

# The DOE E3SM Model Version 2: Overview of the physical model and initial model evaluation

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## Key Points:

- E3SMv2 is nearly twice as fast as E3SMv1 with a simulated climate that is improved in many metrics (e.g. precipitation and clouds).
- Climate sensitivity is substantially lower with a more plausible ECS of 4.0 K (compared to an unlikely value of 5.3 K in E3SMv1).
- E3SMv2 underestimates the warming in the late historical period due to excessive aerosol-related forcing.

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## Abstract

This work documents version two of the Department of Energy’s Energy Exascale Earth System Model (E3SM). E3SMv2 is a significant evolution from its predecessor E3SMv1, resulting in a model that is nearly twice as fast and with a simulated climate that is improved in many metrics. We describe the physical climate model in its lower horizontal resolution configuration consisting of 110-km atmosphere, 165-km land, 0.5° river routing model, and an ocean and sea ice with mesh spacing varying between 60-km in the mid-latitudes and 30-km at the equator and poles. The model performance is evaluated with Coupled Model Intercomparison Project Phase 6 (CMIP6) Diagnosis, Evaluation, and Characterization of Klima (DECK) simulations augmented with historical simulations as well as simulations to evaluate impacts of different forcing agents.

The simulated climate has many realistic features of the climate system, with notable improvements in clouds and precipitation compared to E3SMv1. E3SMv1 suffered from an excessively high equilibrium climate sensitivity (ECS) of 5.3 K. In E3SMv2, ECS is reduced to 4.0 K which is now within the plausible range based on a recent World Climate Research Programme (WCRP) assessment. However, a number of important biases remain including a weak Atlantic Meridional Overturning Circulation, deficiencies in the characteristics and spectral distribution of tropical atmospheric variability, and a significant underestimation of the observed warming in the second half of the historical period. An analysis of single-forcing simulations indicates that correcting the historical temperature bias would require a substantial reduction in the magnitude of the aerosol-related forcing.

## Plain Language Summary

The U.S. Department of Energy recently released version two of its Energy Exascale Earth System Model (E3SM). E3SMv2 experienced a significant evolution in many of its model components (most notably the atmosphere and sea ice models), and its supporting software infrastructure. In this work, we document the computational performance of E3SMv2 and analyze its ability to reproduce the observed climate. To accomplish this, we utilize the standard Diagnosis and Evaluation and Characterization of Klima (DECK) experiments augmented with historical simulations for the period 1850-2015. We find that E3SMv2 is nearly twice as fast as its predecessor and more accurately reproduces the observed climate in a number of metrics, most notably clouds and precipitation. We also find that the model’s simulated response to increasing carbon dioxide (the Equilibrium Climate Sensitivity) is much more realistic. Unfortunately, E3SMv2 underestimates the global mean surface temperature compared to observations during the second half of historical period. Using sensitivity experiments, where forcing agents (carbon dioxide, aerosols) are selectively disabled in the model, we determine that correcting this problem would require a strong reduction in the impact of aerosols.

## 1 Introduction

The U.S. Department of Energy (DOE) Energy Exascale Earth System Model (E3SM) project (<https://e3sm.org>) was conceived from the confluence of energy mission needs and disruptive changes in scientific computing technology. E3SM aims to optimize the use of DOE resources to meet the science needs of DOE. Efficient utilization of emerging computational architectures requires a significant evolution in present programming models in Earth System Models (ESMs), leading DOE to develop a new ESM, initially branching from CESM1 (Community Earth System Model; Hurrell et al., 2013, <http://www.cesm.ucar.edu/models/cesm1.0>). The long-term goal of the E3SM project is to produce robust actionable predictions of Earth system variability and change, with an

96 emphasis on the most critical scientific questions facing the nation and DOE (Leung et  
97 al., 2020).

98 Version one of E3SM (E3SMv1) was first released in 2018 as a physical climate model  
99 with a lower horizontal resolution configuration (110-km atmosphere, 60-to-30 km ocean;  
100 Golaz et al., 2019) followed by a higher resolution configuration (25-km atmosphere, 18-  
101 to-6 km ocean; Caldwell et al., 2019). The lower resolution configuration served as the  
102 starting point for a biogeochemistry configuration (E3SMv1.1; Burrows et al., 2020) and  
103 a cryosphere configuration (E3SMv1.2; Comeau et al., 2022).

104 Although E3SM was originally branched from CESM1, the river routing, ocean,  
105 and sea ice components as well as the atmosphere dynamical core and stratospheric chem-  
106 istry are completely new or significantly different. The atmosphere physics, the land model,  
107 and the coupler retain similarities to current CESM2 (Danabasoglu et al., 2020). E3SMv2  
108 is the second release of a CMIP6-class (Coupled Model Intercomparison Project Phase  
109 6) model for E3SM. E3SMv2 also serves as a foundation for additional upcoming con-  
110 figurations targeting DOE applications: (i) a regionally refined mesh (RRM) configura-  
111 tion with a high resolution region (25-km atmosphere, 14-km ocean) centered over North  
112 America, (ii) a biogeochemistry configuration with interactive carbon, nitrogen and phos-  
113 phorous cycles, and (iii) a cryosphere configuration with RRM over the Southern Ocean  
114 and ice-shelf cavities.

115 Version two of E3SM is a significant evolution from version one. Herein we describe  
116 the changes made in E3SM version 2 (E3SMv2) in each model component and the sup-  
117 porting infrastructure. We further diagnose its performance relative to E3SMv1. E3SMv2  
118 development focused on improving the performance on existing and emerging architec-  
119 tures and improving the physical climate in key metrics. Thus, E3SMv2 includes signif-  
120 icant improvements to component model structure and physical parameterizations. The  
121 result of v2 development is a model that is nearly twice as fast as version one with a sim-  
122 ulated climate that is improved in many metrics. Also new to E3SMv2 is the introduc-  
123 tion of fully coupled RRM configurations, a critical capability to creating actionable pro-  
124 jections of interest to the nation and DOE. Although simulations with the RRM will be  
125 the subject of forthcoming manuscripts, the validation herein will provide a benchmark  
126 for RRM configurations.

127 As with E3SMv1, we focus on the physical climate model at lower resolution with  
128 a 110 km atmosphere, 165 km land,  $0.5^\circ$  river routing model, and an ocean and sea ice  
129 with mesh spacing varying between 60 km in the mid-latitudes and 30 km at the equa-  
130 tor and poles. The vertical grids remain the same as in E3SMv1 with 72 layers and a  
131 top at approximately 60 km in the atmosphere and 60 layers (10 m near-surface reso-  
132 lution) in the ocean. We focus our analysis on the CMIP6 Diagnosis, Evaluation, and  
133 Characterization of Klima (DECK) and historical simulations (Eyring et al., 2016). E3SMv2  
134 DECK simulations reveal a number of improvements in the simulated mean climate and  
135 variability: equilibrium climate sensitivity, precipitation, shortwave cloud radiative ef-  
136 fects, ozone hole, aerosol absorption, and sea ice. Yet despite numerous improvements,  
137 a number of important biases remain including a weak Atlantic Meridional Overturning  
138 Circulation and an inability to appropriately simulate the historical temperature record.  
139 Mitigating these biases will be central to E3SMv3 development. To diagnose the latter  
140 bias we conduct an ensemble of simulations following the Detection and Attribution Model  
141 Intercomparison Project (DAMIP) protocol (Gillett et al., 2016). Using a decomposi-  
142 tion analysis, we find that an overly strong aerosol effect is responsible for this bias and  
143 further that if this effect can be reduced, other reductions in regional radiation, temper-  
144 ature, and other biases can be expected.

145 We begin in Section 2 with a description of the changes in E3SMv2 for each model  
146 component. We also describe important improvements to energy conservation in the cou-  
147 pled system as well as our coupled tuning strategy. Section 3 details computational per-

148 performance and factors leading to the nearly doubling of throughput. Section 4 details the  
 149 simulation campaign and analysis of the simulated climate in each portion of the cam-  
 150 paign. Section 5 presents an examination of the historical temperature record bias and  
 151 the potential impact of altering the contribution of aerosols and greenhouse gases on the  
 152 simulated climate. We end with summary and conclusions in Section 6.

## 153 2 Model description

### 154 2.1 Atmosphere

#### 155 2.1.1 Dynamical core

156 The dynamical core in the E3SM Atmosphere Model (EAM) v2 solves the equa-  
 157 tions of motion in a rotating reference frame with the hydrostatic and shallow atmosphere  
 158 approximations and a hyperviscosity based turbulence closure. It is implemented in the  
 159 High Order Method Modeling Environment (HOMME) (Dennis et al., 2005, 2011; Evans  
 160 et al., 2013), which provides meshing infrastructure and discretization operators for spec-  
 161 tral elements in spherical geometry. The equations are formulated following Taylor et  
 162 al. (2020) using a terrain following mass based vertical coordinate (Kasahara, 1974; Laprise,  
 163 1992). In EAMv2, the nonhydrostatic formulation given in Taylor et al. (2020) is mod-  
 164 ified to be hydrostatic. This modification is straightforward due to the use of the (Laprise,  
 165 1992) mass coordinate. The prognostic equations consist of the time-reversible adiabatic  
 166 terms, a  $\nabla^4$  hyperviscosity (Dennis et al., 2011; Guba et al., 2014), and a sponge layer  
 167 at the model top (described below). The discretization of the adiabatic terms is struc-  
 168 ture preserving, meaning that the discrete equations are in quasi-Hamiltonian form, lead-  
 169 ing to energetic consistency in the sense of (Gassmann & Herzog, 2008). Energetic con-  
 170 sistency is obtained via a term-by-term balance in the discrete kinetic, internal, and po-  
 171 tential energy budgets, ensuring an energy conserving discretization with no spurious en-  
 172 ergy sources or sinks.

173 The horizontal discretization uses the collocated mimetic spectral finite element  
 174 method from Taylor and Fournier (2010). Within each element the prognostic variables  
 175 are represented by degree  $p$  polynomials with  $p = 3$  and order of accuracy  $n_p = 4$ . The  
 176 vertical discretization uses the Lorenz staggered mimetic centered difference from Simmons  
 177 and Burridge (1981). The vertical grid remains the same as in EAMv1 with 72 layers  
 178 and a top at approximately 60 km. For the vertical transport terms, we use a vertically  
 179 Lagrangian approach adapted from Lin (2004). The timestepping algorithm, unchanged  
 180 from EAMv1, is the high-CFL, 5-stage, third-order-accurate Runge-Kutta method from  
 181 (Guerra & Ullrich, 2016).

182 There are several sources of dissipation in the dynamical core. The  $\nabla^4$  hypervis-  
 183 cosity is the largest. It is applied to all prognostic variables and on every model layer.  
 184 For the model-top sponge layer, we apply a  $\nabla^2$  Laplacian operator in the top 6 model  
 185 layers to all prognostic variables. The strength is proportional to the model layer ref-  
 186 erence pressure, following Lauritzen et al. (2011). In addition, vertical dissipation is in-  
 187 troduced by the monotone vertical remap operator. A smaller amount of dissipation is  
 188 also generated by the Runge-Kutta timestepping. In EAMv1, we used additional diver-  
 189 gence damping in order to control noise when running with realistic topography. This  
 190 was implemented by separating the hyperviscosity into compressible and rotational com-  
 191 ponents and using a larger hyperviscosity coefficient for the compressible component. EAMv2  
 192 has a more accurate pressure gradient formulation which improves the treatment of to-  
 193 pography and no longer needs nor uses additional divergence damping.

194 The dynamical core’s passive tracer transport method is a new interpolation semi-  
 195 Lagrangian (ISL) scheme called Islet (Bradley et al., 2021). A high-order ISL method  
 196 using the natural Gauss-Lobatto-Legendre (GLL) element-local interpolant is unstable;  
 197 thus, Islet provides modified element-local interpolation basis functions that obey a nec-

198 necessary condition for stability. EAMv2 uses the lowest-order Islet basis set, the one for  
 199  $n_p = 4$ . Because the model code was frozen before the Islet bases were finalized, the  
 200 formulation of the  $n_p = 4$  stable basis set is slightly different than reported in Bradley  
 201 et al. (2021), but this difference has essentially no impact. To achieve global mass con-  
 202 servation, shape preservation, and mass-tracer consistency, Islet uses element-local and  
 203 global versions of the communication-efficient density reconstructor (CEDR) described  
 204 in Algorithm 3.1 of Bradley et al. (2019). The global version is sometimes called a “mass  
 205 fixer,” but note that, in combination with the element-local version, it also enforces grid-  
 206 point-local, time-dependent lower and upper bounds on tracer mixing ratios to enforce  
 207 shape preservation and mass-tracer consistency. Generally, even a mass-conserving semi-  
 208 Lagrangian method that can take time steps longer than the advective CFL number re-  
 209 quires a method like a CEDR to achieve shape preservation and mass-tracer consistency  
 210 when coupled to a dynamical core that uses a different discretization for the dynamical  
 211 equations. The ISL scheme’s time step can be, and in EAMv2 is, longer than the ver-  
 212 tical remap time step of the dynamics. In integrating from time  $t_1$  to time  $t_2$ , Lagrangian  
 213 levels at time  $t_2$  are reconstructed from data on the reference grid at times  $t_1$  and  $t_2$ . Then  
 214 horizontal velocity at time  $t_2$  is remapped to the Lagrangian levels. Finally, departure  
 215 points within each Lagrangian level are computed at time  $t_1$ . Then 2D advection within  
 216 each level can proceed as usual. In this time step configuration, the CEDR must be ap-  
 217 plied to the 3D data rather than separately to each level because the reconstructed lev-  
 218 els do not conserve mass within each level; thus, corrections must be applied among lev-  
 219 els as well as within each level. In EAMv2’s lower resolution configuration, the vertical  
 220 remap time step is two times larger than the dynamics time step, and the passive tracer  
 221 transport time step is six times larger. Like the rest of the dynamical core, Islet works  
 222 without modification in RRM configurations.

### 223 *2.1.2 Separate dynamics and column parameterizations grids*

224 New in version 2, EAMv2 uses separate grids for dynamics and column parame-  
 225 terizations. The dynamics grid, described in Section 2.1.1, is the same as used in EAMv1.  
 226 For brevity, it is referred to as the “np4 grid”, following the dynamical core’s grid nam-  
 227 ing convention, because each element has a  $4 \times 4$  subgrid of grid points. The column pa-  
 228 rameterizations grid shares the element grid with the dynamics but has a  $2 \times 2$  subgrid  
 229 of quadrilaterals for a total of four columns per element. This grid is referred to as the  
 230 “pg2 grid”, where “pg” refers to what is usually called the “physics grid” but, more gen-  
 231 erally, could be called the “parameterizations grid”, and “2” refers to the  $2 \times 2$  subgrid.  
 232 Thus, the total number of physics columns in a simulation is  $4/9$  the number used in EAMv1  
 233 for a given element grid, where 9 rather than 16 results from sharing of element edge grid  
 234 points among adjacent elements. The dynamics grid has an average grid spacing of 110  
 235 km, while the parameterizations grid and, as a result, the land grid have an average grid  
 236 spacing of 165 km. Hannah et al. (2021) describe the remap algorithms to transfer data  
 237 between the grids and the new topography file format to support these grids. The grids  
 238 are the same as introduced in (Herrington et al., 2019), but in EAMv2, the high-order  
 239 remap method is local to each element except for some halo data for extremal mixing  
 240 ratio values. Thus, EAMv2’s grid remap algorithms work without modification in RRM  
 241 configurations. The motivation for the separate grids is to speed up the physics param-  
 242 eterization computations by up to nearly 2.25 times with little change to the climatol-  
 243 ogy (Hannah et al., 2021). Hannah et al. (2021) discuss the effective resolution differ-  
 244 ences between simulations using just the np4 grid and those using separate np4 and pg2  
 245 grids, where effective resolution is inferred from power spectra. Briefly, first, the dynam-  
 246 ical core’s effective resolution is nearly identical; second, the parameterizations’ effective  
 247 resolution is, as expected, lower when using the pg2 grid. Given the first point, the re-  
 248 sult is a more efficient simulation.

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### 2.1.3 Updated atmosphere physics

As in EAMv1 (Rasch et al., 2019; Xie et al., 2018), EAMv2 represents subgrid turbulent transport and cloud macrophysics by use of the Cloud Layers Unified By Binormals (CLUBB) parameterization (Golaz et al., 2002; V. E. Larson, 2017). In EAMv2, CLUBB represents all stratiform and shallow cumulus clouds, but not deep convective clouds. CLUBB prognoses various subgrid moments of turbulence, heat content, and moisture, and the moments are used to estimate a multivariate subgrid probability density function (PDF). The PDF is then used to diagnose liquid cloud fraction and cloud liquid water via a saturation adjustment. CLUBB is called immediately before the microphysics.

The main update of CLUBB for EAMv2 is that CLUBB’s internal call order has been changed so that CLUBB’s subgrid moments are prognosed first, and the PDF is estimated immediately afterward. This leaves a state for the microphysics that is adjusted with respect to liquid saturation. This call order eliminates the unrealistic pockets of supersaturation that were left for the microphysics to handle in EAMv1. Another update of CLUBB is that its code has been refactored in order to improve computational performance. For instance, arrays were restructured to permit contiguous memory access. Loops were rearranged in order to allow calculations with no data dependencies to be done in parallel. Asymptotic values of functions were approximated analytically in order to avoid the unnecessary calculation of expensive special functions.

The deep convection scheme (G. J. Zhang & McFarlane, 1995, ZM hereafter) in EAMv2 is the same as that in EAMv1, except that ZM adopts two updates described in (Xie et al., 2019) to improve its simulated precipitation, in particular the diurnal cycle. The new ZM feature combines the dynamic Convective Available Potential Energy (dCAPE) trigger proposed in (Xie & Zhang, 2000) with an unrestricted air parcel launch level (ULL) approach used in (Y.-C. Wang et al., 2015) (hereafter the dCAPE-ULL trigger). The dCAPE trigger provides a dynamic constraint for preconditioning of convection-favoring environments and prevents CAPE from being released spontaneously. The ULL trigger removes the constraint that convection is always rooted within the boundary layer, as is often assumed in deep convection schemes. Thus, it captures mid-level convection by detecting atmospheric instability above the boundary layer. As shown in Xie et al. (2019), the use of the dCAPE-ULL trigger helps address the “too frequent, too weak” precipitation issue — a long-standing climate model bias — as well as capture the nocturnal elevated convection systems which are often seen downstream of major mountains associated with the propagation of Mesoscale Convective Systems (MCSs) but missed in most climate models including E3SM. It also significantly improves the phase of the diurnal cycle of precipitation over both land and ocean.

After releasing EAMv1, (Ma et al., 2022) proposed a set of recalibrated atmospheric parameters in the deep convection scheme, the microphysics scheme, and the CLUBB turbulence and macrophysics scheme (hereafter EAMv1p). Many of these parameter changes have been carried over to EAMv2. A new feature in EAMv1p is the inclusion of surface wind speed enhancements from the gustiness associated with turbulence, shallow and deep convection in the surface flux calculations over land and ocean (Ma et al., 2022; Harrop et al., 2018; Redelsperger et al., 2000).

In the deep convection scheme, the parcel buoyancy considers the subgrid temperature perturbation from the CLUBB scheme in addition to a constant value of 0.8 K used in EAMv1. A new tunable parameter with a default value of 2.0, *zmconv\_tp\_fac* (see Table A1), is introduced to scale the square root of the CLUBB subgrid temperature variance to be the subgrid temperature perturbation. Additionally, the parameters related to the autoconversion rate, detrained ice cloud effective radius, and cloud fraction in deep convective clouds are reduced, while the parameters related to the downdraft mass flux

300 fraction and the impact of the surface temperature change are enhanced compared to  
301 EAMv1.

302 A number of tunable parameters in the CLUBB scheme have been updated in EAMv1p  
303 to improve both stratocumulus and shallow cumulus clouds. Briefly, EAMv1p separated  
304 the setting of several coefficients at low skewness ( $X$ , with  $X = \text{clubb}\_{\{c1, c6, c11, \text{gamma\_coef}\}}$ )  
305 and high skewness ( $Xb$ ), recalibrated transition factors between the two regimes ( $Xc$ ),  
306 and adjusted parameters controlling the low cloudiness (e.g.,  $\mu$ ,  $c8$ ,  $c1$ ,  $c.k10$ ) to in-  
307 crease stratocumulus clouds and reduce shallow cumulus clouds. To better represent clouds  
308 and precipitation in subtropical low cloud regimes, the liquid cloud accretion enhance-  
309 ment factor and the exponent coefficient for liquid cloud autoconversion rate in the mi-  
310 crophysics scheme have been updated as well. For ice and mixed-phase clouds, the overly  
311 suppressed scaling factor (0.1) for the Wegener–Bergeron–Findeisen (WBF) process in  
312 EAMv1 has been updated to be 0.7. The Aitken mode sulfate aerosol size threshold for  
313 homogeneous ice nucleation is increased. The minimum subgrid vertical velocity for liq-  
314 uid droplet nucleation is reduced from 0.2 to 0.1 m/s in EAMv2.

315 Based on atmosphere-only and coupled simulations performed during the tuning  
316 process, EAMv2 keeps tunable parameters related to liquid droplet sedimentation, ice  
317 particle fall speed, and the lateral entrainment of deep convection the same as EAMv1  
318 instead of EAMv1p (see Table A1 for details).

319 The effective aerosol radiative forcing ( $\text{ERF}_{\text{aer}}$ ) estimated in E3SMv1 is about  $-1.6$   
320  $\text{W m}^{-2}$  (Golaz et al., 2019), which is relatively large compared to other CMIP6 mod-  
321 els (Smith et al., 2020). After applying the EAMv1p parameter tuning proposed by (Ma  
322 et al., 2022), the simulated magnitude of  $\text{ERF}_{\text{aer}}$  shortwave and longwave components  
323 is reduced significantly, but the change in net  $\text{ERF}_{\text{aer}}$  is small due to the compensation  
324 between longwave and shortwave. Clouds are more susceptible to aerosol perturbations  
325 under relatively clean conditions. Based on analysis of developmental configurations (to  
326 be documented in a separate work), unrealistically-small cloud droplet number concen-  
327 trations (e.g.,  $< 10 \text{ cm}^{-3}$ ) frequently appeared, especially in mid- and high-latitude re-  
328 gions. As a temporary remedy, a lower bound ( $10 \text{ cm}^{-3}$ ) is applied to the simulated cloud  
329 droplet number concentration in EAMv2. Results show that it reduces the net  $\text{ERF}_{\text{aer}}$   
330 magnitude by  $0.3\text{-}0.4 \text{ W m}^{-2}$ , which agrees with findings from previous studies (e.g. Hoose  
331 et al., 2009). The lower bound value is also consistent with other CMIP6 models (e.g.  
332 Mignot et al., 2021). We note however that this is not a cure for the problem. Additional  
333 efforts are planned to improve the simulated aerosol and cloud properties in pristine re-  
334 gions and reduce  $\text{ERF}_{\text{aer}}$  in a more physical manner for future versions of E3SM.

335 EAMv2 employs the same orographic and non-orographic gravity wave (GW) pa-  
336 rameterization as EAMv1, following (Richter et al., 2010), which includes separate rep-  
337 resentation of orographic GWs (McFarlane, 1987), convective GWs (Beres et al., 2004),  
338 and GWs generated by frontal systems (Charron & Manzini, 2002). Tunable param-  
339 eters in the orographic and frontal GW parameterizations remain the same as in EAMv1.  
340 In EAMv1, the period of the quasi-biennial oscillation (QBO) in the tropical stratospheric  
341 zonal mean wind was only 18 months as compared to 28 months in observations (Richter  
342 et al., 2019). In order to arrive at a more realistic representation of the QBO in EAMv2,  
343 several combinations of tunable parameters in the (Beres et al., 2004) parameterization  
344 were explored, focusing on the convective fraction (CF) and efficiency with which con-  
345 vection generates GWs,  $\text{effgw}\_{\text{beres}}$ , starting with the setting that improved the QBO  
346 in EAMv1 described in Richter et al. (2019) ( $\text{effgw}\_{\text{beres}}=0.35$  from 0.4,  $\text{CF}=8\%$  from  
347  $5\%$ ). Based on sensitivity simulations performed in parallel with the pre-industrial spinup  
348 simulation, CF was changed from  $8\%$  to  $10\%$  ( $\text{gw}\_{\text{convect}\_{\text{hcf}}} = 1/\text{CF} = 10$ ), and  $\text{effgw}\_{\text{beres}}$   
349 remained 0.35 (Table A1), resulting in a QBO period of  $\sim 21$  months in the pre-industrial  
350 control. Due to changes in tropical variability (Kelvin and mixed-Rossby gravity waves)  
351 related to the convective parameterization changes described above, the amplitude of the  
352 QBO in E3SMv2 is weaker than in observations.

#### 2.1.4 Atmospheric chemistry

The atmospheric chemistry in EAMv1 was the O3v1 model with prognostic stratospheric ozone by the linearized chemistry (Linoz v2) (Hsu & Prather, 2009) and the prescribed tropospheric ozone with the v1.0 input4MIPS ozone data set (Hegglin et al., 2016). The prescribed tropospheric ozone data only contained decadal monthly zonal climatology of latitude-pressure values. Due to the sharp cross-tropopause ozone gradient, unphysical ozone distributions were simulated in the vicinity of the tropopause when the modelled tropopause was higher than that of the prescribed data, assigning stratospheric ozone abundances to the tropospheric model grid boxes. Since ozone interacts with the radiation transfer code in E3SM, such ozone deficiencies impacted the solar heating and radiative forcing.

