The University of Michigan

Department of Chemical and Metallurgical Engineering

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EXACT SOLUTIONS FOR ANISOTROPIC, MULTIPLE SCATTERING
BY
PARALLEL PLANE DISPERSIONS

bv

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#### NOMENCLATURE

angular distribution coefficients, i = 0, 1,..., N a. coefficients in expansion of  $\cos m \varphi = A_0 + \sum_{i=1}^{m} A_i \cos \varphi$ A; coefficients in expansion I-21 of angular dependence  $b_i(\tau_1,\mu,\mu_0)$ of specific intensity of reflected radiation. Defined in Appendix A-1 fraction of radiation scattered backward. Defined by I-47 В six-flux representation of backward scattering component B.6 height of backward peak of phase function, p(-1.0) ΒP coefficients in expansion I-22 of angular dependence of  $c_{i}(\tau_{1},\mu,\mu_{0})$ diffuse portion of transmitted radiation. Defined in Appendix A-1 diameter of scattering particle D defined by  $\int_{-\infty}^{\infty} \frac{e^{-x}}{x^n} dx$  $E_n(Z)$ angular distribution function, same as  $p(\cos \theta)/4\pi$  $f(\Theta)$ f(x)any function of x any tabular function of x f\*(x) $f^{*(n)}(x)$ approximation to solution of I-39 after n iterations fraction of radiation scattered forward, defined by I-46 F  $F(\mu)$ any function of  $\mu$ six-flux representation of forward scattering component  $F_{6}$  $F_k^m(\tau_1,\mu,\mu')$ function defined by I-3 height of forward peak of phase function, p(1.0) FP  $G_{k}^{m}(\tau_{1},\mu,\mu')$ function defined by I-4 Chandrasekhar's H function, the same as  $\psi^{\bullet}(\infty,\mu)$  $H(\mu)$ 

I <sub>1</sub> ,I <sub>2</sub> ,	discrete components of integrated radiant flux
Ι(Ζ <b>,</b> μ, <b>φ</b> )	specific intensity of radiation at position Z in the solid angle defined by $\mu$ and $\phi$
$I_{O}^{(-\mu_{O}, \varphi_{o})}$	intensity of a parallel beam incident on plane parallel dispersion in the direction (- $\mu_{_{\rm O}}$ , $m{\gamma}$ )
m	index of refraction of scattering particle
mį	constants derived in equation (III-29, 30) from the approximation of the isotropic scattering integral equation
n	number of particles per unit volume
n <sub>i</sub>	coefficients of the approximate kernal derived in equation (III-18)
N	number of terms in the phase function
p	number of ordinates (excluding the zeroth ordinate) used in numerical integration with $I=35$
$p(\mu)$	phase function
P	peakedness of the phase function in the forward direction defined by I-49
$P_{B}$	peakedness of the phase function in the backward direction defined by I-50
$P_1(\mu)$	Legendre polynomial
$P_{\perp}^{m}(\mu)$	associated Legendre polynomials
r <sub>i</sub>	i-th root of $P_n(x)$
R	integrated reflectance
$S(\overrightarrow{r})$	Virtual source function
S (7,; \(\mu, \varphi, \psi_0, \varphi_0\)	scattering function defined by I-11
5 (m) (7; µ, µo)	function defined by I-13

thickness of dispersion t total integrated transmission  $\mathbf{T}$ diffuse portion of integrated transmission  $T_{D}$  $\boldsymbol{\mathcal{J}}(\boldsymbol{\pi};\boldsymbol{\mu},\boldsymbol{\varphi};\boldsymbol{\mu}_{\bullet},\boldsymbol{\varphi}_{\bullet})$  transmission defined by I-16  $\mathcal{I}^{(m)}(\tau_1;\mu,\varphi_0)$ function defined by I-18 weighting coefficients for use in numerical integration by Wi I-33, i = 0, ...., pspecific values of ordinates to be used in numerical  $\mathbf{x}_{\mathbf{i}}$ integration by I-33,  $i = 0, \ldots, p$  $\mu_{\text{O}}\text{-interpolation factor for integrated reflectance}$ х  $\mu_{\text{O}}\text{-interpolation factor for total integrated transmission}$ X six-flux representation of sidewise scattering component Chandrasekhar's X function—the same as  $\Psi^{\circ}_{\bullet}(\tau_{1},\mu)$  $X(\mu)$ Chandrasekhar's Y function--the same as  $\phi_{\bullet}^{\bullet}(\tau,\mu)$  $Y(\mu)$ distance into the dispersion in optical units Z  $\pi D/\lambda$ , circumference of particle expressed in wavelength α positive roots of equation (III-26)  $\alpha_{\mathtt{i}}$ positive roots of the equation  $\beta^{l_1} - \frac{b}{c} \beta^2 + \frac{l}{c} = 0$ , βį where  $a = \frac{5}{21}$ ,  $b = \frac{5}{7}$ ,  $c = \frac{4}{105}$ Snm Kronecker delta ( $\delta_{nm}$ = 0 when m  $\neq$  n;  $\delta_{nm}$ = 1.0 when m = n) error in numerical integration with I-35 દ azimuthal angle of radiation azimuthal angle of incident radiation θ0

wavelength of incident radiation λ cos 0 direction of radiation μ  $\cos \theta_0$  direction of incident radiation  $\mu_{o}$ energy density of isotropic scattering at a position r  $\rho(\mathbf{r})$ optical thickness of dispersion  $\tau_{\gamma}$ Total cross-section of a particle (scattering and absorption)  $\phi_1^{\mathrm{m}}(\tau_1,\mu)$ functions defined by I-2 bearing angle of radiation bearing angle of incident radiation **4**0  $\pmb{\psi_{i}^{\textit{m}}}\!(\tau_{\text{l}},\mu)$ functions defined by I-l albedo for single scattering, ratio of scattered to ω, intercepted radiation coefficients in Legendre polynomial expansion of phase  $\omega_{l}$ function  $\omega_{\ell} \frac{(\ell-m)!}{(\ell+m)!}$ 

#### ABSTRACT

The effect of anisotropic scattering on radiative transfer was investigated theoretically. Exact solutions were obtained for a plane source obliquely incident on a multiply scattering, parallel plane dispersion by numerical integration. The results are utilized to evaluate the various approximate models which have been proposed for radiative transfer and neutron diffusion. A new variable-order, diffusion-type model is proposed as an improved approximation.

#### INTRODUCTION

The long range objective of this research is to develop methods for predicting the transport of thermal radiation through the atmosphere.

The scattering of radiation by dispersed material presents a problem of considerable mathematical difficulty. Exact solutions have been completed only for highly idealized conditions: isotropic or Rayleigh scattering in a one-dimensional configuration (1).

Accordingly many approximate models have been developed (2). Examples are the diffusion model and the various discrete flux models.

In the initial phase of this research a discrete flux model was used to obtain approximate results for a finite, spherical source both above and inside a haze (3). The work was then redirected to the development of exact solutions for particular conditions as discussed below.

The lack of exact solutions for general conditions makes it difficult to compare and evaluate the various approximate models which have been proposed. As a consequence of a general conference on radiant transport (4) and subsequent discussions with Dr. S. Chandrasekhar and AFSWP personnel in which this situation was stressed, the immediate objective of this research was shifted to the development of exact solutions for idealized but significant conditions. Particular attention

has been given to the development of exact solutions for anisotropic scattering. Solutions were first obtained for parallel-plane radiation, obliquely incident on a semi-infinite dispersion. These results are presented in the annual report for the previous year (5). The present report presents similar results for parallel-plane radiation obliquely incident on parallel-plane dispersions of finite thickness. The method of solution follows the lead of Chandrasekhar (1) and involves the reiterative numerical solution of integral equations on a high speed computer. (The University of Michigan IBM-704). The formulation of the method of solution and the numerical results are presented in Part I of the report. In Part II numerical values obtained from various approximate models are compared with the exact values presented in Part I. Suggested formulas for interpolation and extrapolation of the exact values are also included. None of the approximate models which had previously been proposed proved to be entirely satisfactory and a new variable-order, diffusion-type model is developed in Part III.

During the period while this work was being done an approximate but exceedingly detailed incremental model for one-dimensional radiant transport with anisotropic scattering was developed by the Internuclear Corporation (6). Their final results have not been available for comparison with the values obtained in this investigation but comparison of preliminary values for a few cases indicates that their model may be quite accurate.

In the continuation of this research it is planned 1) to refine the exact method of solution to expedite the numerical calculations

2) to extend the new variable order diffusion-type model for anistropic scattering and 3) to develop exact or nearly exact solutions for two-dimensional problems utilizing numerical methods and/or the model.

#### A. Mathematical Formulation

### 1. Geometry and Coordinate System.

This investigation is concerned with the irradiation of a semi-infinite slab by a uniform parallel flux as indicated in Figure I-1. The Z = 0 plane represents the interface between region (1) consisting of free space containing the source and region (2) consisting of a uniform dispersion of multiply scattering and absorbing particles each characterized by the same absorption cross-section, scattering cross-section and phase function,  $p(\cos \theta)$ , for single scattering. The slab has an optical thickness  $\tau_1$  and the Z =  $\tau_1$  plane represents the interface between region (2) and region (3) consisting of free space. The objective of this study is to determine the intensities of transmitted and reflected radiation as a function of position, direction, and the characteristics of the source and dispersion.

The coordinate system used in the analysis is shown in Figure I-2. An incident flux of intensity  $I_0$  strikes the interface at an angle  $\theta_0$  with respect to the normal (Z axis). The direction at Z = 0 of the reflected intensity leaving the dispersion is defined by the azimuthal angle  $\theta$ , measured from the normal, and the bearing angle  $\phi$  measured from the Y axis. For simplicity the coordinate system is

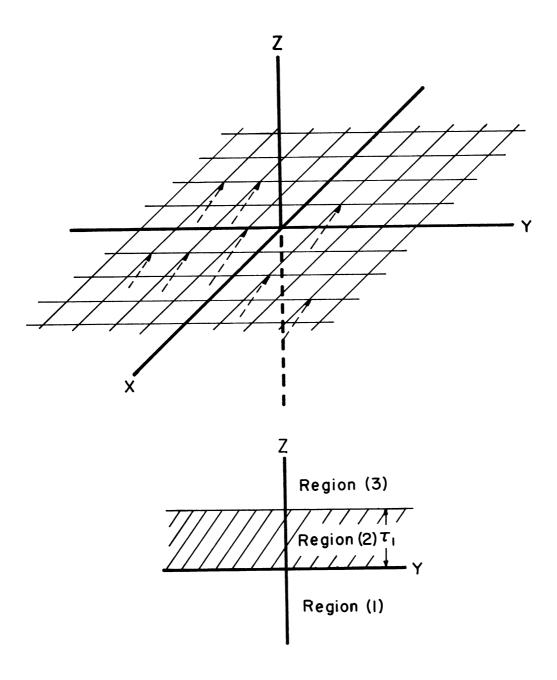
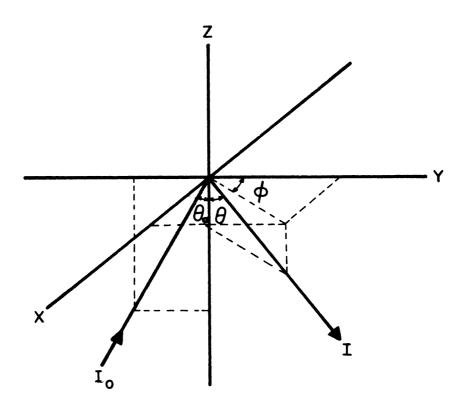


Figure I-l Geometry of the Problem 5



XY plane: media interphase

 $I_{o}$ : incident flux

I : emergent flux under consideration

 $\theta_0 = \cos^{-1} \mu_0$ ,  $\theta = \cos^{-1} \mu$ ,  $\phi = \text{Bearing angle}$ 

Figure I-2 Coordinate system

oriented so that  $\varphi_{\mathbf{0}}$  = 0, i.e.,  $\mathbf{I}_{\mathrm{O}}$  lies in the YZ plane.

2. The  $\psi$  and  $\phi$  Functions.

The following integral equations for the two basic functions  $\psi_1^m$  and  $\phi_1^m$  have been derived by Chandrasekhar<sup>(1)</sup>\*.

$$\begin{split} \psi_{1}^{m}(\tau_{i}, \mu) &= P_{1}^{m}(\mu) + \frac{1}{2} \mu \sum_{k=m}^{N} (-1)^{1+k} \omega_{k}^{m} \int_{0}^{1} \left[ \psi_{k}^{m}(\tau_{i}, \mu) \psi_{k}^{m}(\tau_{i}, \mu') - \psi_{k}^{m}(\tau_{i}, \mu) \psi_{k}^{m}(\tau_{i}, \mu') \right] P_{1}^{m}(\mu') \frac{d\mu'}{\mu + \mu'} \\ &- \psi_{1}^{m}(\tau_{i}, \mu) &= e^{-\tau_{i}/\mu} P_{1}^{m}(\mu) + \frac{1}{2} \mu \sum_{k=m}^{N} \omega_{k}^{m} \int_{0}^{1} \left( \phi_{k}^{m}(\tau_{i}, \mu) \psi_{k}^{m}(\tau_{i}, \mu') - \psi_{k}^{m}(\tau_{i}, \mu) \psi_{k}^{m}(\tau_{i}, \mu') \right] P_{1}^{m}(\mu') \frac{d\mu'}{\mu - \mu'} \end{split}$$

$$I-2$$

$$m = 0, 1, \dots, N$$

 $l = m, m + 1, \dots, N$ 

where,

 $\mu = \cos \theta$ 

 $\tau_1$  = optical thickness of media (number of mean free paths for scattering) =  $n\sigma^{\,t}$ 

 $P_1^m(\mu)$  = associated Legendre polynomials

 $\mathbb{N}$  = number of terms in the expansion of the phase function

$$p(\mu) = \sum_{l=0}^{N} \mathbf{w}_{l} P_{l}(\mu)$$

 $\omega_1$  = coefficients of phase function expansion

$$\omega_{\perp}^{m} = w_{\perp} \frac{(1-m)!}{(1+m)!}$$

\* Hereafter referred to as R. T.

The equations may be written in more convenient form by defining

$$F_{k}^{m}\left(\boldsymbol{\tau}_{i},\boldsymbol{\mu},\boldsymbol{\mu}'\right) = \frac{\omega_{k}^{m}}{\boldsymbol{\mu}+\boldsymbol{\mu}'}\left[\boldsymbol{\psi}_{k}^{m}\left(\boldsymbol{\tau}_{i},\boldsymbol{\mu}\right)\boldsymbol{\psi}_{k}^{m}\left(\boldsymbol{\tau}_{i},\boldsymbol{\mu}'\right) - \boldsymbol{\phi}_{k}^{m}\left(\boldsymbol{\tau}_{i},\boldsymbol{\mu}\right)\boldsymbol{\phi}_{k}^{m}\left(\boldsymbol{\tau}_{i},\boldsymbol{\mu}'\right)\right] \qquad \qquad \mathbb{I}-3$$

$$G_{k}^{m}(z_{i}, \mu, \mu') = \frac{\omega_{k}^{m}}{\mu - \mu'} \left\{ \phi_{k}^{m}(z_{i}, \mu) \psi_{k}^{m}(z_{i}, \mu') - \psi_{k}^{m}(z_{i}, \mu) \phi_{k}^{m}(z_{i}, \mu') \right\}$$
 I-4

The equations then become

$$\Psi_{L}^{m}(\tau_{1}, \mu) = P_{L}^{m}(\mu) + \frac{1}{2}\mu \sum_{k=m}^{N} (-1)^{l+k} \int_{0}^{1} F_{k}^{m}(\tau_{1}, \mu, \mu') P_{L}^{m}(\mu') d\mu' \qquad I-5$$

$$\Phi_{L}^{m}(z_{i}, \mu) = e^{-z_{i}/\mu} P_{L}^{m}(\mu) + \frac{1}{2} \mu \sum_{k=m}^{N} \int_{0}^{1} G_{k}^{m}(z_{i}, \mu, \mu') P_{L}^{m}(\mu') d\mu'$$
I-6

$$m = 0, 1, \dots, N$$

$$l = m, m+1, \cdots, N$$

For any value of N, equations I-1 and I-2 yield N + 1 systems of simultaneous integral equations. Each system consists of 2(N+1), 2N,..., 2 simultaneous equations corresponding respectively to  $m=N,\;N-1,\ldots 0. \quad \text{As an example, using the notation } F_1^m=F_1^m(\tau_1,\mu,\mu^*)$  and  $G_1^m=G_1^m(\tau_1,\mu,\mu^*)$ , the systems of integral equations for N = 2 are for m = 0

$$\psi^{\circ}(z_1, \mu) = 1 + \frac{1}{2}\mu \left( [F^{\circ} - F^{\circ}_1 + F^{\circ}_2] d\mu' \right)$$
 I-7a

$$\Psi_{1}^{\circ}(z_{1}, \mu) = \mu - \frac{1}{2}\mu\int_{0}^{1} \left[F_{0}^{\circ} - F_{1}^{\circ} + F_{2}^{\circ}\right]\mu'd\mu'$$
I-76

$$\psi_{2}^{\circ}(\tau_{i},\mu) = \frac{1}{2}(3\mu^{2}-1) + \frac{1}{2}\mu([F_{i}^{\circ}-F_{i}^{\circ}+F_{2}^{\circ}]\frac{1}{2}(3\mu'^{2}-1)d\mu'$$
 I-7c

$$\phi_{\circ}^{\circ}(z_{1}, \mu) = e^{-z_{1}/\mu} + \frac{1}{2}\mu \left\{ \left[ G_{\circ}^{\circ} + G_{1}^{\circ} + G_{2}^{\circ} \right] d\mu' \right\}$$
 I-7d

$$\begin{split} & \phi_{1}^{\circ}(z_{1}, \mu) = \mu e^{-z_{1}/\mu} + \frac{1}{2}\mu \int_{0}^{1} \left[ G_{0}^{\circ} + G_{1}^{\circ} + G_{2}^{\circ} \right] \mu' d\mu' \\ & \phi_{2}^{\circ}(z_{1}, \mu) = \frac{1}{2} (3\mu^{2} - 1) e^{-z_{1}/\mu} + \frac{1}{2}\mu \int_{0}^{1} \left[ G_{0}^{\circ} + G_{1}^{\circ} + G_{2}^{\circ} \right] \frac{1}{2} (3\mu'^{2} - 1) d\mu' \end{split}$$

# for m = 1

$$\psi'_{1}(\tau_{1},\mu) = (1-\mu^{2})^{1/2} + \frac{1}{2}\mu \left( [F'_{1} - F'_{2}](1-\mu'^{2})^{1/2} d\mu' \right)$$
 I-8a

$$\psi_{2}^{'}(\tau_{1},\mu) = 3(1-\mu^{2})^{\nu_{2}}\mu - \frac{1}{2}\mu \left( [F'_{1} - F'_{2}] 3(1-\mu'^{2})^{\nu_{2}}\mu' d\mu' \right)$$
I-8b

$$\phi_{1}^{+}(z_{1},\mu) = (1-\mu^{2})^{1/2}e^{-z_{1}/\mu} + \frac{1}{2}\mu\int_{0}^{1}[G_{1}^{+}+G_{2}^{+}](1-\mu'^{2})^{0/2}d\mu'$$
I-8c

$$\phi_{2}^{'}(z_{1},\mu) = 3(1-\mu^{2})^{1/2}\mu e^{-z_{1}/\mu} + \frac{1}{2}\mu \int_{0}^{1} \left[G_{1}^{'} + G_{2}^{'}\right] 3(1-\mu'^{2})^{1/2}\mu' d\mu' \qquad I-8d$$

## m = 2

$$\psi_{2}^{2}(\tau_{1}, \mu) = 3(1 - \mu^{2}) + \frac{1}{2}\mu \int_{0}^{1} (F_{2}^{2}) 3(1 - \mu'^{2}) d\mu'$$
I-9a

$$\phi_{2}^{2}(\tau_{1},\mu) = 3(1-\mu^{2})e^{-\tau_{1}/\mu} + \frac{1}{2}\mu\int_{0}^{1}(G_{2}^{2})3(1-\mu'^{2})d\mu'$$
I-9b

### 3. The Intensity of Reflected and Transmitted Radiation

The intensity of radiation reflected diffusely by a semi-infinite slab at the plane Z = 0 in the direction ( $\mu$ ,  $\varphi$ ) may be expressed in terms of a scattering function\*

$$I(o, \mu, \varphi) = \frac{I_{\bullet}(-\mu_{\bullet}, \varphi_{\bullet})}{4\pi\mu} S(\tau_{\bullet}; \mu, \varphi; \mu_{\bullet}, \varphi_{\bullet})$$
I-10

where

I<sub>•</sub>(-
$$μ$$
•,  $φ$ •) = incident flux intensity in the direction  $(-μ$ <sub>O</sub>, $φ$ <sub>O</sub>).

I(
$$\rho$$
,  $\mu$ ,  $\varphi$ ) = intensity of flux reflected in the direction ( $\mu$ ,  $\varphi$ ) at Z = 0.

**S** = Scattering function

The scattering function is\*\*

$$S(\tau_i; \mu, \varphi; \mu_o, \varphi_o) = \sum_{m=0}^{N} S_i^{(m)}(\tau_i, \mu, \mu_o) \cos m (\varphi_o - \varphi)$$
 I-11

or for  $\varphi_{0} = 0$ 

$$S(z_1; \mu, \varphi; \mu_0, \varphi_0) = \sum_{m=0}^{N} S_1^{(m)}(z_1, \mu, \mu_0) \cos m\varphi$$
 I-12

where,

$$S_{i}^{(m)}(\tau_{i}; \mu, \mu_{o}) = (2 - \delta_{o,m}) \mu \mu_{o} \sum_{l=m}^{N} (-1)^{l+m} F_{l}^{m}(\tau_{i}, \mu, \mu_{o})$$
 I-13

<sup>\*</sup> R. T. Page 161

<sup>\*\*</sup> R. T. Pages 180 and 177

The diffuse reflected intensity may then be written as

$$\frac{I(0,\mu,\varphi)}{I_{\bullet}(-\mu_{\bullet},0)} = \frac{\mu_{\bullet}}{4\pi} \sum_{m=0}^{N} \cos m\varphi \left(2 - \delta_{o,m}\right) \sum_{l=m}^{N} (-1)^{m+l} F_{l}^{m} \left(\tau_{l},\mu,\mu_{o}\right)$$
I-14

In a similar manner the intensity of radiation transmitted diffusely through the slab at the Z =  $\tau_1$  plane may be expressed in terms of a transmission function\*

$$I(\tau_{i}, -\mu, \varphi) = \frac{I_{o}(-\mu_{o}, \varphi_{o})}{4\pi\mu} \mathcal{J}(\tau_{ij}, \mu, \varphi; \mu_{o}, \varphi_{o})$$
I-15

where,

I.(- $\mu$ .,  $\varphi$ .) = incident flux intensity in the direction  $(-\mu_{\rm O}, \varphi_{\rm O})$ 

I(τ,- $\mu$ , $\varphi$ ) = intensity of flux at Z =  $\tau_1$  transmitted in the direction  $(-\mu, \varphi)$ 

 $\mathcal{J}$  = transmission function

The transmission function is\*\*

$$\mathcal{J}(\tau_i; \mu, \varphi; \mu_o, \varphi_o) = \sum_{m=0}^{N} \mathcal{J}^{(m)}(\tau_i, \mu, \mu_o) \cos m (\varphi_o - \varphi)$$
 I-16

or for  $\varphi_0 = 0$ 

$$\mathcal{J}(\tau_{i}; \mu, \phi; \mu_{o}, \varphi_{o}) = \sum_{m=0}^{N} \mathcal{J}^{(m)}(\tau_{i}, \mu, \mu_{o}) \cos m \varphi$$
 I-17

where.

$$\mathcal{T}^{(m)}(\tau_{i}, \mu, \mu_{o}) = \mu \mu_{o} (2 - \delta_{o,m}) \sum_{l=m}^{N} G_{l}^{m}(\tau_{i}, \mu, \mu_{o})$$
 I-18

The diffuse portion of the transmitted intensity may then be written as

$$\frac{I(\tau_{1}, -\mu, \varphi)}{I_{o}(-\mu_{o}, 0)} = \frac{\mu_{o}}{4\pi} \sum_{m=0}^{N} (2 - \delta_{o, m}) \cos m\varphi \sum_{l=m}^{N} G_{l}^{m} (\tau_{1}, \mu, \mu_{o})$$
I-19

\* R. T. Page 161

\*\* R. T. Pages 180 and 177

Equations I-14 and I-19 may be written in a slightly more convenient form by noting that

$$\cos m\varphi = A_0 + \sum_{i=1}^{m} A_i \cos^i \varphi$$
 I-20

where, for example, for m equal to 2,  $A_0 = -1$ ,  $A_1 = 0$ , and  $A_2 = 2$ .

Thus, new expressions can be derived for the reflected and the diffuse portion of the transmitted intensities of the following form

$$\frac{\mathrm{I}\left(0,\mu,\varphi\right)}{\mathrm{I}_{\circ}\left(-\mu_{\circ},0\right)}=b_{\circ}\left(\tau_{1},\mu,\mu_{\circ}\right)+\sum_{m=1}^{N}b_{m}\left(\tau_{1},\mu,\mu_{\circ}\right)\cos^{m}\varphi$$
 I-21

and

$$\frac{I(\tau_{i}, -\mu, \varphi)}{I_{\circ}(-\mu_{\circ}, 0)} = c_{\circ}(\tau_{i}, \mu, \mu_{\circ}) + \sum_{m=1}^{N} c_{m}(\tau_{i}, \mu, \mu_{\circ}) \cos^{m}\varphi$$
I-22

This means that for a N term phase function and a given combination of  $\tau_1$ ,  $\mu$ , and  $\mu_0$ , the intensities need be calculated at only N + 1 values of  $\varphi$  in order to obtain the complete angular distributions. The explicit relations for  $b_m(\tau_1,\mu,\mu_0)$  and  $c_m(\tau_1,\mu,\mu_0)$  in terms of  $\psi_1^m(\tau_1,\mu,\mu_0)$  and  $\phi_1^m(\tau_1,\mu,\mu_0)$  for N = 2 are derived in Appendix A-1.

