GNSS-R Non-Local Sea State Dependencies: Model and Empirical Verification

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5 Abstract

- 6 Global Navigation Satellite System Reflectometry (GNSS-R) is an active, bistatic remote sensing
- 7 technique operating at L-band frequencies. GNSS-R signals scattered from a rough ocean surface are
- 8 known to interact with longer surface waves than traditional scatterometery and altimetry signals. A
- 9 revised forward model for GNSS-R measurements is presented which assumes an ocean surface wave
- 10 spectrum that is forced by other sources than just the local near-surface winds. The model is motivated
- 11 by recent spaceborne GNSS-R observations that indicate a strong scattering dependence on significant
- 12 wave height, even after controlling for local wind speed. This behavior is not well represented by the
- 13 most commonly used GNSS-R scattering model, which features a one-to-one relationship between wind
- speed and the mean-square-slope of the ocean surface. The revised forward model incorporates a third
- 15 generation wave model that is skillful at representing long waves, an anchored spectral tail model, and a
- 16 GNSS-R electromagnetic scattering model. In comparisons with the spaceborne measurements, the new
- 17 model is much better able to reproduce the empirical behavior.

18 1 Introduction and Overview

19 Global Navigation Satellite System Reflectometry (GNSS-R) is a relatively young remote sensing 20 technique proposed to measure geophysical quantities such as ocean surface roughness and wind speed. With it quickly gaining momentum [Zavorotny et al., 2014], there has been rapid and ongoing 21 22 development of instrumentation [e.g., Gleason et al., 2016], retrieval algorithms [e.g., Clarizia et al., 23 2014] and scattering models [e.g., Zavorotny and Voronovich, 2000; Lin and Katzberg, 1999]. GNSS-R is a 24 relatively low-cost technique which leverages existing navigation signals as the transmitter half of the 25 bistatic radar system. This technique makes use of a forward scattering geometry, in contrast to 26 conventional monostatic scatterometers and altimeters, which use a back scattering geometry. The 27 frequency of operation is dictated by the transmitters, which are typically L-band (1-2 GHz) navigation 28 satellites. 29

30 The combination of L-band signals and forward scattering geometry has been rarely used in the past by

- 31 remote sensing instruments, and thus brings about new implications for electromagnetic interaction
- 32 with surface features. In particular, bistatic L-band radar return is dominated by quasi-specular
- 33 scattering, which is dictated by waves longer than about 3 times the electromagnetic wavelength [e.g.,
- Valenzuela, 1978; Brown, 1978]. In the ocean, L-band GNSS-R is therefore sensitive to surface waves of
- 35 about 50 cm in wavelength and longer. In contrast, for radar scatterometers, according to two-scale
- models, these 50 cm waves are tilting waves that bring about mostly secondary effects compared to the
 primary Bragg scatterers. Although radar altimeter scattering is also primarily quasi-specular, they
- 37 primary bragg scatterers. Although radar altimeter scattering is also primarily quasi-specular, they
 38 typically operate at higher C-band (5 GHz) or Ku-band (13 GHz) frequencies (e.g., TOPEX/Poseidon ALT)
- (Fu et al., 1994]), which correspond to wavelengths of order one centimeter. For typical ocean
- 40 roughness spectra, these centimeter scale features dominate the roughness, so the ~50 cm scale waves
- 41 in most cases play only a minor role for altimetric sensors.
- 42

This is the author manuscript accepted for publication and has undergone full peer review but has not been through the copyediting, typesetting, pagination and proofreading process, which may lead to differences between this version and the Version record. Please cite this article as doi:10.1002/2016JC012308.

43 Modelling GNSS-R ocean scattering at L-band presents novel challenges. Short gravity waves tens of 44 centimeters long are in a different regime than the millimeter capillary waves, because they are governed by different physics. With surface tension being negligible, these short gravity waves take 45 longer to dissipate and propagate further before decaying. There have been questions [Cardellach, 2014] 46 47 and results showing non-negligible GNSS-R sensitivity to long gravity waves. However, such waves have 48 yet to be taken into account in GNSS-R forward models. For example, the end-to-end simulator for the 49 upcoming Cyclone Global Navigation Satellite System (CYGNSS) mission [Ruf et al., 2016a] uses the 50 Katzberg relationship to model the roughness, which assumes that the scattering cross section is 51 determined by the local, instantaneous wind speed alone [Ruf et al., 2016b]. The limitation of this 52 assumption is illustrated in Section 3, below, in which spaceborne scattering measurements are shown 53 to exhibit large differences from those predicted by the Katzberg-model (e.g., Figure 8 and Figure 9). The 54 differences are most pronounced, and considered significant, at lower wind speeds. 55 56 In this paper, our objective is to develop, and then experimentally validate, a more accurate GNSS-R

for this paper, our objective is to develop, and then experimentally validate, a more accurate GNSS-R
 forward model by incorporating forcing effects other than local winds. In Section 2, we present the
 model, which includes a third-generation wave model that has not previously been incorporated into a
 GNSS-R forward model. The rationale for the choice of the surface model and parameters therein are
 discussed. In Section 3, we compare spaceborne measurements with our model predictions and with the
 predictions produced by the Katzberg model. We conclude with a discussion of some of the non-local
 effects that contribute to the scattering measurements predicted by our model, and consider other,
 second-order, effects that have not been incorporated into the model but could be as future work.