In EAMv2, we implemented the O3v2 model (Tang et al., 2021) to overcome the limitations in the O3v1 model by replacing the prescribed ozone data with a passive ozone tracer in the troposphere. Ozone is transported from the stratosphere into the troposphere and decays within the lowest four model layers (below 1 km) with a 48-hour e-folding to 30 ppb (parts per billion by mole fraction). The choice of 30 ppb is based on observations (Ziemke et al., 2019) and gives a tropospheric ozone mass similar to full chemistry models. O3v2 is capable of interacting with the tropopause changes and hence captures the naturally sharp ozone cross-tropopause gradient. Moreover, the ozone sink at the lower boundary in O3v2 allows us to diagnose the stratosphere-troposphere exchange flux of ozone, an important tropospheric ozone budget term, which was not possible with O3v1. The ozone hole is simulated following Cariolle et al. (1990) to represent the rapid chlorine-induced ozone depletion at cold temperatures, but the polar stratospheric cloud (PSC) temperature threshold is increased to 197.5 K in EAMv2 from 193 K in EAMv1 due to a warmer Antarctic winter pole. More details about O3v2 in E3SM are documented by (Tang et al., 2021).

#### 2.1.5 Aerosol

The aerosol model in EAMv2 is based on EAMv1 (H. Wang et al., 2020) which itself evolved from the four mode version of Modal Aerosol Module (MAM4) in the Community Atmosphere Model version 5.3 (CAM5.3) (Liu et al., 2016) that represents the major aerosol species within four internally mixed size modes, and incorporated the new treatments of aerosol processes related to new particle formation, secondary organic aerosol formation, aerosol convective transport and wet removal, resuspension, and deposition and mixing with snow grain. These new treatments in EAMv1 led to significant improvements in characterizing global distributions of aerosols and interactions with clouds and radiation. The development and evaluation of aerosol representation in the E3SMv1 coupled model simulations with both standard resolution (Golaz et al., 2019) and high resolution (Caldwell et al., 2019) configurations have mainly focused on the global budgets and annual mean constraints of aerosol optical depth (AOD) with observational estimates in the present-day conditions. While the total/speciated AOD and direct radiative effects are constrained to a large extent, further analysis of E3SMv1 simulations suggested that the shortwave absorption of aerosols is too strong in the model compared with observations especially over the dusty regions (Feng et al., 2022). The heating effect in the atmosphere due to the overestimated dust absorption could lead to changes of the lower tropospheric stability and affect the model-simulated clouds and precipitation.

In E3SMv2, we updated dust refractive indices in the shortwave bands with the observationally derived values from the AERONET measurements (Dubovik et al., 2000), which replace the strongly absorbing dust properties used in E3SMv1 (Hess et al., 1998). Additionally, we implemented a different dust particle size distribution (Kok, 2011) in E3SMv2 for calculating fractional dust emission fluxes into the accumulation and coarse modes. Kok et al. (2017) suggests that dust size distributions at emission in current global



404 climate models under-represent the coarse-mode ( $>1 \mu\text{m}$ ) dust particles in the atmosphere.  
405 For the same dust optical depth, coarse-mode dust particles would result in larger long-  
406 wave warming and less shortwave cooling than the fine particles, resulting in a less cool-  
407 ing net effect of dust aerosols. Compared to E3SMv1, the new size distribution imple-  
408 mented to E3SMv2 (Kok, 2011) predicts more particles in larger dust sizes: about 1.1%  
409 in the accumulation mode and 98.9% in the coarse mode, which is consistent with the  
410 recent measurements (Kok et al., 2017) but can substantially change the dust transport  
411 to remote regions (Wu et al., 2020). With these updates in E3SMv2, dust emissions are  
412 re-tuned for the globally constrained dust optical depth of  $0.03 \pm 0.005$  (Ridley et al., 2016).

## 413 2.2 Ocean

414 Due to development priorities, the ocean component (the Model for Prediction Across  
415 Scales-Ocean: MPAS-Ocean) in E3SMv2 is mostly unchanged from E3SMv1 (Petersen  
416 et al., 2018, 2019). The underlying spatial discretization (Thuburn et al., 2009) is ap-  
417 plied to the primitive equations with a free surface (Ringler et al., 2013), with 60 lay-  
418 ers using a z-star vertical coordinate (Petersen et al., 2015; Reckinger et al., 2015). The  
419 parameterizations of unresolved physics, such as the Gent and McWilliams (1990) param-  
420 eterization for mesoscale eddy transport, and K-Profile Parameterization (KPP, Large  
421 et al., 1994; Van Roekel et al., 2018) for vertical mixing, remain largely the same with  
422 minor updates.

423 In E3SMv2, Redi isopycnal mixing is introduced following the triad formulation  
424 from (Griffies et al., 1998). The Gent-McWilliams mesoscale eddy mixing parameteri-  
425 zation continues to utilize a globally constant value for the bolus coefficient. However,  
426 in development of E3SMv2, a series of sensitivity simulations were conducted to find a  
427 more optimal value of the Gent-McWilliams bolus kappa parameter. Based on these sim-  
428 ulations, a value of  $900 \text{ m}^2 \text{ s}^{-1}$  was chosen. This value is half of that used in E3SMv1.  
429 The reduction improved the surface salinity bias and increased Antarctic Circumpolar  
430 Circulation (ACC) transport (not shown). A smaller globally constant value ( $400 \text{ m}^2 \text{ s}^{-1}$ )  
431 is utilized for Redi isopycnal mixing. In the Redi parameterization, slope tapering is a  
432 slightly modified version of (Danabasoglu & Williams, 1995) with a critical slope param-  
433 eter of 0.01. We also implemented the stratification-based tapering from (Danabasoglu  
434 & Marshall, 2007).

435 In addition to the improvements in model physics, a sign error in the high-order  
436 reconstruction of tracer values on cell edges was discovered in the flux corrected tracer  
437 transport advection scheme. A set of simulations was conducted to determine the im-  
438 pact of this bug. The percentage change in ocean heat content (OHC) due to the bug  
439 fix at various levels is shown in Fig. S1. In a broad sense the effect of the bug fix was  
440 to increase ocean heat content, although there are broad swaths of decrease in OHC in  
441 the North Pacific and ACC in the upper ocean (Figs. S1a-c). In the deep ocean, the bug  
442 fix resulted in weaker OHC anomalies (Fig. S1d).

## 443 2.3 Sea ice

444 Improvements have been made to the column physics, coupling, and analysis of E3SM's  
445 sea ice component (MPAS-Seaice) since the E3SMv1 configuration described by Turner  
446 et al. (2021). Here we expand on innovations new to E3SMv2. The core Delta-Eddington  
447 radiative transfer of Briegleb and Light (2007) has been updated to the Dang et al. (2019)  
448 SNICAR-AD model, ensuring radiative consistency across all snow surfaces, including  
449 on land, ice sheets and sea ice. The SNICAR-AD radiative transfer code includes five-  
450 band snow single-scattering properties, two-stream Delta-Eddington approximation with  
451 the adding-doubling technique, and parameterization for correcting the near-infrared (NIR)  
452 snow albedo biases when solar zenith angle exceeds  $75^\circ$  (Dang et al., 2019). However,  
453 radiative coupling with the atmosphere still integrates across just two bands (visible and

454 NIR) separated at 700nm, which does not fully exploit the five-band capability; an  
455 expansion of the coupling bands is planned for E3SMv3.

456 A new snow-on-sea ice-morphology has been added to E3SMv2 that includes the  
457 effects of wind redistribution: losses to leads and meltponds, and the piling of snow against  
458 ridges. Snow grain radius, now a prognosed tracer field on sea ice, evolves according to  
459 temperature gradient and wet snow metamorphisms and feeds back to the SNICAR-AD  
460 radiative model up to a dry maximum of 2800  $\mu\text{m}$ . Fresh snow falls at a grain radius of  
461 54.5  $\mu\text{m}$ , and five vertical snow layers replace the previous single snow layer atop each  
462 of the five sea ice thickness categories retained from E3SMv1. The combined default con-  
463 figurations of the new radiative and snow schemes were found to minimally impact the  
464 climatic state of sea ice reported in this manuscript, but greater parametric sensitivities  
465 are explored in a sister paper in preparation.

466 The most significant improvement to the sea ice climate since E3SMv1 was achieved  
467 with coupling changes associated with mushy-layer thermodynamics. Whereas the basal  
468 temperature of the ice was held fixed at  $-1.8^\circ\text{C}$  in E3SMv1, the new version of the model  
469 assumes the mushy liquidus basal temperature from the sea ice as described by Turner  
470 and Hunke (2015). Conversion of frazil ice from MPAS-Ocean with a fixed reference salin-  
471 ity of 4 PSU to the mushy layer now conserves to computational accuracy over a 500-  
472 year control integration. This was achieved by exchanging additional mass between the  
473 upper ocean and sea ice model to accommodate an assumed 25% mushy liquid content  
474 assumed from heat and mass transferred adiabatically from the MPAS-Ocean frazil scheme  
475 active from a depth of 100 m. In addition to achieving perfect heat and mass conserva-  
476 tion between sea ice and ocean models, this improvement greatly reduces a negative sea  
477 ice thickness bias in the summer Arctic reported by Golaz et al. (2019) for E3SMv1; it  
478 only minimally impacts Southern Ocean sea ice mass that was better simulated as com-  
479 pared to northern hemisphere sea ice in E3SMv1. Note that E3SM does not use virtual  
480 ice-ocean fluxes, but instead full volume and heat flux exchange consistent with a Boussi-  
481 nesq ocean model as described by Campin et al. (2008).

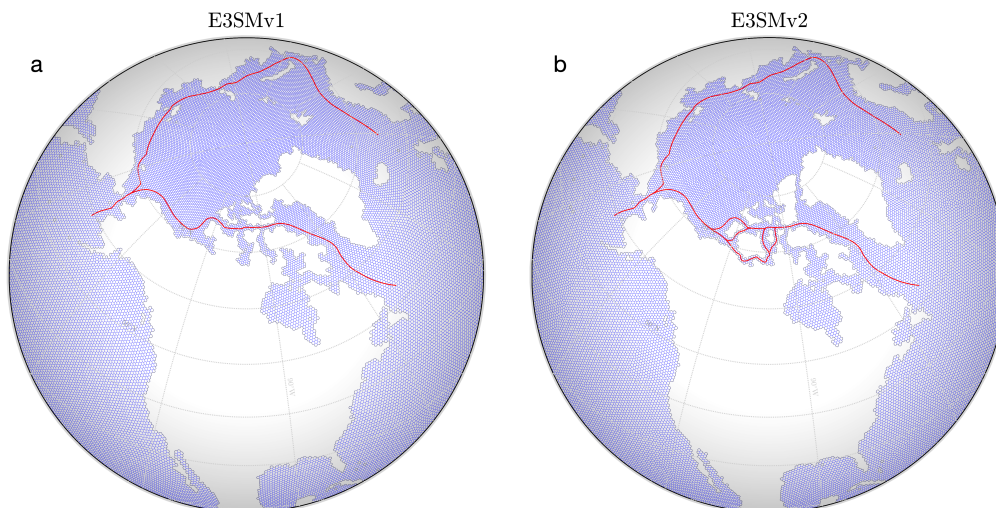
482 In addition to these core physics improvements, E3SMv2 includes a number of struc-  
483 tural additions to the sea ice model. E3SMv2 has significantly increased output to bet-  
484 ter diagnose behavior and compare against seasonal extremes and data. For example,  
485 daily Ice Numerals for Arctic shipping (Aksenov et al., 2017) are easily derived from this  
486 output, commensurate with the new E3SMv2 marine mesh that resolves major Arctic  
487 shipping channels (Section 2.4).

488 E3SMv2 now also includes a prescribed-extent ice mode for MPAS-Seaice based  
489 on that found in the Community Ice Code (CICE) in E3SMv1 and CESM (Bailey et al.,  
490 2011). This mode is needed for AMIP (Atmospheric Model Intercomparison Project) style  
491 simulations where a full prognostic sea ice model is not desired but sea ice surface fluxes,  
492 albedos, snow depth, and surface temperature are needed by the atmosphere model and  
493 are calculated by the vertical thermodynamics module of the sea ice component. The  
494 mode is intended for atmosphere sensitivity experiments and does not conserve energy  
495 or mass. In this mode, sea ice thermodynamics is active but sea ice dynamics is disabled  
496 and at each time step ice area and thickness are reset to specified values. Ice area is in-  
497 terpolated in time and space from an input data set, while ice thickness in grid cells con-  
498 taining sea ice is set to 2 m in the Northern hemisphere and 1 m in the Southern hemi-  
499 sphere. During each adjustment snow volume is adjusted to preserve the snow thickness  
500 prognosed in the previous time step. Snow temperatures are reset to the surface tem-  
501 perature, as prognosed in the previous time step, while ice temperatures are set so that  
502 the ice temperature gradient is linear, with the ice temperature at the top equal to the  
503 prognosed surface temperature, and equal to the sea freezing temperature at the base  
504 of the ice. The vertical ice salinity profile is reset to the profile from Bitz and Lipscomb  
505 (1999).

## 2.4 Unstructured marine mesh generation

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Generation of the unstructured Centroidal Voronoi-type meshes (e.g. Ringler et al., 2008) used in the ocean and sea ice components of E3SMv2 is handled using the JIGSAW library (Engwirda, 2017), enabling the creation of complex, variable-resolution meshes to resolve regional ocean (Hoch et al., 2020), sea ice (Turner et al., 2021), and land-ice (Hoffman et al., 2018) dynamics. Compared to E3SMv1, improvements to the robustness, efficiency, and flexibility of our meshing workflows have been targeted – employing a multi-paradigm mesh generation strategy that combines ‘off-center’ Delaunay-refinement and ‘hill-climbing’ optimization approaches (Engwirda & Ivers, 2016; Engwirda, 2018) to build the Spherical Centroidal Voronoi Tessellations (SCVTs) used in the MPAS-Ocean and MPAS-Seaice dynamical cores. Key to improved robustness in E3SMv2 is the elimination of invalid grid configurations centered around obtuse triangles, in which a lack of geometrical consistency between adjacent computational cells would lead to breakdowns in the numerical discretization used by the ocean dynamical core. Difficulties associated with the generation of valid meshes limited the application of variable mesh resolution in E3SMv1, restricting model configurations to quasi-uniform resolution cases. These effects are remedied in E3SMv2, with our enhanced optimization strategies leading to the generation of valid, well-conditioned meshes in complex, regionally-refined configurations. Equally important are improvements to E3SM’s COMPASS (Configuration Of MPAS Setups) package—a Python-based scripting environment that allows modelers to readily customize mesh and model configurations based on proximity to geographic features, climatological states, and user-defined inputs, with geometric tuning parameters that are easy to adjust on the fly. COMPASS tracks mesh provenance data associated with the creation of each new E3SM configuration to support model regression testing and ensure long-term reproducibility. Overall, improvements to the unstructured meshing workflows in E3SMv2 have led to significantly improved turnaround in the mesh design, simulation, and analysis process, reducing the time required to complete various MPAS mesh-related tasks from days-to-weeks in E3SMv1 to minutes-to-hours in E3SMv2. As a consequence of these improvements, the E3SMv2 coastline is more realistic across the globe. As one example, E3SMv2 includes key shipping routes in the Canadian Archipelago that were missing from E3SMv1 (Figure 1), eliciting improved archipelagic through-flow.



**Figure 1.** Comparison of the (a) old and (b) new standard resolution E3SM unstructured marine mesh, highlighting improved geographic acuity in E3SMv2 including Arctic coastal shipping channels fitting standard routes published by the Arctic Council (2009) (red).

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## 2.5 Land and river

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The physics configuration of E3SM Land Model version 2 (ELMv2) used in E3SMv2 is similar to E3SMv1 (Golaz et al., 2019). ELMv2 simulates hydrologic and thermal processes in vegetation, snow, and soil for different land cover types, which include bare soils, vegetated surfaces, lakes, glaciers, and urban areas. Present-day leaf area index (LAI) is prescribed using satellite data and photosynthesis and is not limited by leaf nutrients. The prescribed vegetation distribution has been updated for E3SMv2 to resolve inconsistencies across computing platforms in translating land use to changes in plant functional types. ELMv2 includes the new shortwave radiation model SNICAR-AD for snow which is also used for sea ice as described in Section 2.3.

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The river routing component in E3SMv2 (Model for Scale Adaptive River Transport, MOSARTv2) takes the runoff produced by ELM and routes it to the river mouth as freshwater input to the ocean component. The physics scheme and configuration are the same as used in E3SMv1 standard resolution (Golaz et al., 2019). Specifically, MOSARTv2 uses the kinematic wave approach to route streamflow across hillslopes, tributaries, and main river stems on an eight-direction-based river network (Li et al., 2013) at  $0.5^\circ$  latitude-longitude spatial resolution.

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## 2.6 Coupled system

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As in E3SMv1, the coupler/driver for E3SMv2 is cpl7 (Craig et al., 2012). The driver of cpl7 performs the integration of the coupled model and provides the “main” for the single executable. cpl7 relies on the Model Coupling Toolkit (MCT; J. Larson et al., 2005) for inter-component communication and remapping operations.

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### 2.6.1 Mapping weights

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The remapping operations are performed using mapping weights precomputed by external tools for each grid pair using two different algorithms. Nearly all maps in both directions use the TempestRemap conservative, monotone map (Ullrich & Taylor, 2015; Ullrich et al., 2016). In the case of the atmosphere’s pg2 grid and the ocean’s Voronoi grid, TempestRemap implements an  $L^2$  projection between the finite-volume grids. The requirement of monotonicity implies the projection must use the constant-function basis rather than a high-order reconstruction. This map type is used for all fluxes and most states in the coupled model. The second map type is bilinear interpolation from ESMF (Hill et al., 2004). This map type is used to transfer state from the atmosphere to the ocean and sea ice.

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### 2.6.2 Energy conservation

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With respect to energy conservation, EAM closely follows the design of CAM (Neale et al., 2012). The formulation of total energy is given in Lauritzen and Williamson (2019), equation 8. Each parameterization is expected to conserve energy and mass; therefore, there are no energy fixers (column fixers) for parameterizations at the interface level in EAM. Two known components of EAM that require energy fixers are the pressure adjustment and the dynamical core. Parameterizations operate under the constant moist pressure assumption, which requires the moist pressure adjustment described in Neale et al. (2012) and Lauritzen and Williamson (2019). In the dynamical core, there are several sources of energy dissipation, as described in Section 2.1.1. Therefore, a global energy fixer in the form of a uniform increment to the temperature field is applied to compensate for these processes as well as pressure adjustment. The fixer’s typical average value is  $-0.07 \text{ W m}^{-2}$ .

583 The change of energy in the atmosphere should be equal to the difference in net  
584 fluxes at the top of the model and the surface. A long-term average of the energy change  
585 should be close to zero if the model conserves energy, since storage in the atmosphere  
586 is minimal. EAMv1 contained a few energy leaks. For example, one source of leaks is  
587 the presence of a limiter for water forms (K. Zhang et al., 2018), but this source is small.  
588 In EAMv1, we recognized the gravity wave drag (GWD) parameterization as the source  
589 of the largest energy leak. In the orographic gravity wave parameterization, the change  
590 in kinetic energy was not properly accounted for. After a fix, the energy imbalance in  
591 the atmosphere is reduced from  $0.07 \text{ W m}^{-2}$  to  $0.01 \text{ W m}^{-2}$ . Figure S2 depicts the en-  
592 ergy imbalances for atmosphere simulations with and without the GWD energy fix.

593 MPAS-Ocean utilizes a fixed two band exponential formulation for penetrating short-  
594 wave radiation. For grid cells with shallow bottom depths, a portion of the penetrating  
595 shortwave radiation reaches the bottom of the ocean. In E3SMv1, this portion of the short-  
596 wave radiation was not accounted for, resulting in a globally averaged energy leak of ap-  
597 proximately  $0.25 \text{ W m}^{-2}$ . In E3SMv2, the shortwave radiation that reaches the bottom  
598 of the ocean is added to the bottom layer. In the development of E3SMv2, we found that  
599 this change had minimal impact on the large scale ocean climate.

600 After these energy conservation errors in the atmosphere and the ocean were ad-  
601 dressed, we realized that the coupled system was no longer in energy balance compared  
602 to E3SMv1. Further investigation led to the energy correction term incorporated in E3SMv1  
603 to account for the inconsistent definition of energy in the ocean and atmosphere (see Go-  
604 laz et al., 2019, Appendix A). While conceptually correct, the computation of that cor-  
605 rection term was based on all the precipitation, when instead it should have included only  
606 precipitation over ocean and ocean runoff. Precipitation over land should not have been  
607 included because the land model ELM does not take into account heat carried by pre-  
608 cipitation. The energy imbalance was corrected by calculating the needed energy to bring  
609 fluxes of water to a common temperature with the ocean, and then pass the globally av-  
610 eraged value as a correction term to be applied in the atmosphere every coupling time  
611 step.

### 612 **2.6.3 Coupled tuning**

613 The coupled tuning objectives for the pre-industrial control simulation were sim-  
614 ilar to Golaz et al. (2019):

- 615 1. Near-zero long-term average net top-of-atmosphere (TOA) energy flux and total  
616 OHC in equilibrium.
- 617 2. Minimum long-term drift in global mean surface air temperature.
- 618 3. Reasonable absolute global mean surface air temperature.

619 Furthermore, spatial root mean square errors (RMSEs) against observations for key  
620 climate variables (e.g., annual mean sea surface temperature (SST), annual and seasonal  
621 precipitation, TOA radiation, cloud radiative effect, sea surface wind stress, etc.) from  
622 the E3SM Diagnostic package (C. Zhang et al., 2022) are also considered. Tuning was  
623 performed iteratively through trial and error at component levels and with the coupled  
624 system under perpetual pre-industrial (1850) forcings.

625 In the atmosphere, we firstly conducted short atmosphere sensitivity tests with re-  
626 peating SST and sea ice annual cycle (“F2010”) to estimate the impact of individual pa-  
627 rameters on the modeled precipitation, cloud radiative forcing, and other climate state  
628 variables. Based on the short sensitivity tests, we designed several groups of parameters  
629 and conducted F2010 simulations in a parallel manner. Then, we evaluated these tun-  
630 ing combinations based on the aforementioned key climate variables from the E3SM Di-  
631 agnostics package. Promising atmospheric configurations were then evaluated with longer

632 AMIP simulations (prescribed SST for the years 1980-2015) before being tested in pre-  
633 industrial coupled mode. Results from the coupled simulation then fed back into another  
634 round of atmospheric tuning. Periodically, we also performed atmospheric simulations  
635 to evaluate cloud feedback and aerosol ERF to inform the atmospheric tuning. Specif-  
636 ically, we estimated the cloud feedback using Cess-like simulations (Cess et al., 1989) by  
637 comparing the differences between an 11-year AMIP standard simulation (years 1980-  
638 1990) and the same simulation except with globally +4K SST (Ringer et al., 2014). The  
639 aerosol ERF was estimated with time slice simulations (e.g. Hansen, 2005) consisting  
640 of a 9-year 2010 simulation vs a 2010 simulation except with 1850 aerosol emissions. To  
641 estimate the aerosol ERF more efficiently, we also used short (1 year after 3-month spin-  
642 up) nudged simulations with 2010 and 1850 aerosol emissions (all other external forc-  
643 ings kept as year 2010 conditions), where the horizontal winds were nudged towards model  
644 output from a baseline simulation. Nudging was used mainly to constrain the large-scale  
645 circulation, so that the noises caused by the atmospheric internal variability are reduced  
646 and short simulations can be used to estimate the aerosol ERF (Kooperman et al., 2012;  
647 K. Zhang et al., 2014). Previous studies (K. Zhang et al., 2022; S. Zhang et al., 2022)  
648 showed good agreement in the global and regional annual mean aerosol ERF estimates  
649 between the free-running and nudged simulations in E3SMv1.

650 Component-level development and tuning for the ocean also relied on simulations  
651 forced with atmospheric reanalyses (Tsuji no et al., 2018) to guide the tuning of the Gent-  
652 McWilliams bolus kappa parameter and the newly implemented Redi isopycnal mixing  
653 scheme.

654 As in E3SMv1, the last step was a final tuning of the CLUBB parameter *clubb\_c14*  
655 in the coupled system to minimize long-term drift by adjusting shortwave cloud radi-  
656 ative effects (SWCRE) in the low-cloud regimes.

657 Pre-industrial simulations were the only coupled simulations performed before the  
658 model was frozen. In particular, no idealized CO<sub>2</sub> or test historical simulations were per-  
659 formed before finalizing E3SMv2.