## 4. Integrated Reflectance and Transmission

The integrated reflectance is defined as the ratio of reflected to incident power passing through a unit area of the interface. The flux,  ${\bf I}_0$ , is defined as the radiant energy passing through a unit area normal to the direction of propagation per unit time. Therefore, the

incident power striking a unit area of the interface is  $I_0 \times \mu_0$ . Similarly, the reflected power passing out of a unit area of the interface is  $\int_0^1 \int_0^{2\pi} I(0,\mu,\Psi) \mu d\mu d\Psi$ . The integrated reflectance

$$R = \frac{1}{I(-\mu_0)\mu_0} \left( \int_{0}^{2\pi} I(0,\mu,\varphi) \mu d\mu d\varphi \right)$$
I-23a

Replacing  $I(0,\mu,\varphi)$  by I-14 and noting that

$$\int_{0}^{2\pi} F(\mu) \cos m\theta \, d\theta = 2\pi F(\mu) \text{ or o } \text{for } m = 0 \text{ or } m \neq 0 \text{ respectively } I-24$$

$$R = \frac{1}{2} \sum_{l=0}^{N} (-1)^{l} \int_{0}^{1} \mu F_{l}^{\bullet}(\tau_{l}, \mu, \mu_{\bullet}) \, d\mu$$

$$I-23b$$

In a similar manner the integrated diffuse transmission  $T_p$  is defined as the ratio of transmitted to incident power passing through a unit area of interface.

Thus

is then expressed as

$$T_{p} = \frac{1}{I_{o}(-\mu_{o})\mu_{o}} \int_{0}^{2\pi} I(z_{i}, -\mu, \varphi) \mu d\mu d\varphi$$

$$I-25a$$

and substituting I-19 for  $I(\tau_1, -\mu, \varphi)$ 

$$T_p = \frac{1}{2} \sum_{k=0}^{N} {\mu_{G_k}(z_i, \mu, \mu_o) d\mu}$$
 I-25b

The total transmission T is the sum of the direct transmission and the diffuse transmission.

Thus 
$$T = e^{-\tau_1/\mu_0} + T_p$$

It has been shown (5) that, since,

$$\psi_{i}^{\circ}(\tau_{i}, \mu) = \mu - \frac{1}{2}\mu \sum_{h=0}^{N} (-1)^{h} \left( F_{h}^{\circ}(\tau_{i}, \mu, \mu') \mu' d\mu' \right)$$
I-27

then 
$$1 - \frac{\Psi_{i}^{\circ}(\tau_{i}, \mu)}{\mu} = \frac{1}{2} \sum_{k=0}^{N} (-1)^{k} \int_{0}^{1} F_{k}^{\circ}(\tau_{i}, \mu, \mu') \mu' d\mu'$$

$$R = 1 - \frac{\psi_i^*(\tau_i, \mu_o)}{\mu_o} \quad \text{for} \quad N \geqslant 1, \quad \mu_o > 0$$

Similarly it can be shown that, since

$$\phi_{i}^{\circ}(z_{i}, \mu) = \mu e^{-z_{i}/\mu} + \frac{1}{2} \mu \sum_{k=0}^{N} \int_{0}^{1} G_{k}^{\circ}(z_{i}, \mu, \mu') \mu' d\mu'$$
I-30

Then

$$\frac{\Phi_{i}^{*}(\tau_{i}, \mu)}{\mu} - e^{-\tau_{i}/\mu} = \frac{1}{2} \sum_{k=0}^{N} \int_{0}^{1} G_{k}^{*}(\tau_{i}, \mu, \mu') \mu' d\mu'$$
I-31

and

$$T_{p} = \frac{\phi_{i}^{\circ}(z_{i}, \mu_{e})}{\mu_{e}} - e^{-z_{i}/\mu_{e}}$$
I-32

For

The total transmission is simply

$$T = \frac{\phi_{i}(z_{i}, \mu_{o})}{\mu_{o}}$$

## B. Numerical Procedure

1. The General Numerical Problem.

As shown in the previous section, the intensities of reflected and transmitted radiation for a particular phase function for single scattering can be expressed in terms of the solutions of sets of simultaneous, non-linear integral equations. The numerical problem, then, is to solve these sets in order to obtain values for the intensity, I, in terms of its parameters.

The most straightforward method of obtaining numerical solutions of integral equations is simple iteration using the original equations. For example, in the case of a single integral equation of the type

$$f(x) = g(x) + \lambda \int_{0}^{1} f(x)f(x')h(x')dx'$$
I-34

where f(x) is the function to be found

g(x) and h(x) are known functions of x and  $\lambda$  is some known constant,

the method involves the following:

- 1. Replacing the continuous function, f(x), by a tabular function,  $f^*(x)$  where  $f^*(x)$  consists of a set of p values of  $f(x_i)$ ,  $\left\{f_0,f_1,\ldots f_p\right\}$ , corresponding to a set of values of  $x_i$ ,  $\left\{x_0,x_1,\ldots x_p\right\}$ , where p is the number of points.
- 2. Choosing an initial  $f^{*}(0)(x)$ .
- 3. Evaluating the right side of I-34 by some mechanical quadrature method--thus, obtaining a new  $f*^{(1)}(x)$ .

- 4. Testing to see if  $f^{*(1)}(x)$  agrees with  $f^{*(0)}(x)$  everywhere within acceptable limits and, if not,
- 5. Repeating the process using  $f^{*}(1)(x)$  in place of  $f^{*}(0)(x)$ .
- 6. The process is continued until the sequence:  $f^{*(0)}(x)$ ,  $f^{*(1)}(x)$ ,...,  $f^{*(n)}(x)$  converges to some limiting value.

Advancing one step in the iteration procedure to obtain  $f^{*(n+1)}(x)$  from  $f^{*(n)}(x)$  requires making p+l numerical integrations.

Integration by mechanical quadrature consists of replacing the integral by a sum, or in other words

$$\int_{a}^{1} f(x) dx = \sum_{i=0}^{p} w_{i} f(x_{i}) + \epsilon$$

where  $\boldsymbol{\epsilon}$  is the error. Thus, selecting a specific method of numerical integration involves selecting the set of values of x,  $\{x_0, x, x_p\}$ , and the set of weighting coefficients,  $\{w_0, w_1, \dots w_p\}$ .

For integration by the trapezoidal rule

$$x_i = i/p$$
 $W_0 = W_p = 1/2p$ 
 $W_i = 1/p \quad (i = 1,...,p-1)$ 

For integration by Simpson's rule

$$p = 2k \quad \text{where k is an anteger}$$

$$x_i = i/p$$

$$W_0 = W_p = 1/3p$$

$$W_i = 4/3p \quad i = 1, 3, ..., p-1$$

$$W_i = 2/3p \quad i = 2, 4, ..., p-2$$

For integration using n-point Gaussian quadrature

$$x_i = (x_{i+1} + 1)/2$$
 I-38

where  $r_i$  is the i-th root of  $P_n(x)$ 

$$W_i = \frac{1}{2 P'_n(x)} \int_{-1}^{1} \frac{P_n(x) dx}{x - x_i}$$

Thus, the definition of a specific numerical procedure involves:

- 1. Selecting the number of points, p+1,  $\{x_i\}$  and  $\{W_i\}$ .
- 2. Selecting the initial function  $f^{*(0)}(x)$ .
- 3. Selecting a test to determine when the sequence  $f^{*(0)}(x)$ ,  $f^{*(1)}(x)$ ,...,  $f^{*(n)}(x)$  has converged.

It should be noted that the fact that the function f\*(x) has converged to some limiting value merely assures that a solution has been found to the following equation:

$$f^*(x_i) = g^*(x_i) + \lambda \sum_{i=0}^{p} W_i f^*(x_i) f^*(x_i) h^*(x_i)$$
I-39

for j = 0, l,..., p. This does not, however, guarantee that f\*(x) = f(x). Convergence in this latter sense may be tested by comparison of the results with (a) results obtained by other methods (b) results obtained using more accurate methods of numerical integration, and (c) solutions for idealized and limiting cases.

Extension of the preceeding comments to the case of N+l simultaneous integral equations is straightforward. The general form of the equations is

$$f_{\ell}(x) = g_{\ell}(x) + \sum_{k=0}^{N} \lambda_{k,\ell} \left( f_{k}(x) f_{k}(x') h_{\ell}(x') dx', (1 = 0, 1, \dots, N) \right)$$

Advancing one step in the iteration procedure involves obtaining  $f_{0}^{\star(n+1)} \, \dots \, f_{N}^{\star(n+1)} \qquad \text{from } f_{0}^{\star(n)} \, \dots \, f_{N}^{\star(n)}.$ 

#### 2. Specific Numerical Techniques Used in Computation.

A computer program was written to solve equations I-l and I-2 in their general form by the method of simple iteration. The programs were written in a form suitable for use on the IBM 704 or 709 and are described in detail in Appendix C. All calculations, however, were performed on the IBM 704.

Every effort was made to develop general techniques suitable for solution of equations with an arbitrary value of N; arbitrary set of constants in the phase function; an arbitrary method of numerical integration; any thickness,  $\tau_1$ ; any albedo for single scattering,  $\omega_{\bullet}$ ; and any choice of starting functions. The same procedure was used for all values of m and N to obtain all of the functions presented in this report. The only limitation on the size of the problem that can be solved is that of computer time required and number of storage locations available.

In practice, almost all of the results were obtained using five point Gausian quadrature. For this case, using numerical values given by Lowan, Davids, and Levenson (7)

$W_{O}$	=	0.11846	$\mu_{O}$	=	0.04691
$\mathbf{W}_{1}$	=	0.23931			0.23077
$\overline{W_{P}^{P}}$	=	0.28444			0.50000
$\overline{W}_3$	=	0.23931			0.76923
$W_{4}$	=	0.11846			0.95309

In addition some values were obtained for comparison using a 10-point Simpson's Rule technique.

The starting functions used for all calculations were

$$\psi_{1}^{m}(\tau_{1}, \mu)^{(0)} = P_{1}^{m}(\mu)$$

$$\psi_{1}^{m}(\tau_{1}, \mu)^{(0)} = e^{-\tau_{1}/\mu}P_{1}^{m}(\mu)$$
I-41b

After considerable experimenting, it was found that for  $m \neq 0$  a simple numerical test was satisfactory to determine convergence. In other words, when

$$\max_{i} \left| f^{*(n+1)}(x_i) - f^{*(n)}(x_i) \right| < \text{Tolerance}$$
 I-42a

the iteration was stopped. For all cases the tolerance was chosen as 0.001.

For  ${\tt m}={\tt O},$  however, it was found that a better test was to stop the iteration whenever

$$[f^{*(n+1)}(x_i) - f^{*(n)}(x_i)][f^{*(n)}(x_i) - f^{*(n-1)}(x_i)] < 0$$

for  $\phi$ , or  $\psi$ , or whenever I-42a was satisfied. This effectively stops the iteration as soon as the incremental change in any function reverses sign. This gave the best agreement with known solutions for special cases (i.e. the  $X(\mu)$  and  $Y(\mu)$  functions of Chandrasekhar (8) which are the same as  $\phi$ ,  $(\mu)$  and  $\psi$ ,  $(\mu)$  in this study).

Some difficulty occurs when evaluating  $G_{i}^{m}$  (7,  $\mu$ ,  $\mu'$ )

for the case  $\mu = \mu'$ 

$$\frac{G_{\ell}^{m}(\tau_{1},\mu,\mu)}{\omega_{\ell}^{m}} = \lim_{\mu \to \mu'} \frac{\psi_{\ell}^{m}(\mu) \phi_{\ell}^{m}(\mu') - \phi_{\ell}^{m}(\mu) \psi_{\ell}^{m}(\mu')}{\mu - \mu'}$$
I-43

However, by L'Hospital's rule the limit is

$$\frac{G_{i}^{m}(z_{i},\mu,\mu)}{\omega_{i}^{m}} = \Psi_{i}^{m}(\mu) \frac{d\phi_{i}^{m}(\mu)}{d\mu'} - \phi_{i}^{m}(\mu) \frac{d\psi_{i}^{m}(\mu)}{d\mu'}$$

$$I-44$$

In practice,  $\frac{d\phi_{\mu}^{m}(\mu)}{d\mu'}$  and  $\frac{d\psi_{\mu}^{m}(\mu)}{d\mu'}$  were evaluated at a given value of  $\mu$  by fitting a second order polynomial to the three nearest values of the function, and differentiating analytically.

In addition, a program was written for calculating the integrated diffuse reflection and transmission for any set of functions using equations I-23b and I-25b.

A discussion of the effectiveness of the numerical procedures will be deferred to section C-5 of this report after the presentation of numerical results.

## C. Solutions to Specific Problems

1. Choice of Independent Variables.

The choice of specific values of the independent variables was based on the following:

- 1. To indicate quantitatively the effect of the independent variables for representative conditions.
- 2. To gain experience in techniques of numerical solution of the problems which would be useful in solving further problems.
- 3. To obtain solutions for specific problems which can be used as a standard for evaluating approximate techniques.

In order to completely specify the problem the following parameters must be chosen: 1) number of terms in the phase function, N; 2) coefficients of the angular distribution function,  $a_0$ ,  $a_1$ ,...aN; 3) albedo for single scattering,  $\omega$ , ; and 4) the optical thickness of the dispersion,  $\tau_1$ .

In order to obtain a reasonable picture of the effect of  $\omega$ , and  $\tau_1$  a minimum of three values of each variable must be used, making a total of at least nine solutions for each phase function. There is an almost endless number of possible phase functions. For instance, for only a two term phase function ( $a_0 = 1$ ) using 10 values of  $a_1$  and  $a_2$  would require 900 solutions. For a five term phase function, 900,000 solutions would be required.

In addition, for a given value of m, the effort in solving one set of integral equations increases with the square of the number of

equations in a set. Also, the number of sets of equations increases with N. For example, in comparable units of computer time, the effort in solving the equations for N = 0, 1, 2, 3, 4, and 5 is

$$N = 0$$

N = 1

$$(1)^2 + (2)^2$$
 = 5 units

N = 2

$$(1)^2 + (2)^2 + (3)^2$$
 = 14 units

N = 3

$$(1)^2 + (2)^2 + (3)^2 + (4)^2$$
 = 30 units

N = 4

$$(1)^2 + (2)^2 + (3)^2 + (4)^2 + (5)^2$$
 = 55 units

N = 5

$$(1)^2 + (2)^2 + (3)^2 + (4)^2 + (5)^2 + (6)^2 = 91 \text{ units}$$

(Note: in the preceeding it is assumed that all coefficients in the phase function are non-zero; in estimating the labor required to solve a set of equations N may be taken as the number of non-zero terms in the phase function)

Thus, it is apparent that the selection of representative phase functions is very important and that the number of non-zero terms in the phase function must be kept small in order to be able to vary  $\omega_{\bullet}$  and  $\tau_1$  sufficiently to achieve the previously stated objectives.

#### a) Choice of the Phase Function

The construction of a phase function for scattering by spherical particles depends on the wave length  $\lambda$  of the radiation, the diameter of the scattering particle D and the index of refraction m. The problem is fully described by Chu and Churchill<sup>(9)</sup>. In summary, the phase function,  $p(\cos \theta)$  or  $p(\mu)$ , can be written

$$p(\mu) = 1 + \sum_{n=1}^{\infty} a_n P_n(\mu)$$
 I-45

where  $P(\mu)/4\pi$  is the fraction of randomly polarized radiation scattered by a spherical particle into a unit solid angle in the direction  $\theta$ .

 $\theta$  is  $\cos^{-1}(\mu)$ 

a<sub>n</sub>(d,m) angular distribution coefficients

 $P_n(\mu)$  are Legendre polynomials

#### $d = \pi D / \lambda$

The phase function (I-45) is for non-aborbing particles  $(\omega_0 = 1.0)$ . For partially absorbing particles with complex indices of refraction, the angular distribution coefficients must be recomputed for each value of  $\omega_0$ . As a good approximation, however, the phase function may be considered to be  $\omega_0$ p ( $\mu$ ) where p( $\mu$ ) is the phase function of a non-absorbing particle.

Probably the most important single characteristic of any given phase function is the fraction scattered forward and backward, F and B,

where

$$F = 2\pi \left( p(\mu) d\mu \right)$$

$$B = 2\pi \int_{-1}^{0} P(\mu) d\mu$$
 I-47

$$B = 1 - F$$
, since  $2\pi \int_{-1}^{1} p(\mu) d\mu = 1$ 

A good measure of how "peaked" is the phase function is the ratio of the second moment to the zeroth moment in the forward and backward direction

$$P = 2\pi \int_{-1}^{1} p(\mu) \mu^{2} d\mu / F$$
 I-49

and

$$P_{B} = 2\pi \left( \int_{-1}^{0} p(\mu) \mu^{2} d\mu / B \right)$$
 I-50

where P = Peakedness in Forward Direction.

 $P_{B}$  = Peakedness in Backward Direction.

In addition, other characteristics are

FP = Height of the Forward Peak, i.e. p(1.0)

BP = Height of the Backward Peak, i.e. p(-1.0)

The equations for these parameters in terms of the distribution coefficients,  $a_0,\;\dots,\;a_N^{}$  are

$$F = \frac{1}{2} + \frac{a_1}{4} - \frac{a_3}{16} + \frac{a_5}{32} - \frac{5a_7}{256} + \frac{7a_9}{512} - \frac{21a_{11}}{2048} + \frac{33a_{13}}{4096} - \dots$$

$$B = 1 - F$$

$$P = \left[ \frac{1}{6} + \frac{a_1}{8} + \frac{a_2}{15} + \frac{a_3}{48} - \frac{a_5}{384} + \frac{a_7}{1280} - \frac{a_9}{3072} + \frac{a_{11}}{6144} - \frac{3a_{13}}{32716} + \cdots \right] / F \qquad I-52$$

$$P_{B} = \left[ \frac{1}{6} - \frac{a_{1}}{8} + \frac{a_{2}}{15} - \frac{a_{3}}{48} + \frac{a_{5}}{384} - \frac{a_{7}}{1280} + \frac{a_{9}}{3072} - \frac{a_{11}}{6144} + \frac{3a_{13}}{32716} - \cdots \right] / B$$
I-53

$$FP = 1 + a_1 + a_2 + a_3 + a_4 + \cdots$$

$$BP = 1 - a_1 + a_2 - a_3 + a_4 - \cdots$$

$$I-55$$

For N = 0 the phase function has only one term,  $a_0 = 1$ . Thus, all quantities are fixed

$$F = 1/2$$

$$B = 1/2$$

$$P = 1/3$$

$$P_{B} = 1/3$$

$$FP = BP = 1$$

For N = 1 the phase function is completely specified by choosing any one characteristic such as F. For n = 2, the phase function is completely specified by choosing F and P (or any other pair of characteristics). As more terms are added, more characteristics of the function may be choosen independently. In order to consider a significant range of  $\omega$ , and  $\tau_1$  with the computer time available a two term phase function was used for all computations including a few special cases in which some of the coefficients were zero.

A careful study was made of all possible two term phase functions satisfying the conditions that  $p(\mu)$  be positive for all values of  $\mu$  and that FP be greater than BP. It was found that the phase function

$$p(\mu) = \omega_0 \left[ 1 + a_1 P_1(\mu) + a_2 P_2(\mu) \right]$$

gave the greatest range of F and P while satisfying the conditions that  $p(\mu) \geqslant 0$  for  $0 \leqslant \mu \leqslant 1$  and FP  $\geqslant$  BP. Figure I-3 is a plot of FP vs F with P and BP as parameters. The shaded area represents all values of F and FP for which  $p(\mu)$  was always positive.

From consideration of Figure I-3, the following phase functions were selected for computation.

TABLE I-1 Phase Functions Used in Computation

$\mathtt{a}_1$	a <sub>2</sub>	F	P	Remarks
0.0000	0.00000	0.50000	0.33333	Isotropic Scattering
0.00000	0.50000	0.50000	0.40000	Rayleigh Scattering
1.00000	1.81650	0.75000	0.5504}	Arbitrary: same F, different P
1.00000	1.04904	0.75000	0.4821	dillerent r
1.00000	0.00000	0.75000	0.3899	
1.73205	1.00000	0.9330	0.4821	Largest F for N = 2
1.73458	2.24153	0.9336	0.5821	Same Phase Function as Reference (5).

The last phase function listed in Table I-l is the same as that in reference (5). Computations with it were made only for  $\tau_1=\infty$ , since  $p(\mu)$  was not always positive for this function. The results were used as a check on the accuracy of the numerical techniques.

All of the phase functions are plotted in figures 4a, 4b, and 4 c. In these figures the function  $f(\theta)$  is plotted which is equivalent to  $p(\cos\,\theta)/4\pi \quad .$ 

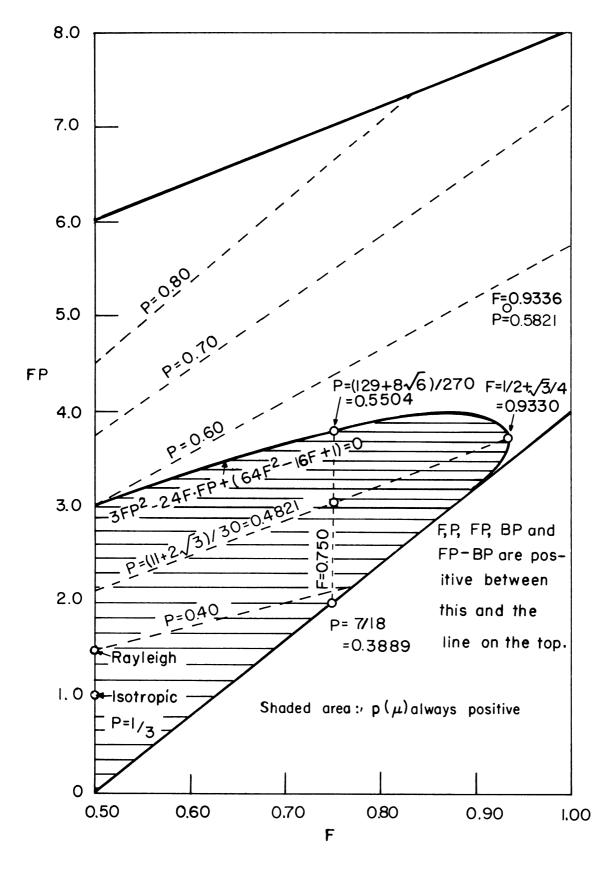


Figure I-3 Characteristics of a two term phase function

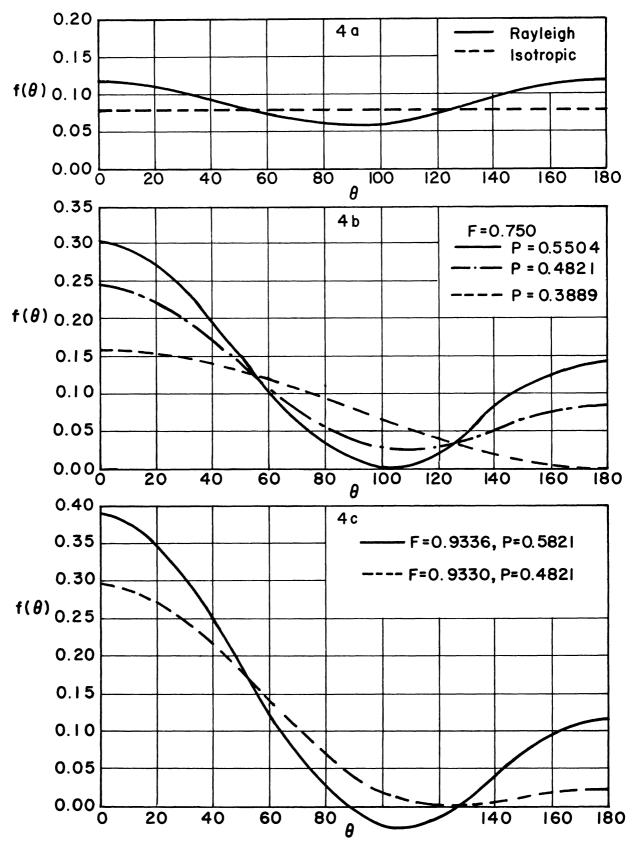


Figure I-4 Phase functions used in calculations, F = 0.50, F = 0.75, and F = 0.933, Parameters of P

# b) Choice of Other Independent Variables

The independent variables other than the phase function (i.e.  $\omega_{\bullet}$  and  $\tau_{1}$ ) were selected to give as good an indication of their effect as possible.

For most of the calculations  $\omega$ , was chosen as 0.90. This was done to be able to compare results for  $\tau_1$  = 00 with previous solutions. For every phase function results were obtained at least for  $\tau_1$  = 1.00. For isotropic scattering, Rayleigh scattering and for the function corresponding to F = 0.9330 and P = 0.4821, results were obtained at several values of  $\tau_1$ .

A total of 36 problems were solved; the values of the parameters for each are listed in Table I-2.