64 2 The Forward Model

65 2.1 Surface Wave Models

For phase-averaging surface wave models, one important goal is to quantify the spectral energy accurately in the form of a wave spectrum, which can range from a one-dimensional directionallyintegrated spectrum in the simplest case to a full three-dimensional frequency-wavenumber-direction spectrum for linear and nonlinear waves. These models can generally be divided into two types: 1. empirical models based on dimensional analysis and parameterized by wind speed and, possibly, wave age, and 2. spectral evolution models based on the energy-balance equation.

73 The first type constrains the shape of the spectrum, which is typically a smooth function of the input parameters. Usually, conditions are classified as duration- or fetch-limited [Hwang and Wang, 2004], and 74 75 the wave age is computed accordingly. The wave age and windspeed are then used to parameterize the 76 wave spectrum. The Pierson-Muskowitz [Pierson and Moskowitz, 1964], JONSWAP [Hasselmann et al., 77 1973], Elfouhaily [Elfouhaily et al., 1997], and Hwang [Hwang et al., 2013] spectra are of this type. The 78 second type of model includes WAVEWATCH3 [Tolman et al., 2014] (denoted by WW3 hereafter), 79 University of Miami Wave Model [Donelan et al., 2012], SWAN [Booij et al., 1999], and WAM [Komen et 80 al., 1994]. These models solve the energy balance equation numerically, a Eulerian form of which in

- 81 simple cases (conditions given below) may be expressed as
- 82

$$\frac{\partial E(k, x, t)}{\partial t} + c_g \frac{\partial E(k, x, t)}{\partial x} = S(k, x, t) , \qquad (1)$$

- 83 where E is the one-dimensional wavenumber-direction spectrum with SI units of m^3 , with the
- 84 wavenumber energy spectrum being $\rho g E$ with units of J/m. ρ is the mass density of sea water, and g is
- 85 the gravitational constant. c_g is the group velocity in the x direction. S(k, x, t) is the collective source
- 86 term combining the effects of wind input, whitecapping dissipation, and non-linear wave-wave

- 87 interaction. Equation (1) models the temporal evolution and spatial propagation of the elevation
- 88 variance of a one-dimensional wave in deep water and neglects the effects of currents. In practice, an
- 89 equation of this type is discretized and integrated in time and space to solve for the wave spectrum at
- 90 each time step and grid point. The source terms, with improved understanding of wave physics, have
- 91 undergone significant development in the last 50 years, and are now in their "3rd generation" [Komen et
- 92 al., 1994].
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For the second type of wave model, rather than having an a-priori form, the individual source terms are crafted, and the spectrum is left free to evolve. Before the 1950s, models of the first kind were used for wave forecasting. However, several aspects are challenging for the parametric models to handle, such as the accounting for swell generated afar, and irregular bathymetry and coastlines [Ardhuin, 2016 p.52]. In addition, Chen et al. [2016] found that two such empirical models show significant errors in modelling the response time of waves to wind in general conditions, while later investigations showed the third-generation model WW3 performs significantly better in comparisons with in-situ measurements.

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102 Despite its shortcomings, the parametric models and the associated experiments that led to them, are 103 widely used when modelling idealized duration and fetched-limited cases. These ideal cases have been 104 invaluable in the development of the spectral-evolution models' source terms, and they continue to 105 serve as reference calibration points for the state-of-the-art third generation models. Moreover, these 106 parametric models are considered the current state-of-the-art models for high frequency waves. The 107 form and shape of the spectral tail assumed in the model is still an area of active research [Ex. Plant, 108 2015, Reichl et al., 2015, Hwang et al., 2013], partly due to the challenges in their accurate 109 measurement [Hwang, 2005]. Many electromagnetic models to-date have incorporated these parametric models [e.g., Voronovich and Zavorotny, 2001; Apel, 1994; Hwang and Fois, 2015] as the 110 111 surface wave model, with the inverse wave age often set to 0.84 for "well-developed" conditions. It 112 should be noted that formulations of source term balance of short Bragg waves have been attempted 113 [e.g. Lyzenga et al., 1988], but much uncertainty remain [Hwang et al., 2013].