### 660 3 Computational performance

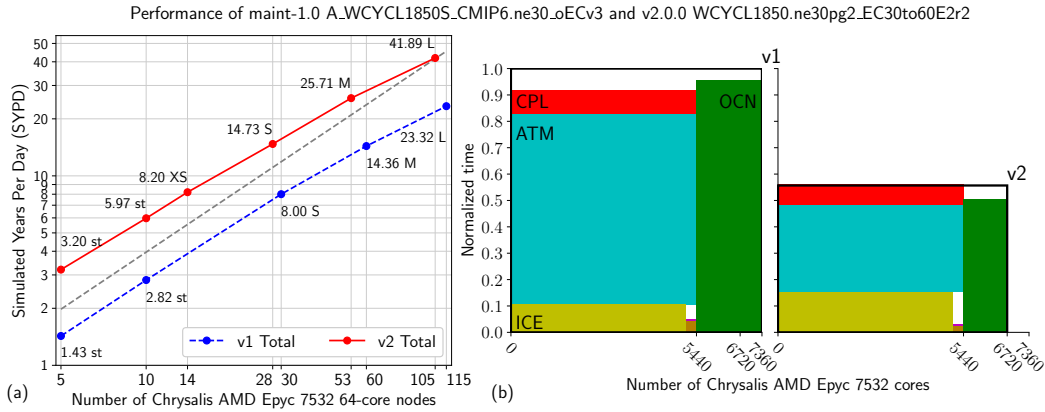
#### 661 3.1 Performance comparison of v1 and v2 simulations

662 This section examines computational performance using a set of atmosphere-only  
663 and fully coupled simulations. Relative to EAMv1, EAMv2 is approximately twice as  
664 efficient for primarily two reasons: faster passive tracer transport and fewer physics columns.  
665 E3SMv2 is also approximately twice as efficient because the ocean dynamics time step  
666 is three times larger than in E3SMv1. The sea ice component is slower in v2 than in v1  
667 because of additional snow layers.

668 Performance benchmark simulations were performed on the ANL Chrysalis clus-  
669 ter. Chrysalis has 512 compute nodes. Each node of the cluster has two AMD Epyc 7532  
670 “Rome” 2.4 GHz processors, and each processor has 32 cores, for a total of 64 cores per  
671 node. Each node has 256GB 16 channel DDR4 3200MHz memory. The interconnect hard-  
672 ware is Mellanox HDR200 InfiniBand and uses the fat tree topology. The model code  
673 was compiled with Intel release 20200925 with GCC version 8.3.1 compatibility and run  
674 with OpenMPI 4.1.1 provided in the Mellanox HPC-X Software Toolkit.

675 All throughput values reported in this section are derived using the maximum time  
676 (minimum throughput) over all MPI processes. Only the total throughput value is fully  
677 accurate, as it is computed using the top-level wallclock time of the simulation, exclud-  
678 ing initialization; component and subcomponent throughput values are approximations  
679 because these lower-level timers are not associated with global synchronization points.  
680 The simulations are run with one MPI process per core and no OpenMP threading. A

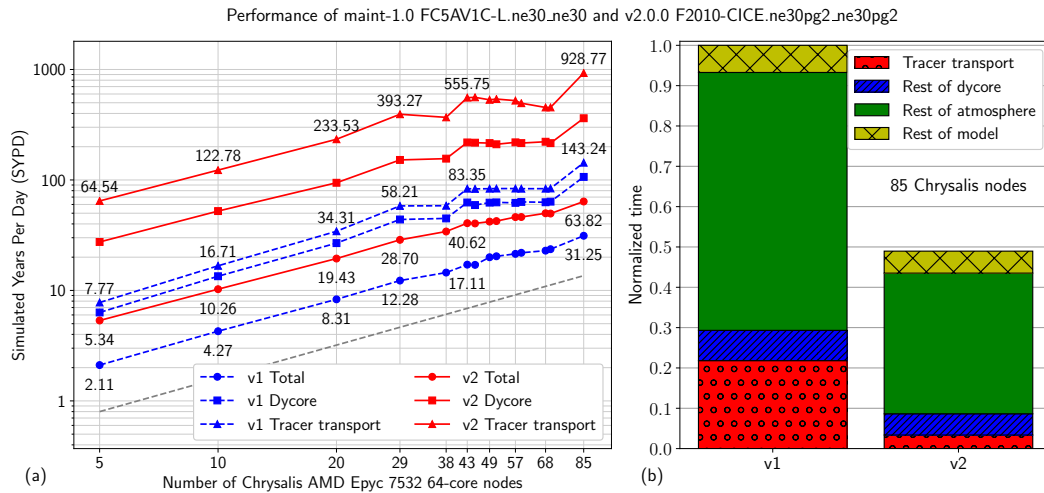
681 throughput data point corresponds to one simulation run for three months with the de-  
 682 fault input/output (I/O) configuration and one restart file at the simulation end. For  
 683 these tests, both v1 and v2 simulations use the new SCORPIO (Software for Caching  
 684 Output and Reads for Parallel I/O) I/O library; thus, performance differences in these  
 685 simulations are due to components' computational and I/O volume differences rather than  
 686 I/O library differences. Performance improvements from SCORPIO are documented sep-  
 687 arately in Section 3.2.



**Figure 2.** Performance of the lower resolution E3SMv1 and E3SMv2 pre-industrial control simulations. (a) Throughput vs. number of nodes. PE layouts XS, S, M, L are provided as part of the models. Points annotated with “st” use a simple stacked layout in which each component runs serially with respect to the others, and all components share the same processors. (b) Throughput-resource product plots. Each component has one rectangle. A rectangle has the area given by the product of throughput and number of nodes. In v2, the atmosphere and ocean components have substantially smaller throughput-resource products.

688 Figure 2 summarizes the performance of E3SMv2 relative to E3SMv1 on the lower  
 689 resolution E3SMv1 and E3SMv2 pre-industrial control simulations. Figure 2a plots total  
 690 throughput versus the number of computer nodes. The models provide a small number  
 691 of optimized layouts, available using the names XS (v2 only), S, M, L. In addition,  
 692 the figure shows small-node-count simulations using a simple stacked layout (“st”) in  
 693 which each component runs serially with respect to the others, and all components share  
 694 the same processors. Each simulation’s data point is annotated with its throughput in  
 695 simulated years per day (SYPD) and layout. Comparing S, M, and L layouts between  
 696 models, v2 is at least 1.97 times more efficient than E3SMv1. Figure 2b illustrates this  
 697 efficiency difference by plotting the throughput-resource product for each component  
 698 as a rectangle for the L layouts. The atmosphere (ATM), sea ice (ICE), coupler (CPL), land  
 699 (LND), and river runoff (ROF; LND and ROF are too small to label) components run  
 700 on one set of nodes, while the ocean (OCN) component runs on another set. An unfilled  
 701 rectangle having “v1” or “v2” at the top-right corner shows the total product; because  
 702 the throughput value of each component is approximate, the filled rectangles do not sum  
 703 to the total throughput value.

704 Figure 3 focuses on just the atmosphere component using prescribed SST and sea  
 705 ice simulations. In E3SMv2, by default MPAS-Seaice now replaces CICE in such con-  
 706 figurations (see Section 2.3). However, we use CICE for this study for three reasons. First,  
 707 MPAS-Seaice requires a partition file for each process decomposition, and one goal of  
 708 this study is to run simulations with a large number of decompositions. With CICE, we  
 709 do not need to generate a decomposition file for each one. Second, MPAS-Seaice is slower



**Figure 3.** Performance of the lower resolution EAMv1 and EAMv2 atmosphere simulations. (a) Throughput vs. number of nodes. PE layouts are simple stacked layouts. (b) Proportion of time spent in each subcomponent, with the total time for v1 normalized to 1.

than CICE, and it must run on an MPAS grid; the combined slowdown from each of these would reduce the precision of our analysis of just the atmosphere component’s performance in this study. Finally, v1 must use CICE, so a comparison of just the changes to the atmosphere component is best done by using CICE in the v2 simulations as well.

Figure 3a shows total throughput of the simulation and approximate throughputs of the dynamical core (“dycore”) and passive tracer transport. A subset of data points are annotated with throughput values. Passive tracer transport is at least six to at least eight times faster in v2 than in v1. Two details are apparent in this plot. First, the dynamical core is sensitive to the element decomposition, while the rest of the model is sensitive to the finer physics column decomposition. Thus, between 43 and 68 nodes, performance of the dynamical core subcomponents plateaus or slightly degrades, since in this range an increase in node count provides no improvement to the most-burdened MPI processes. Nonetheless, total throughput is roughly monotonically increasing even in this node count range. Second, representative node counts are chosen to favor, generally separately, v1 and v2 in roughly equal numbers. Thus, there are closely spaced pairs of points in this same range to show the best available throughputs of both model versions.

Figure 3b decomposes performance of the 85-node simulations into the same sub-components. Only each full-height bar is fully accurate; subcomponent proportions are approximate. Again, tracer transport in v2 is over six times faster than in v1, speeding up the dynamical core by over three times in this case. The total model speedup is a little over two times in this case, with the speedup outside of the dynamical core coming from the reduction in number of physics columns.

### 3.2 File Input/Output

The EAM and E3SM simulations discussed above used the SCORPIO library for reading input data and writing simulation output to the file system. To improve the I/O write performance, the library caches and rearranges output data among MPI processes before using low-level I/O libraries, such as NetCDF, Parallel NetCDF (PnetCDF), and the Adaptable IO System (ADIOS), to write the data to the file system. In all the sim-



738 ulation campaigns we used PnetCDF as the low-level I/O library in SCORPIO, and I/O  
739 accounted for less than 4% of the total runtime of the simulation.

740 To measure the I/O improvements in the model, we compared the old version of  
741 the I/O library, SCORPIO CLASSIC (based on PIO, Dennis et al., 2012), used by E3SMv1,  
742 with the new version of the library, SCORPIO, used by E3SMv2, by running E3SMv1  
743 benchmark simulations on Chrysalis with the S, M, and L configurations. The simula-  
744 tion was run for 90 simulated days and generated  $\sim 30$ GB of history and restart model  
745 output for each configuration. We found that SCORPIO provides a higher write and read  
746 performance than SCORPIO CLASSIC for all the model configurations. SCORPIO pro-  
747 vides a consistent write throughput of 3-3.5 GB/s for all the configurations while the write  
748 throughput of SCORPIO CLASSIC drops from 1.9 GB/s for the S configuration to 356  
749 MB/s for the L configuration. The time to read the model input data stays relatively  
750 constant for SCORPIO with the different model configurations while it increases expo-  
751 nentially with the number of MPI processes for SCORPIO CLASSIC. The time to read  
752 the model input data is  $\sim 40\%$  higher for SCORPIO CLASSIC compared to SCORPIO  
753 for the S and M model configurations, and for the L model configuration the time to read  
754 data with SCORPIO CLASSIC is 3.3 times the time taken with SCORPIO. The total  
755 time, including reads and writes, spent in I/O by both the libraries was less than 8% of  
756 the total runtime for all the model configurations except the L configuration with the  
757 SCORPIO CLASSIC library, where I/O accounted for 25% of the total runtime.

#### 758 4 Simulation Campaign

759 Table 1 summarizes the E3SMv2 simulation campaign. All simulations were con-  
760 figured to adhere to the CMIP6 specifications as closely as possible and rely on the same  
761 boundary files as E3SMv1 (Golaz et al., 2019). The CMIP6 DECK plus historical sim-  
762 ulations (Eyring et al., 2016) include the pre-industrial control (*piControl*) spanning a  
763 total of 500 years, idealized CO<sub>2</sub> simulations (*1pctCO2*, *abrupt-4xCO2*; 150 years each),  
764 and a five-member ensemble of historical simulations (*historical\_N*; 1850-2014). These  
765 simulations were initialized from *piControl* on Jan 1 of various years as indicated in Ta-  
766 ble 1. AMIP simulations (prescribed SST and sea ice extent) were also performed to cover  
767 the entire period for which CMIP6 provides surface boundary conditions (1870-2014).  
768 Atmosphere, land, and river initial conditions for *amip\_N* were taken from year 1870 of  
769 the corresponding *historical\_N* coupled simulations.

770 To understand the relative importance of different forcing agents, a set of DAMIP  
771 (Gillett et al., 2016) historical simulations was performed. They consist of five-member  
772 ensembles with well-mixed greenhouse-gas-only (*hist-GHG*) and anthropogenic-aerosol-  
773 related (*hist-aer*) forcing agents. Instead of natural-only historical simulations as in Gillett  
774 et al. (2016), we opted for a third set with all agents active except well-mixed GHG and  
775 aerosols (*hist-all-xGHG-xaer*). This non-standard choice was motivated by a desire to  
776 include all forcing agents in our decomposition (including land-use and ozone).

777 Finally, we performed a set of simulations following RFMIP (Radiative Forcing Model  
778 Intercomparison Project; Pincus et al., 2016) with slight updates to the protocol ([https://  
779 rfmip.leeds.ac.uk/rfmip-erf](https://rfmip.leeds.ac.uk/rfmip-erf)). These simulations are designed to estimate time-varying  
780 total and aerosol-related ERF. Three sets of prescribed SST and sea ice simulations are  
781 performed with SST and sea ice derived from a 500-year average of *piControl*. *piClim-*  
782 *control* is the control simulation with all forcing agents held at their 1850 values. *piClim-*  
783 *histall* activates all time varying forcing agents, whereas *piClim-histaer* only activates  
784 time varying agents related to anthropogenic aerosols and their precursors.

785 The entire simulation campaign was performed on the DOE-E3SM Chrysalis cluster  
786 located at Argonne National Laboratory. E3SMv2 experienced only a single model  
787 crash during the nearly 3000 simulated years. The failure occurred during year 121 of

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*abrupt-4xCO2* ensemble member 301. The failure was overcome by rerunning and toggling a flag in the coupler (“BFBFLAG”) that changes order of arithmetic operations. This introduces a “butterfly effect” sufficient to alter the weather and avoid the original failure point.

**Table 1.** Summary of E3SMv2 simulations.

Label	Description	Period	Ens.	Initialization
<b>Fully coupled</b> (atmosphere, ocean, sea ice, land, and river)				
<i>piControl</i>	Pre-industrial control	500 years	-	Pre-industrial spinup
<i>1pctCO2_N</i>	Prescribed 1% yr <sup>-1</sup> CO <sub>2</sub> increase	150 years	1	<i>piControl</i> (101)
<i>abrupt-4xCO2_N</i>	Abrupt CO <sub>2</sub> quadrupling	150 years	2	<i>piControl</i> (101, 301)
<i>historical_N</i>	Historical	1850-2014	5	<i>piControl</i> (101, 151, 201, 251, 301)
<i>hist-GHG_N</i>	DAMIP well-mixed greenhouse-gas-only historical	1850-2014	5	<i>piControl</i> (101, 151, 201, 251, 301)
<i>hist-aer_N</i>	DAMIP anthropogenic-aerosol-related historical	1850-2014	5	<i>piControl</i> (101, 151, 201, 251, 301)
<i>hist-all-xGHG-xaer_N</i>	Other forcing historical (all forcings except GHG and aer)	1850-2014	5	<i>piControl</i> (101, 151, 201, 251, 301)
<b>Prescribed SST and sea ice extent</b> (atmosphere, thermodynamic sea ice, land and river)				
<i>amip_N</i>	Atmosphere with prescribed SSTs and sea ice concentration	1870-2014	3	<i>historical_N</i> (1870)
<i>piClim-control</i>	RFMIP baseline control	50 years	-	Pre-industrial spinup
<i>piClim-histall_N</i>	RFMIP time-varying ERF all agents	1850-2014	3	<i>piClim-Control</i> (21, 31, 41)
<i>piClim-histaer_N</i>	RFMIP time-varying ERF aerosols	1850-2014	3	<i>piClim-Control</i> (21, 31, 41)

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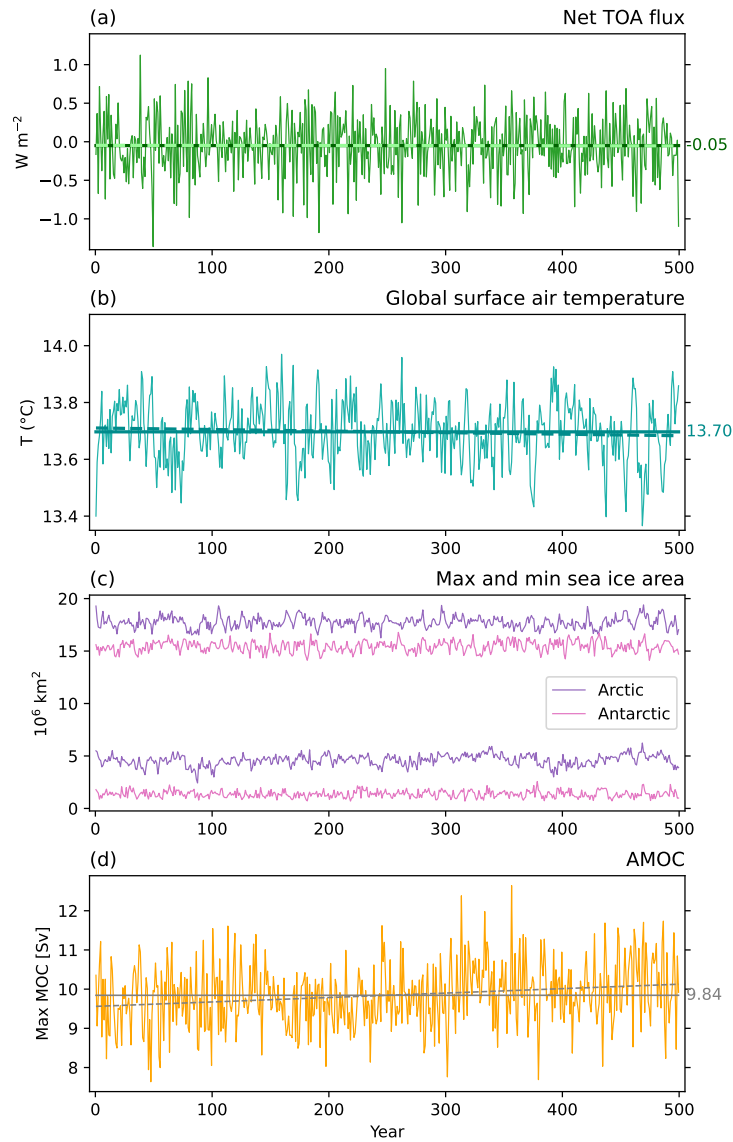
#### 4.1 Pre-industrial control

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The pre-industrial control simulation (*piControl*) was initialized after a 1000-year long spin-up simulation, itself initialized from ocean and sea ice states derived from a one-year forced ocean-sea ice simulation. During the spin-up, the model configuration was final, except for a small retuning of the gravity wave drag parameterization that was introduced at year 800 to improve the period of the QBO as described in Section 2.1.3.

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The climate simulated by E3SMv2 is very stable throughout the 500-year *piControl* as demonstrated in Figure 4. The net TOA radiation (Fig. 4a) averages to  $-0.05$  W m<sup>-2</sup> with no trend. This value is sufficiently close (compared to anthropogenic forcing) to the ideal value of 0 W m<sup>-2</sup> for a fully equilibrated and perfectly energy conserving model.



**Figure 4.** Time evolution of annual (a) global mean net top-of-atmosphere (TOA) radiation (positive down), (b) global mean surface air temperature, (c) maximum and minimum of total sea ice area for the Arctic and Antarctic, and (d) maximum Atlantic Meridional Overturning Circulation (AMOC) at  $26.5^{\circ}N$  below 500-m depth in the piControl simulation. Dashed lines in (a), (b), and (d) represent linear trends. The solid straight line in (a) is the mean TOA energy imbalance of  $-0.05 W m^{-2}$ , while the solid straight line in (d) is the mean annual maximum AMOC of 9.84 Sv.

803 The global mean surface temperature averages to 13.70 °C with a very small down-  
 804 ward trend (dashed line in Fig. 4b). The average temperature is very similar to E3SMv1  
 805 and consistent with observational estimates (e.g. global temperature of 14.0±0.5°C for  
 806 the period 1961-1990 minus estimated warming; Jones et al., 1999). Along with the global  
 807 mean temperature, maximum and minimum seasonal sea ice areas for the Arctic and Antarc-  
 808 tic are stable as well (Fig. 4c).

809 Finally, the maximum AMOC in E3SMv2 is quite weak, similar to that in E3SMv1.  
 810 The AMOC in Fig. 4d is weaker than the value in (Golaz et al., 2019) (~11 Sv). How-  
 811 ever, during the E3SMv2 development it was discovered that the published AMOC did  
 812 not include the contribution of the parameterized mesoscale eddies. In the North Atlantic  
 813 the influence of the Gent-McWilliams parameterization opposes the resolved AMOC. When  
 814 the eddy bolus velocity is included in the v1 calculation, the AMOC is very similar (~9.5  
 815 Sv) to that in E3SMv2.

## 816 4.2 Climate sensitivity and effective radiative forcing

817 Included in the DECK simulations are two idealized CO<sub>2</sub> simulations designed to  
 818 estimate the model response (sensitivity) to CO<sub>2</sub>-forcing at different time horizons. The  
 819 equilibrium climate sensitivity (ECS) is defined as the equilibrium surface temperature  
 820 change resulting from a doubling in CO<sub>2</sub> concentrations. Because it is not practical to  
 821 run a model to equilibrium, ECS is approximated by linear regression of TOA radiation  
 822 vs surface temperature in a 150-year “*abrupt-4xCO2*” simulation (Gregory et al., 2004),  
 823 often referred to as “effective climate sensitivity”. Response on shorter time scales is mea-  
 824 sured by the transient climate response (TCR). TCR is defined as the change in surface  
 825 temperature averaged for a 20-year period around the time of CO<sub>2</sub> doubling from a *1pctCO2*  
 826 simulation. TCR depends on both climate sensitivity and ocean heat uptake rate.

827 Figure 5 illustrates the time evolution of annual-average surface air temperature  
 828 from the E3SMv1 and E3SMv2 idealized CO<sub>2</sub> simulations, as well as their linear regres-  
 829 sions. ECS is reduced from 5.3 K in E3SMv1 to 4.0 K in E3SMv2, a substantial reduc-  
 830 tion (25%). TCR is reduced as well, but by a smaller fraction from 2.93 K to 2.41 K (18%).  
 831 The effective CO<sub>2</sub> radiative forcing is also reduced by 11% (3.34 to 2.98 W m<sup>-2</sup>). To ex-  
 832 plore the impact of the non-linearity, we also calculate separate regressions for the first  
 833 20 and last 130 years as proposed by Andrews et al. (2015a). Doing so leads to only slightly  
 834 larger estimates of the forcing (3.36 vs 2.98 W m<sup>-2</sup>) and ECS (4.25 vs 4.0 K) for E3SMv2,  
 835 and similarly for E3SMv1. The impact on ECS is small compared to some other mod-  
 836 els (Andrews et al., 2015a, their Fig. 2).

837 For comparison, Meehl et al. (2020) evaluated ECS and TCR for 37 CMIP6 mod-  
 838 els. ECS ranged between 1.8 and 5.6 K, with 6 models above 5 K including E3SMv1.  
 839 The multimodel mean ECS was 3.7 K with a standard deviation of 1.1 K. TCR ranged  
 840 from 1.3 to 3.0 K, with E3SMv1 having the largest value. The multimodel mean TCR  
 841 was 2.0 K with a standard deviation of 0.4 K. E3SMv2 is now within one standard de-  
 842 viation of multimodel mean for both ECS and TCR, but still on the high side.

843 World Climate Research Programme (WCRP) researchers conducted a recent as-  
 844 sessment of the equilibrium climate sensitivity following multiple lines of evidence (Sherwood  
 845 et al., 2020). They arrived at a 66% confidence range of 2.6–3.9 K for their baseline cal-  
 846 culation and 2.3–4.5 K under their robustness tests. The broader 5–95% confidence ranges  
 847 were 2.3–4.7 K, respectively 2.0–5.7 K. E3SMv1 with an ECS of 5.3 K is rather unre-  
 848 alistic as it lies outside of most of those ranges. On the other hand, E3SMv2 has a high,  
 849 but plausible ECS of 4.0 K.

850 Although a part of the reduction in ECS stems from the reduced effective radia-  
 851 tive forcing in E3SMv2 (from 3.34 to 2.98 W m<sup>-2</sup>), it is mainly due to the reduced to-  
 852 tal climate feedback. Applying the radiative kernel method (Soden et al., 2008) imple-

853 mented in the E3SM cloud feedback diagnostic package (Qin, 2022) to decompose the  
 854 climate feedback into different components, we find the reduced cloud feedback (E3SMv1:  
 855  $0.93 \text{ W m}^{-2}\text{K}^{-1}$ ; E3SMv2:  $0.72 \text{ W m}^{-2}\text{K}^{-1}$ ), especially over the marine low cloud re-  
 856 gions, contributes the most to the reduction in total climate feedback, whereas the changes  
 857 in other non-cloud feedbacks are negligible. Sensitivity tests on model changes in E3SMv2  
 858 atmosphere physics indicate that the dCAPE-ULL convective trigger in the ZM scheme  
 859 and the updated CLUBB tuning parameters play leading roles in reducing the marine  
 860 low cloud feedbacks in E3SMv2. The new trigger function in ZM deep convection tends  
 861 to be activated more frequently under warming, leading to more cloud water detrain-  
 862 ment to sustain the low clouds. The net impact of CLUBB changes on marine low cloud  
 863 feedback is partly related to the reduced decoupling between boundary layer and free  
 864 troposphere in mean climate. More details will be reported in an upcoming paper.