Table I-2 Values of the Parameters Used in Computation

Problem Number	F	P	$ au_{ extsf{l}}$	ω.
1	0.500	0.333	0.05	0.90
2	0.500	0.333	0.25	0.90
3	0.500	0.333	0.50	0.90
4	0.500	0.333	1.00	0.90
5 6	0.500	0.333	1.50	0.90
	0.500	0.333	2.00	0.90
7	0.500	0.333	00	0.90
8	0.500	0.333	0.05	0.60
9	0.500	0.333	0.50	0.60
10	0.500	0.333	00	0.60
11	0.500	0.333	0.05	0.30
12	0.500	0.333	0.50	0.30
13	0.500	0.333	00	0.30
14	0.500	0.400	0.05	0.90
15	0.500	0.400	0.25	0.90
16	0.500	0.400	0.50	0.90
17	0.500	0.400	1.00	0.90
18	0.500	0.400	1.50	0.90
19	0.500	0.400	2.00	0.90
20	0.500	0.400	00	0.90
21	0.750	0.5504	1.00	0.90
22	0.750	0.4921	1.00	0.90
23	0.750	0.3889	1.00	0.90
21+	0.9330	0.4821	0.05	0.90
25	0.9330	0.4921	0.25	0.90
26	0.9330	0.4921	0.50	0.90
27	0.9330	0.4821	1.00	0.90
28	0.9330	0.4821	2.00	0.90
29	0.9330	0.4821	00	0.90
30	0.9330	0.4821	0.05	0.60
31	0.9330	0.4821	0.50	0.60
32	0.9330	0.4821	00	0.60
33	0.9330	0.4821	0.05	0.30
34	0.9330	0.4821	0.50	0.30
35	0.9330	0.4821	00	0.30
36	0.9336	0.5821	00	0.90

# 2. The $\psi$ and $\phi$ Functions

For each set of parameters as indicated in Table I-2 the complete set of  $\psi_1^m$  and  $\phi_1^m$  functions are calculated at the Gaussian quadrature points. These functions are tabulated in Appendix B. The number of iterations required to evaluate the functions for each value of m is given. In every case except Problem 36 for m = 0, the starting functions are those given by Equations I-41a and I-41b.

3. Sample Values of the Intensity of Reflected and Transmitted Radiation The normalized intensity of reflected radiation and the diffuse portion of the transmitted radiation are, in general, a function of two angles  $\phi$  and  $\theta$  and four parameters:  $\theta_0$ ,  $\omega_0$ ,  $\tau_1$ , and the phase function.

The effect of the variables is illustrated by considering a sample problem with

- 1)  $\theta_0 = \cos^{-1} 0.5$  (i.e.  $\mu_0 = 0.50$ )
- 2)  $\theta = \cos^{-1} 0.76923$  (i.e.  $\mu = 0.76923$ )
- 3)  $\omega_{\bullet} = 0.90$
- 4)  $\tau_{1} = 1.00$
- 5) A two term phase function such that F = 0.9330 and P = 0.4821

The angular distribution of the specific intensity is obtained as a function of the bearing angle  $\varphi$  as well as these six parameters. The effect of each variable is indicated by varying  $\theta_0$ ,  $\theta$ ,  $\omega$ ,  $\tau_1$ , and the phase function, one at a time, while holding the remaining variables at the above, "standard values". (There are exceptions which are noted.)

The specific intensities for each case were calculated on a desk calculator using I-21 or I-22 and the equations in Appendix A-1. The apparently arbitrary values of  $\mu$  and  $\mu_0$  which were used correspond to the Gaussian points for the five-point Gaussian quadrature utilized for integration. The values of the coefficients  $b_0$ ,  $b_1$ ,  $b_2$ ,  $c_0$ ,  $c_1$ ,

and  $c_2$  which appear in equations I-21 and I-22 are tabulated in Table I-3.

The effect of  $\theta_{\rm O}$  (or  $\mu_{\rm O}$ ) is shown in Figures I-5a and I-5b. In this case  $\mu$  was held at 0.50 and four different values of  $\mu_{\rm O}$  were used. It is observed that there is a rather complicated dependence of the intensity of reflected radiation on  $\mu_{\rm O}$ , with some of the curves overlapping. At  $\varphi$  = 180° the maximum intensity of reflected radiation occurs for  $\mu_{\rm O}$  between 0.23079 and 0.76923 which is not surprising since the cosine of the angle of incidence is 0.50.

The effect of  $\theta$  (or  $\mu$ ) is shown in Figures I-6a and I-6b with all other variables at the "standard" values. In this case the dependence of the intensity of both reflected and transmitted radiation on  $\mu$  produce overlapping curves. The maximum intensity of reflected radiation at  $\varphi$  = 180° apparently occurs at a viewing angle almost the same as the angle of incidence.

The effect of  $\boldsymbol{\omega}_{\bullet}$  is shown in Figures I-7a and I-7b. In this case  $\tau_1$  is 0.50. It is apparent that decreasing  $\boldsymbol{\omega}_{\bullet}$  simply decreases the intensity of the reflected and diffuse portion of the transmitted radiation without affecting the nature of the dependence on  $\boldsymbol{\varphi}_{\bullet}$ .

The effect of  $\tau_1$  is shown in Figures I-8a and I-8b. The effect of decreasing  $\tau_1$  is similar to that of decreasing  $\omega_{\bullet}$ : the intensity is decreased but the dependence on  $\varphi$  is not changed very much.

Table I-3 Values of the Coefficients Used in Equations I-21 and I-22 to Calculate the Specific Intensities of Reflected and Transmitted Radiation.

F	Р	ω。	7,	$\mu_{_{ m O}}$	μ	рО	bl	ъ <sub>2</sub>	c <sub>o</sub>	$^{\mathrm{c}}$	c <sub>2</sub>
0.500	0.333	0.90	1.0	0.5000	0.0469	0.1033	0.0000	0.0000	0.0326	0.0000	0.0000
**	***	11	**	**	0.7692	0.0560	0.0000	0.0000	0.0429	0.0000	0.0000
***	11	***	0.25	11	11	0.0219	0.0000	0.0000	0.0174	0.0000	0.0000
11	11	0.60	0.50	11	11	0.0205	0.0000	0.0000	0.0183	0.0000	0.0000
11	0.400	0.90	1.0	11	0.0469	0.0691	0.0020	0.0765	0.0268	0.0035	0.0130
11	11	***	11	11	0.7692	0.0485	0.0091	0.0133	0.0388	0.0061	0.0091
0.750	0.550	11	11	11	11	0.0078	0.0186	0.0581	0.0348	0.0443	0.0420
11	0.482	***	11	11	11	0.0247	0.0014	0.0300	0.0459	0.0285	0.0209
11	0.389	11	11	11	11	0.0423	0.0195	0.0000	0.0571	0.0131	0.0000
0.933	0.482	11	11	**	11	0.0367	0.0225	0.0297	0.0000	0.0000	0.0000
11	11	11	1.0	11	0.0469	0.0214	0.1189	0.1597	0.0304	0.0329	0.0283
***	11	**	11	11	0.2308	0.0230	0.0845	0.1163	0.0432	0.0502	0.0407
11	11	11	11	11	0.7692	0.0132	0.0201	0.0284	0.0601	0.0467	0.0197
11	11	11	11	11	0.9531	0.0126	0.0063	0.0055	0.0719	0.0245	0.0040
11	11	11	11	0.0469	0.5000	0,0020	0.0112	0.0150	0.0028	0.0031	0.0027
11	11	11	11	0.2308	11	0.0106	0.0390	0.0537	0.0197	0.0232	0.0188
11	11	11	**	0.7692	11	0.0204	0.0310	0.0436	0.0924	0.0718	0.0301
11	11	11	11	0.9531	11	0.0241	0.0120	0.0105	0.1371	0.0467	0.0077
11	11	11	0.50	0.5000	0.7692	0.0044	0.0142	0.0232	0.0392	0.0425	0.0210
11	11	11	0.25	11	11	0.0003	0.0076	0.0157	0.0245	0.0298	0.0150
11	11	0.60	0.50	**	11	0.0007	0.0076	0.0149	0.0218	0.0254	0.0135
11	11	0.30	**	11	**	0.0006	0.0030	0.0072	0.0093	0.0117	0.0065

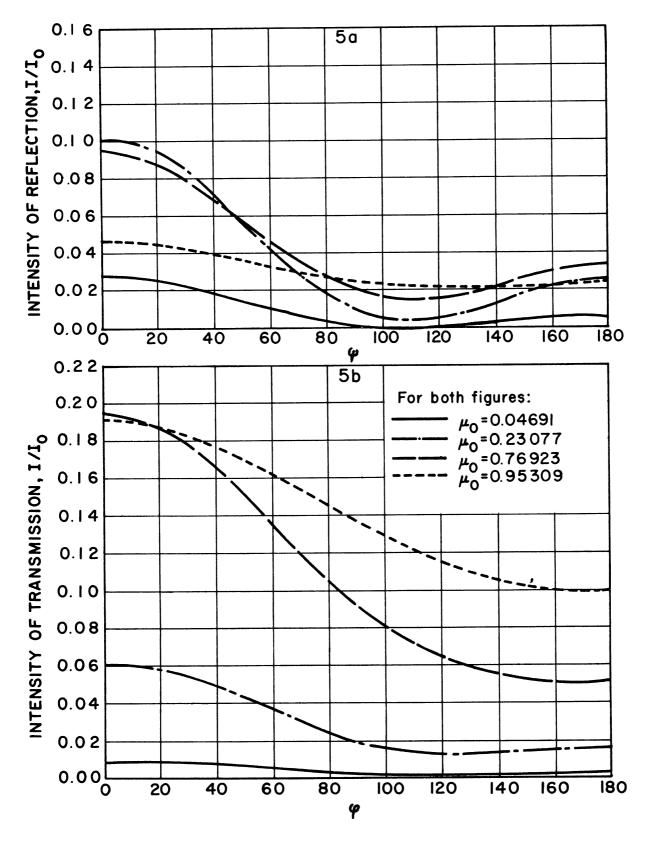


Figure I-5 Intensity of reflected and transmitted radiation as function of bearing angle, I/I $_0$  vs.  $\phi$ , parameters of  $\mu_0$ ,  $\mu$  = 0.50,  $\omega_0$  = 0.90,  $\tau_1$  = 1.00, F = 0.9330, and P = 0.4821

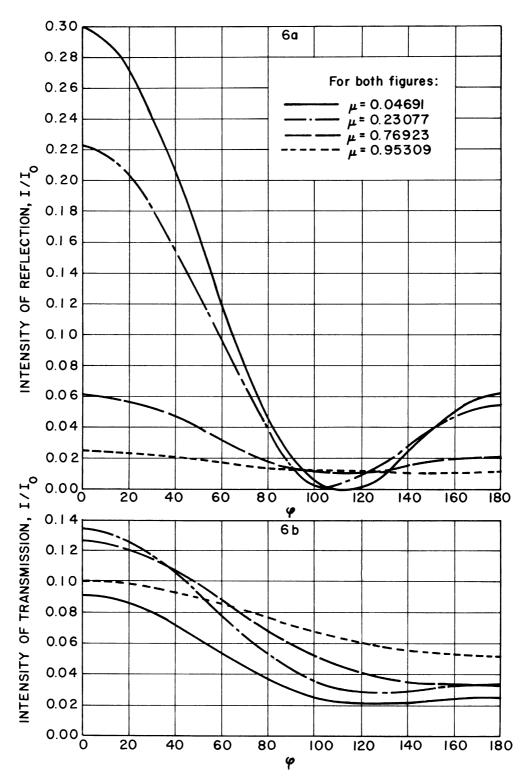


Figure I-6 Intensity of reflected and transmitted radiation as function of bearing angle, I/I $_0$  vs.  $\phi$ , parameters of  $\mu$ ,  $\mu_o$  = 0.50,  $\omega_o$ = 0.90,  $\tau_1$  = 1.00, F = 0.9330, and P = 0.4821

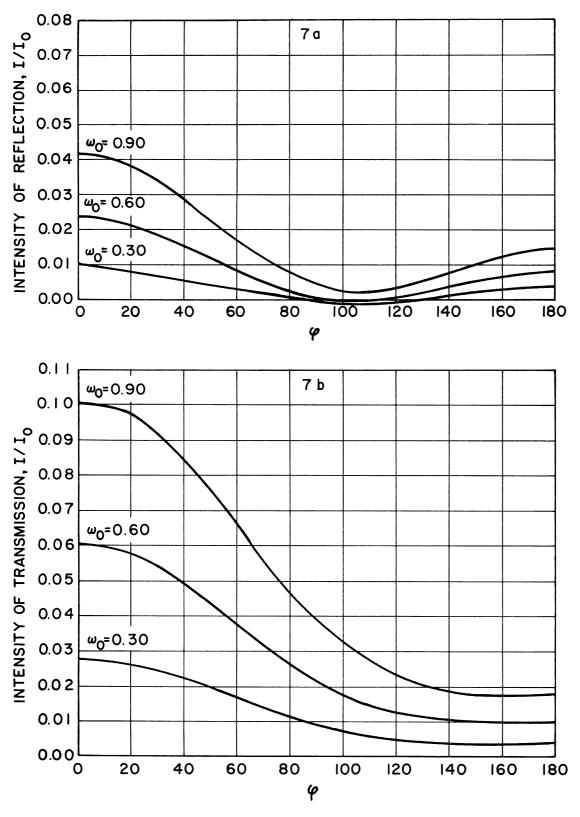


Figure I-7 Intensity of reflected and transmitted radiation as function of bearing angle, I/I $_0$  vs.  $\varphi$  , parameters of  $\omega_o$   $\mu_o$  = 0.50,  $\mu$  = 0.76923,  $\tau_1$  = 0.50, F = 0.9330, P = 0.4821

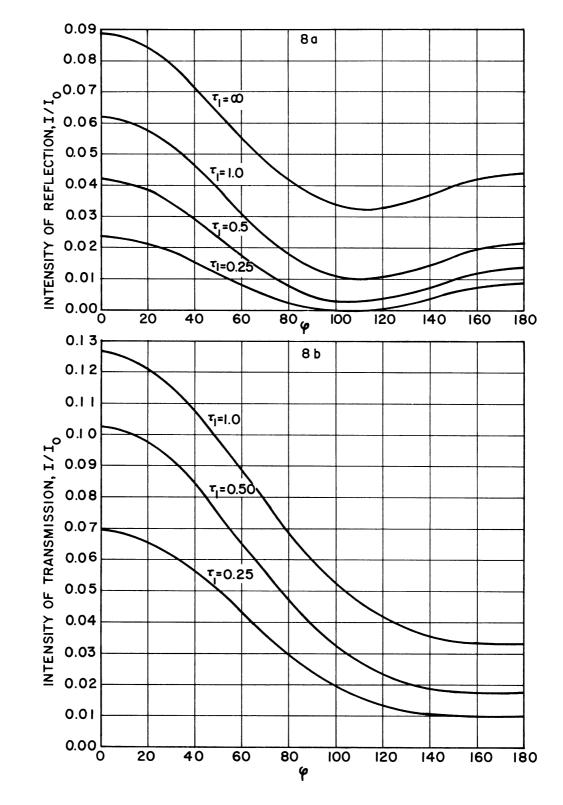


Figure I-8 Intensity of reflected and transmitted radiation as function of bearing angle, I/I $_0$  vs.  $\phi$ , parameters of  $\tau_1$ ,  $\mu_0$  = 0.50,  $\mu$  = 0.76923,  $\omega_{\bullet}$ = 0.90, F = 0.9330, and P = 0.4821

The effect of phase function is indicated in Figures I-9a and I-9b. Results are given for isotropic and Rayleigh scattering in which F is the same (0.50), but P is different (0.3333 and 0.40). An additional set of results was obtained for  $\mu=0.04691$ . It can be seen that the two phase functions produce approximately the same angular distribution of intensity except for reflected radiation with a near grazing viewing angle ( $\mu=0.04691$ ) in which the "dumbell" characteristic of the Rayleigh phase function is accentuated.

Another illustration of the effect of phase function is provided in the polar plots of Figures I-loa and I-lob. In this case the three phase functions for F = 0.750 and P = 0.5504, 0.4821, and 0.3899 are used with the "standard" values of the other variables. It can be seen that the different values of P at the same value of F produce noticeably different angular distributions for the reflected and transmitted intensity.

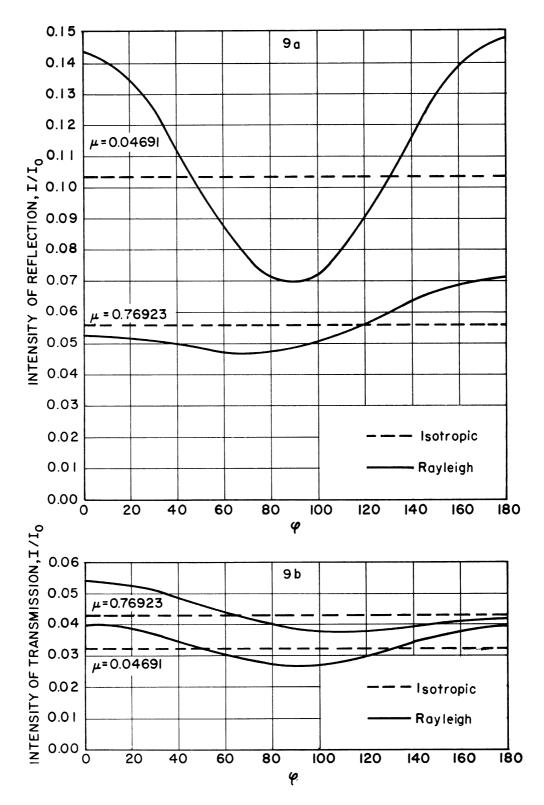


Figure I-9 Intensity of reflected and transmitted radiation as function of bearing angle, I/I $_0$  vs.  $\phi$ , isotropic and Rayleigh scattering,  $\mu$  = 0.04691 and 0.76923,  $\mu_0$  = 0.50,  $\omega_0$ = 0.90,  $\tau_1$  = 1.00

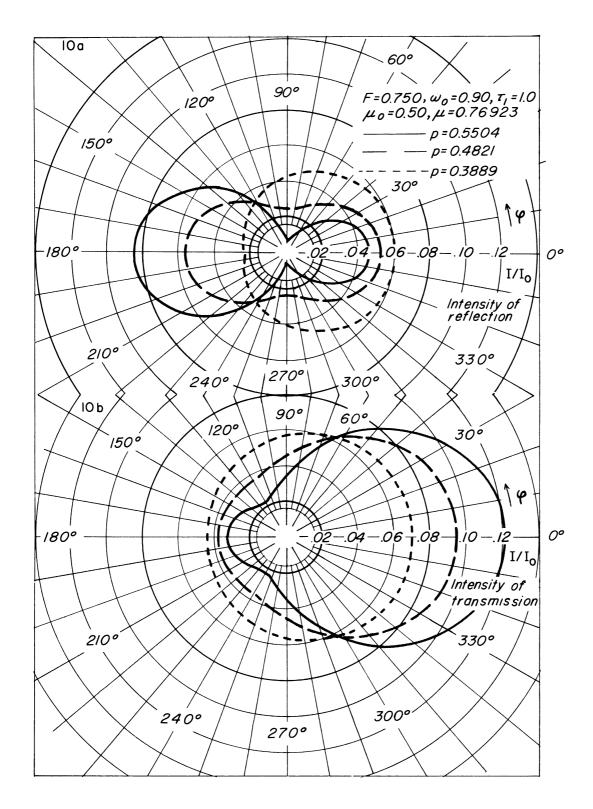


Figure I-10 Intensity of reflected and transmitted radiation as function of bearing angle, I/I $_0$  vs.  $\phi$ , parameters of P,  $\mu_0$  = 0.50,  $\mu$  = 0.76923,  $\omega_e$ = 0.90,  $\tau_1$  = 1.00, F = 0.750

## 4. The Integrated Reflection and Transmission

The integrated reflectance, the diffuse portion of the integrated transmission, and the total transmission were calculated for each set of conditions and the results are tabulated with the  $\psi_{\ell}^{m}$  and  $\phi_{\ell}^{m}$  functions in Appendix B. In addition, these values are presented graphically in Figures I-11 to I-25.

The integrated reflectance for isotropic scattering is shown as a function of angle of incidence in Figures I-11, I-12a, and I-12b for  $\mathbf{w}$ , of 0.90, 0.60, and 0.30 with  $\tau_1$  as a parameter. Similar plots are presented for anisotropic scattering with F = 0.9330 and P = 0.4821 in Figures I-13, I-14a, and I-14b.

The total integrated transmission, including both diffuse and direct components, for isotropic scattering is plotted vs.  $\mu_0$  in Figure I-15 for  $\omega_0$  = 0.90. In Figures I-16 the total transmission for isotropic scattering is presented for  $\tau_1$  = 0.05 and 0.50 with parameters of  $\omega_0$ . Similar plots are given for anisotropic scattering in Figures I-17 and I-18.

Values of the integrated reflectance and transmission were also obtained for Rayleigh scattering for which F is the same (0.50) as for isotropic scattering, but P is different (0.40 compared with 0.333). The results as shown in Figures I-11 and I-15 are almost indistinguishable from those for isotropic scattering. Integrated reflectances for

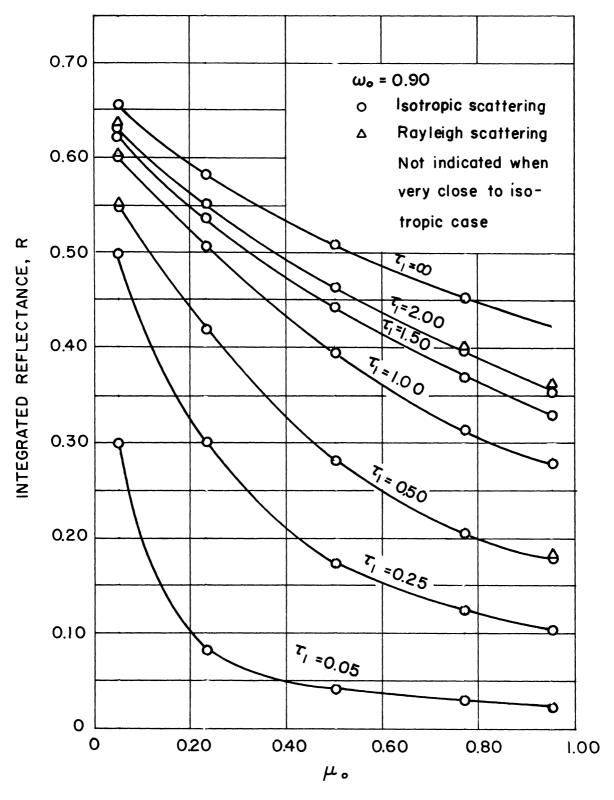
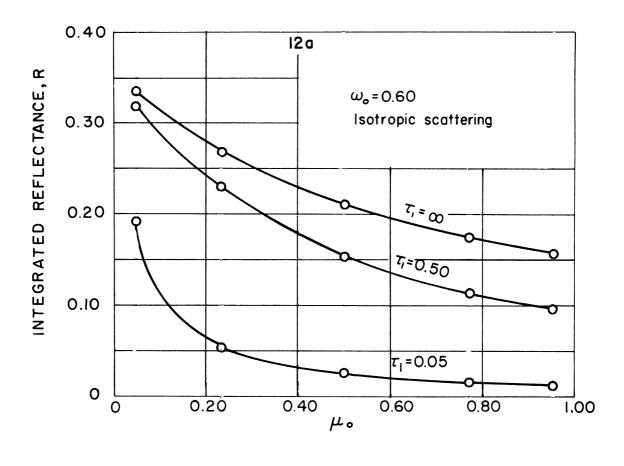


Figure I-ll Integrated reflectance as function of angle of incidence for isotropic and Rayleigh scattering, R vs.  $\mu_0,$  parameters of  $\tau_1,$   $\omega_{\bullet}$  = 0.90



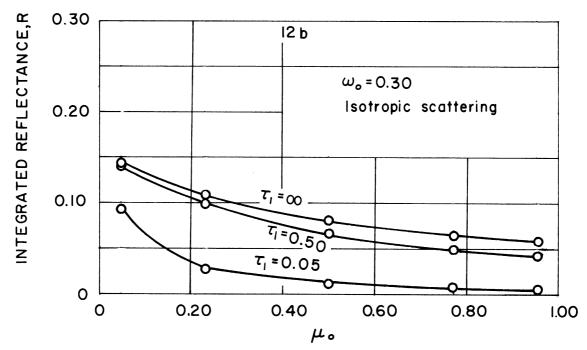


Figure I-12 Integrated reflectance as function of angle of incidence for isotropic scattering, R vs.  $\mu_0$ , parameters of  $\tau_1$ ,  $\omega_0$ = 0.60 and 0.30

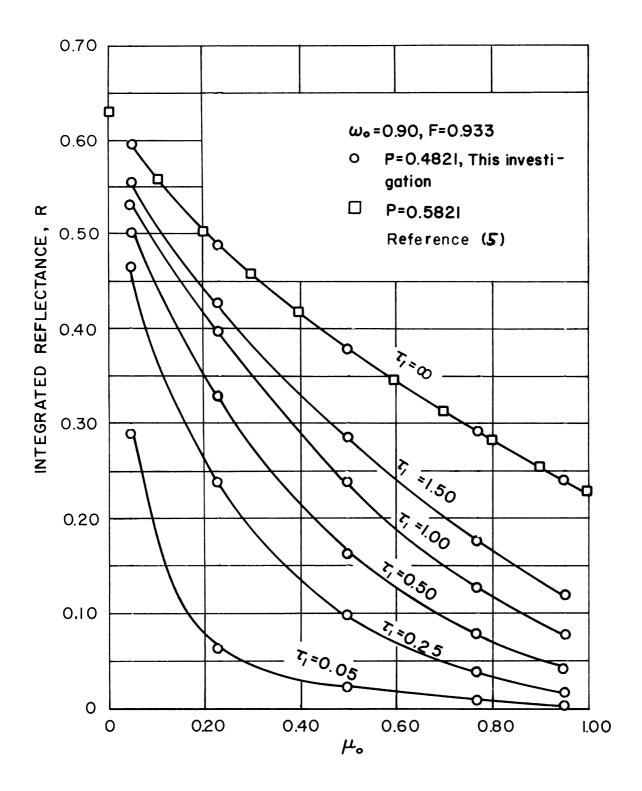


Figure I-13 Integrated reflectance as function of angle of incidence for anisotropic scattering, R vs.  $\mu_0$ , parameters of  $\tau_1$ ,  $\omega_0$  = 0.90, F = 0.9330, P = 0.4821