114

115 For GNSS-R, the surface roughness of relevance is the low-pass-filtered mean square slope (mss)

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 $LP_{mss}(k_u) = \int_0^{k_u} k^2 S(k) dk$ ⁽²⁾

(3)

119 Empirically, Brown [1978] found $k_u = \frac{2\pi \cos \theta}{3\lambda}$ to be a suitable cutoff, with λ being the

electromagnetic wavelength and θ denoting the incidence angle of the observation. For the GPS L1 carrier with a frequency of 1.575 GHz, and typical incidence angles of less than 35 degrees, $k_u \approx 10$ rad/m, so waves of about 60 cm and longer are sensed by GNSS-R. We mention in passing that the quantity significant wave height, usually denoted as Hs and used in our analysis in Section 3, can be computed from the wavenumber spectrum as

- $Hs = 4 \sqrt{\int_{0}^{\infty} S(k) dk}$
- 126 Once the wave spectrum is known, LP_{mss} can be readily calculated. Katzberg et al. [2013] developed a
- semi-empirical, one-to-one relationship between windspeed and mean squared slope by fitting data
- 128 provided by airborne GNSS-R experiments and an adjusted high resolution windspeed model. The

- Katzberg model is even simpler than the parametric wave models because it does not involve the wavespectrum. This relationship is expressed as follows
- 131

$$LP_{mss} = 0.45(0.00316f(U_{10}) + 0.00192f(U_{10}) + 0.003)$$

$$f(U_{10}) = U_{10}, \qquad 0 < U_{10} < 3.49 \ m/s$$

$$f(U_{10}) = 6\ln(U_{10}) - 4, \qquad 3.49 < U_{10} < 46$$

$$f(U_{10}) = 0.411U_{10}, \qquad 46 < U_{10}$$
(4)

132

134 where U_{10} is the windspeed at 10 m height. It is plotted in Figure 1.



- 136 Figure 1 The Katzberg U10-mss relationship
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- The end-to-end simulator for CYGNSS [Ruf et al., 2016a], which ingests windspeed and generates the
 delay-Doppler-map, currently uses the Katzberg relationship.
- 140141According to the Elfouhaily spectra shown in Figure 2, the long waves contribute a considerable portion
- 142 of the *LP*_{mss} sensitivity to wind. Such characteristics are similar to other spectra [e.g., Fig. 6 of Apel,
- 143 1994].
- 144





147 As noted above, the inclusion of a third-generation model, which focuses on the energy-containing long 148 waves, has not been necessary for other sensing techniques. Whether a model would benefit GNSS-R is 149 a question we explore in this work. As third-generation wave models have demonstrated considerable 150 skill in forecasting wave properties near the spectral peak (such as Hs and Tp) [e.g., The WAMDI Group, 151 1988; Ardhuin et al. 2010; Chu et al., 2004], we make use of this type of model in our work. In particular, 152 we select WAVEWATCH III [®] (WW3) as the low-frequency wave model, which is run operationally by the National Weather Service (NWS). The source terms of WW3 include wind input, dissipation, non-linear 153 154 interaction, bottom friction, ice scattering, among others. 155

156 Since our interest is in mss, we use the Ardhuin et al. [2010] source term package, which is the only reported source term package for WW3 validated for mss. Along with WW3, this package is open-source 157 158 to users in most countries. As mentioned earlier, the spectral tail of high frequency waves is not 159 completely resolved at the time of writing. All third-generation waves thus explicitly model the wave 160 spectrum only up to a certain frequency, and attach a high-frequency tail thereafter. We select a simple 161 k^{-3} spectral tail, which is suggested by the work of Banner et al. [1989], Forristall [1981], and Phillips [1958]; it was also used recently by Reichl et al. [2015] in a high-frequency model based on WW3. The 162 163 tail is attached at the last frequency modeled by WW3, and thus is completely determined by the value 164 of the spectrum at that frequency. A more elaborate model may include a high-frequency model like 165 that of Plant [2015], Hwang et al. [2013], or Elfouhaily [1997], but this option is not pursued here. 166 Our WW3 run is driven by the ECMWF operational wind analysis, and has 3-hour temporal output 167 resolution and 0.5-degrees latitude and longitude spatial resolution. The last wavenumber before 168 spectral tail attachment is 2.06 rad/m. The k^{-3} spectral tail ends at k_{μ} , which is determined by the 169

incidence angle of the track under consideration. For our simulation, the model is driven by wind only;
 currents play a minimal role globally [Bidlot, personal communication, 2016] - however, in hurricane

- 172 conditions, currents can have a significant role [Fan et al., 2009]. We limit ourselves to non-hurricane
- 173 conditions in this work, and thus neglect currents. In the following, we refer to the WW3 with spectral174 tail attached as the extended WW3 model.
- 175
- 176 An example of the attachment of the spectral tail is shown in Figure 3.
- 177



 178
 Wavenumber (rad/m)
 Wavenumber (rad/m)

 179
 Figure 3: Example of extended WW3 slope spectrum with a) linear scale (plot upper limit adjusted to 3 rad/m), b) area

 180
 conservative form.