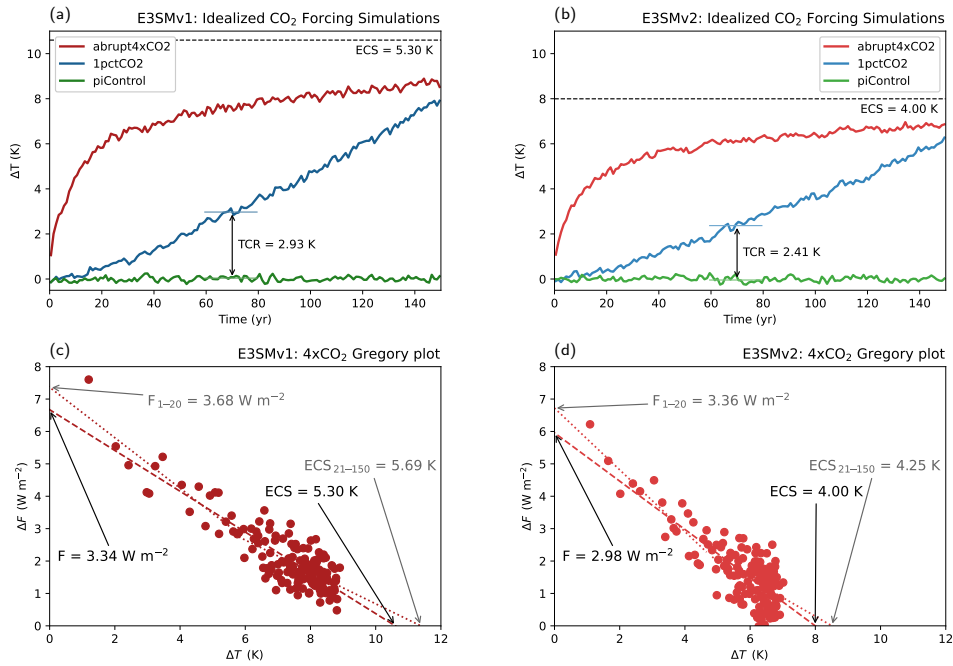
865 We also evaluate the evolution of ERF from pre-industrial to present-day condi-  
 866 tions using RFMIP simulations (Table 1).  $\text{ERF}_{\text{total}}$  is the difference in net TOA radi-  
 867 ation between *piClim-histall* and *piClim-control* and  $\text{ERF}_{\text{aer}}$  the difference between *piClim-*  
 868 *histaer* and *piClim-control*. Their time evolutions are shown in Figure 6 along with their  
 869 counterparts from E3SMv1 (computed with a comparable but slightly different method-  
 870 ology, see Golaz et al., 2019). The time evolutions of  $\text{ERF}_{\text{total}}$  and  $\text{ERF}_{\text{aer}}$  are nearly  
 871 identical between E3SMv2 and E3SMv1.  $\text{ERF}_{\text{total}}$  remains close to zero until the late  
 872 1900's, except for dips during explosive volcanic eruptions. Averaging over the last 20  
 873 years reveals small differences between the two models. The aerosol forcing is slightly  
 874 reduced in magnitude ( $-1.52$  vs  $-1.65 \text{ W m}^{-2}$ ), but the the total forcing does not in-  
 875 crease as a result. In fact it is reduced ( $+1.00$  vs  $+1.10 \text{ W m}^{-2}$ ), likely as a consequence  
 876 of the smaller  $\text{CO}_2$  ERF (Fig. 5).

877 Another assessment was conducted under the auspices of the WCRP with the goal  
 878 of bounding the aerosol radiative forcing (Bellouin et al., 2020). Following multiple lines  
 879 of evidence, the assessment arrived at a 68% confidence interval for the total aerosol ef-  
 880 fective radiative forcing of  $-1.6$  to  $-0.6 \text{ W m}^{-2}$ , or  $-2.0$  to  $-0.4 \text{ W m}^{-2}$  with a 90%  
 881 likelihood. With a forcing of  $-1.52 \text{ W m}^{-2}$ , E3SMv2 is close to the lower bound but within  
 882 the narrower confidence interval.

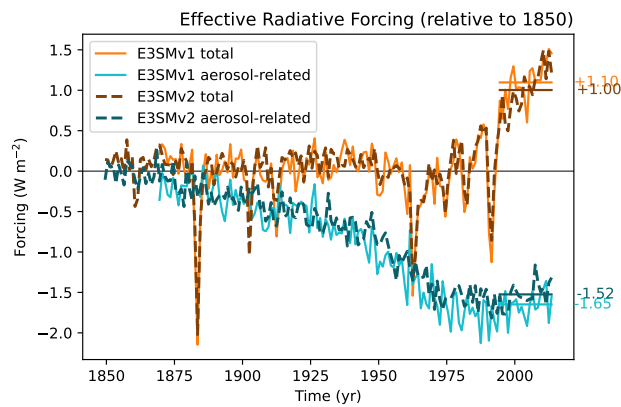
### 883 4.3 Historical ensemble

884 To facilitate comparisons between model and observations, the bulk of the anal-  
 885 ysis focuses on the historical simulations.

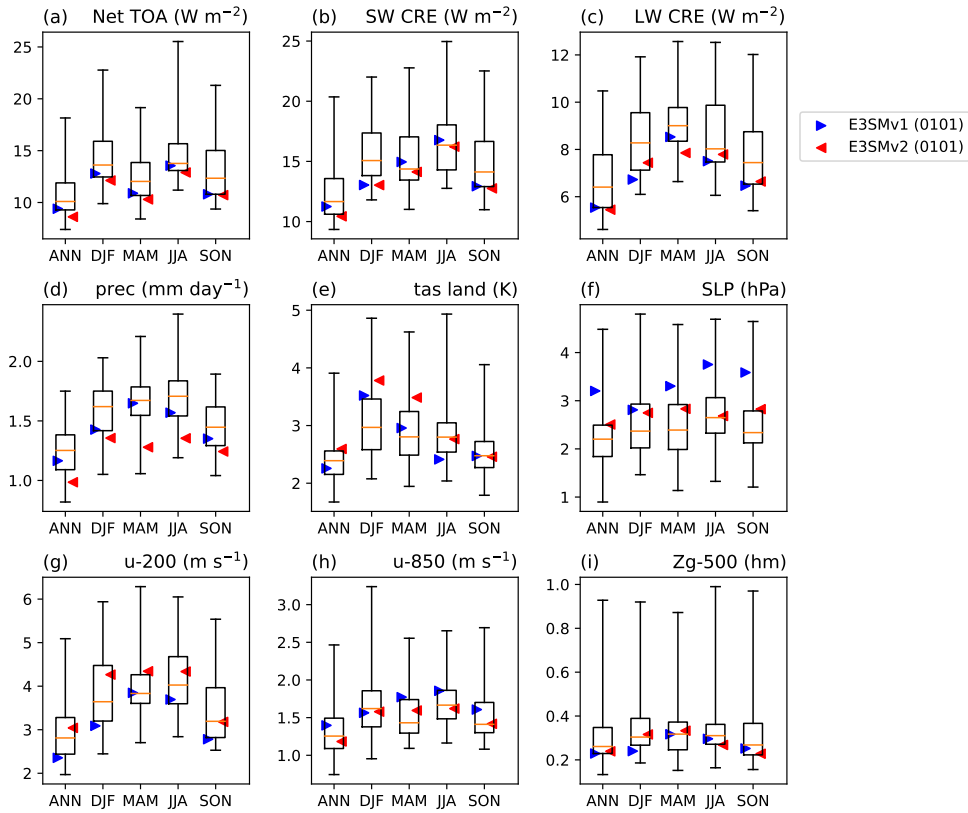
886 Figure 7 provides a broad overview of the model performance. Spatial RMSEs against  
 887 observations or reanalysis products are computed for annual and seasonal averages with  
 888 the E3SM Diagnostics package (C. Zhang et al., 2022). The first historical ensemble mem-  
 889 bers of E3SM are depicted with triangles, blue for E3SMv1 and red for E3SMv2. They  
 890 are compared against 52 CMIP6 models shown with box-and-whisker plots (minimum,  
 891 25th, 75th percentile, maximum). Underlying E3SM Diagnostics comparison figures are  
 892 available online ([https://portal.nersc.gov/project/e3sm/CMIP6\\_E3SMv2\\_Golaz\\_et](https://portal.nersc.gov/project/e3sm/CMIP6_E3SMv2_Golaz_et_al.2022)  
 893 [\\_al.2022](https://portal.nersc.gov/project/e3sm/CMIP6_E3SMv2_Golaz_et_al.2022)). For most fields, E3SMv2 outperforms E3SMv1. Notable improvements in-  
 894 clude precipitation and sea-level pressure. The simulated precipitation in E3SMv2 is now  
 895 competitive with the upper quartile of the CMIP6 ensemble. While sea-level pressure  
 896 is also much improved, it is still only about average compared to CMIP6. Consistent with  
 897 sea-level pressure, zonal wind at 850 hPa also improves. E3SMv2, similarly to E3SMv1,  
 898 has a good representation of TOA radiation fields, moderately improving upon v1 for  
 899 most fields and seasons. Unfortunately, two fields suffer from a degradation in E3SMv2  
 900 as compared to E3SMv1. For the zonal wind at 200 hPa, the degradation is partly as-  
 901 sociated with the change in stratospheric ozone chemistry (i.e., O3v2) (Tang et al., 2021,  
 902 their Figure 10), but the differences between E3SMv2 and E3SMv1 in Figure 7 are larger  
 903 than those between E3SMv1+O3v2 and E3SMv1, suggesting that other factors contribute



**Figure 5.** (a-b): time evolution of annual global mean surface air temperature anomalies for the idealized  $CO_2$  forcing simulations *abrupt-4xCO2* (red), *1pctCO2* (blue), and the control simulation (*piControl*; green) for E3SMv1 and E3SMv2. The transient climate response (TCR) is computed as a 20-year average around time of  $CO_2$  doubling (year 70). (c-d) Gregory regression to estimate effective climate sensitivity (ECS) and effective  $2xCO_2$  radiative forcing ( $F$ ) using the full 150 years as well as the first 20 years for the forcing ( $F_{1-20}$ ) and the last 130 years for the effective climate sensitivity ( $ECS_{21-150}$ ).



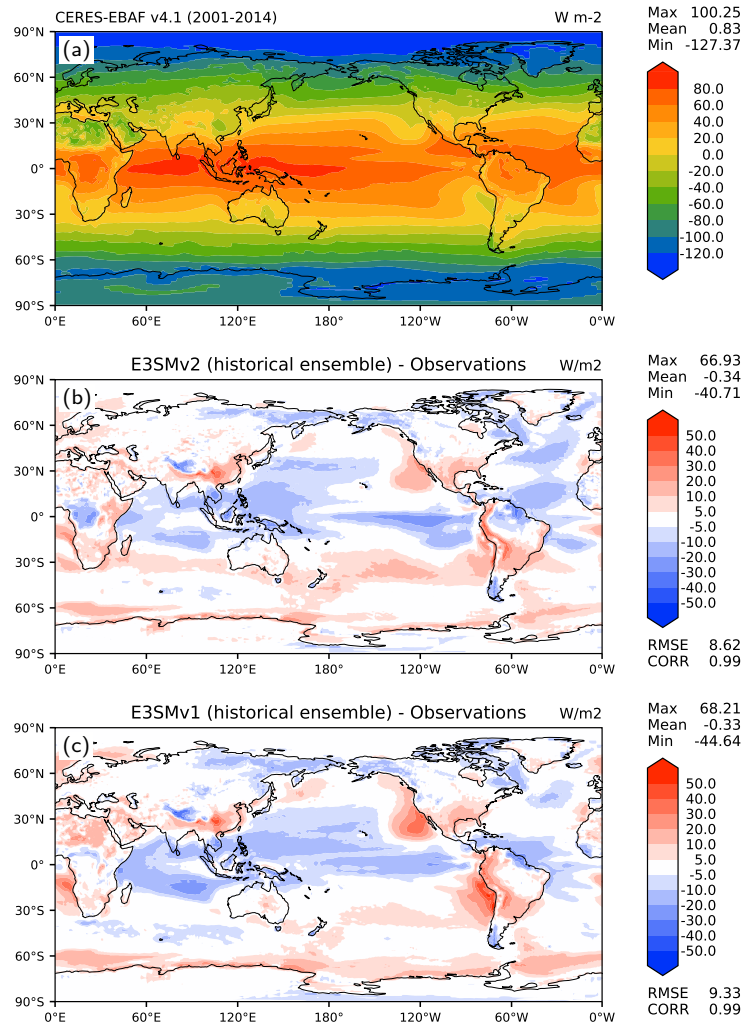
**Figure 6.** Time evolution of annual global mean total ERF (brown) and aerosol-related ERF (blue) for E3SMv1 and E3SMv2. Horizontal lines and adjacent values denote averages from 1995 to 2014.



**Figure 7.** Comparison of RMSEs (1985–2014) of an ensemble of 52 CMIP6 models (first historical members r1i1p1f1) with the first historical members of E3SMv1 (blue triangles) and E3SMv2 (red triangles). Box and whiskers show 25th, median, 75th percentile, minimum, and maximum RMSE for the CMIP6 ensemble. Spatial RMSE against observations are computed for annual and seasonal averages with the E3SM Diagnostics package (C. Zhang et al., 2022). Fields shown include TOA net radiation (a), TOA SW and LW cloud radiative effects (b, c), precipitation (d), surface air temperature over land (e), sea-level pressure (f), 200- and 850-hPa zonal wind (g, h), and 500-hPa geopotential height (i). TOA = top-of-atmosphere; SW = shortwave; CRE = cloud radiative effects; LW = longwave; DJF = December–February; MAM = March–April; JJA = June–August; SON = September–November; RMSE = root-mean-square error. The mean climatology of the reference observational and reanalysis datasets are derived from: CERES-EBAF Ed4.1 (Loeb et al., 2018) (2001–2014) for (a, b, and c), GPCP2.3 (Adler et al., 2018) (1985–2014) for (d) and ERA5 (Hersbach et al., 2020) (1985–2014) for (e, f, g, and h). Due to data availability, not all models are included for every variable. Complete data is available in Table S1.

as well. The degradation in surface air temperature over land is largely attributable to poor simulation of the historical temperature record (see Sections 4.3.6 and 5 below).

Selected fields from Fig. 7 are discussed in more detail in the subsections below. Comparison figures between observations, E3SMv1, and E3SMv2 for the remaining ones are available in the Supporting Information (Figs. S3 to S9).



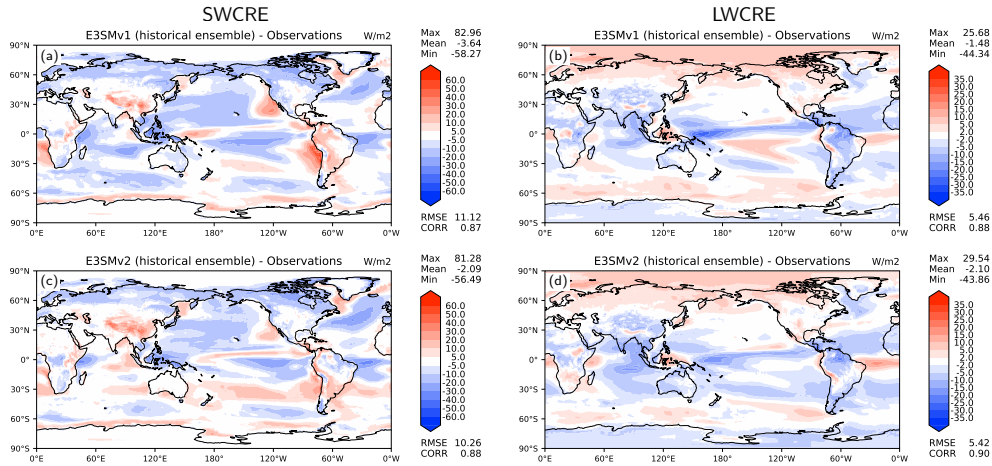
**Figure 8.** Annual net top-of-atmosphere (TOA) radiative flux ( $\text{W m}^{-2}$ ): (a) CERES-EBAF Ed4.1 observational estimate (2001-2014), (b) model bias from the 5-member ensemble of E3SMv2 historical coupled simulations (2001–2014), and (c) model bias from the 5-member ensemble of E3SMv1 historical coupled simulations (2001–2014). RMSE = root-mean-square error. CORR = linear correlation coefficient between observation and model.

#### 4.3.1 Radiation and Clouds

Annual net TOA radiative flux (2001-2014) in E3SMv1 and v2 is depicted in Figure 8 in comparison with observations from CERES-EBAF Ed4.1 (Loeb et al., 2018).



912 The simulated global mean value is nearly identical between the two versions at  $+0.5$   
 913  $\text{W m}^{-2}$ , lower than the observational estimate (but consistent with the smaller warming;  
 914 Figure 6 and Section 4.3.6). Many regional biases are reduced in E3SMv2, includ-  
 915 ing positive biases over stratocumulus regions, as well as negative biases over tropical  
 916 and subtropical Pacific, Indian, and Atlantic oceans, resulting in an overall smaller RMSE  
 917 ( $8.6$  vs  $9.3 \text{ W m}^{-2}$ ).



**Figure 9.** Annual top-of-atmosphere shortwave and longwave cloud radiative effect model biases (historical ensemble, 2001-2014) compared to CERES-EBAF Ed4.1 observational estimate (2001-2014) ( $\text{W m}^{-2}$ ): (a) SWCRE E3SMv1, (b) LWCRE E3SMv1, (c) SWCRE E3SMv2, (d) LWCRE E3SMv2. RMSE = root-mean-square error. CORR = linear correlation coefficient between observation and model.

918 Figure 9a,c demonstrate that the TOA shortwave cloud radiative effect is improved  
 919 in the E3SMv2 historical ensemble compared with E3SMv1 in terms of RMSE and the  
 920 pattern correlation. Overall, the global mean SWCRE in E3SMv2 is weaker and closer  
 921 to observations by  $\sim 1.5 \text{ W m}^{-2}$ . The positive TOA SWCRE bias associated with the  
 922 stratocumulus decks over eastern ocean basins, especially right off the coasts of Califor-  
 923 nia, Peru and Chile, and the southern West Africa, is clearly reduced, while the nega-  
 924 tive SWCRE bias associated with the cumulus regimes over central/western tropical oceans  
 925 is slightly reduced as well. The improvement in the marine boundary layer cloud regimes  
 926 is mainly from the updated CLUBB tuning parameters (Ma et al., 2022). The global mean  
 927 TOA longwave cloud radiative effect (LWCRE) bias is weaker by  $\sim 0.6 \text{ W m}^{-2}$ , slightly  
 928 degraded compared to E3SMv1 while RMSE remains comparable (Fig. 9b,d). Regional  
 929 biases over the equatorial Pacific and the intertropical convergence zone (ITCZ) are re-  
 930 duced associated with the improved precipitation over these areas (described in Section 4.3.2  
 931 below). The positive TOA LWCRE bias is also slightly reduced over the Southern Ocean.

932 The enhanced Wegener-Bergeron-Findeisen (WBF) efficiency and the update to  
 933 the ZM scheme significantly increase ice water in mixed-phase clouds, which also weak-  
 934 ens SWCRE in the Southern Hemisphere (e.g.  $\sim 30^\circ \text{S}$  in Figure 9). The liquid conden-  
 935 sate fraction (LCF) as a function of temperature at all latitudes between  $30^\circ \text{S}$ – $80^\circ \text{S}$  (Fig-  
 936 ure 10) from both E3SMv1 and E3SMv2 historical coupled simulations demonstrate that  
 937 the updated atmosphere features and tuning parameters in E3SMv2 significantly increase  
 938 ice cloud mass fraction in the temperature range between  $-50^\circ \text{C}$  and  $-10^\circ \text{C}$ , which is  
 939 closer to the general pattern of the observational estimate (Y. Zhang et al., 2019). Note  
 940 that the LCFs from the E3SMv1 historical coupled simulation and the E3SMv2 histor-

941 ical coupled simulation are both calculated throughout the whole cloud layer with the  
 942 monthly model output at  $1^\circ$  resolution, while the observation from (Y. Hu et al., 2010)  
 943 is based on cloud top estimates at 5-km resolution. Given the mismatch between the model  
 944 diagnostics and observational estimates and the broad range of observed cloud phase par-  
 945 tition from previous observational estimates (McCoy et al., 2016), it is more reasonable  
 946 to compare the v2 LCF with the v1 LCF.

947 We further quantify the improvements in the subtropical stratocumulus decks com-  
 948 pared to E3SMv1 following Brunke et al. (2019). We define the decks as the areas within  
 949  $30^\circ$  latitude by  $35^\circ$  longitude boxes in the Northeast Pacific (NEP;  $10^\circ$ - $40^\circ$ N,  $110^\circ$ - $145^\circ$ W),  
 950 Northeast Atlantic (NEA;  $0^\circ$ - $30^\circ$ N,  $15^\circ$ - $50^\circ$ W), Southeast Pacific (SEP;  $5^\circ$ - $35^\circ$ S,  $70^\circ$ -  
 951  $105^\circ$ W), Southeast Atlantic (SEA;  $5^\circ$ - $35^\circ$ S,  $20^\circ$ W- $15^\circ$ E), and the Southern Indian Ocean  
 952 (SIO;  $10^\circ$ - $40^\circ$ S,  $80^\circ$ - $115^\circ$ E) where low cloud cover  $> 45\%$ , the LCC45+ decks, over both  
 953 the land and ocean portions of the boxes. E3SMv2 LCC from the Cloud Feedback Model  
 954 Intercomparison Project Observation Simulator Package (COSP) Cloud-Aerosol Lidar  
 955 and Infrared Pathfinder (CALIPSO) satellite simulator is generally improved, falling more  
 956 within the observational spread represented by three satellite and in-situ based clima-  
 957 tologies [the CALIPSO satellite GCM-Oriented CALIPSO Cloud Product (GOCCP),  
 958 the International Satellite Cloud Climatology Project (ISCCP) D2 product, and the Ex-  
 959 tended Edited Cloud Reports Archive (EECRA)] (Figure 11).

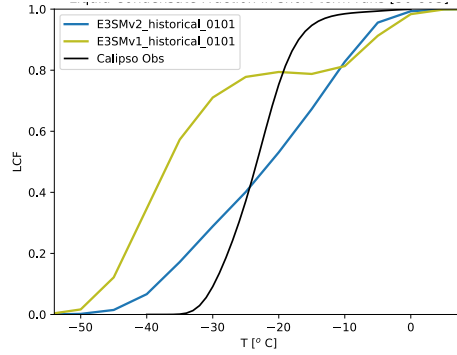
960 The cloud changes that lead to the SWCRE improvements can be explained by the  
 961 spatial errors in the simulated LCC45+ cloud decks with respect to GOCCP which are  
 962 defined as in Brunke et al. (2019). An example of these for the seasons of maximum LCC  
 963 for each region in Figure 11 is given in Figure 12. For “apples-to-apples” comparisons,  
 964 the model output from the COSP CALIPSO satellite simulator is used. Centroid dis-  
 965 tances (Figure 12a) measure the distance between the centroid of the seasonal mean cloud  
 966 deck in GOCCP and the model. Smaller centroid distances are better than large ones.  
 967 Area ratios (Figure 12b) are the ratio of the area of the model’s deck to that of the satel-  
 968 lite to measure cloud deck size errors. Finally, overlap ratios (Figure 12c) are the frac-  
 969 tion of the union of the model and satellite cloud decks in which there is overlap. This  
 970 synthesizes the effects of location, size, and shape errors in the simulated cloud decks.  
 971 Both of these ratios should be close to 1 for minimal errors.

972 Figure 12 shows that E3SMv2 improves most the representation of the widely stud-  
 973 ied subtropical stratocumulus cloud decks in the NEP, NEA, and SEP. In these regions,  
 974 centroid distances are decreased and overlap ratios are similar to or increased to values  
 975 closer to 1. Area ratios are improved in all regions with values closer to 1 except NEA.  
 976 Similar results are found in all other seasons.

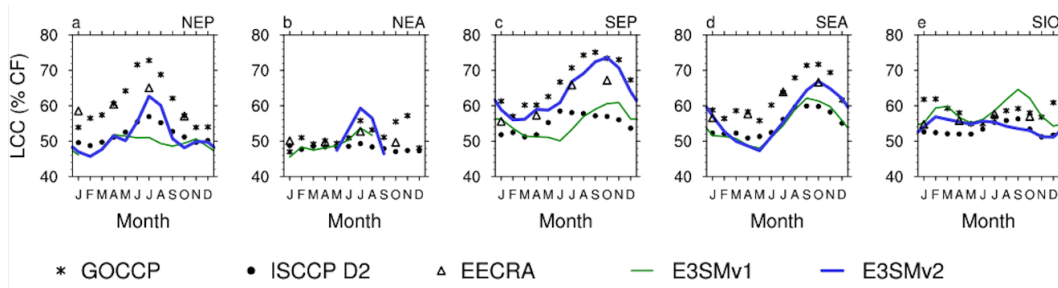
#### 977 *4.3.2 Precipitation*

978 The model bias in annual precipitation from E3SMv2 shows notable improvement  
 979 compared with that in E3SMv1 (Figure 13). The biases are clearly reduced in the Trop-  
 980 ical Pacific ocean, Maritime continent, Central America, and the Amazon. The updated  
 981 ZM tuning parameters, the dCAPE-ULL convective trigger, and the inclusion of the gusti-  
 982 ness effects and the subgrid temperature variance are found to reduce the regional bi-  
 983 ases of annual mean precipitation (Xie et al., 2019; Ma et al., 2022).