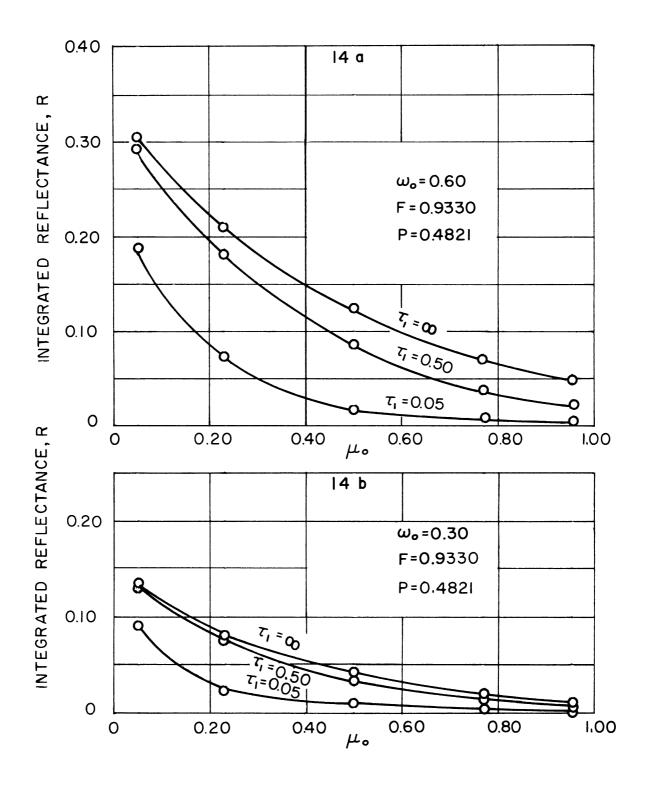


Figure I-14 Integrated reflectance as function of angle of incidence for anisotropic scattering, R vs.  $\mu_0$ , parameters of  $\tau_1$ ,  $\omega_0$  = 0.60 and 0.30, F = 0.9330, P = 0.4821

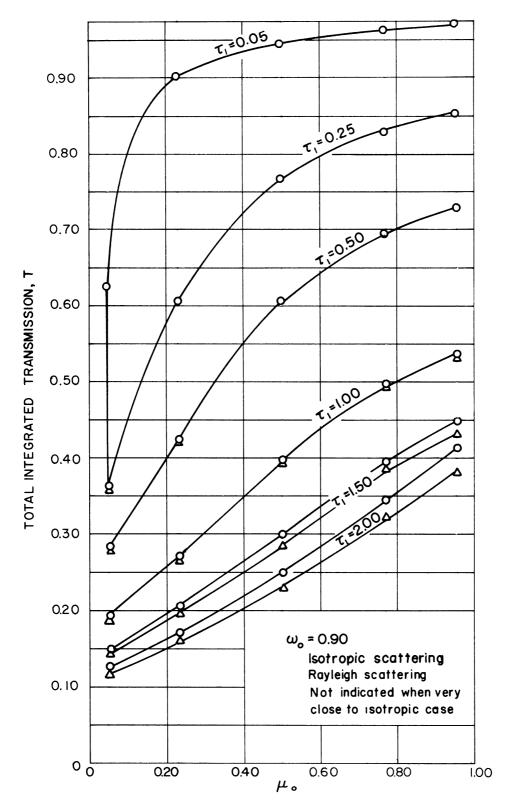


Figure I-15 Total integrated transmission as function of angle of incidence for isotropic and Rayleigh scattering, T vs.  $\mu_{0},$  parameters of  $\tau_{1},\;\omega_{0}\text{=}$  0.90

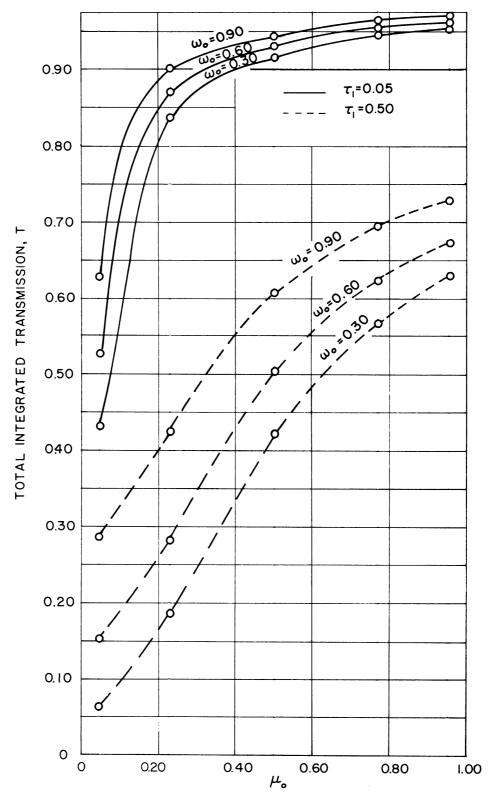


Figure I-16 Total integrated transmission as function of angle of incidence for isotropic scattering, T vs.  $\mu_0$ , parameters of  $\omega_{\bullet}$ ,  $\tau_1$  = 0.05 and 0.50

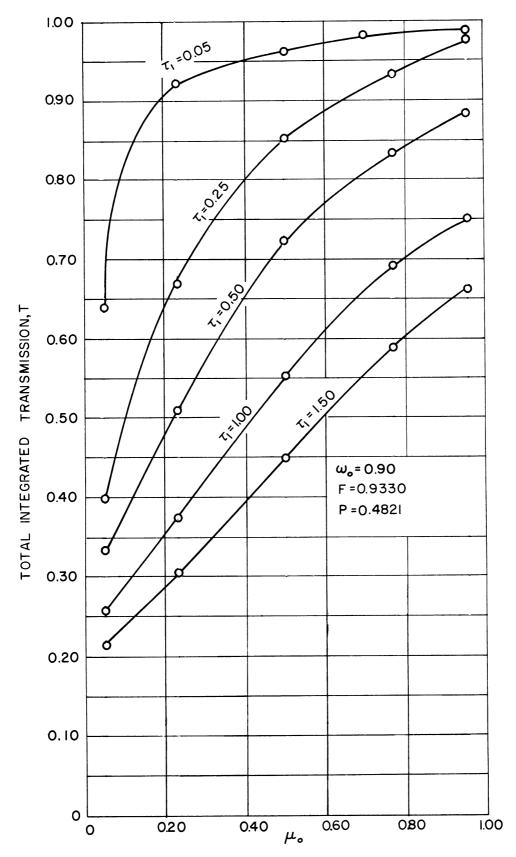


Figure I-17 Total integrated transmission as function of angle of incidence for anisotropic scattering, T vs.  $\mu_0$ , parameters of  $\tau_1$ ,  $\omega_0$ = 0.90, F = 0.9330, P = 0.4821

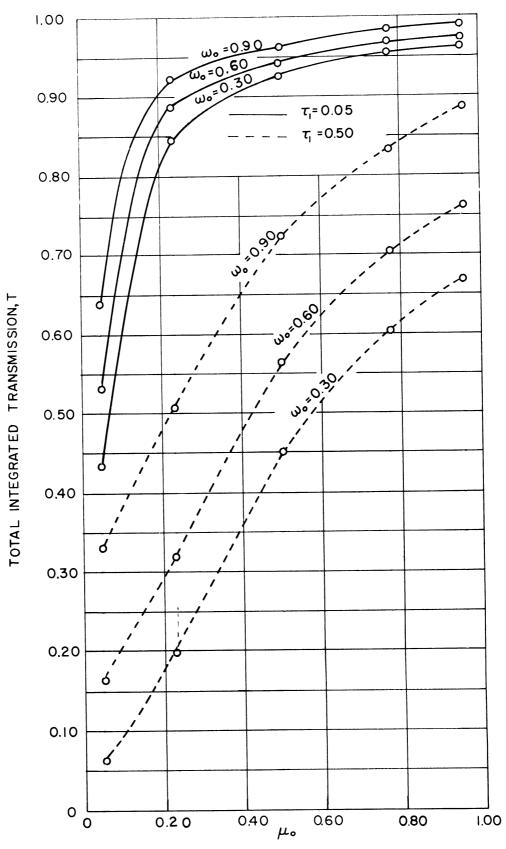
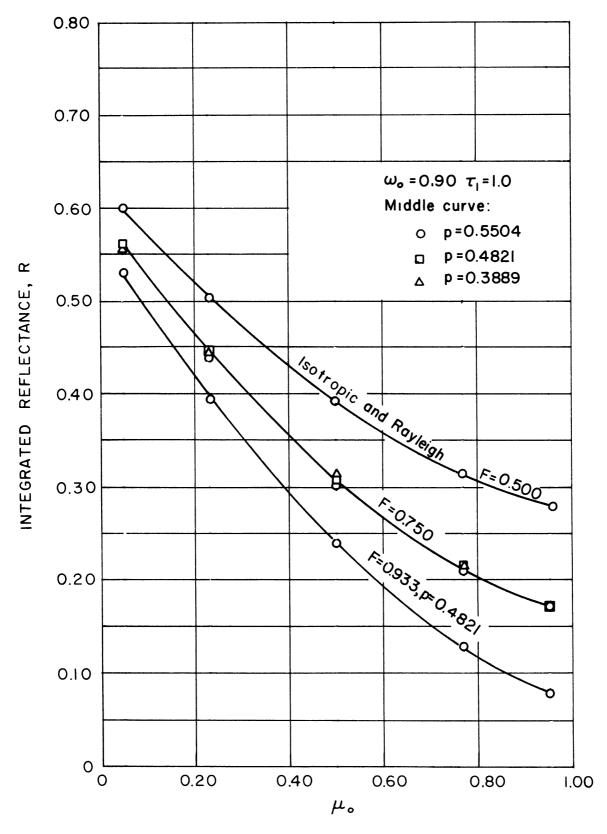


Figure I-18 Total integrated transmission as function of angle of incidence for anisotropic scattering, T vs.  $\mu_0$ , parameters of  $\omega_o$ ,  $\tau_1$  = 0.05 and 0.50, F = 0.9330, P = 0.4821



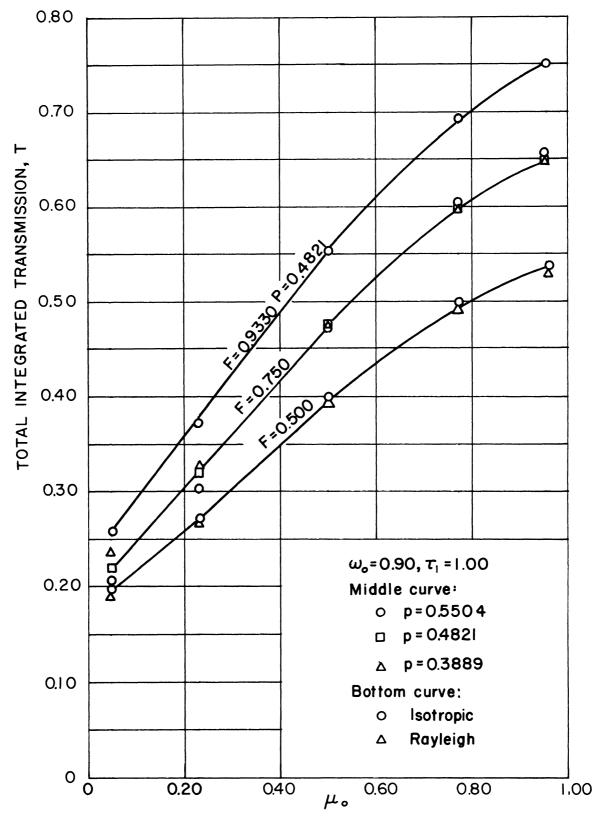


Figure I-20 Total integrated transmission as function of angle of incidence, T vs.  $\mu_0$ , parameters of F,  $\omega_0$ = 0.90 and  $\tau_1$  = 1.00

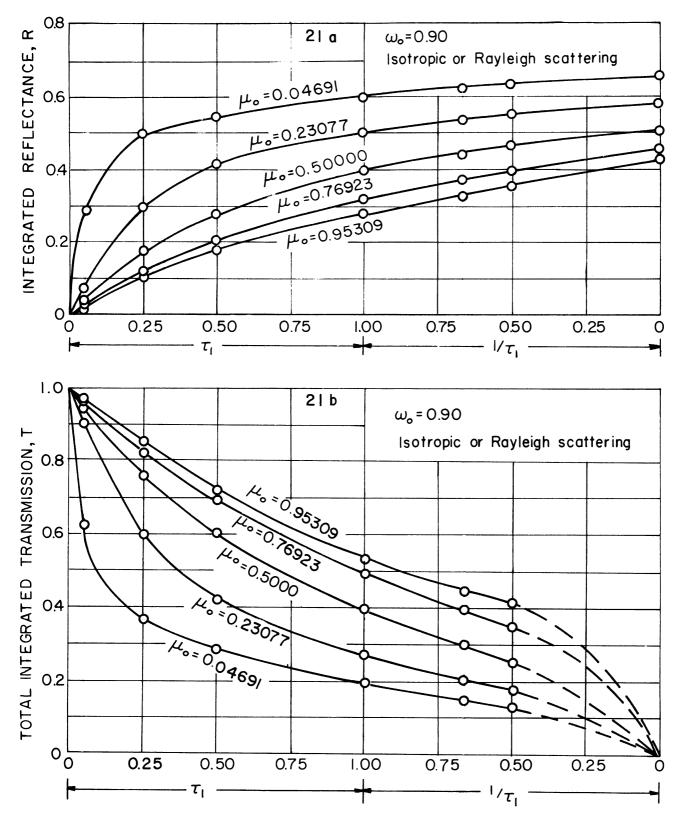


Figure I-21 Integrated reflectance and total integrated transmission as function of optical thickness of dispersion, R and T vs.  $\tau_1$ , parameters of  $\mu_0$  isotropic or Rayleigh scattering F = 0.50

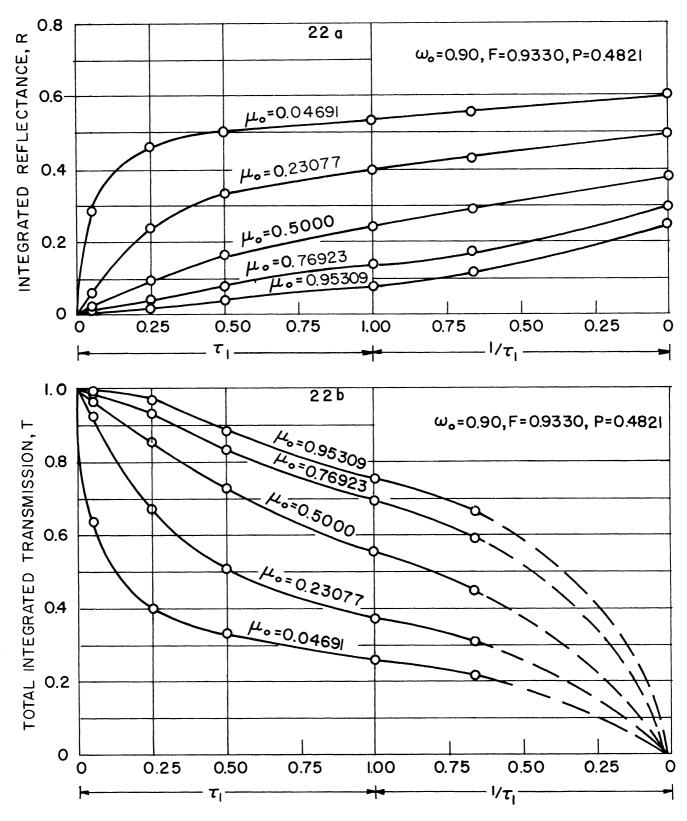


Figure I-22 Integrated reflectance and total integrated transmission as function of optical thickness of dispersion, R and T vs.  $\tau_1$ , parameters of  $\mu_0$ , anisotropic scattering F = 0.9330 and P = 0.4821

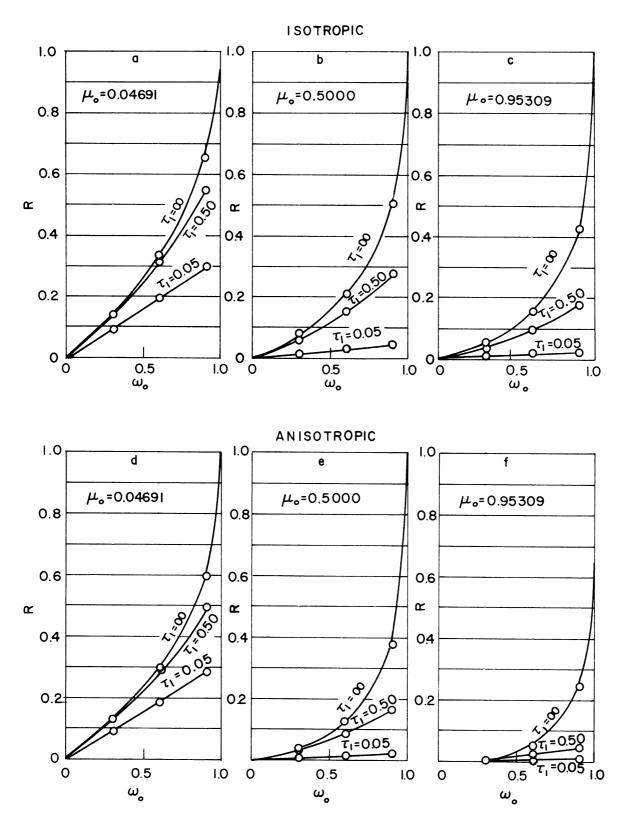


Figure I-23 — Integrated reflectance as function of albedo for single scattering, R vs.  $\omega_o$ , parameters of F,  $\mu_o$ , and  $\tau_1$ .

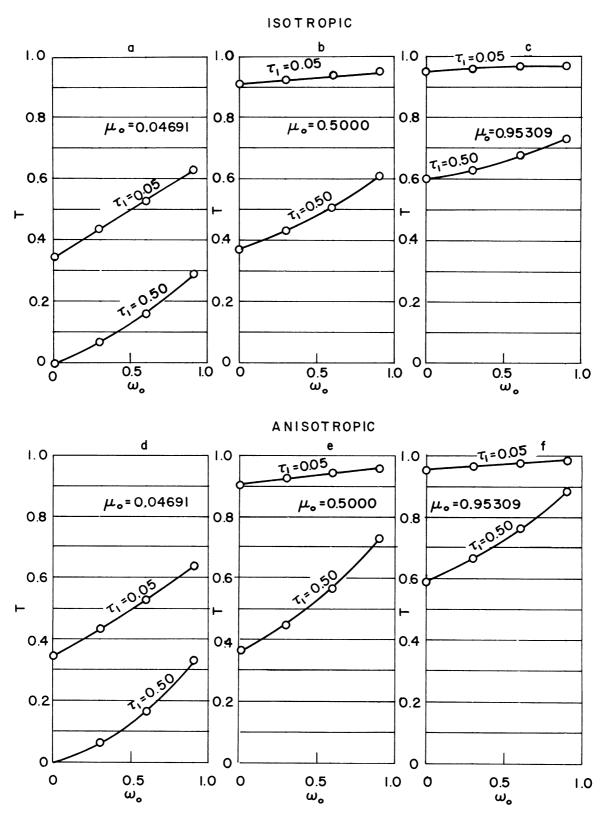
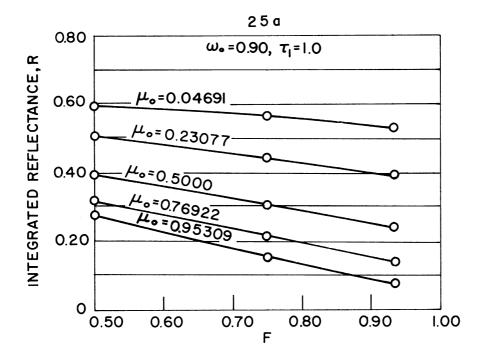


Figure I-24 Total integrated transmission as function of albedo for single scattering, T vs.  $\omega_{\rm o}$ , parameters of F,  $\mu_{\rm O}$ ,  $\tau_{\rm l}$ .



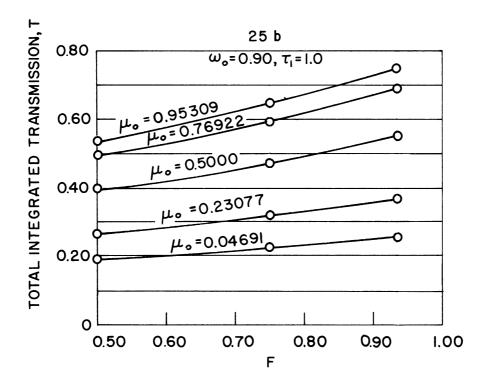


Figure I-25 — Integrated reflectance and total integrated transmission as function of phase function, R and T vs. F, parameters of  $\mu_0,~\omega_0=$  0.90,  $\tau_{\gamma}$  = 1.00

the anisotropic phase function of reference (5) for  $\tau_1 = \infty$  in which F was 0.9336 fall on the same curve as the results for this investigation in which F is almost the same (0.9330), but P is different (0.4821 compared with 0.5821).

The effect of anisotropy is best indicated in Figures I-19 and I-20 where the integrated reflectance and the total transmission are plotted versus angle of incidence with  $\tau_1$  = 1.00 for each of the five phase functions considered in this investigation. Again, it is noted that the results for the three different phase functions with F = 0.750, but with different values of P are almost identical.

The integrated reflectance and total transmission are cross-plotted versus  $\tau_1$  with  $\mu_0$  as a parameter for isotropic scattering in Figures I-21a and I-21b and for anisotropic scattering in Figures I-22a and I-22b. This should provide a satisfactory method of interpolating for different values of  $\tau_1$ . As is expected, the dependence on  $\tau_1$  is almost linear for small values of  $\tau_1$ .

The effect of  $\boldsymbol{\omega}_{\bullet}$  on the integrated reflectance is indicated in Figures I-23a and I-23f for isotropic and anisotropic scattering. The effect of  $\boldsymbol{\omega}_{\bullet}$  on the total transmission is shown in the same manner in Figures I-24a and I-24f. The dependence on  $\boldsymbol{\omega}_{\bullet}$  is almost linear for small values of  $\boldsymbol{\omega}_{\bullet}$ . For small values of  $\tau_{1}$  it is linear over the whole range of  $\boldsymbol{\omega}_{\bullet}$ .

The effect of phase function—in particular, the effect of F—is illustrated by cross plots of R and T versus F in Figures I-25a and I-25b where  $\mu_0$  is a parameter. This indicates the approximate error of neglecting anisotropy and provides a means of estimating the reflection and transmission for an arbitrary dispersion on the basis of the value of F for the phase function. As can be seen from Figures I-19 and I-20, the effect of P is not large. In fact, except for the integrated transmission at the two largest angles of incidence, the complete variation of reflectance and transmission with the range of P considered in this study is indicated by the size of the circles used to plot the points in Figures I-25a and I-25b.

# 5. Discussion of Numerical Technique

It is very difficult to estimate the error involved in the numerical solution of the integral equations. Probably the best indication is comparison with previous solutions and with other methods of solution. The following solutions are available for comparison.

- 1. Chandrasekhar's H functions (1) for  $\omega_{\bullet}$  = 0.90, 0.60, and 0.30. The H( $\mu$ ) functions are the same as the  $\psi_{\bullet}^{\bullet}(\mu)$  functions of this report for  $\tau_{1} = \infty$  and isotropic scattering. The comparison is given in Table I- $\mu$ . The values generally agree to within 0.1%.
- 2. Chandrasekhar's X and Y functions <sup>(8)</sup> for  $\omega_{\bullet}$  = 0.90 and  $\tau_{1}$  = 0.05, 0.25, 0.50, and 1.00. The  $X(\mu)$  and  $Y(\mu)$  are the same as the  $\psi_{\bullet}^{0}(\mu)$  and  $\phi_{0}^{0}(\mu)$  functions of this report for isotropic scattering. The comparison is given in Table I-5. The values generally agree to within a few percent.
- 3. The  $\psi_0^0(\mu)$ ,  $\psi_1^0(\mu)$ ,  $\psi_2^0(\mu)$ ,  $\psi_1^1(\mu)$ ,  $\psi_2^1(\mu)$  and  $\psi_2^2(\mu)$  functions calculated by Churchill, et.al. (5) for  $a_0 = 1.00000$ ,  $a_1 = 1.73458$ ,  $a_2 = 2.24153$ ,  $\omega_{\bullet} = 0.90$  and  $\tau_1 = 00$ . with iteration using a 20-point Simpson's rule technique. The comparison is given in Table I-6. The values generally agree to within 1%.
- 4. The integrated reflectance calculated for Rayleigh scattering,  $\omega_{\bullet}=\text{0.90 and }\tau_{1}=\infty \text{ using the method described in Appendix A-2.}$  The comparison is given in Table I-7. The agreement is very good.

TABLE I-4 Comparison of  $\pmb{\psi}_O^o(\mu)$  Functions for Isotropic Scattering and with Chandrasekhar's H Functions.

ω	0.90		0.	.60	0.30	
μ 0.04691	Η(μ) 1.0937	Ø8(μ) 1.0936	Η(μ) 1.0520	Ø <sub>0</sub> (μ) 1.0520	Η(μ) 1.0233	Ø <sub>O</sub> (μ) 1.0233
0.23077	1.3233	1.3236	1.1584	1.1587	1.0661	1.0663
0.50000	1.5560	1.5554	1.2485	1.2456	1.0976	1.0975
0.76923	1.7299	1.7289	1.3011	1.3009	1.1160	1.1159
0.95309	1.8273	1.8259	1.3291	1.3288	1.1249	1.1248

Table I-5 Comparison of the  $\psi_0^O(\mu)$  and  $\phi_0^O(\mu)$  Functions for isotropic Scattering and  $\omega_0=0.90$  with Chandrasekhar's X and Y Functions.