181 2.2 GNSS-R Electromagnetic Scattering and Signal Processing Model

The ZV model developed by Zavorotny and Vornovich [2000] is a widely used scattering and signal 182 183 processing model for the GNSS-R received signal. This model is based on geometrical optics (GO) and is 184 valid for a sufficiently rough surface and non-grazing incidence. In practice, the ocean surface can be 185 considered sufficiently rough at wind speeds above about 3 m/s and non-grazing incidence angles are 186 those below about 70 degrees. The received signal is a function of delay and frequency. A two-187 dimensional plot of the signal power is known as a Delay Doppler Map (DDM) (explained in greater detail in Section 3). Because the ZV model connects the wave model and the observables and is 188 189 pertinent to our signal processing methods, we discuss it here in some detail. 190

191 The signal power intercepted by the receiver antenna can be expressed as

$$P_{s}(\tau, f) = CP_{t} \iint \frac{G_{t}G_{rant}}{R_{t}^{2}R_{r}^{2}} \chi^{2}(\Delta\tau, \Delta f)\sigma_{0}(\vec{s})dA$$
(5)

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195 where $P_s(\tau, f)$ is the signal power for delay τ and frequency f. C is a constant that depends on the 196 electromagnetic wavelength and coherent integration period of the receiver. P_t is the GPS transmitter 197 power and is assumed to be constant, as is G_t , the product of transmitter antenna and instrument gains. 198 G_{rant} is the receiver antenna gain. and R_t , R_r are the distances from the dummy integration position 199 on the grid to the transmitter and receiver, respectively. The surface integral is performed over an area 200 large enough for the desired τ and f ranges, and is known as the glistening zone. For us, τ ranges over 201 about 30 us and f ranges over about 10 kHz. The glistening zone is chosen to be 200 km by 200 km

centered at the specular point, which is sufficient for most scattering geometries of the TDS-1 202 203 instrument considered in Section 2.3.

204

205 $\Delta \tau = |\tau_e - \tau|$, with τ_e being the delay associated with the location of the differential surface element, dA. $\tau_e = (R_t + R_r)/c$ and for a given geometry, it is a constant for a given surface location, 206 independent of τ and f. Similarly, $\Delta f = f_g - f$, and $f_g = -f_{CW} / c(\overrightarrow{u_R} \cdot \overrightarrow{v_R} + \overrightarrow{u_T} \cdot \overrightarrow{v_T})$. f_{CW} is the 207 frequency of the carrier wave; for the GPS L1 carrier, it is 1.575 GHz. $\vec{u_R}$ is the unit vector from the 208 specular point to the receiver, $\vec{v_R}$ is the receiver velocity vector, $\vec{u_T}$ is the unit vector from the specular 209 point to the transmitter, and $\overrightarrow{v_T}$ is the transmitter velocity vector. 210 211 $\chi^2(\Delta\tau,\Delta f)$ is known as the ambiguity function and models the selectivity of the radar system. Letting 212 au_0 and f_0 be the delay and Doppler shift corresponding to the specular point, respectively, if the 213 selectivity is sufficiently high such that $\frac{G_t G_{rant}}{R_r^2 R_r^2} \sigma_0$ is constant for some small area ΔA around the 214 specular point, then, because $\chi^2(0,0)=1$ 215 216 217 $P_{s0} = P_S(\tau_0, f_0) = CP_t \frac{G_t G_{rant}}{R_t^2 R_r^2} \sigma_0 \Delta A$ 218 (6) 219 We make use of this equation in Section 2.3. 220 221 Similar to τ_g , f_g , G_t , G_{rant} , R_t , and R_r , \vec{s} is also a constant for a given location (independent of τ , f) 222 - it specifies the favorable orientation (two perpendicular slope components) of a facet that reflects the 223 incident ray toward the receiver. The scattering cross section $\sigma_0(\vec{s})$ is where the surface roughness 224 enters – under geometric optics, σ_0 is proportional to the PDF of slopes as well as the square of the 225 226 Fresnel surface reflectivity. The PDF of slopes and its measurement remain an active area of research [e.g., Cardellach and Rius, 2008; Liu et al., 1997]. To a first order, the PDF of slopes can be approximated 227 228 by a bivariate Gaussian 229 230

231

$$pdf(s_u, s_c) = \frac{1}{2\pi\sqrt{mss_u mss_c}} \exp\left\{-\frac{1}{2}\left(\frac{s_u^2}{mss_u} + \frac{s_c^2}{mss_c}\right)\right\}$$
(7)

232 where the subscripts u and c denote the upwind and crosswind components. This assumption is also used in Zavorotny and Voronovich [2000]. At the specular point, pdf(0,0) is proportional to the inverse 233 of the geometric mean of the mss components. It should be noted that more complex PDFs have also 234 been considered. 235