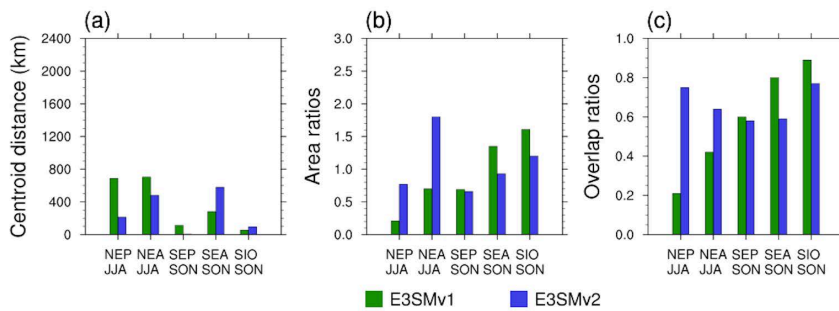
984 As described in Section 2.1.3, the dCAPE-ULL convective trigger is expected to  
 985 broadly improve the simulation of diurnal precipitation. This can be clearly seen in Fig-  
 986 ure 14, which shows the comparison of the time phase (color) and amplitude (color den-  
 987 sity) of diurnal precipitation between TRMM, and E3SMv2 and E3SMv1 historical simu-  
 988 lations over the tropics. Note that the precipitation from the models are the sum of con-  
 989 vective and large-scale precipitation. The diurnal characteristics are dictated by the con-  
 990 vective precipitation and the dCAPE-ULL trigger has little impact on the diurnal cy-  
 991 cle of large-scale precipitation. The improvements from E3SMv1 to E3SMv2 are most



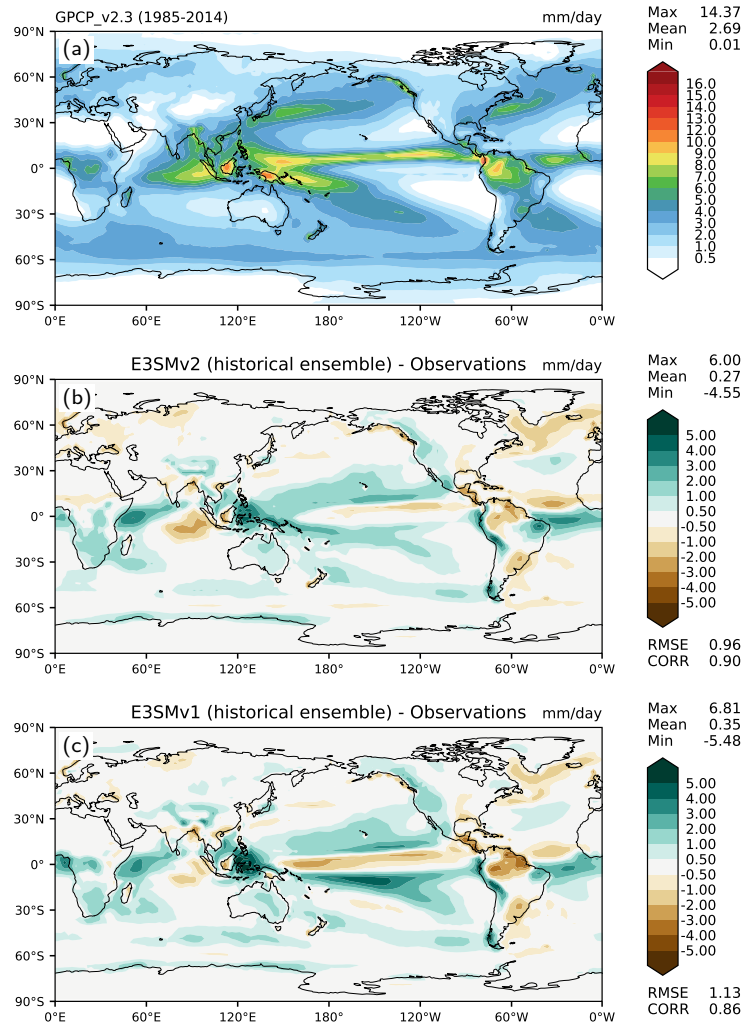
**Figure 10.** Diagnosed liquid cloud fraction (LCF) based on the monthly model output in the 30–80°S latitude band from (blue line) the E3SMv2 historical coupled simulation (1985–2014), (olive line) the E3SMv1 historical coupled simulations (1985–2014), and (black line) observations from (Y. Hu et al., 2010).



**Figure 11.** The mean low cloud cover (LCC) for each of the LCC45+ cloud decks (see text for definitions) for the 30° latitude by 35° longitude boxes over the Northeast Pacific (NEP), Northeast Atlantic (NEA), Southeast Pacific (SEP), Southeast Atlantic (SEA), and the Southern Indian Ocean (SIO).

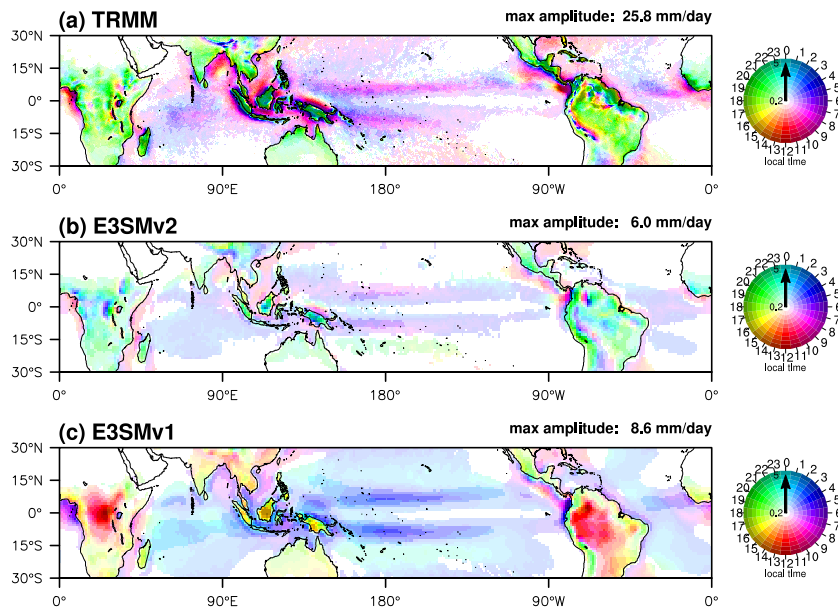


**Figure 12.** Centroid distances, area ratios, and overlap ratios of the LCC45+ decks in June–August (JJA) for the two Northern Hemisphere regions and in September–November (SON) for the Southern Hemisphere regions.



**Figure 13.** Annual precipitation rate (mm/day): (a) Global Precipitation Climatology Project v2.3 observational estimate (1985–2014), (b) model bias from the 5-member ensemble of E3SMv2 historical coupled simulations (1985–2014), and (c) model bias from the 5-member ensemble of E3SMv1 historical coupled simulations (1985–2014). RMSE = root-mean-square error. CORR = linear correlation coefficient between observation and model.

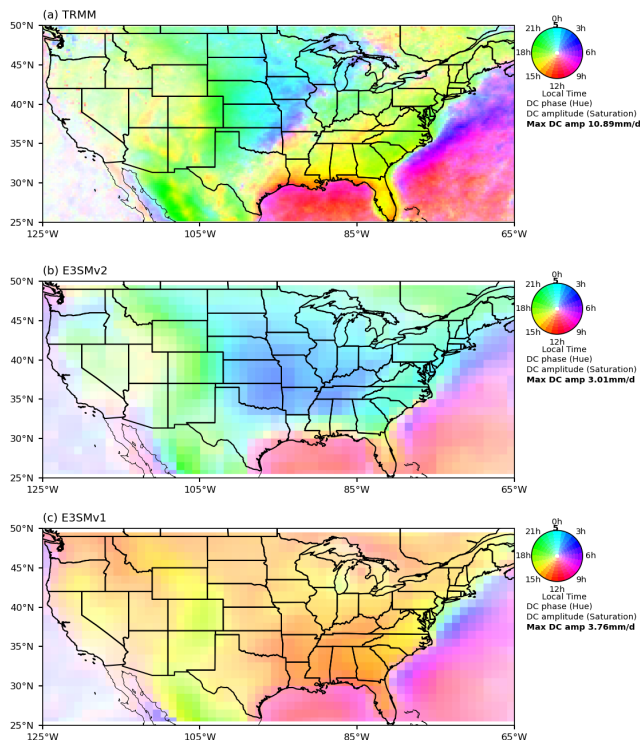
992 evident in the diurnal peak phase. Over the oceans, E3SMv2 captures the observed widespread  
 993 morning peaks, particularly along the primary precipitation bands, where the peak pre-  
 994 cipitation occurs several hours too early (closer to midnight) in E3SMv1. Over the Mar-  
 995 ritime continent region, E3SMv2 closely reproduces the observed early evening peaks over  
 996 land and the transition to morning peaks towards the coasts and open oceans, while E3SMv1  
 997 has too-early diurnal precipitation peaks from noon to early afternoon over land and sim-  
 998 ilarly much earlier peaks around midnight in the coastal regions. Over the tropical contin-  
 999 tents, including Africa, South America, and South Asia, the observed diurnal peaks  
 1000 occur from late evening to midnight. While the diurnal precipitation peaks in E3SMv1  
 1001 are nearly phase-locked to maximum insolation over these land masses, the phase-locking  
 1002 behaviors are avoided in E3SMv2, which sees the peak phases near midnight or shortly  
 1003 after. However, the improvement in simulating diurnal timing phases does not translate  
 1004 to diurnal amplitude. This is presumably due to lack of skill in simulating meso-scale  
 1005 convective systems in coarse resolution models. Furthermore, while the diurnal ampli-  
 1006 tudes are weaker in both models compared to observations, they are somewhat degraded  
 1007 from E3SMv1 to E3SMv2 particularly over weakly precipitating subtropical oceans.



**Figure 14.** Annual mean time phase (color) and amplitude (color density) of the first diurnal harmonic of 3-hourly total precipitation (mm/day) from (a) TRMM (1998-2013), and historical simulations (1985-2014) of (b) E3SMv2 and (c) E3SMv1. Note that the diurnal properties are computed at 0.25 degree resolution for TRMM and 1 degree resolution for the models. Remapping the TRMM data to 1 degree resolution does not lead to noticeable changes in the diurnal phases, though the amplitudes would be reduced somewhat due to spatial averaging (maximum amplitude over the tropics reduced to 18.9 mm/day from 25.8 mm/day). To ease comparison, the same amplitude upper bound is used for the plots, though the maximum amplitudes, as printed above each panel, are different. Amplitudes exceeding the upper bound are shown in the highest color density. Areas with diurnal amplitude less than 0.2 mm/day are left blank.

1008 The improvements in summertime diurnal precipitation peak phase in E3SMv2 over  
 1009 mid-latitude landmass such as the continental United States (Fig.15) is largely consis-  
 1010 tent with what were shown in (Xie et al., 2019) where the dCAPE-ULL trigger was first  
 1011 implemented in E3SMv1. Notably with the new trigger, E3SMv2 is able to produce the

1012 successive delay of precipitation peak phases from the high mountains to the plains across  
 1013 the central U.S.. In particular, over the U.S. Great Plains, the new trigger successfully  
 1014 captures the nocturnal precipitation peak, which has been missed by most climate mod-  
 1015 els. However, the observed late afternoon peak over the Eastern and Southeastern U.S.  
 1016 is missed in E3SMv2. Instead, it produces a late evening peak over the regions. This is  
 1017 in contrast with the well simulated late afternoon peak over the regions as shown in (Xie  
 1018 et al., 2019). Sensitivity experiments indicate this degradation is primarily caused by the  
 1019 re-tuning of a parameter that effectively changes the air parcel launch level from 2 lev-  
 1020 els above the bottom model level to just one level above (E3SMv1p; Ma et al., 2022).  
 1021 The issue is being further investigated and addressed by the E3SM development team.  
 1022 An initial test with further enhancements to the ZM scheme has indicated that a much-  
 1023 improved diurnal cycle of precipitation could be achieved in the next version of E3SM.



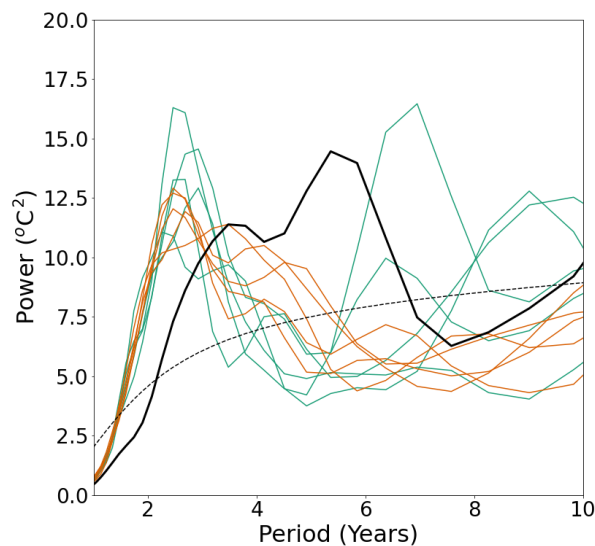
**Figure 15.** Same as Fig. 14 except for the contiguous United States in June-July-August season.

1024 After the model was finalized, it was observed that the dCAPE trigger, independent  
 1025 of the ULL trigger and other model settings, including the new physics pg2 grid,  
 1026 induces a checkerboard grid-level noise pattern in a number of output fields, including  
 1027 total grid-box cloud water liquid and ice paths, when these fields are temporally instan-  
 1028 taneous or averaged over not more than several days (Hannah et al., 2022). Figure S10  
 1029 illustrates this issue by comparing a daily average output of the total grid-box cloud liq-  
 1030 uid water path in four lower resolution atmosphere simulations over two binary experi-  
 1031 mental settings: with the dCAPE trigger on and off, on the new pg2 physics grid and  
 1032 on the original np4 physics grid. These checkerboard noises occur mostly in the sub-tropics  
 1033 where convection is naturally weak and sparse. As indicated in many earlier studies, the  
 1034 ZM scheme with its default CAPE trigger frequently produces spurious weak precipita-  
 1035 tion in the subtropical region. Such spurious precipitation is mostly suppressed with  
 1036 the use of the dCAPE trigger, which adds a dynamical constraint to suppress convec-

1037 tion when the large-scale environment is not favorable. As a result, the weak convection  
 1038 that is still triggered when using dCAPE becomes more isolated, leading to the appar-  
 1039 ent checkerboard noise when visualized spatially. Although this is largely an indication  
 1040 that the dCAPE trigger is working as intended, we are testing a few approaches to re-  
 1041 duce the checkerboard problem, including setting a threshold for dCAPE (currently zero)  
 1042 and considering the impact of the large-scale dynamic forcing on cloud base mass fluxes.  
 1043 We will attempt to address this problem in our next release of the model.

#### 1044 4.3.3 Tropical variability

1045 As in (Golaz et al., 2019) we examine the E3SMv2 variability of the El Niño South-  
 1046 ern Oscillation (ENSO) via wavelet analysis (Torrence & Compo, 1998) of the Niño 3.4  
 1047 SST for the *piControl* and historical simulations in Fig. 16. In this figure the *piControl*  
 1048 has again been divided into five 100-year intervals. The 90% confidence interval is shown  
 1049 as the dashed black line. ENSO variability in E3SMv2 shows a number of similarities  
 1050 to E3SMv1 (compare to Golaz et al., 2019, their Fig. 20). Again E3SMv2 shows a very  
 1051 robust peak of variability at short periods ( $\sim 2.5$  years), which is similar to E3SMv1 and  
 1052 shorter than ERSSTv4 (thick black line). While a longer period (6-9 years) remains in  
 1053 the *piControl*, the mean for the five 100-year intervals has reduced relative to E3SMv1.  
 1054 This longer term variability is weaker than simulated in other CMIP5 and CMIP6 mod-  
 1055 els (see Orbe et al., 2020, their Fig. 10a) and observations (black line in Fig. 16). The  
 1056 intermediate periods (3-6 years) seen in ERSSTv4 are not well captured in E3SMv2. The  
 1057 spatial SST response to ENSO is shown in Fig. S11. The magnitude of the SST response  
 1058 (approximately  $2.5^{\circ}\text{C}$ ) in the *piControl* and historical ensemble mean (panels b and c)  
 1059 is consistent with E3SMv1, other CMIP models, and observations (Golaz et al., 2019;  
 1060 Brown et al., 2020). However, the center of the response is shifted too far westward, which  
 1061 is consistent with other models.



**Figure 16.** El Niño–Southern Oscillation (ENSO; Niño3.4) variability of the pre-industrial (PI) control simulation and historical ensemble. The Morlet wavelet of degree 6 is used (e.g., Torrence & Compo, 1998). The PI control (green lines) has been divided into five 100-year sections; each Historical ensemble member is shown as an orange line. ERSSTv4 data (W. Liu et al., 2015) is shown as the thick black line. The 90% confidence interval is shown as the dashed black line.

1062 The Madden-Julian oscillation (MJO; Madden & Julian, 1971), the dominant mode  
 1063 of tropical variability on subseasonal (10-100 day) scales, is a key contributor to ENSO  
 1064 events (C. Zhang & Gottschalck, 2002), monsoon activity (Wheeler & McBride, 2012),  
 1065 extratropical atmospheric blocking episodes (Henderson et al., 2016), tropical cyclone  
 1066 formation (Maloney & Hartmann, 2000), and weather extremes (Higgins et al., 2000; Mat-  
 1067 sueda & Takaya, 2015; Mundhenk et al., 2016). Its accurate representation in numer-  
 1068 ical models is essential for weather and climate prediction (Vitart & Robertson, 2018),  
 1069 yet a satisfactory depiction of the MJO remains elusive (Jiang et al., 2015; Ahn et al.,  
 1070 2020). Figure 17 shows the distribution of tropical precipitation spectral power, normal-  
 1071 ized by a smoothed background spectrum, in zonal wavenumber-frequency space (Wheeler  
 1072 & Kiladis, 1999). Model-observation comparisons span 2001-2010 within the TRMM satel-  
 1073 lite era (Huffman et al., 2001), but comparisons between E3SM versions use the extended  
 1074 period 1985-2015; our conclusions do not change whether the shorter or longer time win-  
 1075 dows is used. Results from an E3SMv2 historical simulation (Fig. 17b) indicate slightly  
 1076 lower power values for equatorial Rossby waves and the MJO and a MJO peak that is  
 1077 at a higher frequency compared to observations (Fig. 17a). Relative to an E3SMv1 his-  
 1078 torical simulation (see Golaz et al. (2019) and Orbe et al. (2020) for details), precipita-  
 1079 tion normalized power in the broad MJO spectral region has increased and shifted to higher  
 1080 frequencies (Fig. 17c). Both E3SMv2 and E3SMv1 dramatically underestimate precip-  
 1081 itation variability associated with atmospheric Kelvin waves and other synoptic-scale dis-  
 1082 turbances. In AMIP simulations, MJO spectral power is again larger in E3SMv2 com-  
 1083 pared to E3SMv1 (Fig. 17e) yet, unlike the historical runs, it is not shifted to higher fre-  
 1084 quencies and thus it is more realistic; for other wave types, intraseasonal variability bi-  
 1085 ases in E3SM AMIP and historical simulations are generally similar.

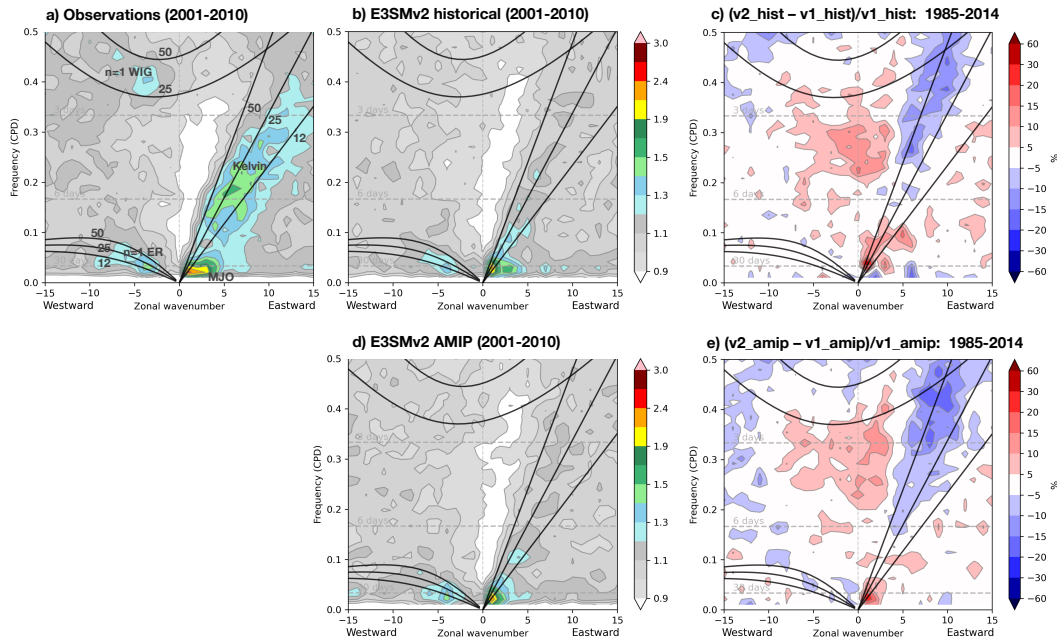
1086 Lag correlations of equatorial precipitation and 850 hPa zonal wind with Indian Ocean  
 1087 precipitation (Figure 18) suggest some improvement in MJO propagation across the Mar-  
 1088 itime Continent in E3SMv2 compared to E3SMv1, as evidenced by more consistent red  
 1089 shading eastward to 125°E. MJO eastward propagation in AMIP simulations (Figs. 18d,e),  
 1090 particularly across the Indian Ocean and Maritime Continent between 50°–120°E, is much  
 1091 improved in E3SMv2. In both E3SMv2 and E3SMv1, the quadrature phasing of precip-  
 1092 itation and zonal wind resembles that in observations, but the MJO phase speed begins  
 1093 to exceed the observed 5.5 m s<sup>-1</sup> reference value (dashed green line) east of 120°E and  
 1094 especially in E3SMv2. E3SMv2 historical simulation results are similar to those for CESM2  
 1095 reported in Danabasoglu et al. (2020). A more detailed evaluation of tropical subseasonal  
 1096 variability in E3SMv2 will be presented in a forthcoming manuscript.

#### 1097 4.3.4 Ozone

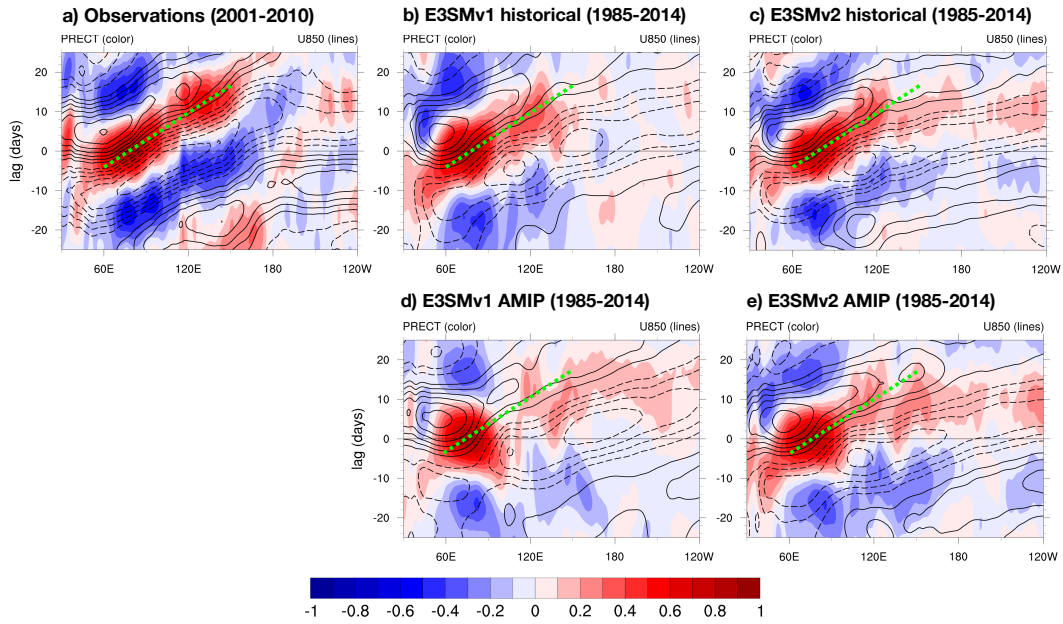
1098 The stratospheric column ozone (SCO) of the historical ensemble mean of E3SMv2  
 1099 is compared with the satellite observations from the Ozone Monitoring Instrument (OMI)  
 1100 and the Microwave Limb Sounder (MLS) at 60°S to 60°N, where the satellite observa-  
 1101 tions have good quality all year round. Figure 19 shows the climatology of SCO zonal  
 1102 mean annual cycle from years 2005–2014 of E3SMv2 historical simulations and years 2005–  
 1103 2014 of the OMI+MLS observations. The E3SMv2 historical simulations match the ob-  
 1104 served SCO seasonal phase and pattern, but generally overestimate the SCO magnitude  
 1105 except at 40°N – 60°N in all months and near 30°S from March to September. Compar-  
 1106 ing to the E3SMv1 SCO in Fig. 1d of Tang et al. (2021), the E3SMv2 SCO better matches  
 1107 observations in the SH mid-latitudes, but is worse in the NH mid-latitudes. This E3SMv1-  
 1108 E3SMv2 difference in the SCO is likely associated with the QBO and GW retuning for  
 1109 the E3SMv2.