	<u> </u>			
μ 0.04691 0.23077 0.50000 0.76923 0.95309	x 1.0560 1.0778 1.0814 1.0826 1.0831	T <sub>1</sub> = 0.05 1.0546 1.0726 1.0783 1.0800 1.0806	Y 0.3914 0.8805 0.9852 1.0190 1.0314	\$\psi_0\$ 0.3838 0.9049 0.9824 1.0138 1.0252
		$T_1 = 0.25$		
μ 0.04691 0.23077 0.50000 0.76923 0.95309	x 1.0810 1.1967 1.2440 1.2610 1.2668	11 - 0.2)  **  1.0797  1.1993  1.2413  1.2611  1.2694	Y 0.0441 0.5017 0.8285 0.9672 1.0228	\$0 0.0467 0.4990 0.8317 0.9501 1.0001
		$T_1 = 0.50$		
μ 0.04691 0.23077 0.50000 0.76923 0.95309	X 1.0861 1.2553 0.3528 1.3948 1.4122	1.0842 1.2530 1.3465 1.3896 1.4089	Y 0.0269 0.2769 0.6506 0.8636 0.9583	ø° 0.0275 0.2636 0.6497 0.8462 0.9249
μ 0.04691 0.23077 0.50000 0.76923 0.95309	X 1.0902 1.2933 1.4574 1.5472	T <sub>1</sub> = 1.00 <b>\psi</b> : 1.0888 1.2915 1.4541 1.5427 1.5852	Y 0.0171 0.1888 0.4120 0.6646 0.8000	9° 0.0165 0.1231 0.4012 0.6476 0.7553

Table I-6 Comparison of the  $\psi$  and  $\emptyset$  Functions of this Report with Similar Functions of Reference (4) Calculated Using a 20-point Simpson's Rule Technique for  $a_1=1.73458$ ,  $a_2=2.24153$ ,  $\pmb{\omega_\bullet}=0.90$  and  $\tau_1=\infty$ .

μ	$\Psi_{O}^{O}(1)$	u)	Ψο	
	Ref.	This Rept.	Ref.	This Rept.
0.04691	1.1186	1.1185	0.0187	0.0185
0.23077	1.3538	1.3550	0.1182	0.1179
0.50000	1.4865	1.4868	0.3121	0.3121
0.76923	1.4693	1.4699	0.5413	0.5412
0.95309	1.3946	1.3948	0.7094	0.7094
	ψ <sub>0</sub>		₩ <del>]</del>	
μ	Ref.	This Rept.	Ref.	This Rept.
0.04691	0.5368	0.5370	1.0515	1.0516
0.23077	0.5110	0.5124	1.0888	1.0896
0.50000	0.2078	0.2079	0.9564	9.0561
0.76923	0.3782	0.3771	0.6578	0.6583
0.95309	0.9347	0.9348	0.2822	0.2905
	$\psi_{\overline{1}}^2$		ψ2	
μ	Ref.	This Rept.	Ref.	This Rept.
0.04691	0.1084	0.1080	3.1399	3.1376
0.23077	0.5908	0.5911	3.1955	3.1955
0.50000	1.2591	1.2596	2.6409	2.6406
0.76923	1.5187	1.5179	1.4697	1.4695
0.95309	0.8840	0.9173	0.3330	0.3331

Table I-7 Comparison of Integrated Reflectance for Rayleigh Scattering,  $\tau_1$  =  $\infty$  and  $\pmb{\omega}_\bullet$  = 0.90 with Values Calculated using Method Described in Appendix A-2.

$\mu_{O}$	Appendix	This Report
0.04691	0.660	0.656
0.23077	0.581	0.580
0.50000	0.507	0.507
0.76923	0.453	0.453
0.95309	0.424	0.424

It must be noted that in all the preceeding comparisons (except for the case of  $\mu=0.50$ ) the results from other investigators were linearly interpolated to get values for comparis**on** with the results at the Gaussian quadrature points of this report. From the comparisons, however, it is estimated that the functions are generally accurate to within  $\pm 0.02$  which in most cases in an error of less than two to three percent. The intensities of reflected and transmitted radiation as well as the integrated reflectance and transmittance are of the same general accuracy as the functions. This is evident from equations I-15, I-19, I-29 and I-32.

The use of Gaussian quadrature seems preferable to Simpson's Rule or to the trapezoidal rule in that five point quadrature afford the same accuracy as the use of ten equally spaced ordinates. In addition, a better approximation may be had by adding only one additional point where with Simpson's rule, in order to obtain a better approximation, it is necessary to double the number of points. The disadvantage of the use of quadrature is that results are obtained at odd values of  $\mu$ . On the other hand, in using the results they are to be integrated again, this may be an advantage, since quadrature can be used to perform the integration.

The time in seconds required for one iteration on the IBM  $70^4$  is Time = 0.025 (no. Points) $^2$ x(no. equations) $^2$ x(no. iterations) I-56

For example, using five-point quadrature; for m=0, N=2; performing four iterations

Time = 
$$0.025 (5)^2 (3)^2 (4)$$
= 22.5 seconds

Considering the cost of 704 time at \$5.00 per minute

Cost =  $$2.08 \times 10^{-3}$  (no. points) $^2$ x(no. equations) $^2$ x(no. iterations) I-58

To scale the cost to some other machine, simply multiply the factor  $$2.08 \times 10^{-3}$  by (cost per operation on new machine)/(cost per operation on the IBM 704).

The 704 performs 40,000 operations on the average per second so its cost per operation is

 $$5/(60x40,000) = $2.08 \times 10^{-6} \text{ per operation}$ It is estimated that the  $$2.08 \times 10^{-3} \text{ factor in I-58 could be reduced}$ by a factor of five through the use of efficient hand-programming.

#### PART IT EVALUATION OF APPROXIMATE MODELS

Exact solutions in closed form are not attainable by analytical methods for problems of multiple scattering. Exact solutions are attainable by numerical methods, as in Part I, but the required calculations are lengthy and expensive, and the results are for specific cases. Simple, approximate solutions therefore have considerable value if reliable. In this section the reflectance and transmission obtained from various approximate models are compared with the "exact" values obtained in Part I.

## A. Approximate Models

1. Two-flux Model(11,12)

The two-flux formulation of the slab problem with an obliquely incident plane-parallel source is

$$\mu_o \frac{dI_i}{dz} = -(1 - \omega_o F)I_i + \omega_o BI_2$$
 II-la

$$\mu_{\circ} \frac{dI_2}{dz} = (1 - \omega_{\circ} F) I_2 - \omega_{\circ} B I_1$$
II-lb

Where  $I_1$  is the integrated radiant flux in the forward direction and  $I_2$  the integrated radiant flux in the backward direction, and F and B are the forward and the backward components respectively of the phase function for single scattering.

Equations II-la and II-lb can be solved with the boundary conditions  $I_1 = 1$  at Z = 0, and  $I_2 = 0$  at  $Z = \tau_1$  to give

$$R = \frac{(I_2)_{z=0} \mu_0}{(I_1)_{z=0} \mu_0} = \frac{(I - \omega_0 F - p)(I - e^{-p \tau_1/\mu_0})}{\omega_0 \beta (I - \frac{I - \omega_0 F - p}{I - \omega_0 F + p} e^{-2p \tau_1/\mu_0})}$$
II-2

and

$$T = \frac{(I_2)_{z=z_1} \mu_o}{(I_1)_{z=0} \mu_o} = \frac{2 p e^{-p \tau_1 / \mu_o}}{1 - \omega_o F + p - (1 - \omega_o F - p) e^{-2 p \tau_1 / \mu_o}}$$
II-3

where

$$p = \sqrt{(1 - \omega_0 F)^2 - (\omega_0 B)^2}$$

# 2. Six-flux Model (11,12)

The six-flux formulation of the slab problem with an obliquely incident parallel-plane source is

$$\frac{d\mathbf{I}_1}{d\mathbf{z}} = -C_1 \sec \theta_0 \mathbf{I}_1 + C_2 \sec \theta_0 \mathbf{I}_2 + C_3 \sec \theta_0 (\mathbf{I}_3 + \mathbf{I}_4)$$
 II-5a

$$-\frac{dI_2}{dz} = -C_1 \sec \theta_0 I_2 + C_2 \sec \theta_0 I_1 + C_3 \sec \theta_0 (I_3 + I_4)$$
II-5b

$$\frac{dI_3}{dz} = -C_1 \csc \theta_0 I_3 + C_2 \csc \theta_0 I_4 + C_3 \csc \theta_0 (I_1 + I_2)$$
III-5c

$$-\frac{dI_4}{dZ} = -C_1 \csc\theta_1 + C_2 \csc\theta_2 I_3 + C_3 \csc\theta_2 (I_1 + I_2)$$
III-5d

$$I_5 = I_6 = \frac{\omega_0 \chi}{(-\omega_0 F_6 - \omega_0 B_6)} (I_1 + I_2 + I_3 + I_4)$$
II-6

Where  $I_1$ ,  $I_2$ ,  $I_3$ ,  $I_4$ ,  $I_5$  and  $I_6$  are the components of the radiant flux in the forward, backward and sidewise directions, and

$$C_1 = 1 - \omega_0 F_6 - \frac{2\omega_0 X^2}{1 - \omega_0 F_6 - \omega_0 B_6}$$
II-7a

$$C_2 = \omega_0 B_6 + \frac{2\omega_0 X^2}{1 - \omega_0 F_6 - \omega_0 B_6}$$
II-76

$$C_3 = \omega_0 X + \frac{2\omega_0 X^2}{1 - \omega_0 F_6 - \omega_0 B_6}$$
II-7c

where  $F_6$ ,  $B_6$  and X are the six-flux representations of forward, backward and sidewise scattering components.  $F_6 = (F)(P)$ ,  $B_6 = (B)(P)$   $X = (1 - F_6 - B_6)/4$ .

The solution for Equations II-5a to II-6 with the boundary conditions  $I_1$  = 1 and  $I_3$  = 0 at Z = 0, and  $I_2$  =  $I_4$  = 0 at Z =  $\tau_1$ , is reproduced in Appendix A-3.

# 3. Richard's Modified Diffusion Theory (13)

The transport equation is approximated as

$$\frac{d^2q}{dz^2} - K^2q = -(1 - K^2)\ell_d$$
 II-8

where **z** is normal distance in mean free paths for scattering, **K** =  $\sqrt{3(1-\omega_0)/(3-2\omega_0)}$ ,  $\rho_d$  is the density of unscattered photons,  $\rho$  is the photon density and q is defined by

$$q = \left(1 - \frac{2}{3}\omega_{\bullet}\right) \ell - \ell_{\delta}$$

The source strength is  $\rho_d = e^{-z/\mu_o}$ .

Equation II-8 can be solved for the conditions  $\omega_{\bullet}=0$  in vacuum,  $\rho$  finite, and q and  $\frac{dq}{dz}$  both continuous at  $\tau_{1}=0$  and  $\tau_{1}$ .

Giving
$$R = \frac{1}{\mu_0} \frac{dq}{dz}\Big|_{z=0} = \frac{-2K(1-\mu_0)e^{-z_1/\mu_0}}{(1+K)^2e^{K\tau_1} - (1-K)^2e^{-K\tau_1}} \frac{(1-K^2)\mu_0}{1-K^2\mu_0^2} = \frac{-1-K^2\mu_0^2}{1-K^2\mu_0^2}$$
II-10

and

$$T = -\frac{1}{\mu_0} \frac{dq}{dz}\Big|_{z=z_1} + e^{-z_1/\mu_0}$$

$$= \frac{2K(1+\mu_0) - (1+K)(1+K\mu_0)e^{-Kz_1-z_1/\mu_0}}{(1+K)^2e^{-kz_1} - (1-K)^2e^{-kz_1}} \frac{1-K^2}{1-K^2\mu_0^2} + e^{-z_1/\mu_0} II-11$$

The above two equations are for  $\omega_{\bullet} < 1$ . The solutions for the case  $\omega_{\bullet} = 1$  can be found easily and are omitted here.

#### 4. Numerical Results

Numerical values were computed from the above approximate solutions for comparison with the "exact" results obtained in Part I. The computed values are given in Tables B-1 to B-5 in Appendix B and are illustrated in Figures II-1 through II-5.

Table B-l gives reflectances for isotropic scattering as obtained from Richard's model, the two-flux model and the six-flux model. The exact values are included for comparison. The six-flux results for Rayleigh scattering are also included. The exact values for Rayleigh scattering did not vary significantly from the values for isotropic scattering and hence are not reproduced here. The two-flux model

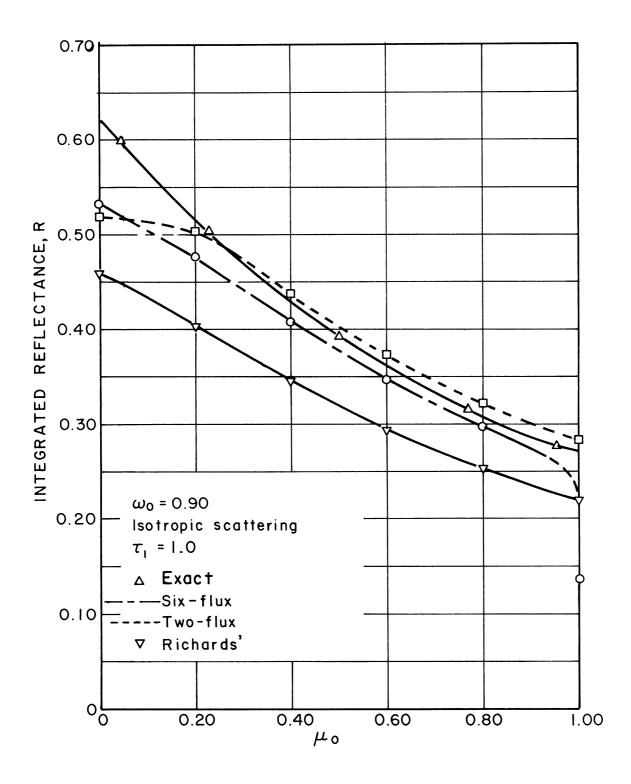


Figure II-l Integrated reflectance for various solutions as functions of angle of incidence for isotropic scattering, R vs.  $\mu_0$ ,  $\omega_0$  = 0.90,  $\tau_1$  = 1.0

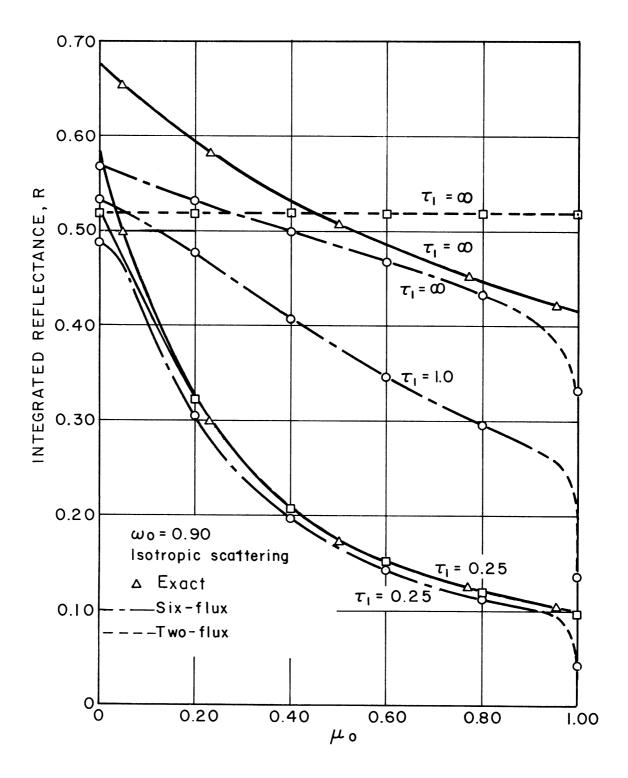


Figure II-2 Integrated reflectance for various solutions as function of angle of incidence for isotropic scattering, R vs.  $\mu_{0},$  parameter of  $\tau_{1}$ ,  $\omega_{\text{o}}\text{=}$  0.90

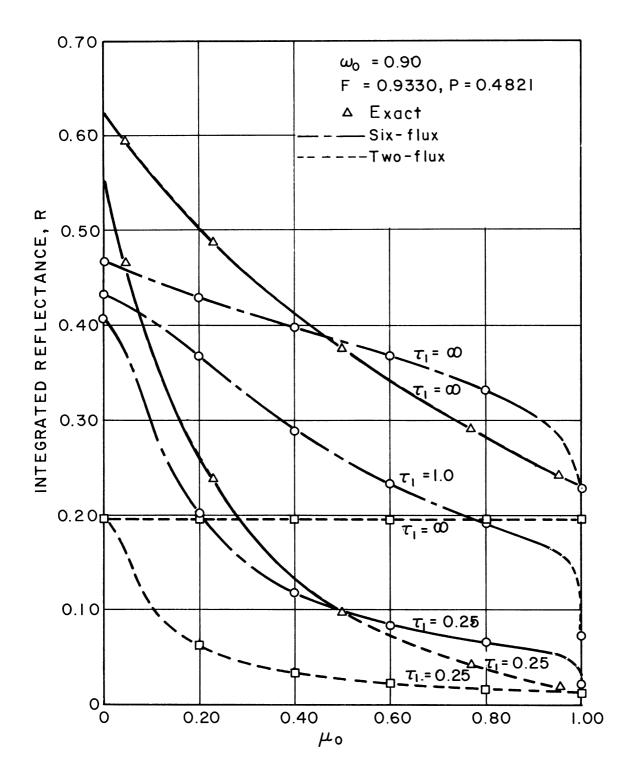


Figure II-3 Integrated reflectance for various solutions as function of angle of incidence for anisotropic scattering, R. vs.  $\mu_0$ , parameter of  $\tau_1$ ,  $\omega_0$ = 0.90, F = 0.9330, P = 0.4821

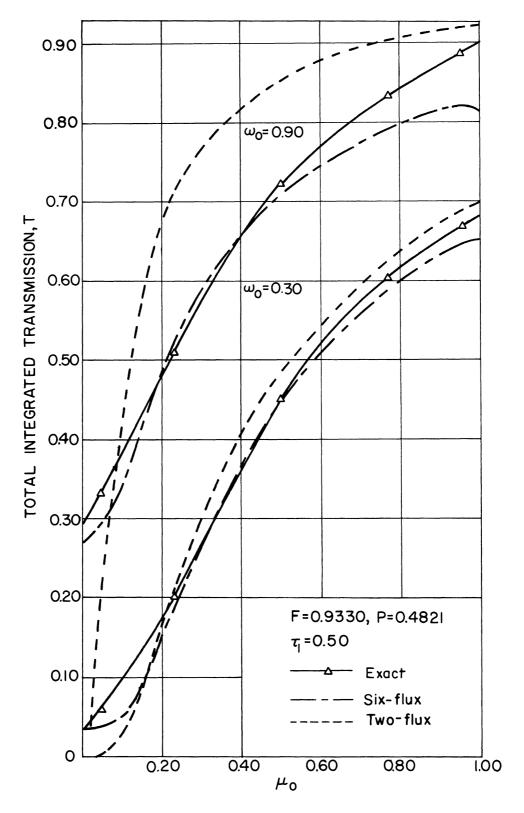


Figure II-4 Total integrated transmission, various solutions as function of angle of incidence for anisotropic scattering, R vs.  $\mu_0$ , parameter of  $\omega_0$ ,  $\tau_1$  = 0.50, F = 0.9330, P = 0.4821

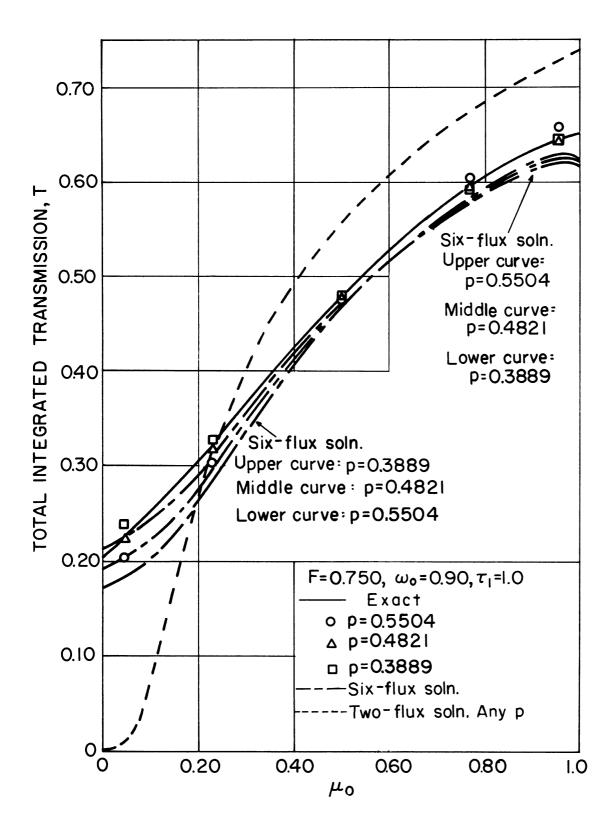


Figure II-5 Total integrated transmission, various solutions as function of angle of incidence for anisotropic scattering, R vs.  $\mu_0$ , parameter of P,  $\omega_0$ = 0.90,  $\tau_1$  = 1.0, F = 0.750

obviously gives identical values for Rayleigh and isotropic scattering. Representative values are plotted in Figures II-1 and II-2. From the tabulated and plotted values it can be observed that the six-flux solution is generally better than the two-flux solution which is in turn generally better than the Richard's solution. The simple diffusion solution which is not given is poorer than Richard's solution. Exceptions are for normal incidence where the six-flux solution falls off and for infinite thickness for which the two-flux reflectance is invariant with angle of incidence and inferior to Richard's solution. All of the approximations improve as the albedo,  $\omega_{\bullet}$ , increases and as the optical thickness,  $\tau_{1}$ , decreases. This behavior can be attributed to the decreased contribution of multiple scattering to the reflectance.

Table B-2 gives transmissions for the same conditions and models as in Table B-1, except for the omission of the infinitely thick medium. Again the values for Rayleigh scattering are included for the six-flux model, are identical for the two-flux model and are essentially equal for the exact method. These values are not illustrated graphically but study of the tabulated values indicates clearly that 1) the six-flux values are better than Richard's values which are this time better than the two-flux values. The two flux values are two low for small  $\mu_0$  but improve for  $\mu_0 > 0.5$ . Richard's values and the six-flux values are remarkably equal for normal incidence. The greater the albedo,  $\omega_0$ , the

better are the approximate values. No clear cut trend in accuracy is apparent with thickness. The predictions of transmission by Richard's and the six-flux model are better than the predictions of reflectance but the opposite is the case with the two-flux model.

Table B-3 and B-4 give reflectances and transmissions respectively for anisotropic scattering with F = 0.933 and P = 0.4821 as computed from the six-flux and two-flux models. The exact values are also reproduced for comparison. Representative values are plotted in Figures II-3 and II-4. It can be observed that the six-flux model is more accurate than the two-flux model for both reflection and transmission. The six-flux model is most reliable near  $\mu_0$  = 0.5 and for large  $\omega_0$  and appears to give consistently low reflectances for normal incidence.

Table B-5 gives reflectances and transmissions for anisotropic scattering and with F = 0.750 and P = 0.5821, 0.4821 and 0.3889 **as** computed exactly and from the six-flux and two-flux models. The computed transmissions are illustrated in Figure II-5. The six-flux model yields a dependence or P which is in accordance with the exact values. The two-flux model of course indicates no dependence on P.

#### B. Interpolation Factors

Examination of the approximate solutions and the trend of the exact values provided the basis for the development of several simple relationships for approximate interpolation and extrapolation of the computed values.

The integrated reflectance and the total transmission are observed to be strong functions of  $\mu_0$ ,  $\omega_{\bullet}$ , F and  $\tau_1$  but only weakly dependent on P. Interpolation equations for  $\mu_0$  and  $\tau_1$  were developed for some conditions as follows.

For isotropic scattering by a half-infinite medium simple diffusion theory suggests the interpolation formula

$$R = \frac{R_o}{1 + \left(\frac{R_o}{R} - 1\right)\mu_o}$$

where R and R are the integrated reflectance for grazing ( $\mu_0$  = 0) and normal ( $\mu_0$  = 1) incidence, respectively. This equation is found to be satisfactory for interpolation of R with respect to  $\mu_0$  even though the absolute values of R indicated by diffusion theory are seriously in error. The success of this interpolation is indicated by the top curve (for  $\tau_1$  =  $\infty$ ) in Figures II-6 and II-7, and in Table II-1.