236 In this work, we further assume that the seas are isotropic and the two components of mss are equal.

- Equivalently, the two-dimensional PDF is rotation-invariant in the sense that it only depends on the 237
- magnitude of \vec{s} . Note that the mss is obtained from WW3 using equation (2). 238

240	2.3 Model Configuration and Post-Processing for TDS-1
241	TechDemoSat-1 (TDS-1) is a technology demonstration mission operated by Surrey Satellite Technology
242	Limited (SSTL) [Unwin et al., 2016]. One of its payloads is the Space GNSS Receiver Remote Sensing
243	Instrument (SGR-ReSI), the GNSS-R instrument of interest. TDS-1 has a circular orbit with an altitude of
244	about 630 km. Because there are other instruments on the TDS-1 mission, the SGR-ReSI has limited
245	operating time, so the data it collects are limited. In this paper, all references to TDS measurements
246	refer to data collected by the SGR-ReSI.
247	
248	From Section 2.2, several pieces of information are required to compute the received signal. The GPS
249	transmitted power is not published, so it is assumed to be constant. The other parameters needed are:
250	- transmitter position and velocity.
251	- receiver position and velocity.
252	- receiver antenna and instrument (RF and IF) gains, and
253	- mss.
254	
255	All these quantities are functions of time. For a moving receiver, the specular point traces out a
256	trajectory in time across the ground known as a track. In this work, two surface models are used for
257	computing the mss: the Katzberg model and the extended WW3 model. The GPS transmitter and TDS
258	receiver positions and velocities, along with the TDS receive antenna pattern, are furnished by SSTL.
259	However, the instrument gain is not available; in fact, the receiver has automatic gain control (AGC)
260	turned on, so the instrument gain changes with signal level, and this time-varying gain is not recorded.
261	We therefore process the DDMs in a way that is not sensitive to the gain value, by forming the ratio
262	between their signal and noise regions. The resulting DDMs are of relative received power, normalized
263	by the noise floor of the measurements They are still sufficiently sensitive to changes in the surface
264	conditions, provided variations in the receiver noise floor are small enough over relevant time scales.
265	
266	The glistening zone is set to 200 km by 200 km. This determines the surface area over which the
267	numerical integration is taken in the model. The wind and mss are assumed to be constant over the area
268	of integration.
269	
270	In addition to the contribution to the received signal power by scattering from the ocean surface, $P_{ m s}$ in
271	Equation (5) also contains, other components due to radiometric thermal emission by the scene, noise
272	due to the receiver instrumentation (including the antenna), and radio-frequency interference (REI) [e.g.,
273	Chen et al. 2015]. We neglect RFI in this paper. The total received signal (in uncalibrated units of counts)
274	can then be modelled as
275	
276	$C(\tau, f) = G(P + P) \tag{8}$
270	$C_T(v, j) = O_{ri}(V_N + V_s) $
277	
278	where G_{ri} is the receiver instrument gain (excluding the antenna gain) and P_N is the total noise power.
279	$P_{\!\scriptscriptstyle N}$ includes the radiometric thermal emission from the scene referred to the output of the antenna and
280	the noise due to receiver instrumentation. $P_{\scriptscriptstyle s}$ is the GNSS-R signal power, given by the ZV model in
281	Equation (5). To be precise, $P_{\!s}$ is the ensemble mean of the signal power. In practice, there will also be
282	speckle noise present in the measurements. Our model neglects the speckle noise and estimates the
283	ensemble mean.

