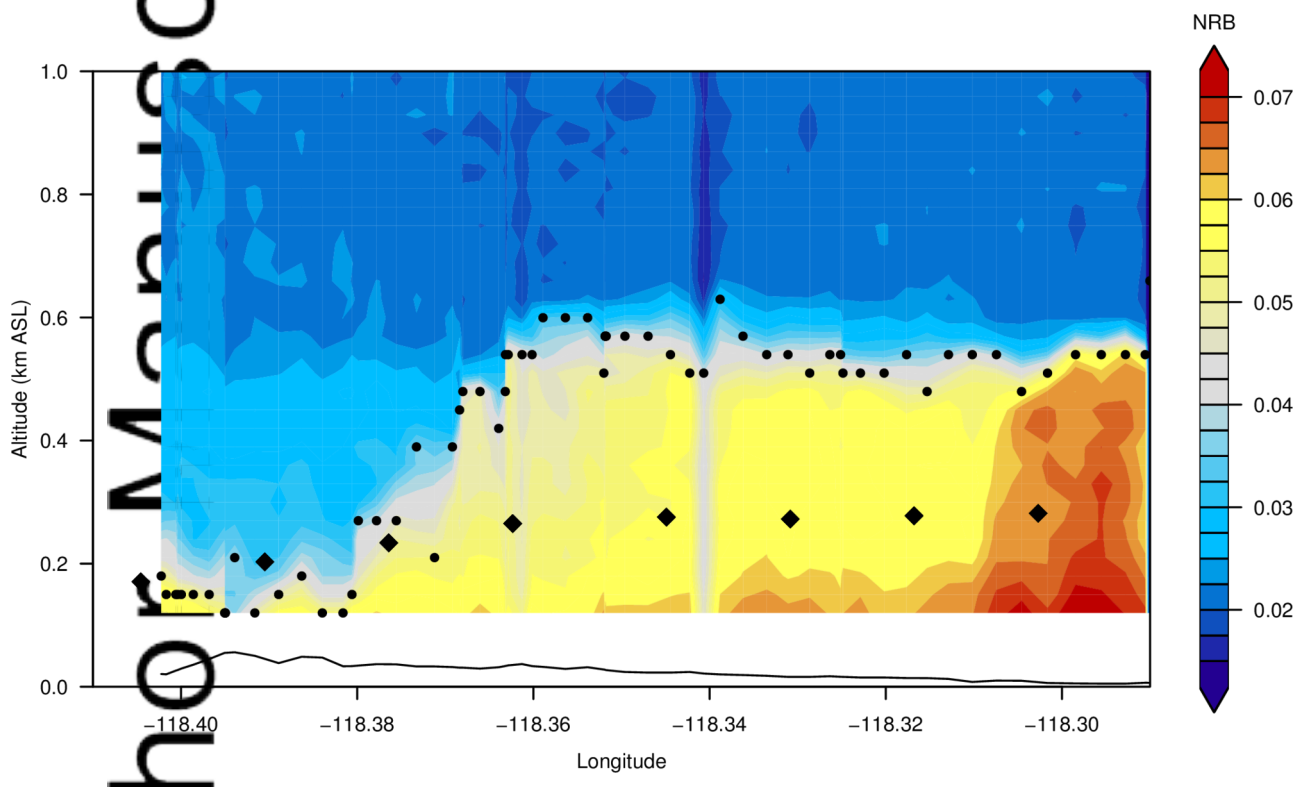


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