Richard's model does not give accurate values of R for isotropic scattering by finite slabs but the indicated dependence of R on  $\tau_1$ , is

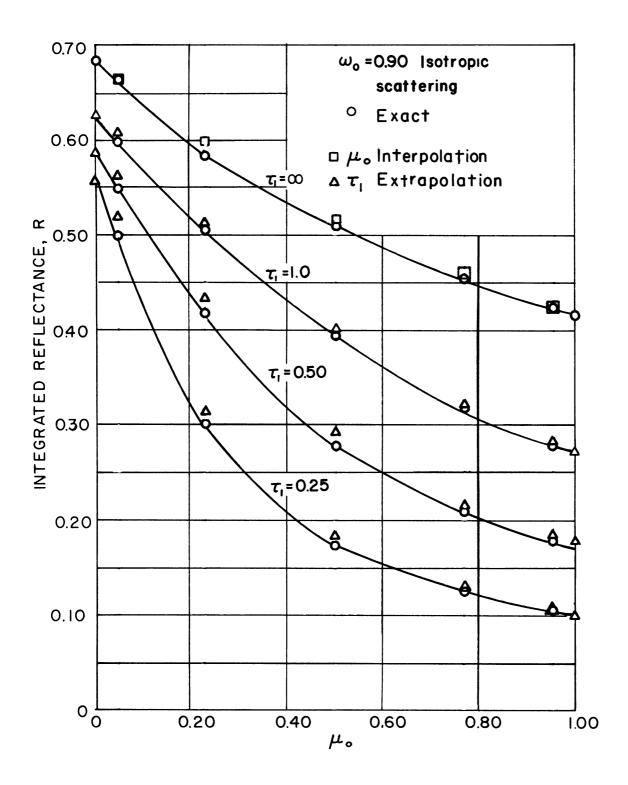


Figure II-6 Integrated reflectance, exact, interpolated and extrapolated as function of angle of incidence for isotropic scattering, R vs.  $\mu_{o}$ , parameter of  $\tau_{1}$ ,  $\omega_{o}$ = 0.90

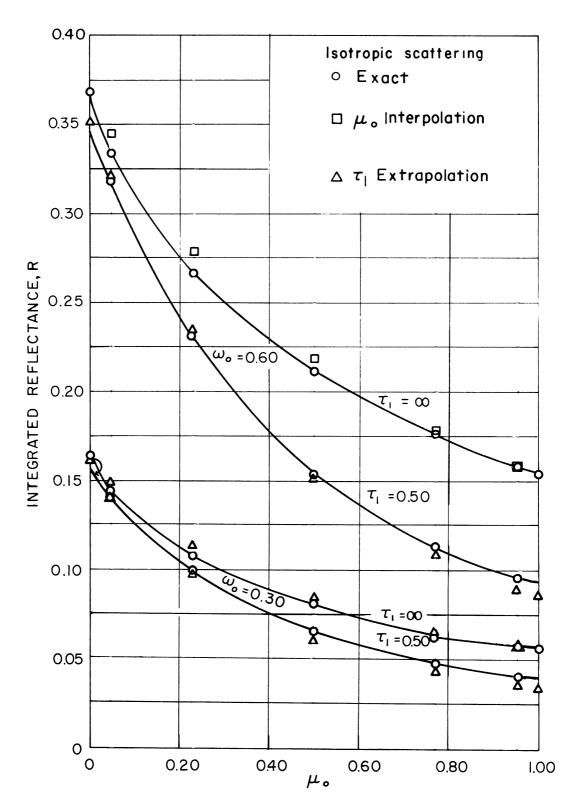


Figure II-7 Integrated reflectance, exact, interpolated and extrapolated as function of angle of incidence for isotropic scattering, R vs.  $\mu_{o}$ , parameters of  $\omega_{o}$  and  $\tau_{1}$ 

Table II-l Interpolation and Extrapolation of R for Isotropic Scattering with Respect to  $\mu_{\text{O}}$  and  $\tau_{\text{l}}$ 

 $\omega_{o} = 0.90$ :  $\tau_1 = \infty$ **7**; = 0.5  $\tau_i = 0.25$  $\tau_i = 1.0$ Exact Interp. Exact Extrap. Exact Extrap. Exact Extrap.  $\mu_{O}$ 0.00000 0.684 0.684 0.624 0.625 0.584 0.586 0.571 0.556 0.04691 0.664 0.664 0.598 0.606 0.548 0.561 0.498 0.516 0.23077 0.582 0.595 0.504 0.511 0.417 0.433 0.300 0.316 0.50000 0.508 0.517 0.394 0.401 0.279 0.293 0.173 0.185 0.76923 0.452 0.456 0.316 0.322 0.209 0.215 0.125 0.129 0.95309 0.423 0.423 0.279 0.284 0.179 0.184 0.105 0.105 1.00000 0.415 0.415 0.271 0.274 0.170 0.176 0.100 0.100  $\omega_o = 0.30$ :  $\omega_o = 0.60$ :  $\tau_i = \infty$  $\tau_{i} = 0.50$ 7; = ∞  $\tau_{\rm i} = 0.50$ Exact Interp. Exact Extrap. Exact Extrap. Exact Extrap.  $\mu_{o}$ 0.00000 0.368 0.368 0.345 0.352 0.163 0.163 0.156 0.161 0.04691 0.334 0.346 0.318 0.321 0.144 0.150 0.142 0.141 0.23077 0.267 0.279 0.232 0.236 0.108 0.114 0.100 0.098 0.212 0.218 0.154 0.152 0.082 0.085 0.50000 0.066 0.061 0.76923 0.177 0.179 0.114 0.109 0.066 0.067 0.048 0.045 0.95309 0.159 0.159 0.097 0.091 0.059 0.059 0.041 0.036

0.087

0.057

0.057

0.040

0.035

0.093

1.00000

0.155

0.155

<sup>\*</sup> Extrapolated by Lagrange's formula for  $\mu_{0}$  = 0 and  $\mu_{0}$  = 1.0

excellent suggesting the equation

$$R_{\mu_{o}, \tau_{i}} = R_{\mu_{o}, \infty} \left( \frac{R_{\mu_{o}, \tau_{i}}}{R_{\mu_{o}, \infty}} \right)_{R:chards}$$
II-13

for the extrapolation of exact values of R for  $\tau_1 = \infty$  and any  $\mu_0$  and  $\omega_0$  to finite values of  $\tau_1$ . The success of this formula is indicated in Figures II-6 and II-7 and in Table II-1 for the finite values of  $\tau_1$ . The ratios of R for Richard's solution were taken from Tables B-1 and B-2. The success of the interpolation equations indicated that exact calculations for isotropic scattering could be limited to normal and grazing incidence on a half-infinite medium.

For anisotropic scattering the expression

$$R = xR_1 + (1-x)R_0$$

where x is a function of  $\omega_{\bullet}$ ,  $\tau_{1}$  and  $\mu_{0}$  but not of F, can be used to interpolate for  $\mu_{0}$ . Values of  $\pi$  can therefore be determined from computed values of R for isotropic scattering. A plot of  $\pi$  as a function of  $\mu_{0}$  for several values of  $\tau_{1}$  and  $\omega_{\bullet}$  are shown in Figures II-8 and II-9. These figures were prepared from the isotropic values. (The six-flux solution erroneously indicates independence of  $\pi$  from  $\tau_{1}$  and  $\omega_{\bullet}$  as well as from F and P.) Values of R for intermediate values of  $\mu_{0}$  were computed from the exact values of  $\pi_{0}$  and  $\pi_{1}$  and the values of  $\pi$  in Figures II-8 and II-9 using Equation II-14 and are given in Table II-2 and in Figures II-10 and II-11 (triangles). The agreement with the exact values is seen to be reasonably good.

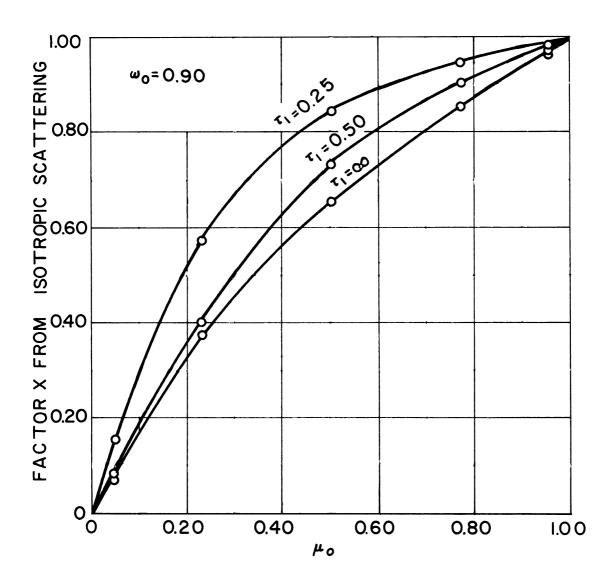


Figure II-8 Factor x in equation II-14 calculated from isotropic R and R , as function of angle of incidence, x vs.  $\mu_{0}$  , parameter of  $\omega_{o}$ 

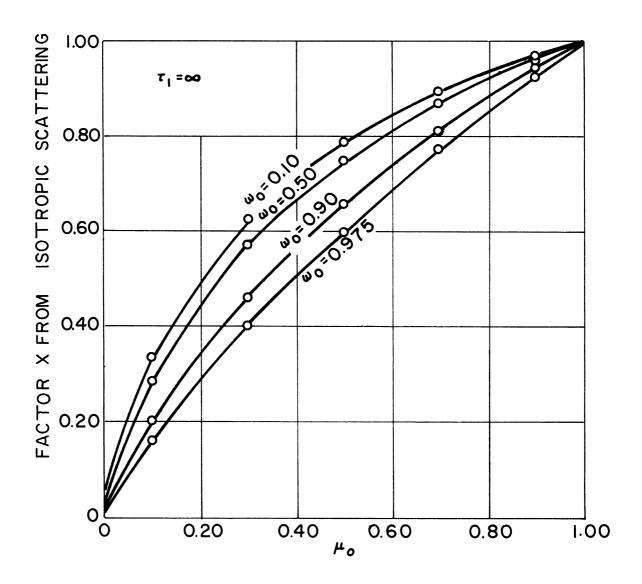


Figure II-9 Factor x in equation II-14 calculated from isotropic R and R as function of angle of incidence, x vs.  $\mu_0$  , parameters of  $\tau_1$ 

Table II-2 Interpolation of R with Respect to  $\boldsymbol{\mu}_O$  for Anisotropic Scattering

F = 0.9330, P = 0.4821,  $\omega_0 = 0.90$ :

	<i>℃,≖</i> ∞		$\tau_i = 1.0$		T; = 0.5		$T_i = 0.25$	
$\mu_{o}$	Exact*	Interp.	Exact*	Interp.	Exact*	Interp.	Exact*	Interp.
0.00000	0.628	0.628	0.567	0.567	0.547	0.547	0.552	0.552
0.04691	0.595	0.599	0.530	0.530	0.500	0.502	0.466	0.468
0.23077	0.489	0.478	0.396	0.397	0.333	0.338	0.238	0.241
0.50000	0.377	0.370	0.239	0.241	0.163	0.166	0.097	0.096
0.76923	0.292	0.299	0.131	0.131	0.078	0.079	0.043	0.041
0.95309	0.244	0.245	0.079	0.078	0.041	0.041	0.018	0.019
1.00000	0.233	0.233	0.067	0.067	0.030	0.030	0.013	0.013

F = 0.9330, P = 0.4821,  $\omega_0 = 0.60$ :

	T <sub>1</sub> = 00		$\tau_i = 0.50$		$\tau_i = \infty$		$\tau_1 = 0.50$	
$\mu_{O}$	Exact*	Interp.	Exact*	Interp.	Exact*	Interp.	Exact*	Interp.
0.00000	0.336	0.336	0.330	0.330	0.157	0.157	0.151	0.151
0.04691	0.303	0.289	0.293	0.296	0.144	0.139	0.131	0.134
0.23077	0.206	0.197	0.180	0.189	0.108	0.105	0.075	0.082
0.50000	0.124	0.121	0.085	0.092	0.082	0.081	0.034	0.040
0.76923	0.073	0.073	0.040	0.042	0.066	0.066	0.015	0.018
0.95309	0.048	0.049	0.021	0.021	0.059	0.059	0.008	0.009
1.00000	0.043	0.043	0.016	0.016	0.058	0.058	0.007	0.007

 $\omega_o = 0.30$ :

F = 0.7500, P = 0.4821,  $\omega_{o} = 0.90$ ;  $z_{i} = 1.0$ 

Exact\* Interp.  $\mu_{O}$ 0.00000 0.598 0.598 0.564 0.566 0.04691 0.23077 0.445 0.451 0.317 0.50000 0.310 0.76923 0.218 0.221 0.95309 0.175 0.176 1.00000 0.166 0.166

<sup>\*</sup> Extrapolated by Lagrange's formula for  $\mu_{O}$  = 0 and  $\mu_{O}$  = 1.0

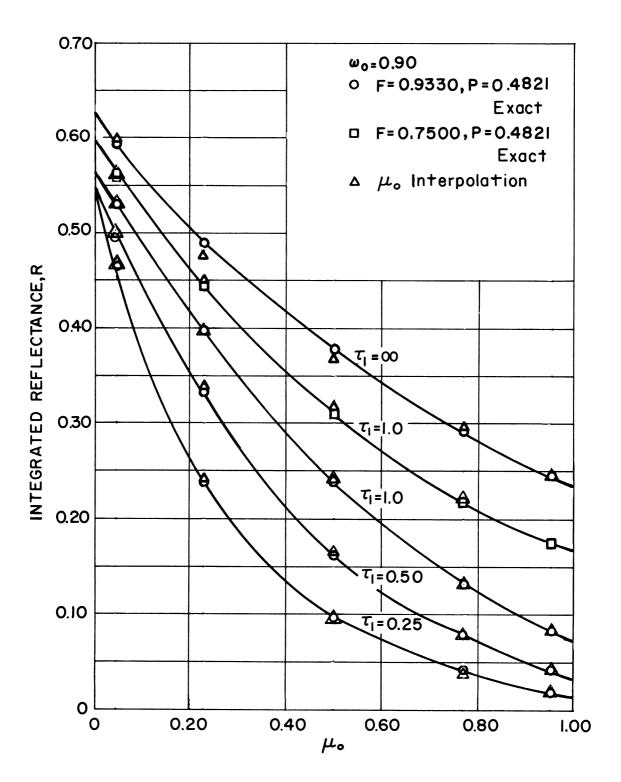


Figure II-10 Integrated reflectance, exact and interpolated as function of angle of incidence for anisotropic scattering, R vs.  $\mu_{0}$ , parameter of  $\tau_{1}$ ,  $\omega_{0}$ = 0.90, F = 0.9330, P = 0.4821

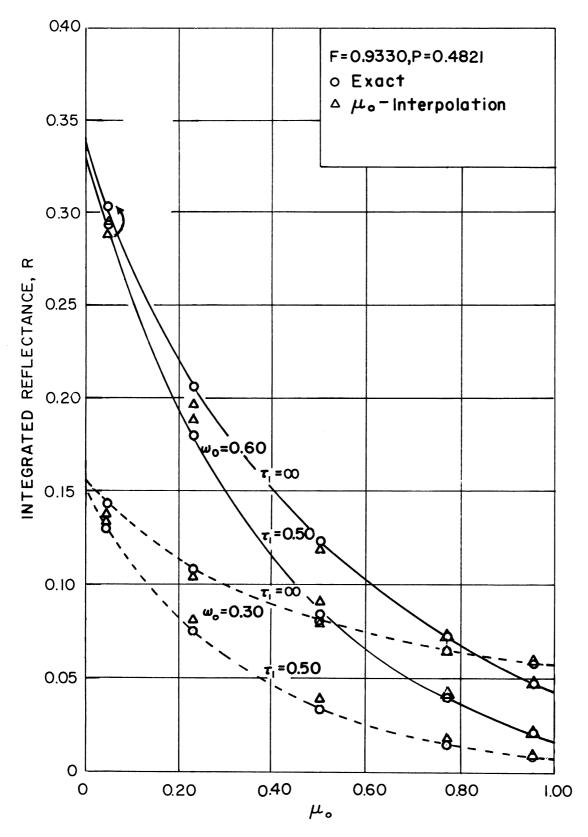


Figure II-ll Integrated reflectance, exact and interpolated as function of angle of incidence for anisotropic scattering, R vs.  $\mu_{0},$  parameters of  $\omega_{o}\,\&\,\,\tau_{1}$ 

The same type of interpolation, i.e.,

 $T = yT_0 + (1-y)T_1$  II-15

was also successful for the total transmission as indicated in Table II-3.

Table II-3 Interpolation of T with Respect to  $\boldsymbol{\mu}_O$  for Anisotropic Scattering

F = 0.9330, P = 0.4821,  $\omega_0 = 0.90$ :

් = 1.0		<b>℃</b> ; =	0.50	$7_i = 0.25$		
μ	$\mathtt{Exact}^{igstar}$	Interp.	Exact*	Interp.	$\mathtt{Exact}^ullet$	Interp.
0.00800	0.761	0.761	0.904	0.904	0.991	0.991
0.04691	0.750	0.754	0.885	0.886	0.979	0.983
0.23077	0.692	0.698	0.833	0.841	0.933	0.951
0.50000	0.553	0.551	0.723	0.733	0.853	0.881
0.76923	0.372	0.365	0.507	0.505	0.669	0.689
0.95309	0.259	0.256	0.331	0.328	0.396	0.402
1.00000	0.236	0.236	0.291	0.291	0.297	0.297

	F = 0	F = 0.7500, $P = 0.4821\omega_0 = 0.90:$					
	<b>7</b> 1 = 1	0.50	₹, =	$r_i = 0.50$		$z_i = 1.0$	
$\mu_{O}$	Exact*	Interp.	Exact*	Interp.	Exact*	Interp.	
0.00000	0.783	0.783	0.692	0.692	0.652	0.652	
0.04691	0.761	0.761	0.668	0.665	0.645	0.646	
0.23077	0.701	0.701	0.601	0.600	0.598	0.598	
0.50000	0.565	0.561	0.452	0.433	0.476	0.473	
0.76923	0.319	0.309	0.198	0.189	0.318	0.313	
0.95309	0.165	0.161	0.063	0.059	0.222	0.221	
1.00000	0.143	0.143	0.052	0.052	0.203	0.203	

<sup>\*</sup> Extrapolated by Lagrange's formula for  $\mu_{0}$  = 0 and  $\mu_{0}$  = 1.0



### A. Introduction

It is well known in classical physics that multiple scattering processes can be described approximately by means of the classical diffusion equation. The major difficulty in the application of classical diffusion equation to the problems of multiple scattering however, is in the specification of the correct boundary condition. Various authors have investigated the use of the simple diffusion equation in multiple scattering problems, and formulated modified (13,14,15) . In this investigation versions of the diffusion equation it is shown that the various modified diffusion equations can be derived in a unified manner from the integral equation of transport, and therefore can be not only refined indefinitely, but also, in principle, can have appropriate boundary conditions derived corresponding to each step of refinement. On this basis a new modified diffusion equation is proposed and solutions compared numerically with known exact solutions. Better agreement is found than with previous models. The boundary conditions are obtained from purely mathematical reasoning and a good physical interpretation is yet to be found. To simplify the mathematical formulation, isotropic scattering has been considered as an example, but the extension to anisotropic scattering is trivial, and will be carried out in the continuation of this work.

## B. Approximate Differential Equations

For isotropic scattering, the integral equation of transport for the photon density  $\rho(\vec{r})$  is

$$P(\vec{r}) = S(\vec{r}) + \omega_o \iiint \frac{e^{-|\vec{r}-\vec{r}'|}}{4\pi |\vec{r}-\vec{r}'|^2} P(\vec{r}') dv'$$
Volume of dispersion

III-1

where

 $S(\vec{r})$  is the source term,

 $\omega_{o}$  is the albedo of single scattering, and

 $\vec{\gamma}$  represents the space coordinates measured in mean free paths.

Equation (1) is a linear integral equation of the Fredholm type in three dimensions.

$$\rho(\vec{r}') = S(\vec{r}) + \omega_o \iiint K(\vec{r}, \vec{r}') \rho(\vec{r}') d\sigma'$$
III-2

For some special forms of the kernal  $K(\vec{r},\vec{r})$  this equation can be reduced to a simple differential equations. In this section, it is shown that the various forms of diffusion equation can be obtained by such reduction after approximating the kernal

$$K(\vec{r}, \vec{r}') = \frac{e^{-|\vec{r}-\vec{r}'|}}{4\pi |\vec{r}-\vec{r}'|^2}$$

by comparatively simple forms.

Consider the class of  $K(\vec{r},\vec{r})$  whose Fourier transform can be represented by ratio of two polynomials\*:

$$\frac{N(k^2)}{D(k^2)} = \frac{a_0 + a_1 k^2 + a_2 k^4 + \dots + a_n k^{2n}}{b_0 + b_1 k^2 + b_2 k^4 + \dots + b_m k^{2m}}$$

\* The same derivation applied to a two dimensional case was given in Ref. (16)

so that,

$$K(\vec{r},\vec{r}') = \frac{1}{(2\pi)^3} \iiint_{D(k^2)} \frac{N(k^2)}{D(k^2)} e^{i\vec{k}\cdot(\vec{r}-\vec{r}')} dk_x dk_y dk_y$$
 III-3

Then, since

$$\vec{\nabla}e^{i\vec{k}\cdot(\vec{r}-\vec{r}')} = -\vec{k}^2e^{i\vec{k}\cdot(\vec{r}-\vec{r}')}$$
III-4

Equation (3) may be reduced to:

$$K(\vec{r}-\vec{r}i) = N(-\vec{r}^2) \frac{1}{(2\pi)^3} \iiint_{-\infty} \frac{e^{i\vec{k}\cdot(\vec{r}-\vec{r}')}}{D(k^2)} dk_x dk_y dk_y$$
 III-5

Substituting Equation (5) into (2) yields

Substituting Equation (5) into (2) yields 
$$\rho(\vec{r}) = S(\vec{r}) + w_0 N(-\nabla^2) \iiint \rho(\vec{r}') dv' \frac{1}{(2\pi)^{3/2}} \iiint \frac{e^{i\vec{k}\cdot(\vec{r}-\vec{r}')}}{D(\vec{k}')} dk_x dk_y dk_z \text{ III-6}$$
Now, operating on both sides of (6) by  $D(-\nabla^2)$ ,

 $D(-\nabla^2)\rho(\vec{r}) = D(-\nabla^2)S(\vec{r}) + \omega_0 N(-\nabla^2) \int \int \rho(\vec{r}') \frac{dv'}{(2\pi)^{3/2}} \int \int e^{i\vec{k}\cdot(\vec{r}-\vec{r}')} dk_x dk_y dk_y = 111-7$ 

It is recognized that

$$\overline{E_{\pi}})^{2} \int \int \int e^{\lambda \vec{k} \cdot (\vec{r} - \vec{r}')} dk_x dk_y dk_z = \delta(\vec{r}, \vec{r}')$$

the three dimensional & -function. Since Equation (7) may be

reduced to:

$$\left[D(-\nabla^2) - \omega_0 N(-\nabla^2)\right] \rho(\vec{r}) = D(-\nabla^2) S(\vec{r}) \qquad , \qquad \qquad \text{III-8}$$

Equation (8) cannot be used directly for Equation (1), because the kernal in Equation (1) cannot be exactly expressed in the form given by Equation (3). In fact, it is easy to see that

$$\frac{e^{-|\vec{r}-\vec{r}'|}}{4\pi |\vec{r}-\vec{r}'|^2} = \frac{1}{(2\pi)^{\frac{1}{2}}} \iiint\limits_{-\infty} e^{i\vec{k}\cdot(\vec{r}-\vec{r}')} \frac{+a\vec{n}'k}{k} dk_x dk_y dk_z$$

Thus a whole class of approximate solutions of Equation (1) can be obtained from approximations of the function

$$F(k) = \frac{\tan^{-1}k}{k} = 1 - \frac{1}{2}k^{2} + \frac{1}{2}k^{4} + \dots$$
 III-9

in terms of ratio of two polynomials.

equation (14), the modified diffusion equation proposed by Richards (13), and of the modified diffusion equation proposed by Grosjean (15), were obtained from various physical and/or mathematical arguments, they can also be obtained in a unified manner by different approximations of Equation (9). These are respectively

(i) 
$$F(k) \stackrel{\sim}{=} \frac{1}{1+\frac{1}{3}k^2} = |-\frac{1}{3}k^2 + \frac{1}{9}k^4 + \dots$$
 III-10

so that Equation (8) takes the form

$$\nabla^2 \rho(\vec{r}) - 3(1-\omega_0) \rho(\vec{r}) = -(3-\nabla^2) S(\vec{r})$$
 III-11

(ii) 
$$F(k) \simeq \frac{1+\frac{2}{3}k^2}{1+k^2} = 1-\frac{1}{3}k^2+\frac{1}{3}k^4+\dots$$
 III-12

so that Equation (8) takes the form:

$$\nabla^2 \rho(\vec{r}) - \left(\frac{1-\omega_o}{1-\frac{2}{3}\omega_o}\right) \rho(\vec{r}) = \frac{-(1-\nabla^2)}{1-\frac{2}{3}\omega_o} S(\vec{r})$$
III-13

(iii) 
$$F(k) \cong \frac{1+\frac{1}{3}k^2}{1+\frac{2}{3}k^2} = [-\frac{1}{3}k^2 + \frac{2}{9}k^4 + \dots]$$
 III-14

so that Equation (8) takes the form:

$$\nabla^{2}\rho(\vec{r}) - \frac{3(-\omega_{0})}{2-\omega_{0}}\rho(\vec{r}) = -\frac{(3-2\nabla^{2})}{2-\omega_{0}}S(\vec{r})$$
 III-15

The left hand sides of Equations (11), (13) and (15) are the equivalent terms of the classical diffusion equation, the modified diffusion equation proposed by Richards and the modified diffusion equation proposed by Grosjean respectively.

It is mathematically difficult to estimate the errors involved in these approximations, perhaps the only valid criterion is to compare the results obtained in various simple problems with exact solutions. Intuitively, however, it seems to be reasonable to require any approximation of F(k) to be as close as possible to the exact function at least in the limiting case of k=0, and  $k\to\infty$ . Inspection of Equation (9), (10), (12) and (14) shows that as  $k\to0$  all three approximations are exact to the coefficient  $k^2$  of the expansion of F(k).

However, as  $k \rightarrow \infty$ ,

$$\lim_{k \to \infty} \frac{\tan^{1}k}{k} = 0$$

and only the simple diffusion equation has the correct limit.