- 285 The noise contributions to the measurements are estimated by examining pixels of the DDM at delay 286 values that correspond to altitudes higher than the surface. As such, these pixels contain no scattered 287 surface signal and P_s=0 can be assumed. In that case, the uncalibrated measurements can be written as 288 $C_N = G_{ri}(P_N)$ 289 (9) 290 The pixel in the DDM with the highest power is assumed to correspond to the specular point location. 291 This is only approximate, as the peak power originates from a region near but not necessarily at the 292 specular point. With two equations $C_T(\tau_0, f_0) = G_{ri}(P_N + P_s(\tau_0, f_0))$ and $C_N = G_{ri}(P_N)$, we cannot 293 completely resolve the three unknowns, $P_s(\tau_0, f_0)$, G_{ri} , and P_N . It should be noted that the upcoming 294 295 CYGNSS mission carries an augmented version of the receiver that incorporates calibration targets and 296 fixed receiver gain, so these unknowns can be determined. For TDS, no absolute calibration can be 297 performed and the DDMA observable [Clarizia et al., 2014] is not easily computed. (An observable is a 298 single number characterization of the DDM.) 299 300 Because of this, a proxy for the DDMA, known as the SNR [Jales, 2015], is now being used in the TDS community. It is defined by 301 302 $SNR = \frac{C_T(\tau_o, f_0) - C_N}{C_N} = \frac{G_{ri}(P_N + P_S) - G_{ri}(P_N)}{G_{ri}(P_N)} = \frac{G_{ri}(P_S)}{G_{ri}(P_N)} = \frac{P_S}{P_N}$ 303 (10)304 305 We see that the SNR observable is independent of gain as desired, but depends on the noise power. 306 Gain varies much faster than the noise power – the dominant factor is changes due to instrument 307 temperature and AGC adjustments. 308 For our simulations, we only model P_s and do not model the thermal noise. To estimate P_N , we 309 compute the ratio between the measured SNR and the modeled P_s over an entire track. Thus: 310 311 $P_N = \frac{E[P_s^{sum}]}{E[SNR^{TDS}]}$ 312 (11)313 where E[.] is the time average operator, P_s^{sim} is the simulated signal power, and SNR^{TDS} is the TDS-314 measured SNR. This assumes that P_N is constant over the track, and there are no biases to P_s . With P_N 315 known, the simulated SNR can then be computed. 316 317 318 The computation of modelled SNR requires the extraction of a single parameter from the measurements. 319 Note that there are other observables that could alternately be used, such as the DDM volume 320 observable [Marchan-Hernandez et al, 2008]. This observable, fundamentally, makes use of the ratio of 321 the signal powers from DDM bins far away from the specular point to those at or near the specular point. We have considered this observable in our analysis and the results are similar in character to those using 322 323 the SNR presented in Section 3, but they are found to exhibit a large noise level than the SNR observable.
- For this reason, we will use an SNR-related observable in the following discussions.
- 325

- To focus on the effects of sea state, we define the Scaled SNR as: 326
- 327

$$SSNR = SNR \frac{R_{tSP}^2 R_{rSP}^2}{G_{rant}}$$
(12)

5

329

330 where R_{rSP} and R_{rSP} are the distances from the specular point to the transmitter and receiver,

- 331 respectively.
- 332

333 We neglect scaling corrections for scattering area and incidence angle-dependent Fresnel reflectivity for 334 simplicity and because the measurement geometries present in the TDS sample population do not 335 exhibit significant variations.

336

337 Lastly, we note that for a given geometry, higher mss values (greater roughness) correspond to smaller 338 SNR values.

3 Results and Discussion 339

340 We analyze one TDS track in this work: Track 407 in RD 17 of SSTL's Version 0.3 dataset. This track

341 contains about 16 minutes of continuous data, collected by a single receiver channel and a single GPS

transmitter (GPS PRN #10 and Receiver Channel #2, per SSTL's numbering conventions). One DDM is 342

343 produced every second. This track exhibits a good variation of coastal and oceanic conditions, as well as

344 a variety of sea states. The track of the specular point is plotted in Figure 4.

345



346 347 Figure 4: The specular point track for RD17 TR407 is shown as the bold yellow line running from Antarctica into the South Pacific. 348 The data is numbered from 1 to 1007, which we call sample number (SN). The transition from land to ocean occurs at SN 264.

349 The receive antenna gain along the track is plotted in Figure 5. The variation in gain results from the

350 progression of transmitter and receiver locations, and the resulting change in measurement geometry, 351 over time.



 352
 Sample Number (SN)

 353
 Figure 5: Along-track antenna gain for RD17 TR407 for specular points in the ocean; the SN ranges from 264 to 1007.

354 Because antenna gain can affect the signal quality, we consider only measurements with gain greater 355 than 3 dB in our analysis. In addition, since our interest is in ocean GNSS-R, we filter out any data with its 356 specular point located less than 100 km away from the coast. The resulting dataset has sample numbers 357 ranging from 293 to 1007. This is the rationale for restricting the grid size to be 200 km by 200 km as 358 mentioned in Section 2. Relative to the specular point, delay and Doppler bins with less than 18 us and 359 5000 Hz in either direction are considered, and this is the range plotted in the DDMs shown below. The average incidence angle for the track under consideration is 13.8 degrees, which results in a cut-off 360 wavenumber, k_u , of 10.59 rad/m or 59 cm in wavelength. 361

362

363 3.1 Empirical Evidence of Measurement Sensitivity to Significant Wave Height

As seen from Equations (2) and (3), significant wave height, Hs, is much more sensitive to long waves than the mean square slope. These long waves include swell that is not correlated with wind. In this subsection, we explore the dependence of SSNR (and thus mss) on Hs using TDS measurements.

In Figure 6, U₁₀ and Hs are plotted against sample number (SN) for Track 407. Each SN is separated by approximately one second, and, for this track, the specular points of two consecutive measurements are spaced about 6000 m apart. Hs is obtained from spatial interpolation of the same WW3 model run, as WW3 is skillful in modelling Hs. U₁₀ comes from the same ECMWF wind reanalysis product that is used to force the WW3 model.