1110 The evolution of the Antarctic ozone hole during the historical time period reflects  
 1111 the combined effect of dynamics, physics, and chemistry. The NASA Ozone Watch web-  
 1112 site (<https://ozonewatch.gsfc.nasa.gov>, last access: October 11, 2021) archives the daily  
 1113 records of the Antarctic ozone hole area (where the total column ozone (TCO) is less than

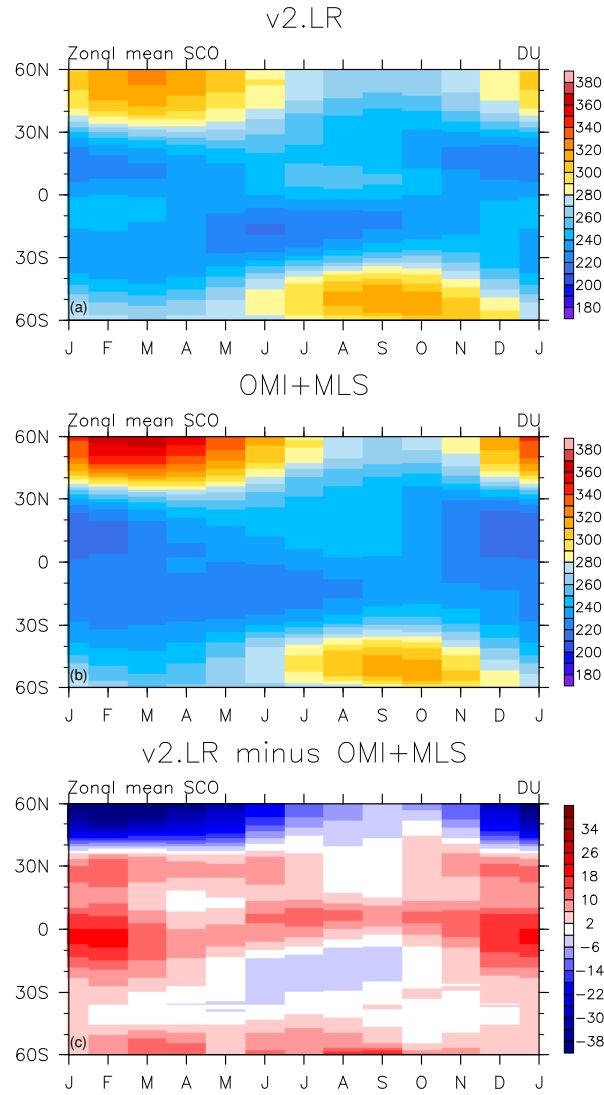




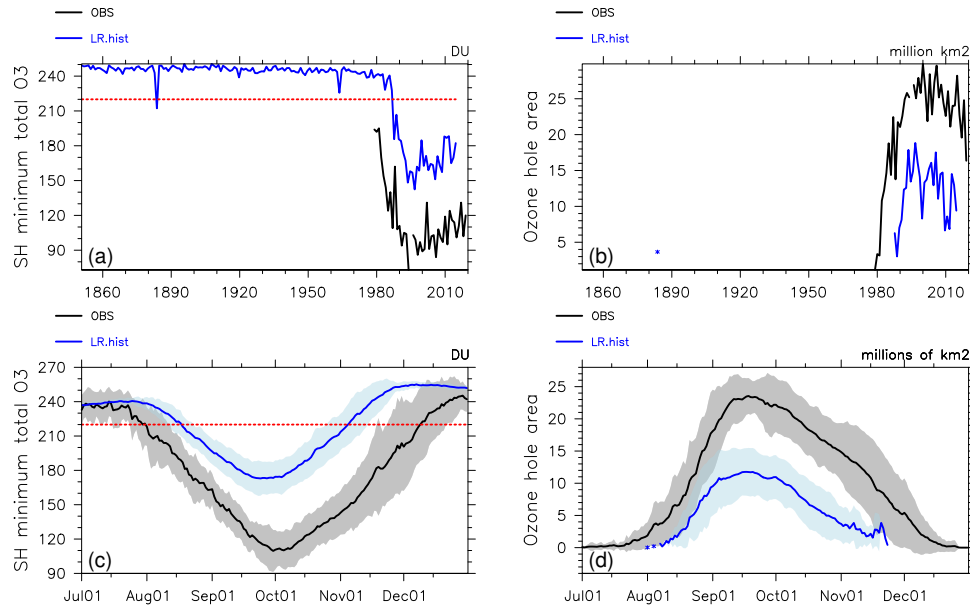
**Figure 17.** Tropical zonal wavenumber-frequency power spectra of the component of precipitation that is symmetric about the Equator for the period 2001-2010 for (a) observations (Tropical Rainfall Measuring Mission product 3B42v7) from 2001-2010 and (b) an E3SMv2 historical simulation, and (d) an E3SMv2 AMIP simulation. Plotted values represent the summed power from  $15^{\circ}\text{S}$ – $15^{\circ}\text{N}$  divided by the smoothed background power (the “normalized” power). Solid black lines indicate shallow water dispersion curves for equivalent depths of 12, 25, and 50 m. Prominent wave types are labeled: westward inertia-gravity ( $n=1$  WIG), Kelvin, equatorial Rossby ( $n=1$  ER), and the Madden-Julian oscillation (MJO). (c,e) The change, expressed as a percent difference, in the normalized spectral power between E3SMv2 and E3SMv1 (c) historical simulations and (e) AMIP simulations for the extended period 1985–2014.



**Figure 18.** Latitudinally averaged ( $10^{\circ}\text{S}$ – $10^{\circ}\text{N}$ ) precipitation (colors) and 850 hPa zonal wind (lines) anomalies lag correlated with precipitation in the Indian Ocean region ( $60^{\circ}$ – $90^{\circ}\text{E}$ ,  $10^{\circ}\text{S}$ – $10^{\circ}\text{N}$ ) for (a) observations from 2001–2010 (Tropical Rainfall Measuring Mission [TRMM] precipitation and Modern-Era Retrospective Analysis for Research and Applications [MERRA] wind), and the 1985–2014 period for (b) E3SMv1 historical, (c) E3SMv2 historical, (d) E3SMv1 AMIP, and (e) E3SMv2 AMIP simulations. The dashed green line in (a) represents the observed Madden-Julian oscillation phase speed ( $5.5\text{ m s}^{-1}$ ) in precipitation and is copied to panels (b) and (c) for reference. The line contour interval is 0.1, solid lines indicate positive correlations, dashed lines indicate negative correlations, and the zero correlation line is omitted. Anomalies, defined as departures from the smoothed seasonal cycle, are bandpass filtered to retain 20–100 day signals prior to correlation.



**Figure 19.** Climatology of zonal mean annual cycle of stratospheric column ozone (SCO, in Dobson units (DU)). The panels show data of years 2005–2014 from (a) E3SMv2 ensemble mean of historical simulations, (b) OMI+MLS observations, and (c) the differences of E3SMv2 minus OMI+MLS.



**Figure 20.** Ozone hole results as shown in the historical time series (top) and daily mean climatology and variance (bottom) of the SH minimum total column ozone (left, unit: DU) and the SH maximum ozone hole area (right, area with total ozone < 220 DU, unit: million km<sup>2</sup>) based on the daily data from July 1 to December 31. In the bottom panels, the lines indicate the multi-year (1990–2014) average (observations in black and models in blue), and shading covers  $\pm 1$  standard deviation.

1114 220 DU) and minimum TCO in the SH based on daily TCO observational data. Figures  
 1115 20a and b compare the yearly E3SMv2 historical ensemble mean time series with the yearly  
 1116 Ozone Watch observations for the SH minimum TCO and the ozone hole area, respec-  
 1117 tively. Both the yearly model and observational results are based on the daily data from  
 1118 July 1 to December 31 of each year.

1119 The Antarctic ozone hole emerges about 1980 after the buildup of anthropogenic  
 1120 chlorofluorocarbons (CFCs) reach a threshold that initiates rapid, catalytic destruction  
 1121 of ozone within the Antarctic stratospheric polar vortex (Molina & Rowland, 1974; Far-  
 1122 man et al., 1985). The ozone hole simulation in E3SMv2 is weaker than observed in terms  
 1123 of minimum TCO (Figure 20a,c) and areal extent of the ozone hole (Figure 20b,d). Given  
 1124 the 50 DU high bias for ozone-hole minimum TCO (Figure 20c), the temporal history  
 1125 of the ozone hole, from onset to partial recovery, is well matched in E3SMv2 (Figure 20a).  
 1126 In terms of seasonality, the E3SMv2 ozone hole begins almost a month later and recov-  
 1127 ers almost a month earlier. The cause of this is not the ozone chemical model, as it works  
 1128 well in other atmospheric models, but is likely related to the formation and persistence  
 1129 of the wintertime vortex. The ozone hole is created chemically, but its size and duration  
 1130 depend on the vortex remaining isolated from the mid-latitude stratosphere throughout  
 1131 most of the lower stratosphere. The E3SMv2 ozone hole interannual variability (IAV,  
 1132 shaded areas in Figure 20c,d), scaled to the size of the ozone hole, matches the obser-  
 1133 vations, indicating that the vortex IAV is similar to observations. It is possible that the  
 1134 weaker ozone hole in E3SMv2 could be improved with a colder stratosphere, or paramet-  
 1135 rically, by increasing the PSC temperature threshold.

**Table 2.** Global and annual mean AOD at 550 nm for total aerosol and major aerosol types.

AOD (2000-2014)	Total	Dust	Sea salt	Sulfate	POM <sup>a</sup>	BC <sup>a</sup>	SOA <sup>a</sup>
E3SMv1 (DECK)	0.146	0.032	0.049	0.024	0.007	0.0049	0.029
E3SMv2 (historical)	0.166	0.028	0.049	0.033	0.009	0.0063	0.040

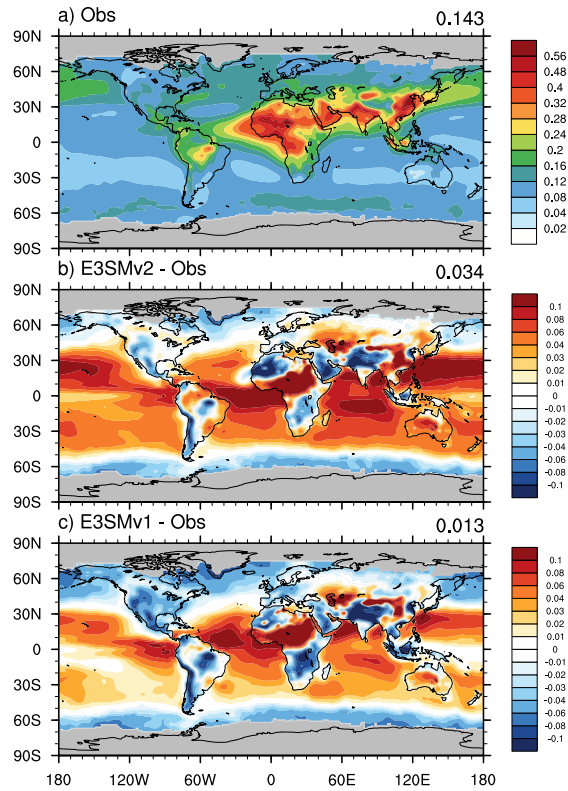
<sup>a</sup>POM (particulate organic matter), BC (black carbon), and SOA (secondary organic aerosol)

#### 4.3.5 Aerosols

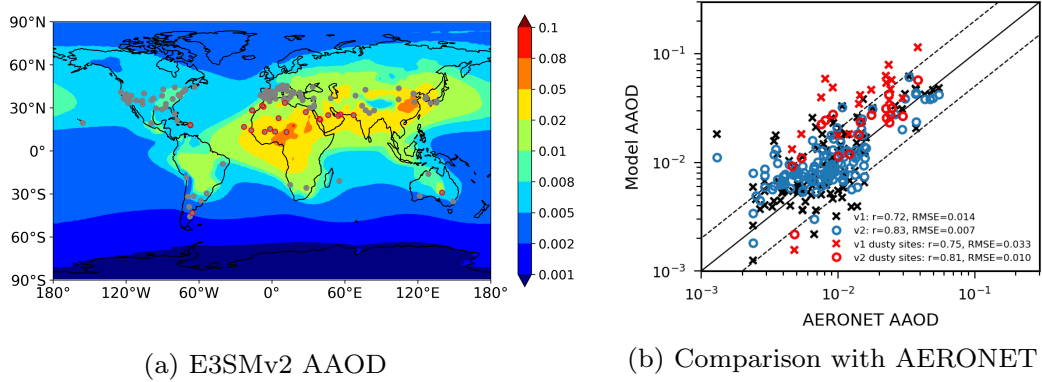
The global distribution of annual mean AOD at 550 nm from E3SMv2 and E3SMv1 historical simulations (2000-2014) is compared with observational composite (Kinne et al., 2013) in Figure 21. Model results are not included for this comparison over regions where the observations are not available, e.g., in the high latitudes. E3SMv1 and v2 realistically capture the broad regional distribution in AOD, but E3SMv2 has a stronger positive bias than E3SMv1 in the global mean (0.034 vs. 0.013) compared to the observational composite, although the low bias over mid-latitude source regions is improved in E3SMv2. Larger positive biases in E3SMv2 than E3SMv1 are found over tropical and subtropical oceans. Decomposition of the total AOD into major aerosol species is provided in Table 2. The positive biases are mostly due to an increase in anthropogenic aerosol species, particularly sulfate and secondary organic aerosol (SOA). The global annual mean burdens of sulfate and SOA have an increase of 1.03 and 0.95 Tg, respectively, in the E3SMv2 historical simulations (2000-2014) compared to E3SMv1 (Fig. S12). The global annual mean burdens of other anthropogenic aerosol species are also larger in E3SMv2 than those in E3SMv1, although both model simulations use the same set of CMIP6 emissions, indicating that the aerosol removal in E3SMv2 is weaker than in E3SMv1. This might be an unintended consequence of intensive cloud and precipitation parameter tuning for EAMv2. Salzmann et al. (2022) recently reported on the impact of moist convection parameter settings on wet deposition and as a result on AOD.

Natural aerosols (e.g., dust and sea salt) are also affected, as shown in Fig. S12, although their global burdens have small changes, contributing less than the anthropogenic aerosols to the positive bias in global mean AOD in E3SMv2 (Table 2). This is because the emissions of dust and sea salt are scaled to match the same global constraints of their optical depth in both E3SMv1 and E3SMv2. On the regional scales, sea salt burden increases in E3SMv2 over some portions of tropical and subtropical oceans, e.g., Indian Ocean, which may contribute to the higher AOD in those regions. Similar to sulfate and SOA, the increase in sea salt is related to the changes in aerosol wet deposition, as those regions with increased burdens are associated mostly with tropical convection and precipitation. With the global constraint, the positive changes in sea salt are offset by the negative changes between E3SMv2 and E3SMv1 such as over the marine stratus or stratocumulus regions in the mid-latitude and sub-tropical oceans. Regional changes are also found in dust aerosol burden. Different from the anthropogenic aerosols, since the emissions of sea salt and dust depend strongly on the simulated surface winds, the regional changes in their burdens (and AOD) may also be related to the changes in the simulated meteorology between E3SMv2 and E3SMv1. On the global mean basis, however, higher AOD in E3SMv2 may still be largely attributable to the anthropogenic aerosols, especially sulfate and SOA.

In addition to AOD, aerosol absorption of sunlight is also an important parameter in determining the aerosol radiative impacts. As discussed in Section 2.1.5, dust refractive indices in the shortwave were updated in E3SMv2. This leads to better agreement in the simulated aerosol absorption optical depth (AAOD) at 550 nm, as shown in Fig. 22, compared with the 10-year AAOD climatology derived from the ground-based



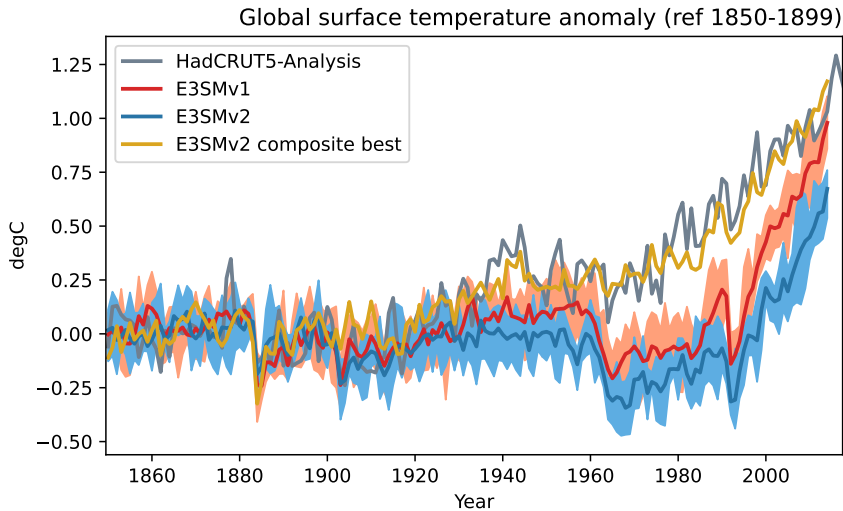
**Figure 21.** Spatial distributions of global annual mean (a) aerosol optical depth (AOD) from observational composite (Kinne et al., 2013) and the AOD difference between (b) E3SMv2 and (c) E3SMv1, respectively, from the historical simulations (2000-2014) and the observational composite. Areas with gray shading in polar regions indicate missing value. The number at the top-right of each panel represents the global mean.



**Figure 22.** (a) Annual mean aerosol absorption optical depth (AAOD) at 550 nm in E3SMv2 for the time period 2000-2014. The gray dots overlaid on top denote the locations of 139 AERONET stations, of which those circled in red denote the 19 dusty sites. (b) Comparison of the E3SMv2 AAOD with the observations derived from AERONET between 2006-2015 (Holben et al., 1998).  $r$  is the calculated correlation coefficient, and RMSE is the root-mean-square error.

1179 AERONET measurements (Holben et al., 1998) between 2006-2015. This 10-year time  
 1180 period is selected as it overlaps with most of the model-simulated 15 years representa-  
 1181 tive of a present-day climatology, but also during this time there are continuous obser-  
 1182 vations over a large number of sites globally. The compiled AERONET data for AAOD  
 1183 are available at a total of 139 stations mostly over land, and 19 of them with aerosol Ångström  
 1184 exponent  $< 0.8$  are denoted as the dusty sites, which are located near the major dust source  
 1185 regions. Compared to E3SMv1, E3SMv2 simulates smaller AAODs over all the dusty  
 1186 sites, and the calculated multi-site mean is 0.024, reducing the overestimation of E3SMv1  
 1187 (0.044) by nearly a factor of two against the observations (0.017). Over the other AERONET  
 1188 sites, AAODs in E3SMv2 are generally larger than those in E3SMv1 mainly due to the  
 1189 increased BC. Overall, E3SMv2 improves from E3SMv1 (0.017) by predicting a smaller  
 1190 mean AAOD (0.014) averaged over all the AERONET sites, similar to the observed mean  
 1191 (0.012) and with a smaller RMSE. The spatial correlation between the modeled and ob-  
 1192 served AAOD is noticeably improved in E3SMv2, for a larger correlation coefficient (0.83)  
 1193 with AERONET than that of E3SMv1 (0.72). Stronger correlation and smaller RMSE  
 1194 with the AERONET observations are also found in AOD (Fig. S13), implying a better  
 1195 representation of AOD and AAOD in E3SMv2 than E3SMv1 over the land area near those  
 1196 AERONET sites.

1197 The AAOD improvement in E3SMv2 compared with the AERONET data implies  
 1198 less aerosol heating in the atmosphere over the dust-influenced regions as a result of the  
 1199 decreased AAOD (Fig. S14) mainly due to the less absorbing dust optical properties.  
 1200 Additionally, we also updated the representation of dust size distribution in emission by  
 1201 accounting for more coarse particles in E3SMv2, which would decrease the net cooling  
 1202 effect of dust, but the impact is less than the enhanced cooling due to the lowered dust  
 1203 absorption (Feng et al., 2022). Over the other regions, E3SMv2 generally predicts higher  
 1204 AAOD than E3SMv1 resulting from the increased AOD (Fig. S14) especially due to BC,  
 1205 which would cause more atmospheric heating by aerosols regionally. Despite a higher global  
 1206 mean AOD, the global mean AAOD in E3SMv2 (0.0084) is slightly lower than in E3SMv1  
 1207 (0.0089), suggesting a possibly more negative direct radiative effect by aerosols.



**Figure 23.** Time evolution of annual global mean surface temperature anomalies (with respect to 1850-1899). Comparison between observations from HadCRUT5-Analysis (grey), E3SMv1 ensemble mean (red) and range (orange) and E3SMv2 ensemble mean (dark blue) and range (light blue). Also shown (gold) is a best fit estimate obtained by scaling E3SMv2 GHG and aerosol as discussed in Section 5.

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#### 4.3.6 Historical temperature record

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We now compare the time evolution of the global mean blended surface temperature (defined as SST over ice-free ocean and 2-m surface air temperature over land and sea ice) in E3SM with the observed historical record. We select the HadCRUT5-Analysis product (Morice et al., 2021); other products are available but the differences are minor compared to the differences with E3SM. Figure 23 shows the temperature anomalies normalized with respect to 1850-1899. As discussed previously (Golaz et al., 2019), E3SMv1 failed to accurately simulate the record by underestimating the warming starting around 1930 but eventually caught up to the observed record near 2010 because it overestimated the pace of warming from 1990 onward. This was attributed to excessively strong aerosol-related forcing and high climate sensitivity. While both have improved in E3SMv2 – slightly for the aerosol-related forcing and significantly for the sensitivity – E3SMv2 further underestimates the global mean surface temperature during the second half of the record. E3SMv2 diverges from E3SMv1 around 1930 and remains colder for the remainder of the record. A more in-depth analysis of this shortcoming is provided in Section 5.