Based upon the above observations it is clear that a class of differential equations based on diffusion type of approximation can be systematically obtained as follows:

(i) First approximation:

$$F(k) \cong \frac{1}{1+3k^2} = 1-\frac{1}{3}k^2 + \frac{1}{9}k^4 - \dots$$

For which, the approximated kernal is,

$$K(\vec{r},\vec{r}') = \frac{1}{(2\pi)^3} \iiint_{1+\frac{1}{3}k^2} e^{i\vec{k}\cdot(\vec{r}-\vec{r}')} dk_x dk_y dk_z = \frac{3}{4\pi |\vec{r}-\vec{r}'|} e^{-i\vec{3}|\vec{r}-\vec{r}'|} \text{ III-16}$$

and the resulting differential Equation is:

$$\nabla \dot{\rho}(\vec{r}) - 3(1-\omega_0)\rho(\vec{r}) = (\nabla^2 - 3)S(\vec{r})$$

(ii) Second approximation

$$F(k) \approx \frac{1 + ak^2}{1 + bk^2 + ck^4}$$

For best approximation at small values of k,

$$a = \frac{8}{21}$$

$$b = \frac{5}{7}$$

$$c = \frac{4}{10^5}$$

and

so that

$$F(k) = \frac{1+\frac{8}{2}-k^2}{1+\frac{5}{2}k^2+\frac{4}{100}k^4} = 1-\frac{1}{3}k^2+\frac{1}{5}k^4-\frac{1}{7}k^6+\frac{2}{21}k^8+\dots$$
 III-17

corresponding to this approximation.

$$K(\vec{r},\vec{r}') = \frac{1}{(2\pi)^3} \iiint_{1+bk'+ck'} \frac{1+ak^2}{1+bk'+ck'} e^{i\vec{k}\cdot(\vec{r}-\vec{r}')} dk_x dk_y dk_g$$

$$= \frac{1}{4\pi |\vec{r}-\vec{r}'|} \left[ n_i e^{-\beta_i |\vec{r}-\vec{r}'|} + n_2 e^{-\beta_2 |\vec{r}-\vec{r}'|} \right] \qquad \text{III-18}$$

where

$$\beta_1 = 1.2344$$
  
 $\beta_2 = 4.1505$ 

are the

positive roots of the equation

$$c\beta^4 - b\beta^2 + 1 = 0$$

while

$$n_1 = \frac{\beta_1^2 a - 1}{c(\beta_1^2 - \beta_2^2)} = 0.70/3$$

and

$$n_2 = \frac{\beta_2^2 a - 1}{c(\beta_2^2 - \beta_1^2)} = 9.2987$$

The resulting differential equation is:

$$\left[\nabla^{4} + \frac{a\omega_{\overline{b}}b}{c}\nabla^{2} + \frac{1-\omega_{\overline{b}}}{c}\right]\rho(\vec{r}) = \left[\nabla^{4} - \frac{b}{c}\nabla^{2} + \frac{1}{c}\right]S(\vec{r})$$
 III-19

In principle, the approximations can be extended indefinitely. But numerical calculations given in the subsequent sections indicate that with the correct form for the boundary conditions, the first approximation yields very good results for the photon density, and the second approximation is almost indistinguishable from the corresponding exact solution.

# C. Point Source in Infinite Medium

For the problem of multiple scattering involving sources in an infinite medium, reduction of integral Equation (1) into differential Equations (11) and (19) enables one to find the approximate solution by use of Green's functions. It is well known that the solution of the equation

$$\nabla^2 G_{i}(\vec{r}, \vec{r}') - \kappa^2 G_{i}(\vec{r}, \vec{r}') = -\delta(\vec{r}, \vec{r}')$$
III-20

subject to the condition that 
$$G(\vec{r},\vec{r}')$$
 is zero as  $r\to\infty$  is, 
$$G(\vec{r},\vec{r}') = \frac{1}{4\pi} \frac{e^{-K/\vec{r}-\vec{r}'/}}{|\vec{r}-\vec{r}'|}$$
 III-21

This Green's function can be used to solve the equation of the 1st approximation

$$\nabla^2 \rho(\vec{r}) - 3(1-\omega_0) \rho(\vec{r}) = -(3-\sqrt{2}) S(\vec{r})$$

$$K = \sqrt{3(1-\omega_0)}$$

multiply Equation (20) by  $-(3-7)^2S(\overrightarrow{F})$ \* and integrate. It is easy to see that the solution of (11) is

$$\rho(\vec{r}) = \iiint_{\text{all space}} G(\vec{r}, \vec{r}') \left[ 3 - \nabla^2 \right] S(\vec{r}') dr'$$

<sup>\*</sup> means the appropriate differentiation with respect to the primed coordinate variables

This yields\*

$$\rho(\vec{r}) = S(\vec{r}) + (3 - \kappa^2) \iiint G(\vec{r}, \vec{r}') S(\vec{r}') d\sigma'$$
III-22

or more explicitly:

$$\rho(\vec{r}) = S(\vec{r}) + \frac{3\omega_o}{4\pi} \iiint \frac{e^{-\sqrt{3(-\omega_o)}|\vec{r}-\vec{r}'|}}{|\vec{r}-\vec{r}'|} S(\vec{r}') dv'$$
III-23

Similarly, the solution of the equation

$$\left[\nabla' + \frac{a\omega_0 - b}{c}\nabla^2 + \frac{1 - \omega_0}{c}\right]G_2(\vec{r}, \vec{r}') = -\delta(\vec{r}, \vec{r}')$$
III-24

subject to the condition that  $G_{\mathbf{r}}(\vec{r},\vec{r}') \rightarrow 0$  as  $\mathbf{r} \rightarrow \infty$  is,

$$G_{2}(\vec{r},\vec{r}') = \frac{1}{4\pi |\vec{r}-\vec{r}'|} \left( \frac{1}{\alpha_{1}^{2} - \alpha_{1}^{2}} \right) \left[ e^{-\alpha_{1}|\vec{r}-\vec{r}'|} - e^{-\alpha_{2}|\vec{r}-\vec{r}'|} \right] \qquad \text{III-25}$$

where  $\alpha_1$ ,  $\alpha_2$  are the positive roots of the equation\*\*

$$\alpha'' + \left(\frac{a\omega_o - b}{c}\right)\alpha^2 + \frac{1-\omega_o}{c} = 0$$
 III-26

\* Equation (22) is obtained by using the second Green's identity, so that  $-\iiint G(\vec{r},\vec{r}') \, \nabla'^2 S(\vec{r}') \, dv' = -\iiint S(\vec{r}') \, \nabla'^2 G(\vec{r},\vec{r}') \, dv'$  But in view of Equation (19)

$$-\iiint S(\vec{r}') \nabla'^2 G(\vec{r}, \vec{r}') dv' = \iiint S(\vec{r}') \left[ S(\vec{r}, \vec{r}') - K^2 G(\vec{r}, \vec{r}') \right] dv'$$

$$= S(\vec{r}) - K^2 \iiint S(\vec{r}') G(\vec{r}, \vec{r}') dv'$$

\*\* For the present problem,  $\omega_0 \le 1$ ,  $\alpha_1$ ,  $\alpha_2$  exist, and are given in Table III-1.

Using the Green function  $G_{\mathbf{r}}(\vec{r},\vec{r}')$  it is possible to write down

formally the solution of the equation in an infinite medium.

$$\left[\nabla^{4} + \frac{a\omega_{\circ} - b}{c} \nabla^{2} + \frac{1 - \omega_{\circ}}{c}\right] p(\vec{r}) = \left[\nabla^{4} - \frac{b}{c} \nabla^{2} + \frac{1}{c}\right] S(\vec{r})$$
III-27

The result is:

$$\rho(\vec{r}) = S(\vec{r}) + \frac{\omega_0}{4\pi} \iiint S(\vec{r}') \left[ \frac{m_1 e^{-\alpha_1 |\vec{r} - \vec{r}'|}}{|\vec{r} - \vec{r}'|} + \frac{m_2 e^{-\alpha_2 |\vec{r} - \vec{r}'|}}{|\vec{r} - \vec{r}'|} \right] dv' \quad \text{III-28}$$

where

$$m_1 = \frac{1 - \alpha d_1^2}{(\alpha_2^2 - \alpha_1^2)C}$$
III-29

and

$$m_z = \frac{|-\alpha x_i^2|}{(\alpha_i^2 - \alpha_i^2)c}$$
III-30

Table III-l The explicit values of  $\alpha_1$ ,  $\alpha_2$ ,  $\mathbf{m}_1$ ,  $\mathbf{m}_2$  as a function of

_w <sub>o</sub>	م <sub>ا</sub>	×2	m,	M2
0	1.2344	4.1505	1.7013	9.2987
0.1	1.2039	4.0374	0.7916	9.2084
0.2	1.1683	3.9 <b>22</b> 4	0.8987	9.1013
0.3	1.1264	3.8054	1.0264	8.9736
0.4	1.0765	3.6866	1.1793	8.8207
0.5	1.0159	3.5662	1.3632	8.6368
0.6	0.9408	3.4446	1.5847	8.4153
0.7	0.8447	3.3221	1.8517	8.1483
0.8	0.7161	3.1996	2.1721	7.8280
0.9	0.5264	3.0778	2.5532	7.4468
1.0	0	2.9580	3.0000	7.0000

Equation (23) and (28) can be directly applied to the simple problem of an isotropic point source of unit strength at the origin in an infinite dispersion. The source term may be taken as the unscattered radiation,

$$S(\vec{r}) = \frac{e^{-r}}{4\pi r^2}$$
 III-31

Substituting Equation (31) into Equation (23) and (28) yields, respectively, for the first approximation:

$$\rho(\vec{r}) = \frac{e^{r}}{4\pi r^{2}} + \frac{\omega_{o}}{8\pi r} \frac{3}{\sqrt{3(-\omega_{o})}} \left\{ e^{\sqrt{3(-\omega_{o})}r} E_{1}[1+\sqrt{3(-\omega_{o})}r] + e^{\sqrt{3(-\omega_{o})}r} + e^{\sqrt{3(-\omega_{o})}r} \right\} + e^{\sqrt{3(-\omega_{o})}r} \left[ -\sqrt{3(-\omega_{o})}r + e^{\sqrt{3(-\omega_{o})}r} \right] + e^{\sqrt{3(-\omega_{o})}r} \left[ -\sqrt{3(-\omega_{o})}r + e^{-\sqrt{3(-\omega_{o})}r} \right] + e^{\sqrt{3(-\omega_{o})}r} \left[ -\sqrt{3(-\omega_{o})}r + e^{-\sqrt{3(-\omega_{o})$$

and the second approximation,

$$\rho(r) = \frac{e^{r}}{4\pi r^{2}} + \frac{\omega_{0}}{8\pi r} \left\{ \frac{\omega_{1}}{\alpha_{1}} \left[ e^{\alpha_{1}r} E_{1}[1+\alpha_{1}]r] - e^{\alpha_{1}r} E_{1}[1+\alpha_{1}]r] + e^{\alpha_{1}r} \ln \frac{1+\alpha_{1}}{1-\alpha_{1}} \right] + \frac{m_{2}}{\alpha_{2}} \left[ e^{\alpha_{1}r} E_{1}[1+\alpha_{1}]r] - e^{\alpha_{1}r} E_{1}[1+\alpha_{2}]r] + e^{\alpha_{2}r} \ln \frac{1+\alpha_{2}}{1-\alpha_{2}} \right] e^{-111-33}$$

where E(x) is the exponential integral defined by

$$E_{x}(x) = \int_{x}^{\infty} \frac{e^{-t}}{t} dt$$
III-34

Some numerical results calculated from Equation (32) and Equation (33) are compared with the known exact solution in Figure III-1 for  $\omega_0 = 0.3$ . Results of both the first and the second approximations are very close to the exact solution whereas the solution of the simple diffusion differs significantly.

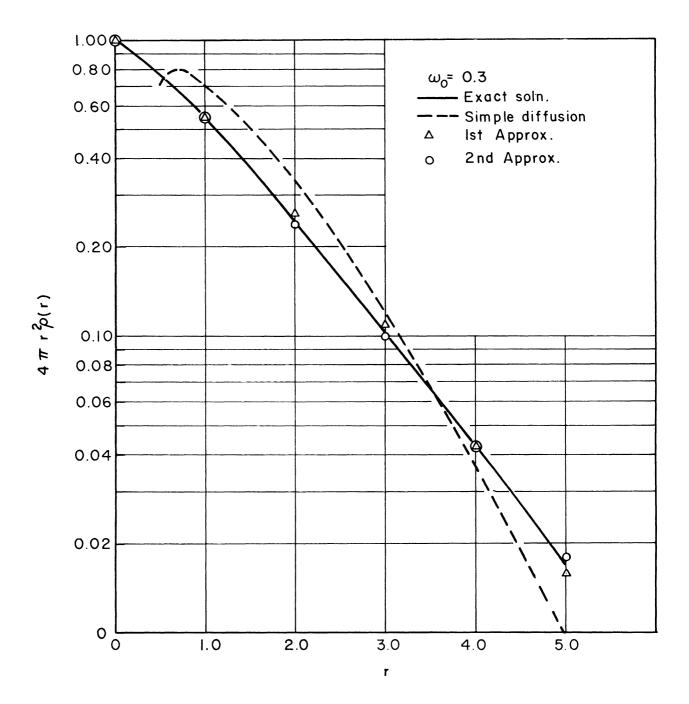


Figure III-l  $4\pi r^{2} \rho(r)$  as a function of r for an infinite medium with point source

## D. Boundary Conditions

The excellent numerical results obtained from the approximate solutions in the case of an infinite medium suggests the possibility of applying the same approximation to a finite region of dispersion.

It is well known that the classical diffusion approximation yields very poor results near the boundaries of a dispersion. This inaccuracy probably results from poor representation of the actual boundary conditions. In this section, instead of using physical arguments, a rigorous mathematical derivation for the boundary conditions appropriate to each order of approximation is derived. It will then be shown that the resulting boundary conditions yield results that are very close to the exact solutions.

Mathematically, for a finite region, the integral equation to be solved is

$$\rho(\vec{r}) = S(\vec{r}) + \omega_o \iiint_{r} K(|\vec{r}-\vec{r}'|) \rho(\vec{r}') dv'$$
III-35

where V is the volume of the dispersion.

In one dimensional cases such equations are known as Weiner -Hopf equations and has various application in electrical engineering problems (17)(18). For the present problem, the kernal is of the exponential type, and may be represented by

$$K(\vec{r}-\vec{r}') = \frac{1}{4\pi} |\vec{r}-\vec{r}'| \sum_{i} n_{i} e^{\beta_{i}|\vec{r}-\vec{r}'|} = \frac{1}{(2\pi)^{3}} \iint_{1+bk^{2}+ck^{2}} e^{i\vec{k}\cdot(\vec{r}-\vec{r}')} dk_{x}dk_{y}dk_{y}$$
III-36

The integral Equation (2) may be transformed to the differential Equation

Note that if c = a = o, and b = 1/3, Equation (36) and (19) represents the first approximation, so that the discussion that follows is applicable to both the first and second approximation. The complete solution of Equation (19) is in general made up of a particular solution depending on  $S(\vec{r})$  and a complementary function which is the general solution of the homogeneous Equation, i.e. Equation (19) with  $S(\vec{r}) = 0$ . The complimentary function would of course involve arbitrary constants, which are to be determined by the boundary conditions. Since Equation (19) is actually derived from the integral Equation (36), the boundary conditions can be derived directly by mathematical manipulation without resorting to physical arguments.

Let V be the volume of the dispersion,  $\Gamma$  the boundary of V (assume convex) and  $\hat{\gamma}$  be the outward normal to  $\Gamma$  from V, then the value of  $\rho(v)$  at the boundary can be directly obtained by use of Green's identity. Values of  $\rho(v)$  at the boundary  $\Gamma$  must satisfy the condition

$$b \iint \left\{ \left[ \rho(\vec{r}) - S(\vec{r}) \right] \nabla K(|\vec{r} - \vec{r}'|) - K(|\vec{r} - \vec{r}'|) \nabla \left[ \rho(\vec{r}) - S(\vec{r}) \right] \right\} \cdot \hat{n} \, da$$

$$+ c \iint_{\Gamma} \left\{ \nabla \left[ \rho(\vec{r}) - S(\vec{r}) \right] \nabla K(|\vec{r} - \vec{r}'|) - \nabla \left[ \rho(\vec{r}) - S(\vec{r}) \right] \nabla K(|\vec{r} - \vec{r}'|) \right\} \cdot \hat{n} \, da$$

$$+ c \iint_{\Gamma} \left\{ K(|\vec{r} - \vec{r}'|) \nabla \nabla^{2} \left[ \rho(\vec{r}) - S(\vec{r}) \right] - \nabla^{2} \left[ \rho(\vec{r}) - S(\vec{r}) \right] \nabla \nabla^{2} K(|\vec{r} - \vec{r}'|) \right\} \cdot \hat{n} \, da$$

$$+ a w_{o} \iint_{\Gamma} \left[ K(|\vec{r} - \vec{r}'|) \nabla \rho(\vec{r}) - \rho(\vec{r}) \nabla K(|\vec{r} - \vec{r}'|) \right] \cdot \hat{n} \, da$$

$$= 0$$
III-37

for all  $\overrightarrow{r}'$  inside V.

The proof of Equation (37) is straight forward, and is outlined in the following steps:

(i) Multiply Equation (19) by  $K(I\vec{r}-\vec{r}'I)$  and integrate over v yielding:

$$\iiint_{K(|\vec{r}-\vec{r}'|)} [1-b\vec{v}^2+c\vec{v}''] \rho(\vec{r}) d\vec{v}$$

$$-\iiint_{K(|\vec{r}-\vec{r}'|)} [1-b\vec{v}^2+c\vec{v}''] S(\vec{r}) d\vec{v}$$

$$-\omega_o \iiint_{K(|\vec{r}-\vec{r}'|)} [1-a\vec{v}^2] \rho(\vec{r}) d\vec{v} = 0$$
III-38

(ii) Using Green's identity, Equation (38) is reduced to:

$$\iint_{V} \rho(\vec{r}) [1-b\vec{V}^{2}+c\vec{V}^{4}] K(|\vec{r}-\vec{r}'|) d\vec{v} \\
-\iint_{V} S(\vec{r}) [1-b\vec{V}^{2}+c\vec{V}^{4}] K(|\vec{r}-\vec{r}'|) d\vec{v} \\
-\omega_{0} \iint_{V} \rho(\vec{r}) [1-a\vec{V}^{2}] K(|\vec{r}-\vec{r}'|) d\vec{v} \\
= -b \iint_{V} K(|\vec{r}-\vec{r}'|) \nabla \rho(\vec{r}) - \rho(\vec{r}) \nabla V^{2} K(|\vec{r}-\vec{r}'|) \int_{V} \hat{n} da \\
+c \iint_{V} [\nabla^{2} K(|\vec{r}-\vec{r}'|) \nabla \rho(\vec{r}) - \rho(\vec{r}) \nabla V^{2} K(|\vec{r}-\vec{r}'|)] \cdot \hat{n} da \\
+c \iint_{V} [K(|\vec{r}-\vec{r}'|) \nabla \rho(\vec{r}) - \nabla^{2} \rho(\vec{r}) \nabla K(|\vec{r}-\vec{r}'|)] \cdot \hat{n} da \\
+c \iint_{V} [K(|\vec{r}-\vec{r}'|) \nabla S(\vec{r}) - S(\vec{r}) \nabla K(|\vec{r}-\vec{r}'|)] \cdot \hat{n} da \\
-c \iint_{V} [K(|\vec{r}-\vec{r}'|) \nabla \nabla^{2} S(\vec{r}) - S(\vec{r}) \nabla V^{2} K(|\vec{r}-\vec{r}'|)] \cdot \hat{n} da \\
-c \iint_{V} [K(|\vec{r}-\vec{r}'|) \nabla \nabla^{2} S(\vec{r}) - \rho(\vec{r}) \nabla V^{2} K(|\vec{r}-\vec{r}'|)] \cdot \hat{n} da \\
+a \omega_{0} \iint_{V} [K(|\vec{r}-\vec{r}'|) \nabla \rho(\vec{r}) - \rho(\vec{r}) \nabla K(|\vec{r}-\vec{r}'|)] \cdot \hat{n} da , \quad \text{III-39}$$

(iii) From the expression of  $K(\vec{r}-\vec{r}')$  given in Equation (36), it can be shown that the left hand side of Equation (39) is identically zero. Collecting the terms given in the right hand side of Equation (39) produces the boundary condition given by Equation (37).

Equation (39) is the general boundary condition that can be derived. It can, in principle, be applied to any geometry and source distribution. However, in applying it to simple problems some simplified version will be considered.

Before considering the simplified version of the boundary conditions, it is to be noted that Equation (35) is a linear integral

equation of Fredholm type, and therefore can be solved in terms of orthogonal expansion without resorting to boundary conditions (19). In principle, for the particular type of K( $\vec{r}$ - $\vec{r}'$ ) it is possible to find a complete set of eigenfunctions  $\phi_i(\vec{r})$  corresponding to eigenvalues  $\lambda_i$ , satisfying the monogeneous equation

$$\Phi_{k}(\vec{r}) = \lambda_{k} \iiint K(\vec{r} - \vec{r} \cdot \vec{r}) \, d\sigma'$$
III-40

If such a set can be found, then, it can be shown that the orthogonal to each other in the sense that

$$\iiint \phi(\vec{r}, t_j(\vec{r}) d\sigma = 0 \qquad i \neq j$$

This immediately leads to the solution of Equation (35) in a series form and the result is:

$$\rho(\vec{r}) = \sum_{i} A_i f_i(\vec{r})$$
III-41

where

$$A_{\lambda} = \frac{\beta_{\lambda} \lambda_{\lambda}}{\lambda_{\lambda} - \omega_{0}}$$
III-42

and B is related to the source terms by

$$\beta_{x} = \frac{\iiint S(\vec{r}) \, \phi_{x}(\vec{r}) \, dv}{\iiint \phi_{x}^{2}(\vec{r}) \, dv}$$
III-43

The approximated kernal given by Equation (36) enables one to find  $\lambda$  and  $\phi(\vec{r})$  in a systematic way. Since  $\phi(\vec{r})$  must satisfy the homogeneous differential equation

$$\left[ \left( \nabla^{4} + (a\lambda_{i} - b) \nabla^{2} + (1 - \lambda_{i}) \right] \Phi (\vec{r}) = 0$$

$$III - 44$$

the forms of  $\phi_i(\vec{r})$  corresponding to each  $\lambda_i$ , are known in terms of arbitrary constants and  $\lambda_i$ . Substituting the expression for  $\phi_i(\vec{r})$ 

thus obtained into Equation (40) yields the characteristic equation for  $\lambda_{i}$ , and the constants for  $\phi_{i}(\vec{r})$ . In this one dimensional case, such a procedure has frequently been used e.g. (18). The extension to three dimensional case is straight forward.

In the present work, the primary purpose is to test the accuracy of the approximation, so that only simple problems such as parallel dispersion of finite and infinite thickness are considered. For such one dimensional cases, both the integral Equation and the boundary condition can be greatly simplified. The boundary condition used in the present work therefore combines the idea of the above two methods to minimize the numerical calculations. In the next section, such simplified boundary conditions are illustrated.

### E. Solution of One-Dimensional Problem

## 1. Second Approximation

## a) Medium bounded by parallel planes

To test the accuracy of the proposed approximation consider the problem of a parallel plane flux of unit intensity obliquely incident on a parallel dispersion bounded by two planes Z=0 and  $Z=\mathcal{T}_1$ . If the direction of the incidence is inclined at an angle  $\theta=\cos^{-1}\mu_0$  from the Z-axis, then  $\rho(\vec{r})$  is a function of Z only so that Equation (19) is reduced to:

$$\left[c\frac{d^{4}}{dz^{4}} + (aw_{o} - b)\frac{d^{2}}{dz^{2}} + (1 - w_{o})\right]\rho(z) = \left[c\frac{d^{4}}{dz^{4}} - b\frac{d^{2}}{dz^{2}} + 1\right]S(z)$$

where S(z) is obviously of the form

$$S(z) = e^{-\frac{2}{h}}.$$

The general solution of Equation (19) can be written in the form:

$$\rho(z) = A_1 e^{\alpha_1 z} + B_1 e^{\alpha_2 z} + A_2 e^{\alpha_1 z} + B_2 e^{\alpha_2 z} + (1+D)e^{-3\mu_0}$$
III-45

where

$$D = \frac{\omega_o(1-\frac{a}{\mu_o})}{(1-\omega_o)+(a\omega_o-b)\frac{1}{\mu_o}+9\mu_o^4}$$
III-46

The constants  $A_1$ ,  $A_2$ ,  $B_1$ ,  $B_2$ , can be obtained from the original integral equation. In the present one dimensional case, this integral equation takes the form:

$$\rho(z) = e^{-3/\mu_0} + \omega_0 \iiint_V \rho(z') K(|\vec{r} - \vec{r}'|) dv'$$
III-47

where  $K(\vec{r}-\vec{r}')$  is given by Equation (18) as

By carrying out the integrations that are independent of Z,

Equation (47) is reduced to the form:

$$\rho(z) = e^{-3/\mu_0} + \frac{\omega_0}{2} \int_0^{\infty} \left[ \frac{n_1}{\beta_1} e^{-\beta_1 |z-z'|} + \frac{n_2}{\beta_2} e^{-\beta_2 |z-z'|} \right] \rho(z') dz' \text{ III-48}$$

Now, substituting Equation (45) into Equation (48), and collecting

terms involving the same exponential forms of Z, the following

boundary conditions for A and B are obtained

# (i) For ω,+1

$$\frac{A_{1}}{-\alpha_{1}+\beta_{1}} + \frac{A_{2}}{\alpha_{1}+\beta_{1}} + \frac{B_{1}}{-\alpha_{2}+\beta_{1}} + \frac{B_{2}}{\alpha_{2}+\beta_{1}} + \frac{1+D}{-\frac{1}{\mu_{0}}+\beta_{1}} = 0$$

$$\frac{A_{1}}{-\alpha_{1}+\beta_{2}} + \frac{A_{2}}{\alpha_{1}+\beta_{2}} + \frac{B_{1}}{-\alpha_{2}+\beta_{2}} + \frac{B_{2}}{\alpha_{2}+\beta_{2}} + \frac{1+D}{-\frac{1}{\mu_{0}}+\beta_{2}} = 0$$

$$\frac{A_{1}e^{-\alpha_{1}\tau_{1}}}{\alpha_{1}+\beta_{1}} + \frac{A_{2}e^{\alpha_{1}\tau_{1}}}{-\alpha_{1}+\beta_{1}} + \frac{B_{1}e^{\alpha_{2}\tau_{1}}}{\alpha_{2}+\beta_{1}} + \frac{B_{2}e^{\alpha_{2}\tau_{1}}}{-\alpha_{2}+\beta_{1}} + \frac{(1+D)e^{-\tau/\mu_{0}}}{\frac{1}{\mu_{0}}+\beta_{1}} = 0$$

$$\frac{A_{1}e^{-\alpha_{1}\tau_{1}}}{\alpha_{1}+\beta_{2}} + \frac{A_{2}e^{\alpha_{1}\tau_{1}}}{-\alpha_{1}+\beta_{2}} + \frac{B_{1}e^{-\alpha_{2}\tau_{1}}}{\alpha_{2}+\beta_{2}} + \frac{B_{2}e^{\alpha_{2}\tau_{1}}}{-\alpha_{2}+\beta_{2}} + \frac{(1+D)e^{-\tau/\mu_{0}}}{\frac{1}{\mu_{0}}+\beta_{2}} = 0$$

The set of equations can be solved for the four constants, so that the values of  $_{\mathbb{Q}}(Z)$  are completely determined.