373 Sample Number (SN)
374 Figure 6: Along track U10 (left axis) and Hs (right axis). The green band denotes a narrow range of U10 values, the relevance of
375 which is discussed in the text.

- 376 Although U₁₀ exhibits some correlation with Hs, there are many points where they deviate from one
- another. To control for U₁₀ and examine the variance of the SSNR explained by Hs alone, we restrict our
- analysis to measurements for which U₁₀ lies in the narrow range between 5.7 and 6.2 m/s. This region is
- 379 shaded by a horizontal green band in Figure 6. A scatterplot of the measured SSNR vs. Hs values in this
- 380 region is shown in Figure 7.
- 381
- 382





- 386 Hs is seen to have a strong effect on SSNR that cannot be accounted for solely by windspeed. This
- behavior has been noted previously [Soisuvarn et al., 2016]. Some scatter is also seen, indicating that
- 388 SSNR has additional variability explained by neither Hs nor windspeed. In the figure, we have picked
- three representative measurements; these are circled in red with their SNs indicated. We examine their
- 390 DDMs in this and the next subsections.
- 391

392 The three DDMs measured by TDS are presented in Figure 8. Both the magnitude and shape of the 393 DDMs change significantly. The magnitude decreases monotonically as Hs increases, which is consistent 394 with theoretical expectations. The mss corresponding to each of the DDMs can be estimated using 395 either the Katzberg or WW3 model. In the case of Katzberg, all three wind speeds are nearly the same, 396 so the mss is, too. It is 0.0172. With the WW3 model, mss is not solely dependent on wind speed and 397 the mss is found to be 0.00028, 0.00063, and 0.0122 for SN 293, 301, and 386, respectively. The 398 significant differences in mss with the WW3 model are due to other influences on the local sea state 399 than simply the wind speed there. In particular, note that the significant wave height varies significantly between the three cases.

400 401

402



Figure 8: TDS Measurements. Top panel: SN 293 with Hs=1.34 m and U10=6.19 m/s. Middle panel: SN 301 with Hs=1.57 m and 407 U10=6.18 m/s. Bottom panel: SN 386 with Hs=2.13 m and U10=5.70 m/s.

Modeling the Effect of Significant Wave Height on the Measurement 408 3.2

- In this subsection, we examine modelled results and compare them to the measurements in the 409
- previous subsection. First, we look at the modelled DDMs of the three cases considered. Second, we 410
- 411 look at the dependence on Hs predicted by the models. Lastly, we look at the along-track plots of the
- SSNR. 412
- 413
- 414 Because the windspeed is essentially the same in all three cases, the Katzberg DDMs should all look
- 415 about the same. This is indeed the case, as seen the modelled DDMs in Figures 9-11 (right panels). The

416 left panels show the results of the extended WW3 model. Comparing to the TDS measurements in

417 Figure 8, it is seen that the WW3-based model is much better able to represent the behavior of the

418 measurements, compared to the Katzberg model, in both the magnitude and shape of the DDMs.

419



420 421









431 Doppler (Hz) Doppler (Hz)
 432 Figure 11: DDMs predicted by the two forward models: WW3 (left) and Katzberg (right) given ocean conditions Hs=2.13 m and
 433 U10=5.70 m/s consistent with observation SN 386. Compare to the bottom panel in Figure 8. Both models are both in good
 434 agreement with the observations.

+5+ ugreement with the observations.

- 435 We now plot modelled SSNR vs. Hs in Figure 12. These plots reaffirm WW3's skill over the Katzberg
- 436 model. In particular, significant improvement is seen for low Hs values; these were found to occur at the
- 437 beginning of the track near the coast. In addition, the Katzberg model demonstrates deficiencies in the
- 438 "branch" near Hs=3 m and SSNR=3e25; these correspond to very low windspeeds of less than 3 m/s.







Figure 12: SSNR vs. SWH, with U10 colorcoded. The figures in the right column are zoomed in versions of the ones on the left.a)
TDS – these two plots are characteristically the same as Figure 7, but no filtering is done based on U10 b) Extended WW3 c)
Katzberg. Because of the inverse dependence of mss, SNR is much more sensitive to mss changes when mss is small.

449 To gain additional insight, we plot the SSNR vs. along-track SN for the TDS measurements and both 450 models in Figure 13. This figure should be used in conjunction with Figure 6, which shows the along-451 track U₁₀ and Hs. Using the variance of the difference between simulations and measurements as the 452 metric, the extended WW3 model shows a 68.7% improvement over the Katzberg model over the entire 453 track. The improvements in the coastal region at the start of the track is one significant contributor. If we consider only SN 342 and higher, we still see a 30.2% improvement in the skill of the extended WW3 454 455 model. This improvement can largely be attributed to the SNs 850 to 900, for which the windspeed is 456 very low. 457 458

459





Figure 13. a. Along track plot of measured SSNR, extended WW3 SSNR, Katzberg SSNR, and scaled Hs and U10. b. Zoomed in version of a.