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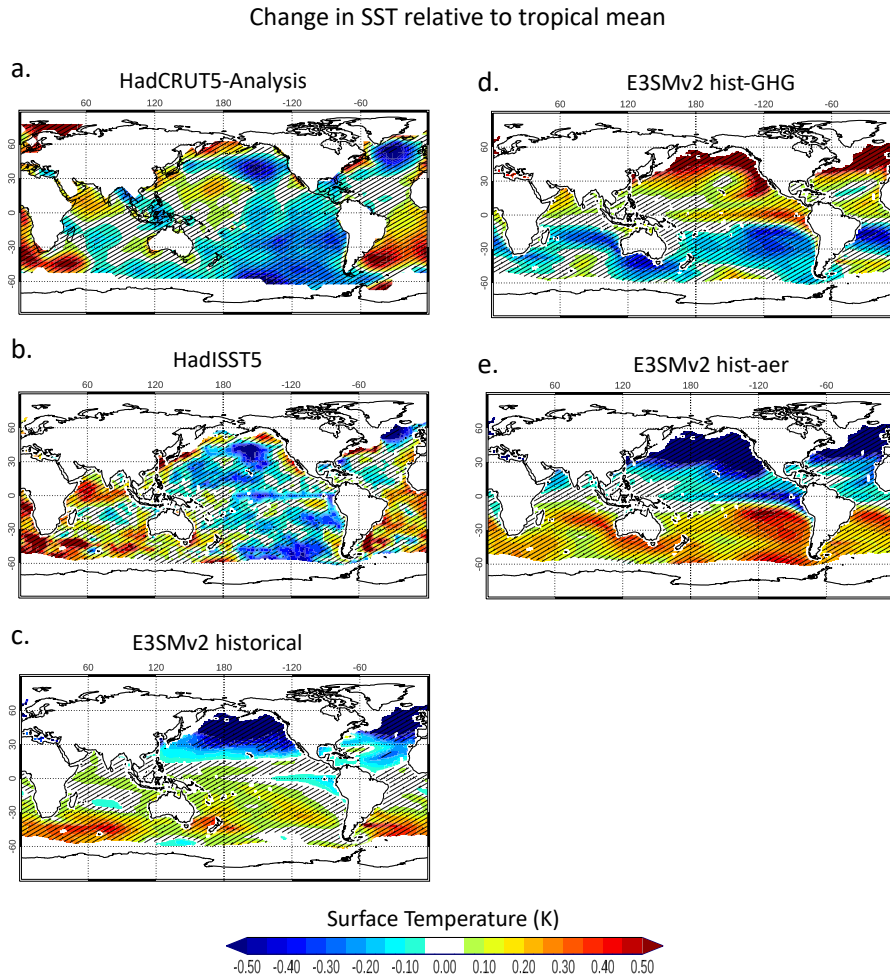
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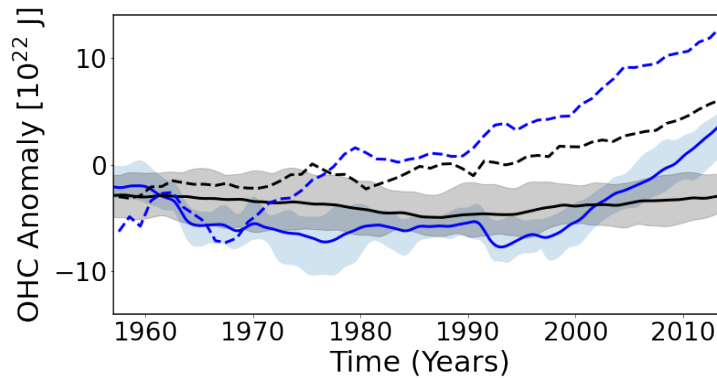
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The temporal evolution of the magnitude of radiative feedbacks, and hence effective climate sensitivity, is known to be influenced by the spatial patterns of the change in SSTs, particularly tropical Pacific SSTs, in climate simulations (e.g. Andrews et al., 2015b; Gregory & Andrews, 2016; Dong et al., 2019, 2020; Stevens et al., 2016). This is due to the impact of the east-west SST gradients in the tropical Pacific on convection, low clouds, humidity, and lapse rates affecting the different radiative feedback mechanisms over the region (e.g. Gregory & Andrews, 2016; Dong et al., 2020). Figure 24 shows the difference in SSTs between the 1980-2014 and the 1870-1904 period with respect to the tropical (25N-25S) mean SST change between the two periods for HadISST data, historical simulation, historical with GHG-only forcing (hist-GHG), and historical with aerosol-only forcing (hist-aer) simulation ensembles. The historical simulation shows a weak La





**Figure 24.** Difference in SST between 1980-2014 and 1870-1904 relative to the tropical mean change between the two periods for (a) HadCRUT5 data, (b) HadISST data, (c) E3SMv2 historical ensemble, (d) E3SMv2 hist-GHG ensemble, and (e) hist-aer ensemble. Hatched areas indicate regions where the difference is statistically significant at the 95% confidence level based on a two-tailed Student's t-test.



**Figure 25.** Ocean heat content from the E3SMv2 historical ensemble. Blue shows the 0-700m integrated ocean heat content, and black shows the 700-2000m heat content. The solid lines are the ensemble average and shading shows the ensemble spread. The dashed lines are pentad-averaged ocean heat content derived from the World Ocean Atlas 2009 (Levitus et al., 2012).

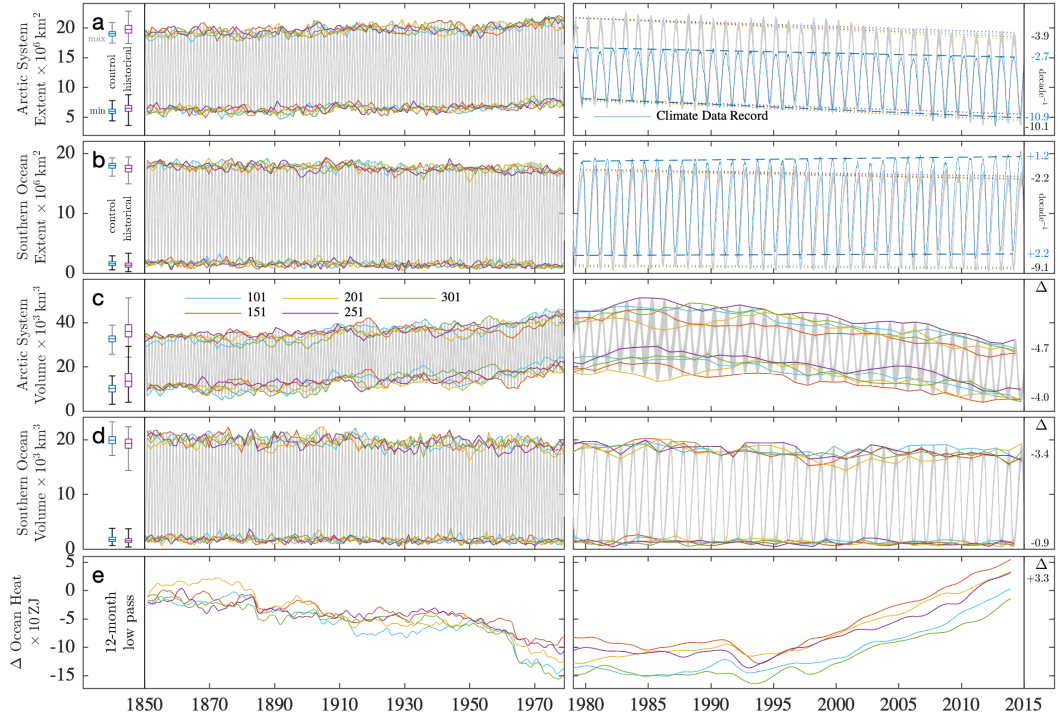
1235 Niña-like spatial pattern over the tropical Pacific with a weaker westward extension of  
 1236 cold anomalies as compared to HadCRUT5-analysis and HadISST dataset. The stronger  
 1237 than observed relative warming over the central Pacific region suggests stronger posi-  
 1238 tive cloud radiative feedbacks over the region (e.g. Andrews et al., 2015b) in E3SMv2,  
 1239 which also may be contributing to the strong climate sensitivity in E3SMv2 over the his-  
 1240 torical period. The spatial pattern of the SST change in the hist-GHG and hist-aer sim-  
 1241 ulation ensembles generally oppose each other, with the hist-GHG (hist-aer) exhibiting  
 1242 a El Niño (La Niña)-like pattern.

1243 The E3SMv2 simulated Ocean Heat Content (OHC) is shown in Fig. 25 relative  
 1244 to the OHC derived from the 2009 World Ocean Atlas (Levitus et al., 2012). The near  
 1245 surface (0-700m) OHC does increase in E3SMv2 (solid blue line) but is much delayed  
 1246 relative to observations. This is consistent with the late warming in the E3SMv2 tem-  
 1247 perature anomalies (Fig. 23). The mid-depth (700-2000m) OHC is very stable through-  
 1248 out the period analyzed, whereas the observations show an increase in the late 20th  
 1249 century. The lack of increasing OHC at that depth in E3SMv2 is likely attributable to the weak  
 1250 simulated AMOC (Figure 4b), which impacts transport of heat anomalies to depth (e.g.,  
 1251 A. Hu et al., 2020).

1252 As mentioned above, no historical test simulations were performed prior to final-  
 1253 izing E3SMv2. Instead, we made a decision to rely on atmosphere-only simulations to  
 1254 estimate effective radiative forcing and feedback during the development. Once the model  
 1255 development was concluded and the first historical simulation complete, the E3SM project  
 1256 made a pragmatic decision to be transparent and release the model version and accom-  
 1257 panying simulations, rather than delay in an attempt to correct the problem with the  
 1258 simulation of the global mean temperature in the historical record. However, future ver-  
 1259 sions of E3SM will include test historical simulations as part of their development cycle  
 1260 to avoid a repeat of this problem.

#### 1261 4.4 Impacts on Polar Climate

1262 In the historical ensemble (Fig. 26), Northern Hemisphere sea ice extent and vol-  
 1263 ume both increase over the time period 1850-1978, and decrease after the mid-1980s, as  
 1264 observed. This behavior is consistent with changes in the ocean heat content (Fig. 26e)



**Figure 26.** Daily sea ice extent (a, b) and volume (c, d) evolution across five ensemble members for the Northern and Southern Hemispheres, respectively, divided at the beginning of the core passive-microwave observation period in 1979 and compared to the change in 12-month filtered total ocean heat content from the start of the historical period in (e). Box plots in the left column compare annual extremes from daily values of the 500-year pre-industrial control (blue) with the industrial-era 5-member ensemble (purple). Trace colors for the year of the control simulation from which the ensemble members were spawned are indicated in (c; 101, 151, 201, 251, 300). Linear decadal trend in annual maximum and minimum daily extent is indicated in the right column for the ensemble mean of each ensemble trend line from 1979 to 2015, as compared to the Meier et al. (2017) NOAA Climate Data Record for (a) and (b). The right column in (c), (d), and (e) indicates the change ( $\Delta$ ) in the ensemble mean of volume extremes and non-filtered ocean content between 1850 and 2015.

1265 and surface air temperature anomalies for the historical simulations (Figs. 23 and B2b  
1266 in Appendix B).

1267 The maximum ice extent in the Arctic is larger in v2 than in v1, while it is smaller  
1268 in the Southern Hemisphere. The minimum ice extent is similar for v1 and v2 in both  
1269 hemispheres. E3SMv1’s large, cold SST bias in the North Atlantic and associated anomalous  
1270 sea ice in the Labrador Sea remain in v2. Unlike v1, which featured both warm and  
1271 cold SST biases in the Northern Hemisphere, the Northern Hemisphere in v2 is too cold  
1272 over its entirety, and so greater sea ice extent in v2 is not surprising. The Southern Hemisphere  
1273 is still biased warm, but not as badly as in v1, and sea ice in the Southern Ocean  
1274 is not extensive enough compared with the climate data record, year-round, in v2.

1275 Trends during the satellite era (Fig. 26, right panels) indicate that the model extent  
1276 is decreasing faster than observed in the Arctic, consistent with the faster increase  
1277 in surface air temperatures than observed (Fig. B2b in Appendix B). The ice extent trend  
1278 has the opposite sign compared with observations in the Antarctic, as in many other models,  
1279 and the change in volume extremes ( $\Delta$ ) between 1850 and 2015 is decreasing.

1280 A counter-intuitive result is that extremes in the ice extent and volume in the historical  
1281 simulations (left column of Fig. 26) are generally larger than in the pre-industrial control,  
1282 with a greater range of variability. However, this behavior is consistent with the aerosol  
1283 forcing biases discussed in Section 5.

1284 The net effect of improvements to the radiative and snow schemes in v2 only minimally  
1285 impacts the climatic state of sea ice, and thus does not ameliorate biases in prior  
1286 v1 simulations. Lack of conservation in the ice-ocean mass coupling scheme played a much  
1287 more important role; the correction of mass exchanges between the upper ocean and sea  
1288 ice models to account for brine content in the sea ice thickens the Arctic ice pack in summer,  
1289 reducing a bias from v1 (Fig. 26c, left column), while minimally impacting ice in the  
1290 Southern Ocean (Fig. 26d). With this mass-conserving scheme, the maximum and  
1291 minimum sea ice areas are now stable in both hemispheres for the 500-year pre-industrial  
1292 simulations, as shown in Fig. 4c.

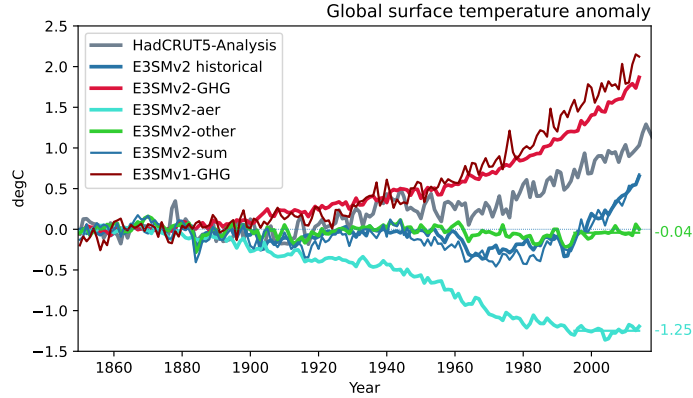
## 1293 5 Historical record: role of GHG vs aerosols

1294 To understand why E3SMv2 fails to accurately simulate the second half of the historical  
1295 temperature record, we analyze an ensemble of coupled simulations spanning 1850-  
1296 2014, but selectively activating only certain time varying forcing agents:

- 1297 • well-mixed greenhouse gases only (“GHG”),
- 1298 • aerosol and aerosol precursors only, including interactions with clouds (“aer”),
- 1299 • everything-else, all forcing agents except well-mixed GHG and aerosol (“other”).

1300 This decomposition is similar to the DAMIP protocol (Gillett et al., 2016), except  
1301 for the everything-else configuration, which is similar to natural forcing but includes additional  
1302 forcing terms (in particular land-use and ozone). We chose this particular decomposition  
1303 so that all the forcing agents are accounted for within the set. Five ensemble members  
1304 were run for each decomposition, initialized identically to the five-member ensemble  
1305 of historical simulations.

1306 The time evolution of annual global mean blended surface temperature is depicted  
1307 in Fig. 27. As expected, the dominant forcings are GHG (red) and aerosol-related (turquoise).  
1308 The remaining forcings (green) show inter-annual variations (mostly from volcanic eruptions  
1309 and the solar cycle) with little long term trend. A summation over the decomposition  
1310 (thin blue) recovers the original historical ensemble (thick blue) very well, indicating  
1311 that the decomposition is linear. The GHG and aerosol contributions almost per-



**Figure 27.** Global annual surface air temperature anomalies for model and observations (gray). For E3SMv2, the decomposition includes contributions from only GHG (red), only aerosol (turquoise), and other (green). The E3SMv2 historical is in blue, with the sum of individual terms in thin blue. Also shown is E3SMv1 with GHG only forcing (dark red). Horizontal lines indicate averages between 1995 to 2014. Observations from HadCRUT5-Analysis are normalized with respect to 1850-1899. Model results are normalized with respect to the 500-year *piControl* simulation.

1312 fectly mirror each other until approximately 1960, thus explaining the lack of net warm-  
 1313 ing until then (Fig. 23). It is only after the aerosol-related forcing stabilizes around 1990  
 1314 due to pollution control in North America and Europe that the GHG starts to dominate  
 1315 and E3SMv2 warms as a whole. As discussed previously, E3SMv2 has a lower TCR and  
 1316 ECS compared to E3SMv1. As a result, the warming from GHG alone is weaker than  
 1317 in v1 (dark red; Zheng et al., 2021). The two models diverge mostly after 1960 which  
 1318 helps explain why E3SMv2 remains colder longer. It is also informative to compare E3SMv2  
 1319 with an ensemble of CMIP6 models (Tokarska et al., 2020, their Fig. 1A). E3SMv2-GHG  
 1320 warms a little faster than the CMIP6 ensemble (reaching  $\sim 1.7^\circ\text{C}$  compared to  $\sim 1.5^\circ\text{C}$ ),  
 1321 but the largest difference is observed for E3SMv2-aer, which has an average cooling of  
 1322  $1.25^\circ\text{C}$  over the last 20 years, twice as large as the CMIP6 ensemble.

1323 Equipped with this decomposition, we can investigate hypothetical configurations  
 1324 with different relative strengths of GHG and aerosol. We can write any variable  $\psi$  as:

$$\psi_{\text{all}} = \psi_{\text{piControl}} + \alpha_{\text{GHG}} (\psi_{\text{GHG}} - \psi_{\text{piControl}}) + \alpha_{\text{aer}} (\psi_{\text{aer}} - \psi_{\text{piControl}}) + (\psi_{\text{other}} - \psi_{\text{piControl}}) \quad (1)$$

1325 This reconstruction is conceptually similar to Neelin et al. (2010), but applied to  
 1326 different forcing terms rather than physics parameter perturbations. Gillett et al. (2012)  
 1327 and Winton et al. (2020) also used a similar decomposition to reconstruct the global mean  
 1328 surface temperature anomalies. Eq. 1 is applicable to any variable. Varying  $\alpha_{\text{GHG}}$  mod-  
 1329 ulates the model response to GHG (akin to modulating the TCR), while  $\alpha_{\text{aer}}$  modulates  
 1330 the model response to aerosols. Setting  $\alpha_{\text{GHG}} = \alpha_{\text{aer}} = 1$  (“composite base”) recov-  
 1331 ers the all-forcing configuration as long as the decomposition is linear. Linearity is a very  
 1332 good approximation for annual global averages (Fig 27). Furthermore, it holds well for  
 1333 two-dimensional and three-dimensional climatological fields as demonstrated in Fig. S15:  
 1334 RMSEs for the composite base configuration (red stars) and E3SMv2 (red triangles) are  
 1335 very similar for most fields and seasons.

Optimal scaling factors  $\alpha_{\text{GHG}}$  and  $\alpha_{\text{aer}}$  are derived in Appendix B by minimizing the difference between simulated and observed surface temperature, leading to  $\alpha_{\text{GHG}} = 0.75$  and  $\alpha_{\text{aer}} = 0.17$  with a probable co-linear range from (0.60, 0.04) to (0.81, 0.23). This hypothetical configuration (“composite best”) matches the historical temperature record much better than E3SMv2 (Fig. 23; gold).

The E3SMv2 scaling factors are comparable to previously reported values. Gillett et al. (2012) derived factors of 0.65 for GHG and 0.3 for aerosols for CanESM2 (TCR= 2.3 K;  $\text{ERF}_{\text{aer}} = -0.8 \text{ W m}^{-2}$ ), while Winton et al. (2020) derived factors of 0.78 and 0.34 for GFDL-CM4 (TCR= 2.05 K;  $\text{ERF}_{\text{aer}} = -0.73 \text{ W m}^{-2}$  from Smith et al. (2020)). The GHG scaling factors are consistent for all three models given their similar TCR. The E3SMv2 aerosol scaling factor is smaller, likely due to its stronger  $\text{ERF}_{\text{aer}} = -1.52 \text{ W m}^{-2}$ .

Based on Shindell (2014) and Rotstayn et al. (2015), the temperature change over the historical record can be approximated as:

$$\Delta T = \frac{\text{TCR}}{F_{2x}} (F_{\text{GHG}} + F_{\text{other}} + E_{\text{aer}} F_{\text{aer}}) \quad (2)$$

where  $F$  denotes effective radiative forcings and  $E_{\text{aer}}$  is an enhancement factor to account for the fact that the aerosol forcing has a different impact on global mean surface temperature than GHG. It is the same as the concept of efficacy first introduced by Hansen (2005).

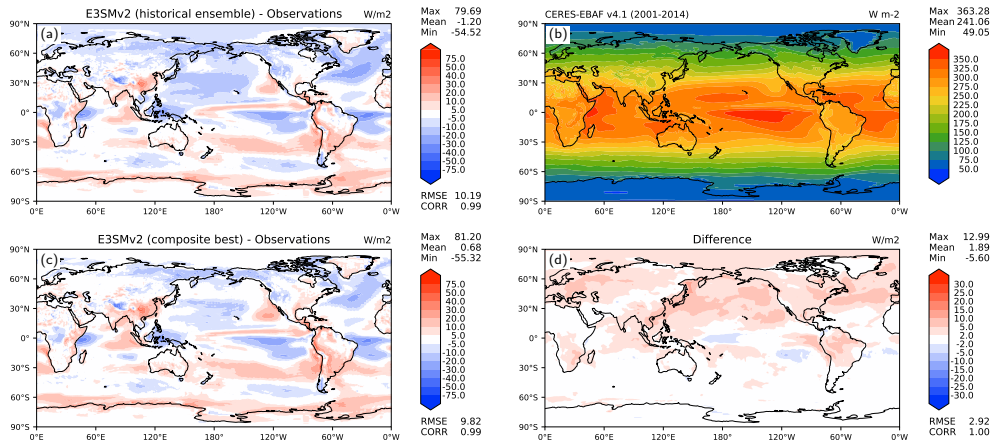
Based in Eq. 2,  $\alpha_{\text{GHG}} = 0.75$  implies a reduction in E3SMv2 TCR from 2.4 to 1.8 K (assuming no change in  $F_{2x}$  and  $F_{\text{GHG}}$ ), close to the CMIP6 observationally constrained TCR of 1.6 K from Tokarska et al. (2020).  $\alpha_{\text{aer}} = 0.17$  implies a scaling of factor of  $0.17/0.75 = 0.23$  for  $E_{\text{aer}} F_{\text{aer}}$ , which could come from either a reduction in the aerosol enhancement factor or effective forcing, or a combination of both. The E3SMv2 aerosol enhancement factor can be estimated as (e.g. Winton et al., 2020):

$$E_{\text{aer}} = \frac{\Delta T_{\text{aer}}}{F_{\text{aer}}} \frac{F_{2x}}{\text{TCR}} = \frac{-1.25}{-1.52} \times \frac{2.98}{2.41} = 1.01 \quad (3)$$

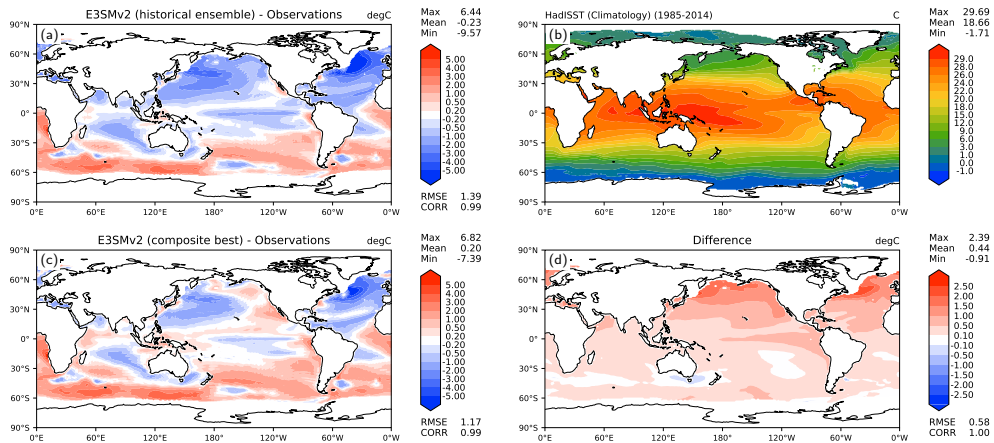
It is smaller than the value of 1.4 for an ensemble of CMIP5 models (Rotstayn et al., 2015). If we use their  $2x\text{CO}_2$  radiative forcing of  $3.70 \text{ W m}^{-2}$ , the enhancement factor increases to 1.26, still lower than 1.4. We therefore postulate that the 0.23 reduction in  $E_{\text{aer}} F_{\text{aer}}$  is more likely to come from a reduction in the aerosol effective forcing than a reduction in the enhancement factor. If that is the case, the resulting  $\text{ERF}_{\text{aer}}$  could be as small as  $-0.35 \text{ W m}^{-2}$ . This would be much smaller than any other CMIP6 model analyzed by Smith et al. (2020, their Table 2). Stevens (2015) argued for a weaker aerosol forcing between  $-0.3$  and  $-1.0 \text{ W m}^{-2}$ , although its methodology based on a box model has been criticized as overly simplified (Kretzschmar et al., 2017).

Finally, we also reconstruct climatological fields using Eq. 1. Figure 28 shows the top-of-atmosphere SW net radiation (2001-2014). Remarkably, the NH negative bias in E3SMv2 (blue shading in Fig 28a) is greatly reduced in composite best (Fig. 28c) which becomes much closer to observations regionally, especially over the N Atlantic and N Pacific oceans. Global metrics also improve with a reduced mean bias and RMSE. A similar picture emerges for the sea-surface temperature (1985-2014; Fig. 29) with substantial reductions in regional cold biases in the NH. SH SST biases are essentially unchanged, pointing to a different cause.

Taken together, our results indicate that a substantial reduction in the aerosol forcing would not only improve the match with the historical temperature record, but also improve aspects of the present-day climatology. Other fields, for example precipitation, exhibit much smaller impact as seen in Figure S15 by comparing the gold (composite best) and red stars (composite base). This is reassuring in the sense that E3SMv2, despite its shortcomings, can still serve as a useful model for many studies.



**Figure 28.** Net TOA SW radiation: observations (CERES-EBAF 4.1; b), model error for E3SMv2 (a), composite best configuration from Fig. B1 in Appendix B (c), and difference between E3SMv2 and composite best configuration (d). Model and observations averaged over 2001-2014.



**Figure 29.** Same as Fig. 28 but for SST. Sea ice covered regions are excluded from the averaging. Model and observations averaged over 1985-2014.