- Another insight is that despite the attachment of a diagnostic tail, we see WW3 is also responsive to
- local wind: at SNs from about 900 to 950, Hs is decreasing but windspeed is increasing (see Figure 6).
- WW3 is able to model the decreasing behavior of the observable correctly.

- 472 This analysis shows that the extended WW3 model has considerable skill modelling the GNSS-R
- observable, derived from its ability to take non-local long waves into account, and in modelling the sea
- 474 state in low windspeed conditions.
- 475
- 476 One implication of our results is that much of the sensitivity of the GNSS-R observable to the sea state 477 derives from long and intermediate-scale waves of wavenumber 2 rad/m and lower. This is consistent 478 with predictions of the parametric Elfouhaily model shown in Figure 2. However, it should be noted that 479 this track does not contain winds that change quickly in time. A track with rapid changes in wind 480 temporally and spatially will be able to better evaluate whether the diagnostic tail should be replaced 481 one that has an explicit wind speed dependence. Fast changes in wind may also necessitate that the 482 model be run at a higher spatial and temporal resolution with the corresponding wind speed products. 483 484 Lastly, we note that both models show overly low SSNRs between serial numbers 350 and 500, while a
- 485 slightly positive bias is seen between 600 and 850. These discrepancies can be the result of an overall,
- 486 constant bias that is not removed before determining and applying the SSNR scale factor in Equation
- 487 (11). Such a bias may be due to errors in the cutoff k_{μ} , or the spectral level. This bias may also
- 488 contribute to the difference in shapes of the measured and WW3 DDMs shown in Figure 8 and Figure 10.

489 4 Conclusions and Future Work

- 490 In this work, we have developed a GNSS-R forward model that incorporates a third-generation surface 491 wave model. The analysis of one track of TDS measurements, with over 700 consecutive DDMs, shows 492 that this model can account for observable dependencies on the local wind as well as other, non-local 493 effects. In contrast to conventional remote sensing techniques, the non-local effects are significant for 494 GNSS-R due to frequency and geometry. The model demonstrates improved skill over the widely used 495 Katzberg one-to-one windspeed-mss model. Significant improvements are seen in low wind conditions, 496 in particular. The novelty and strength of the model is derived from the WW3 model, the source terms 497 of which are the result of decades of work by the wave modeling, experimental, and remote sensing 498 communities. Conversely, given the demonstrated sensitivity of GNSS-R to ocean surface wave spectra, 499 the assimilation of its measurements into numerical wave models may also provide valuable constraints on the derived sea state. 500 501 502
- The use of a third generation wave model in GNSS-R forward modelling has great potential for futurework. Some ideas include:
- Retrieval of mss from the measurements, and taking into account scattering area and Fresnel
 reflectivity,
- 506 Modelling of anisotropic seas with two mss components and a more sophisticated pdf of slopes, in 507 effect creating a tighter coupling between the scattering and wave models,
- Relaxation of assumption of uniformity of wind and mss fields over the 200 km by 200 km glistening
 zone,
- Augmenting the scattering model by taking the coherent scattering component into account for lowwindspeeds,
- 512 Usage of CYGNSS data when it becomes available; with absolute calibration, better signal quality can
 513 be achieved, and
- 514 Addition of wave-current interactions in the wave model.
- 515
- 516 To understand the underlying physical phenomena modelled by WW3 that allows it to produce better
- 517 long-wave mss, it would be helpful to examine the two-dimensional wave spectra, as well as the source

- term spectra. Those insights may lead to the development of ancillary parameters that could be helpful
- 519 in constructing better wind retrieval algorithms for GNSS-R.
- 520

521 The model presented here can also be used to improve our understanding of surface waves with GNSS-R

522 measurements. Possibilities include the tuning of the spectral tail and development of appropriate

523 source terms. The model is also expected to be helpful in the design of future GNSS-R missions and

524 experiments.

525 5 Acknowledgements

526

527 This work was funded in part by the National Aeronautics and Space Administration (grant NNX13AP93H 528 and contract NNL13AQ00C). TechDemoSat-1 data used in this study are available from the publically

and contract NNL13AQ00C). TechDemoSat-1 data used in this study are available from the publically
 accessible MERRByS web site at http://www.merrbys.co.uk. Wavewatch III data used in this study are

- 530 available from the publically accessible NOAA web site at
- 531 http://polar.ncep.noaa.gov/waves/ensemble/download.shtml.

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Figure 1.



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Figure 2.



Figure 3.



Figure 4.



Figure 5.



Figure 6.



Figure 7.



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Figure 8.









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Figure 10.





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Figure 11.





Figure 12.



Figure 13.



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