1382 It is important to note that many CMIP6 models can simulate the historical tem-  
 1383 perature record with stronger aerosol forcing than suggested by our present analysis, for  
 1384 example, CNRM-CM6-1 (Voldoire et al., 2019, their Fig. 23) with a forcing of  $-1.15 \text{ W}$   
 1385  $\text{m}^{-2}$  (Smith et al., 2020). CESM2 (Danabasoglu et al., 2020, their Fig. 7), from which  
 1386 E3SM inherited most of its atmospheric physics parameterizations, also simulates the  
 1387 temperature record more realistically than E3SMv2 and this with an aerosol forcing of  
 1388  $-1.37 \text{ W m}^{-2}$  (Smith et al., 2020). CESM2 does however overestimate the rate of warm-  
 1389 ing over the last three decades due its high sensitivity (TCR = 2.0 K, ECS = 5.3 K).  
 1390 Tokarska et al. (2020, their Fig. 2A) identified a dozen CMIP6 models that satisfy their  
 1391 TCR constraints and have a realistic 1981-2014 temperature trend. A subset of five [CNRM-  
 1392 ESM2-1 (Sfrian et al., 2019); GFDL-ESM4 (Dunne et al., 2020); MIROC6 (Tatebe et  
 1393 al., 2019); MRI-ESM2-0 (Yukimoto et al., 2019); NorESM2-LM (Seland et al., 2020)] had  
 1394 their aerosol forcing assessed in Smith et al. (2020). Their aerosol forcings range from  
 1395  $-0.7$  to  $-1.21 \text{ W m}^{-2}$ . Therefore, it is possible that our results may not be generaliz-  
 1396 able beyond E3SMv2. Other factors not encapsulated in the simple framework repre-  
 1397 sented by Eqs. 1 and 2 may play a role in a successful simulation of the historical tem-  
 1398 perature record.

## 1399 6 Summary and conclusion

1400 By design, E3SMv2 represents an evolution from E3SMv1 and as such resembles  
 1401 E3SMv1 in many aspects. There are nevertheless notable differences that justified a new  
 1402 model release and associated simulation campaign.

- 1403 • E3SMv2 is approximately twice as fast (or efficient if measured in terms of power)  
 1404 compared to v1 (Fig. 2). The efficiency gains are achieved in the atmosphere and  
 1405 ocean components. In the atmosphere, they arise from a new semi-Lagrangian tracer  
 1406 transport method and a new grid for physics calculations (Fig. 3). The gain in the  
 1407 ocean is due to a longer timestep.
- 1408 • The atmospheric physics, while based on the same basic set of parameterizations  
 1409 as v1, underwent significant retuning in v2. Many improvements from the inter-  
 1410 mediate EAMv1p configuration (Ma et al., 2022) are incorporated with additional  
 1411 changes to further improve clouds and precipitation (e.g. Figs. 9, 10, 11, 12, 13).
- 1412 • A new convective trigger function for the deep convection (Xie et al., 2019) sig-  
 1413 nificantly improves the phase of the diurnal cycle of precipitation, but the ampli-  
 1414 tude remains weaker than observed (Figs. 14, 15).
- 1415 • E3SMv2 captures important modes of variability such as ENSO (Fig. 16) and MJO  
 1416 (Fig. 18). However, the ENSO spectrum has excessive energy at short periods ( $\sim 2.5$   
 1417 years) and is too weak for longer periods (6-9 years). MJO phase speed is real-  
 1418 istic west of  $125^\circ\text{E}$ , but then exceeds observations east of it. Tropical variability  
 1419 is significantly too weak (Fig. 17).
- 1420 • A more realistic treatment of ozone is implemented (Tang et al., 2021). It cap-  
 1421 tures the seasonal cycle of stratospheric column ozone (Fig. 19) and the ozone hole  
 1422 in the historical period, although the size is underestimated (Fig. 20).
- 1423 • Dust aerosol optical properties and particle size distributions are revised, result-  
 1424 ing in a better prediction of mean AAOD over dusty AERONET sites (Fig. 22).  
 1425 Burdens of sulfate and SOA aerosols increase likely as an unintended consequence  
 1426 of cloud tuning efforts, giving rise to a slightly overestimated global mean AOD  
 1427 despite regional improvements (Fig. 21). Increases of sea salt aerosol burden re-  
 1428 lated to the changes in wet deposition also contribute to the higher AOD over some  
 1429 portions of tropical and subtropical oceans, but the global burden of sea salt (and  
 1430 dust) has small changes.
- 1431 • E3SMv2 is less sensitive to GHG forcing (Fig. 5). ECS is reduced significantly com-  
 1432 pared to v1 (4.0 K vs 5.3 K) which is mostly attributable to a smaller cloud feed-  
 1433 back. The ECS value of 4.0 K is more plausible as assessed by WCRP (Sherwood



1434 et al., 2020). This is a substantial achievement compared to the unrealistically high  
1435 sensitivity of E3SMv1. On shorter time scales, TCR is also reduced to 2.4 from  
1436 2.9 K.

- 1437 • The effective aerosol forcing ( $\text{ERF}_{\text{aer}} = -1.5 \text{ W m}^{-2}$ ) remains essentially unchanged  
1438 in E3SMv2 (Fig. 6). This value is within the likely range assessed by WCRP (Bellouin  
1439 et al., 2020). Some changes, mainly the introduction of the minimum cloud droplet  
1440 number concentration, were made in v2 that reduced the magnitude of  $\text{ERF}_{\text{aer}}$ ,  
1441 but their impact was likely negated by changes elsewhere in the cloud physics (con-  
1442 vection).
- 1443 • E3SMv2 significantly underestimates the global mean temperature in the second  
1444 half of the historical temperature record (Fig. 23). An analysis of single-forcing  
1445 simulations indicates that correcting the historical record in E3SMv2 would re-  
1446 quire a substantial reduction in the magnitude of  $\text{ERF}_{\text{aer}}$  ( $\sim 75\%$ ), and a moder-  
1447 ate reduction in the TCR ( $\sim 25\%$ ). Reducing  $\text{ERF}_{\text{aer}}$  would furthermore improve  
1448 regional biases in TOA radiative fluxes and SST (Figs. 28, 29). Other fields are  
1449 less impacted (e.g. precipitation; Fig. S15), indicating that E3SMv2 can still serve  
1450 as a useful tool despite its shortcomings.
- 1451 • Proper conservation of mass in ocean/sea-ice exchanges increases Arctic sea ice  
1452 volume, improving a low-thickness bias from v1, while impacting the Southern Ocean  
1453 ice pack very little. Changes to the radiation and snow physics parameterizations  
1454 have little net effect, highlighting the importance of coupled interactions over in-  
1455 ternal sea ice processes in the climate system (Hunke, 2010). The sea ice simu-  
1456 lations shown here are largely consistent with the overall climatic environment,  
1457 including excessively cool surface air and ocean temperatures.

1458 This release of E3SMv2 serves as a starting point for additional configurations. They  
1459 include regionally refined configurations with higher resolution over North America and,  
1460 separately, the Southern Ocean. A configuration with interactive biogeochemistry is also  
1461 under development. While E3SMv2 improves upon its predecessor in many aspects, sig-  
1462 nificant work remains. The highest priorities for future releases of E3SM are address-  
1463 ing the weak AMOC, the biases in interannual and intraseasonal atmospheric tropical  
1464 variability, and the poor historical temperature record.

**Appendix A Atmosphere configuration**

Table A1: List of the atmospheric tuning parameters. Note: the value of *microp\_aero\_wsubmin* was set to 0.001 for v1p and v2 based on Ma et al. (2021). However, an additional lower bound is present in the code that effectively resets it to 0.1 consistent with Ma et al. (2022).

Scheme	Parameter	v2	v1	v1p	Short Description
CLUBB	<i>clubb_c14</i>	2.5	1.06	2.0	Dissipation of $w'^2$ and $v'^2$
	<i>clubb_c1</i>	2.4	1.335	2.4	Low-skewness value of dissipation of $w'^2$
	<i>clubb_c1b</i>	2.8	1.335	2.8	High-Skw value of dissipation of $w'^2$
	<i>clubb_c1c</i>	0.75	1.0	0.75	Smoothness of transition between high-Skw and low-Skw for the dissipation of $w'^2$
	<i>clubb_c6rtb</i>	7.5	6.0	7.5	High-Skw value of pressure damping of water flux
	<i>clubb_c6rtc</i>	0.5	1.0	0.5	Smoothness of transition between high-Skw and low-Skw for the pressure damping of water flux
	<i>clubb_c6thlb</i>	7.5	6.0	7.5	High-Skw value of pressure damping of heat flux
	<i>clubb_c6thlc</i>	0.5	1.0	0.5	Smoothness of transition between high-Skw and low-Skw for the pressure damping of heat flux
	<i>clubb_c8</i>	5.2	4.3	5.2	Pressure damping of $w'^3$
	<i>clubb_c11</i>	0.7	0.8	0.7	Buoyancy damping of $2'^3$ at low Skw
	<i>clubb_c11b</i>	0.2	0.35	0.2	Buoyancy damping of $2'^3$ at high Skw
	<i>clubb_c11c</i>	0.85	0.5	0.85	Smoothness of transition between high Skw and low Skw for the buoyancy damping of $2'^3$
	<i>clubb_c_k10</i>	0.35	0.3	0.35	Coefficient of momentum diffusivity, Kh_zm
	<i>clubb_c_k10h</i>	0.35	0.3	0.35	Coefficient of thermodynamic diffusivity, Kmh_zm
	<i>clubb_gamma_coef</i>	0.12	0.32	0.12	Constant of the width of PDF in w-coordinate
	<i>clubb_gamma_coefb</i>	0.28	0.32	0.28	High-skw value of gamma coefficient
	<i>clubb_gamma_coefc</i>	1.2	5.0	1.2	Smoothness of transition between values of gamma coefficient

	<i>clubb_mu</i>	$5e^{-4}$	$1e^{-3}$	$5e^{-4}$	Fractional parcel entrainment rate per unit height [1/m]
	<i>clubb_wpxp_l_thresh</i>	100.0	60	100	Threshold in length scale below which extra damping is applied to C6 and C7 functions [m]
	<i>clubb_ice_deep</i>	$14e^{-6}$	$16e^{-6}$	$14e^{-6}$	Radius of ice particles detrained from deep convection [m]
	<i>cldfrc_dp1</i>	0.018	0.045	0.018	parameter for deep convective cloud fraction
	<i>clubb_use_sgv</i>	True	False	True	Enables subgrid features gustiness, tpert, and thv fix
	<i>clubb_ipdf_call_placement</i>	1	2	1	Select the placement of the call to CLUBB's PDF: 1 - before advancing CLUBB's predictive fields, 2 - after, 3 - both before and after
ZM	<i>zmconv_alfa</i>	0.14	0.1	0.14	Maximum downdraft mass flux fraction
	<i>zmconv_c0_lnd</i>	0.002	0.007	0.002	Autoconversion coefficient over land for deep convection
	<i>zmconv_c0_oce</i>	0.002	0.007	0.002	Autoconversion coefficient over ocean for deep convection
	<i>zmconv_dmpdz</i>	$-0.7e^{-3}$	$-0.7e^{-3}$	$-1.2e^{-3}$	Parcel fractional mass entrainment rate
	<i>zmconv_mx_bot_lyr_adj</i>	1	2	1	Bottom layer adjustment for setting "launching" level of maximum moist static energy
	<i>zmconv_tp_fac</i>	2	0	2	Tpert scale factor in ZM deep convection scheme
MG2	<i>cld_sed</i>	1.0	1.0	1.8	Scale factor for cloud droplet sedimentation
	<i>ice_sed_ai</i>	500	500	1200	Cloud ice fall speed parameter
	<i>micro_mg_berg_eff_factor</i>	0.7	0.1	0.7	Efficiency factor for WBF processes
	<i>micro_mg_accre_enhan_fac</i>	1.75	1.5	1.75	Accretion enhancement factor
	<i>prc_expl</i>	-1.4	-1.2	-1.4	Tunable exponent coefficient for autoconversion
	<i>micro_mincdnc</i>	10.D6	0.0	0.0	Minimum cloud droplet number concentration imposed when $micro\_mincdnc > 0$ [ $m^{-3}$ ]

nucleate	<i>so4_sz_thresh_icenuc</i>	$0.08e^{-6}$	$0.05e^{-6}$	$0.08e^{-6}$	Aitken mode SO2 size threshold for ice nucleation
microp aero	<i>microp_aero_wsubmin</i>	0.1	0.2	0.1 See note in caption	Minimum subgrid vertical velocity
aerosol	<i>seasalt_emis_scale</i>	0.6	0.85	0.6	Tuning factor for sea salt aerosol emission
dust	<i>dus_emis_fact</i>	1.5	2.05	2.8	Tuning parameter for dust emissions
Linoz	<i>linoz_psc_t</i>	197.5	193.0	193.0	Tunable Linoz PSC ozone loss temperature threshold (K)
Gravity wave drag	<i>gw_convect_hcf</i>	10.0	20.0	20.0	Heating rate conversion factor associated with convective gravity waves
	<i>effgw_beres</i>	0.35	0.40	0.40	Efficiency associated with convective gravity waves from the Beres scheme
	<i>effgw_oro</i>	0.375	0.25	0.25	Efficiency associated with orographic gravity waves

## Appendix B Composite configurations

In order to estimate the role of GHG and aerosols in the mismatch of the historical temperature record, we construct hypothetical composite configurations by varying  $\alpha_{\text{GHG}}$  and  $\alpha_{\text{aer}}$  scaling factors in Eq. 1 (with  $\psi$  set to surface temperature).  $\alpha_{\text{GHG}}$  modulates the model response to GHG (akin to modulating TCR), while  $\alpha_{\text{aer}}$  modulates the model response to aerosols (akin to modulating the magnitude of the aerosol-related forcing and feedback).

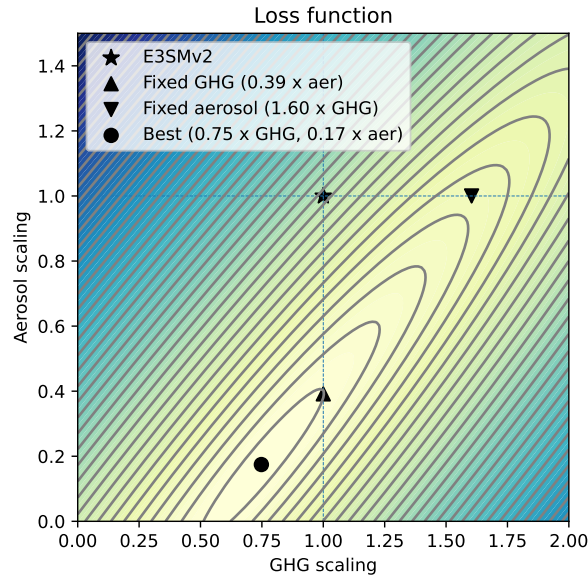
We construct a loss function that quantifies the mismatch between modeled and observed surface air temperature:

$$F = \sum_{SH, NH} \left( \sum_{yr=1950}^{2014} (\bar{T}_{\text{model}} - \bar{T}_{\text{obs}})^2 \right)^{1/2} \quad (\text{B1})$$

We opt to separately account for SH and NH due to the strong asymmetry in aerosol forcing. We also select the latter part of the historical record (1950-2014) when observational uncertainties are smaller and the impact of GHG and aerosols are larger. We first optimize by using five-member ensemble means for each term in Eq. 1. This reconstruction is not expected to realistically capture natural multidecadal variability which plays an important role (e.g. Zeng & Geil, 2016).

The loss function  $F$  is shown in Fig. B1 as a function of scaling parameters  $\alpha_{\text{GHG}}$  and  $\alpha_{\text{aer}}$ . The surface depicts a broad valley oriented diagonally. This shape can be more easily interpreted by imagining 45° rotated axes. The direction perpendicular to the valley floor ( $x' \propto \alpha_{\text{GHG}} - \alpha_{\text{aer}}$ ) measures the scaling of the total warming (since aerosol is always negative) and the direction parallel ( $y' \propto \alpha_{\text{GHG}} + \alpha_{\text{aer}}$ ) the hemispheric distribution of the warming.

The global minimum (best) is situated at  $\alpha_{\text{GHG}} = 0.75$  and  $\alpha_{\text{aer}} = 0.17$ , indicating that improving the historical temperature record simulated by E3SMv2 would require a moderate reduction in impact from GHG (TCR), but a very substantial one from the aerosols. Also shown in Fig. B1 are two local minima: one holding GHG scaling (fixed



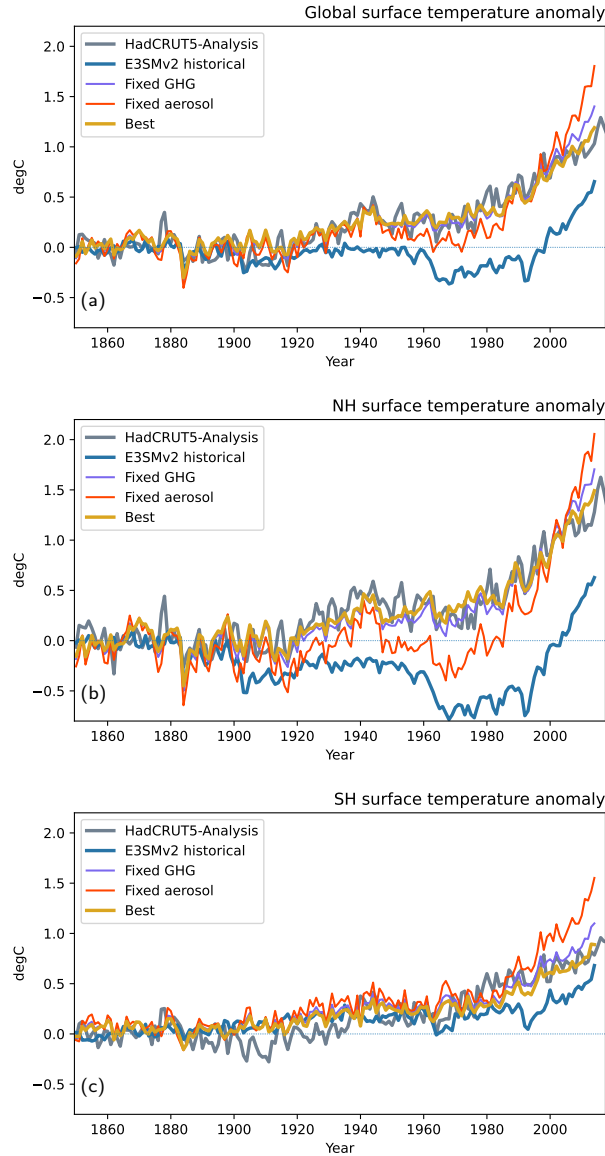
**Figure B1.** Loss function from Eq. B1. Star represents E3SMv2, circle global minimum, and triangles local minima by minimizing along a single dimension.

1491 GHG;  $\alpha_{\text{GHG}} = 1$  and  $\alpha_{\text{aer}} = 0.39$ ) and one holding aerosol scaling (fixed aerosol;  $\alpha_{\text{GHG}} =$   
 1492  $1.60$  and  $\alpha_{\text{aer}} = 1$ ). Both of these local minima reside on the valley floor (thus match-  
 1493 ing the total warming), but the first one is much closer to the global minimum compared  
 1494 to the second one, confirming that aerosols are the dominant source of the mismatch.

1495 The shape of the loss function also helps explain why the simulated historical record  
 1496 became worse in E3SMv2 than E3SMv1. Aerosol forcing is almost unchanged between  
 1497 the two, so both models would lie on a horizontal line. But TCR is larger in E3SMv1  
 1498 and thus closer to the valley floor. In other words, E3SMv2 moved uphill compared to  
 1499 E3SMv1 due to the lack of colinear change between GHG and aerosol impacts.

1500 This can be further illustrated by constructing global and hemispheric tempera-  
 1501 ture time series corresponding to these composite configurations (Fig. B2). The com-  
 1502 posite best solution (gold) corresponding to the global minimum improves considerably  
 1503 upon E3SMv2 and matches the historical record best for each region (global, NH, SH).  
 1504 Composite fixed-GHG (purple) also does an adequate job, but with some indication of  
 1505 excessive warming in the 2000s due to its higher response to GHG. Composite fixed-aerosol  
 1506 (orange), which increases the response of GHG to balance the strong aerosol cooling, fails  
 1507 to match the historical record well. This confirms the argument that higher sensitivity  
 1508 cannot adequately compensate for excessive aerosol forcing owing to the presence of a  
 1509 plateau in the aerosol forcing and hemispheric asymmetry (e.g. Zhao et al., 2018; Al-  
 1510 bright et al., 2021).

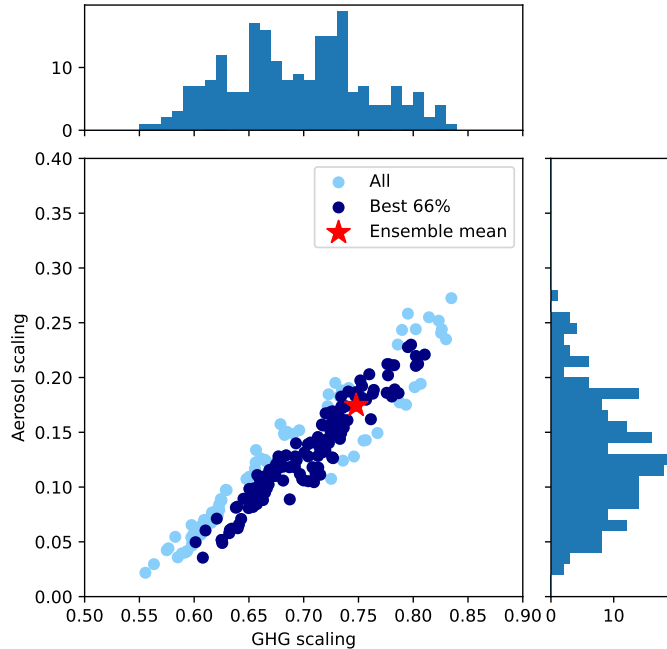
1511 To estimate the impact of multidecadal variability, we repeat the minimization in  
 1512 Eq. B1 including individual ensemble members. This gives a total of 216 individual re-  
 1513 alizations (5 ensemble member plus ensemble mean for each of GHG, aer, and other in  
 1514 Eq. 1). While this is not expected to realistically reproduce natural variability, it does  
 1515 provide a useful range. Optimal solutions are shown in Figure B3. Individual loss func-  
 1516 tions (not shown) have similar shape as Figure B1, but with slight variations in the place-  
 1517 ment of the valley. We estimate a probable range for  $(\alpha_{\text{GHG}}, \alpha_{\text{aer}})$  from (0.60, 0.04) to



**Figure B2.** Surface temperature anomalies (with respect to 1850-1899) for (a) global, (b) northern hemisphere, and (c) southern hemisphere. Lines shown include observations (HadCRUT5-Analyses, grey), E3SMv2 (blue), and composite configurations from Fig. B1 (red, purple, red, gold).

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(0.81, 0.23) by retaining the best 66% solutions. There is a high degree of co-linearity between variations in  $\alpha_{\text{GHG}}$  and  $\alpha_{\text{aer}}$  that originates from the elongated valley present in the loss functions.



**Figure B3.** Scatter and distributions of optimal solutions  $\alpha_{\text{GHG}}$  and  $\alpha_{\text{aer}}$  for the loss function in Eq. B1 using all 216 combinations of individual ensemble members or ensemble mean. Dark blue dots represent the best 66% solutions. Red star is the solution with ensemble mean for every term.

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Additionally, we also investigated the impact of changing assumptions in the loss function by using a global average instead of separate hemispheric averages and including the entire record since 1850 instead of starting in 1950. While the optimal solutions change a little as a result, the probable ranges are broadly consistent. We also attempted to construct a loss function based on the gridded  $5^\circ \times 5^\circ$  HadCRUT5-Analysis product. However, this led to considerably poorer fits when reconstructing the global mean time series. We speculate that this is caused by the loss function becoming too sensitive to regional model biases not related to GHG or aerosols.

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1550 E3SM-Project/E3SM](https://github.com/E3SM-Project/E3SM). A maintenance branch (`maint-2.0`; [https://github.com/E3SM-  
1551 -Project/E3SM/tree/maint-2.0](https://github.com/E3SM-Project/E3SM/tree/maint-2.0)) has been specifically created to reproduce these sim-  
1552 ulations. Bit-for-bit results with the original simulations on identical machines will be  
1553 maintained on that branch for as long as the computing environment supports it.

1554 Complete native model output is accessible directly on NERSC at [https://portal.  
1555 .nersc.gov/archive/home/projects/e3sm/www/WaterCycle/E3SMv2/LR](https://portal.nersc.gov/archive/home/projects/e3sm/www/WaterCycle/E3SMv2/LR) with accom-  
1556 panying documentation at [https://e3sm-project.github.io/e3sm\\_data\\_docs](https://e3sm-project.github.io/e3sm_data_docs). A sub-  
1557 set of the native output is also available through the DOE Earth System Grid Federa-  
1558 tion (ESGF) at [https://esgf-node.llnl.gov/search/e3sm/?model\\_version=2.0](https://esgf-node.llnl.gov/search/e3sm/?model_version=2.0). Data  
1559 reformatted following CMIP conventions will also be available through ESGF at [https://  
1560 esgf-node.llnl.gov/projects/e3sm](https://esgf-node.llnl.gov/projects/e3sm).

1561 Performance data and scripts for Figures 2 and 3 are available at [https://github  
1562 .com/E3SM-Project/perf-data/tree/main/v2-overview/chrysalis-perf-study](https://github.com/E3SM-Project/perf-data/tree/main/v2-overview/chrysalis-perf-study); see  
1563 the `readme.txt` file there for further details.

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