The above equations are valid for  $\omega_{\circ} \neq 1$  . (ii) For  $\omega_{\circ} = 1$ , (See table 1)

$$d_1 = 0$$
  $d_2 = 2.9580$ 

So that the solution of Equation (19) is

$$f^{(2)} = A_1 + A_2 Z + B_1 e^{-\alpha_2 Z} + B_2 e^{\alpha_2 Z} + (1+D) e^{-\frac{2}{3}\mu_0}$$
 III-49

The "boundary conditions" for determining  $\mathbf{A}_1$ ,  $\mathbf{A}_2$ ,  $\mathbf{B}_1$ ,  $\mathbf{B}_2$  are

slightly modified to:

$$\frac{A_{1}}{\beta_{1}} - \frac{A_{2}}{\beta_{1}^{2}} + \frac{B_{1}}{-\alpha_{2}+\beta_{1}} + \frac{B_{2}}{\alpha_{2}+\beta_{1}} + \frac{1+D}{-\frac{1}{\mu_{0}}+\beta_{1}} = 0$$

$$\frac{A_{1}}{\beta_{2}} - \frac{A_{2}}{\beta_{2}^{2}} + \frac{B_{1}}{-\alpha_{2}+\beta_{2}} + \frac{B_{2}}{\alpha_{2}+\beta_{2}} + \frac{1+D}{-\frac{1}{\mu_{0}}+\beta_{2}} = 0$$

$$\frac{A_{1}}{\beta_{1}} + \frac{(1+\beta_{1}\tau_{1})}{\beta_{1}^{2}} A_{2} + \frac{B_{1}e^{-\alpha_{2}\tau_{1}}}{\alpha_{2}+\beta_{1}} + \frac{B_{2}e^{\alpha_{2}\tau_{1}}}{-\alpha_{2}+\beta_{1}} + \frac{(1+D)e^{-\tau/\mu_{0}}}{\frac{1}{\mu_{0}}+\beta_{1}} = 0$$

$$\frac{A_{1}}{\beta_{2}} + \frac{(1+\beta_{2}\tau_{1})}{\beta_{2}^{2}} A_{2} + \frac{B_{1}e^{-\alpha_{2}\tau_{1}}}{\alpha_{2}+\beta_{2}} + \frac{B_{2}e^{\alpha_{2}\tau_{1}}}{-\alpha_{2}+\beta_{2}} + \frac{(1+D)e^{-\tau/\mu_{0}}}{\frac{1}{\mu_{0}}+\beta_{2}} = 0$$

### b) Semi-infinite medium

For a semi-infinite medium,  $\tau_i \rightarrow \infty$  the results can be simplified considerably. Following the same procedure as in the finite case,

(i) For 
$$w_0 \neq 1$$

$$\rho(z) = A_1 e^{-\alpha_1 z} + B_1 e^{-\alpha_2 z} + (1+D) e^{-\frac{2}{3}\mu_0}$$
III-50

where D is given by Equation (46) and  $A_1$  and  $B_1$  are determined from:

$$\frac{A_{1}}{-\alpha_{1}+\beta_{1}} + \frac{B_{1}}{-\alpha_{2}+\beta_{1}} + \frac{1+D}{-\frac{1}{\alpha_{0}}+\beta_{1}} = 0$$

$$\frac{A_{1}}{-\alpha_{1}+\beta_{2}} + \frac{B_{1}}{-\alpha_{2}+\beta_{2}} + \frac{1+D}{-\frac{1}{\alpha_{0}}+\beta_{2}} = 0$$

(ii) For  $\omega_0 = 1$ 

$$\rho(z) = A_1 + B_1 e^{-\alpha_2 z} + (1+D) e^{-\frac{2}{3}\mu_0}$$
III-51

where  $A_1$  and  $B_1$  are determined from:

$$\frac{A_1}{\beta_1} + \frac{B_1}{\alpha_2 + \beta_1} + \frac{1+D}{\lambda_0 + \beta_1} = 0$$

$$\frac{A_1}{\beta_2} + \frac{B_1}{-\alpha_2 + \beta_2} + \frac{1+D}{\lambda_0 + \beta_2} = 0$$

#### 2. First Approximation

The formula above were derived for the second approximation. The corresponding formulas for the first approximation can simply be obtained by letting a=c=o and b=1/3. The explicit results for the one dimensional problem, using the first approximation are given below:

(i) Integral Equation

$$\rho(z) = e^{-\frac{1}{2}\omega_c} + \frac{13}{2}\omega_c \int_0^{\tau_c} e^{-\frac{13|z-z'|}{2}} \rho(z') dz'$$
III-48

(ii) Differential Equation

$$\left[\frac{1}{3}\frac{d^2}{dz^2} - (1-\omega_0)\right]\rho(z) = \left(\frac{1}{3\mu_0^2} - 1\right)e^{-\frac{2}{3\mu_0}}$$
III-19'

(iii) General solution of the differential equation

$$\rho(z) = A_1 e^{-\alpha_1 z} + A_2 e^{\alpha_1 z} + (1+D) e^{-\frac{7}{\mu_0}}$$
III-45'

where

$$D = \frac{\omega_{\circ}}{(1-\omega_{\circ}) - \frac{1}{3\mu_{\circ}}}$$
III-46'

and  $\alpha_1 = [3(1-\omega_0)]^2$ 

(iv) Semi-infinite dispersion,  $\omega_{\circ} \pm 1$ 

$$\rho(z) = A_1 e^{-\alpha/2} + (I+D) e^{-3/4}.$$
 III-50'

where

$$A_{i} = \frac{-(1+D)}{-\frac{1}{4}a_{i} + \sqrt{3}} \left(\sqrt{3} - \alpha_{i}\right)$$
III-52

(v) Semi-infinite dispersion,  $\omega_0 = 1$ 

$$\rho(z) = A_1 + (1+D)e^{-3/\mu}.$$
 III-51:

$$A_{1} = \frac{-(1+D)}{\sqrt{3}-\frac{1}{4}}\sqrt{3}$$
III-53

(vi) Parallel plane dispersion  $\omega_o \neq 1$ 

$$\rho(z) = A_1 e^{-\alpha_1 z} + A_2 e^{\alpha_1 z} + (HD) e^{-3/u}.$$
 III-45

where A and A are determined from,

$$\frac{A_1}{-\alpha_1+\sqrt{3}} + \frac{A_2}{\alpha_1+\sqrt{3}} + \frac{(1+D)}{-\alpha_0} = 0$$

$$\frac{A_1e^{\alpha_1T}}{\alpha_1+\sqrt{3}} + \frac{A_2e^{\alpha_1T}}{\alpha_1+\sqrt{3}} + \frac{(1+D)e^{-\tau/\mu_0}}{-\tau} = 0$$

(vii) Parallel plane dispersion,  $W_b = 1$ 

$$\rho(z) = A_1 + A_2 z + (1+7) e^{-\frac{\pi}{2} u_0}$$
III-49:

where  $A_1$  and  $A_2$  are determined by

$$\frac{A_{1}}{\sqrt{3}} - \frac{A_{2}}{3} + \frac{1+D}{-\frac{1}{4}} = 0$$

$$\frac{A_{1}}{\sqrt{3}} + \frac{(1+\sqrt{3}t)}{3} A_{2} + \frac{(1+D)e}{\sqrt{3}+\frac{1}{4}} = 0$$

## F. Calculation of Albedo

Available results on the exact solutions of the parallel plane problem are generally expressed in terms of the reflectance of a half space, and the transmittance T or reflectance R of a parallel plane dispersion. One may calculate these parameters by direct integration of the expression for  $\rho(Z)$ . The resulting expressions are given below:

#### 1. First Approximation

(i) Semi-infinite medium, 
$$\omega_0 \neq 1$$

$$\rho(z) = A_1 e^{-\alpha_1 z} + (|+D|) e^{-2/\omega_0}$$
III-50\*

$$R = \frac{\omega_0 A_1}{2 \alpha_1 \mu_0} \left\{ \left[ -\frac{1}{\alpha_1} \ln \left( \left[ + \alpha_1 \right] \right) \right] + \frac{\omega_0 \left( \left[ + D \right)}{2} \left\{ \left[ -\mu_0 \ln \left( \left[ + \frac{1}{\mu_0} \right] \right] \right] \right]$$
 III-54.

(ii) Semi-infinite medium,  $\omega_o = 1$ 

$$\rho(z) = A_1 + (1+D) e^{-\frac{z}{\mu}}.$$
III-51:

$$R = \frac{\omega_o A_1}{4\mu_o} + \frac{\omega_o (HD)}{2} \{1 - \mu_o \ln(H \pm i_o)\}$$
 III-55'

(iii) Parallel plane dispersion,  $\omega_0 \neq 1$ 

$$\rho(z) = A_1 e^{-\alpha/2} + A_2 e^{\alpha/2} + (1+D) e^{-\alpha/2}$$
III-45'

Then

$$R = A_1 R_{\alpha_1} + A_2 R_{\alpha_2} + (I+D) R_{\perp}$$
III-56.

and

$$T = A_1 T_{\alpha_1} + A_2 T_{-\alpha_1} + (1+D) T_{\frac{1}{2}}$$
 III-57'

where

 $R_{\alpha_i} = \frac{\omega_o}{2\alpha_i \mu_o} \left\{ 1 - \frac{1}{\alpha_i} \ln \left| H \alpha_i \right| + e^{-\frac{\pi}{4} \alpha_i} \int_{\alpha_i} E_i(\tau_i) - e^{-\frac{\pi}{4} \alpha_i} \int_{\alpha_i} E_i(\tau_i) - \frac{E[\tau_i(H \alpha_i)]}{\alpha_i} \right\}$ 

and

 $T_{\alpha_{i}} = \frac{-\omega_{o}}{2\alpha_{i}} \left\{ e^{-d_{i}T_{i}} + \frac{e^{-d_{i}T_{i}}}{2\alpha_{i}} \ln\left|1-\alpha_{i}\right| - \frac{1}{\alpha_{i}} E_{i}(T_{i}) - E_{i}(T_{i}) + \frac{e^{-a_{i}T_{i}}}{2\alpha_{i}} \left[ T_{i}(1-\alpha_{i}) \right] \right\}$ 

and

$$E_n = \int_T^\infty \frac{e^{-x}}{x^n} dx$$

(iv) Parallel plane dispersion,  $\omega_o = 1$ 

$$P(z) = A_{1} + A_{2}z + (1+D)e^{-\frac{2}{3}\mu_{0}}$$

$$R = A_{1} \frac{\omega_{0}}{2\mu_{0}} \left[\frac{1}{2} - E_{3}(\tau_{1})\right] + A_{2} \frac{\omega_{0}}{2\mu_{0}} \left[\frac{1}{3} - E_{4}(\tau_{1}) - \tau_{1}E_{3}(\tau_{1})\right] + (1+D) R_{\frac{1}{2}\mu_{0}}$$

$$T = A_{1} \frac{\omega_{0}}{2\mu_{0}} \left[\frac{1}{2} - E_{3}(\tau_{1})\right] + A_{2} \frac{\omega_{0}}{2\mu_{0}} \left[-\frac{1}{3} + \frac{1}{2}\tau_{1} + E_{4}(\tau_{1})\right] + (1+D) T_{\frac{1}{2}\mu_{0}}$$

$$III-61$$

- Second Approximation
  - (i) Semi-infinite medium  $\omega$ .  $\Rightarrow$  \

$$\rho(z) = A_1 e^{-\alpha_1 z} + B_1 e^{\alpha_2 z} + (HD) e^{-\frac{\gamma}{\mu_0}}$$
III50

$$R = \frac{\omega_o A_i}{2\alpha_i \mu_o} \left\{ \left[ -\frac{1}{\alpha_i} \ln(H\alpha_i) \right] + \frac{\omega_o B_i}{2\alpha_i \mu_o} \left\{ \left[ -\frac{1}{\alpha_i} \ln(H\alpha_i) \right] + \frac{\omega_o (HT)}{2} \left\{ \left[ -\mu_o \ln(H\alpha_i) \right] \right\} \right\}$$
III-54

(ii) Semi-infinite medium

$$\rho(z) = A_1 + B_1 e^{-4z} + (1+D) e^{-2/u_0}$$

$$R = \frac{\omega_0 B_1}{2 d_2 \mu_0} \left\{ 1 - \frac{1}{d_2} \ln(1+\alpha_0) \right\} + \frac{\omega_0 A_1}{4 \mu_0} + \frac{\omega_0 (1+D)}{2} \left\{ 1 - \mu_0 \ln(1+\frac{1}{\mu_0}) \right\} \quad \text{III-55}$$

(iii) Parallel plane medium

$$\rho(z) = A_1 e^{-x/2} + A_2 e^{x/2} + B_1 e^{-x/2} + \beta_2 e^{-x/2} + (1+D)e^{-x/4}.$$
 III-45

$$R = A_1 R_{\alpha_1} + A_2 R_{\alpha_1} + B_1 R_{\alpha_2} + B_2 R_{\alpha_2} + (I+D) R_{\alpha_2}$$
 III-56

$$T = A_1T_{\alpha_1} + A_2T_{-\alpha_1} + B_1T_{\alpha_2} + B_2T_{-\alpha_2} + (I+D)T_{-\alpha_2}$$
III-57

where  $R_{\alpha}$  and  $T_{\alpha}$  are defined as Equations (58) and (59) respectfully.

(iv) Parallel plane dispersion  $\omega_o =$ 

$$\rho(z) = A_1 + A_2 z + B_1 e^{-\alpha_2 z} + B_2 e^{\alpha_2 z} + (1+D) e^{-\frac{2}{2} \mu_0}$$
 III-49

$$R = B_{1}R_{4} + B_{2}R_{4} + \frac{\omega_{0}}{2\mu_{0}} \left[ \frac{1}{2} - E_{3}(\tau_{1}) \right] A_{1} + \frac{\omega_{0}}{2\mu_{0}} \left[ \frac{1}{3} - E_{4}(\tau_{1}) - \tau_{1}E_{3}(\tau_{1}) \right] A_{2} \text{ III-60}$$

$$+ (HD) R_{\frac{1}{2}}$$

$$T = \beta_{1}T_{4} + \beta_{2}T_{4} + \frac{\omega_{0}}{2\mu_{0}} \left[ \frac{1}{2} - E_{3}(\tau_{1}) \right] A_{1} + \frac{\omega_{0}}{2\mu_{0}} \left[ -\frac{1}{3} + \frac{1}{2}\tau_{1} + E_{4}(\tau_{1}) \right] A_{2} \text{ III-61}$$

$$+ (I+D)T_{4}$$

#### G. Numerical Results

Some typical numerical results calculated by using the appropriate approximate formulas derived in the previous section are plotted in Figures III 2-8.

In Figure III-2, the integrated reflectance R for a semi-infinite medium is plotted versus  $\mu_0$  for  $\kappa_0=0.9$  and 0.5. The results of the second approximation are indistinguishable from the exact values. In Figures III 3-5 the integrated reflectance R is plotted versus  $\kappa_0$  for several values of  $\mu_0$ . The numerical values were calculated from the first and second approximations. Again, the second approximations are indistinguishable from the exact values and the first approximation are acceptably accurate.

In Figure III-6, the reflectance R for a finite medium is plotted against  $\mu_0$  for several  $\omega_0$ . Again the results of the second approximation are very good. The deviation of the first approximation from the exact solution becomes bigger as  $\omega_0$  increases. In Figures III-7 and 8, the diffusive transmission and the integrated transmission are plotted against  $\mu_0$  for several  $\omega_0$ . Only second approximations are compared with the exact solutions.

From all these figures, it is obvious that the results of the second approximation are practically indistinguishable from the exact solution for all conditions, and the simple diffusion equation with correct boundary conditions (i.e. the first approximation) is an acceptable approximation.

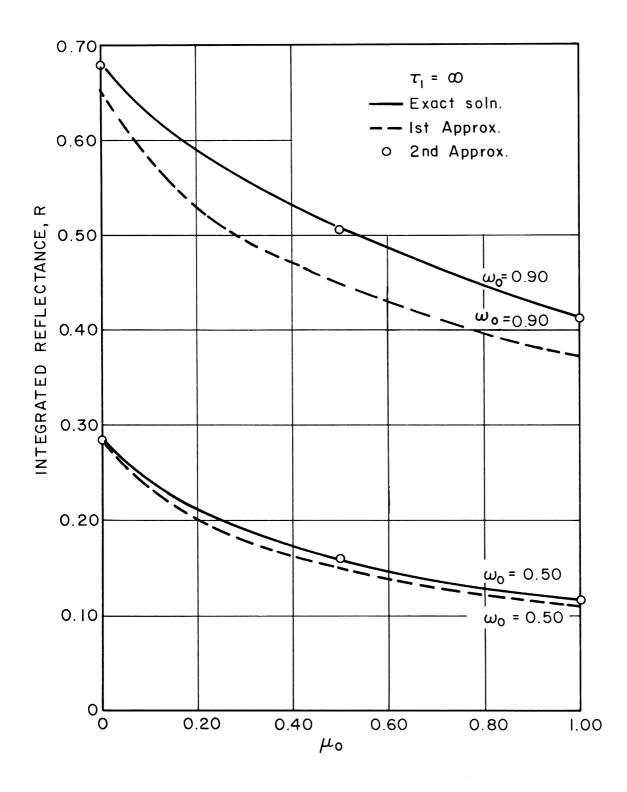


Figure III-2 Integrated reflectance as a function of incident angle, R vs  $\mu_{0}$  , parameters of  $\omega_{o}$  , and  $\tau_{1}$  = 00

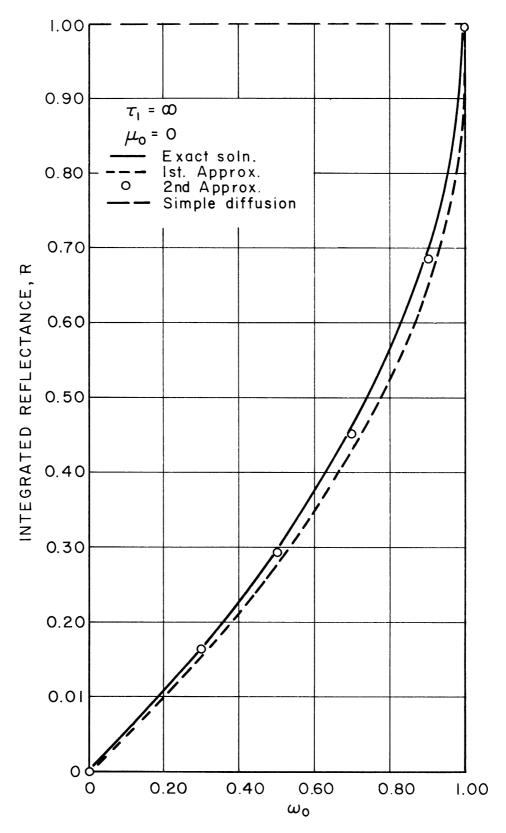


Figure III-3 Integrated reflectance as a function of albedo for single scattering, R vs  $\omega_{\rm o}$ ,  $\mu_{\rm o}$  = 0, and  $\tau_{\rm l}$  =  $\infty$ 

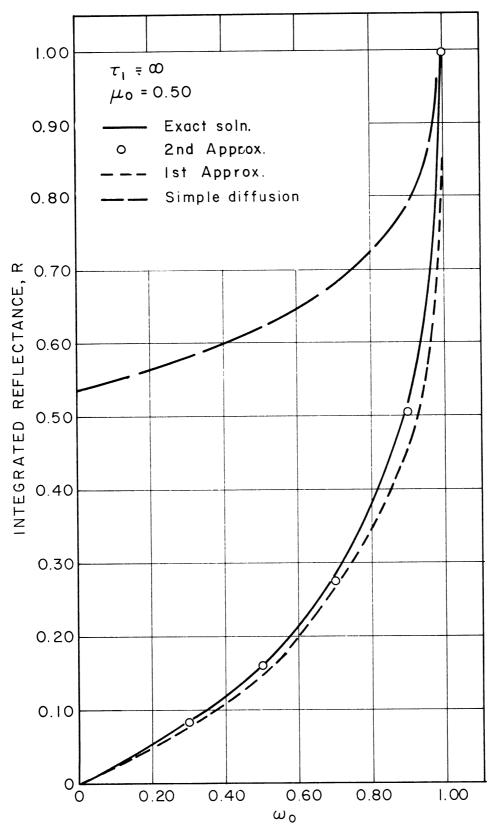


Figure III-4 Integrated reflectance as a function of albedo for single scattering, R vs.  $\omega_{o}$  ,  $\mu_{o}$  = 0.5, and  $\tau_{1}$  = 00

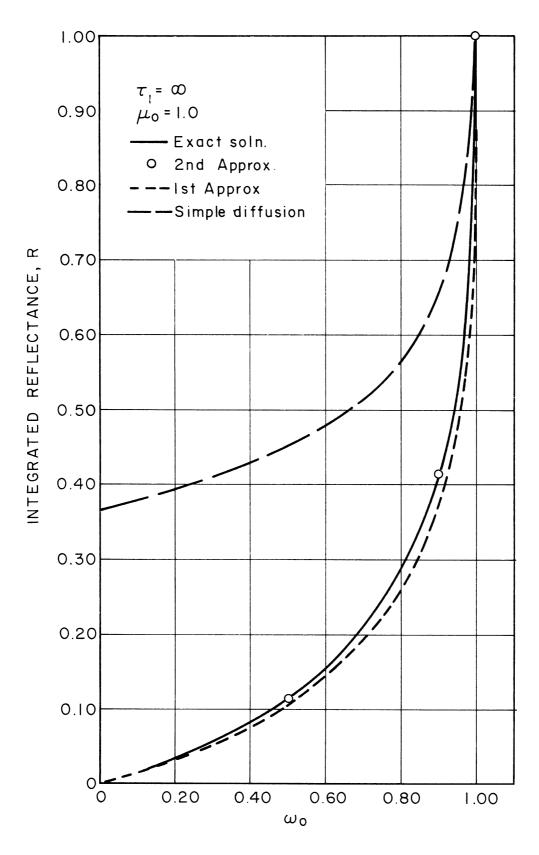


Figure III-5 — Integrated reflectance as a function of albedo for single scattering, R vs.  $\omega_{o}$   $\mu_{O}$  = 1.0, and  $\tau_{1}$  =  $\infty$ 

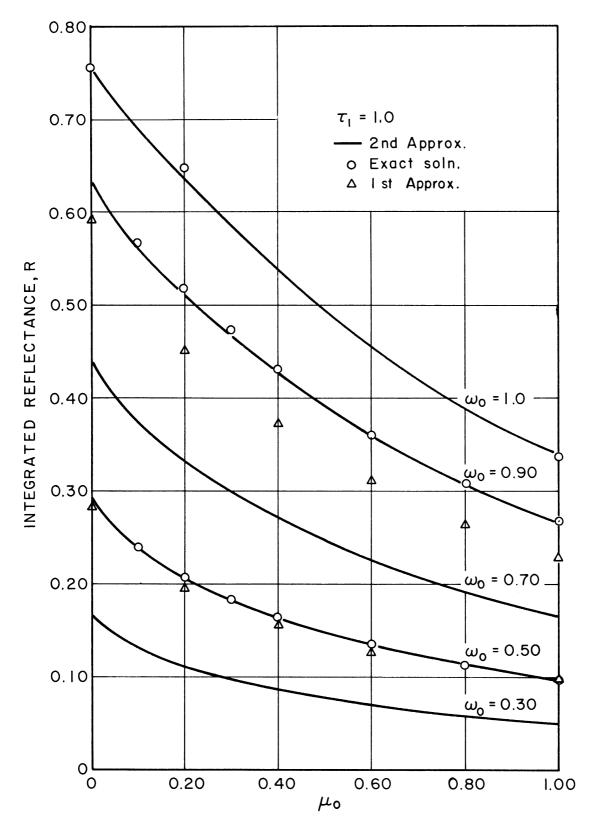


Figure III-6 Integrated reflectance as a function of incident angle, R vs.  $\mu_0,$  parameters of  $\omega_{\text{o}}$  , and  $\tau_1$  =  $\infty$ 

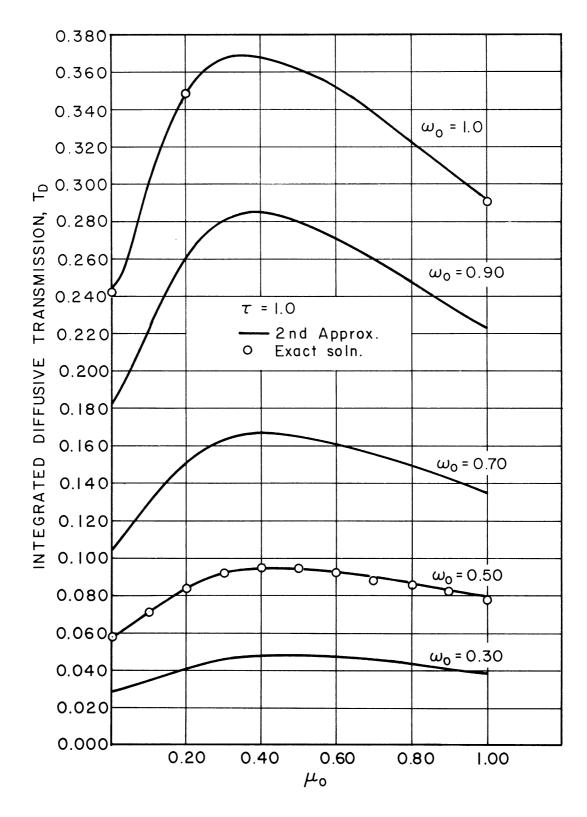


Figure III-7 Integrated diffusive transmission as a function of incident angle, T<sub>D</sub> vs.  $\mu_{o}$ , parameters of  $\omega_{o}$ , and  $\tau_{1}$  = 1

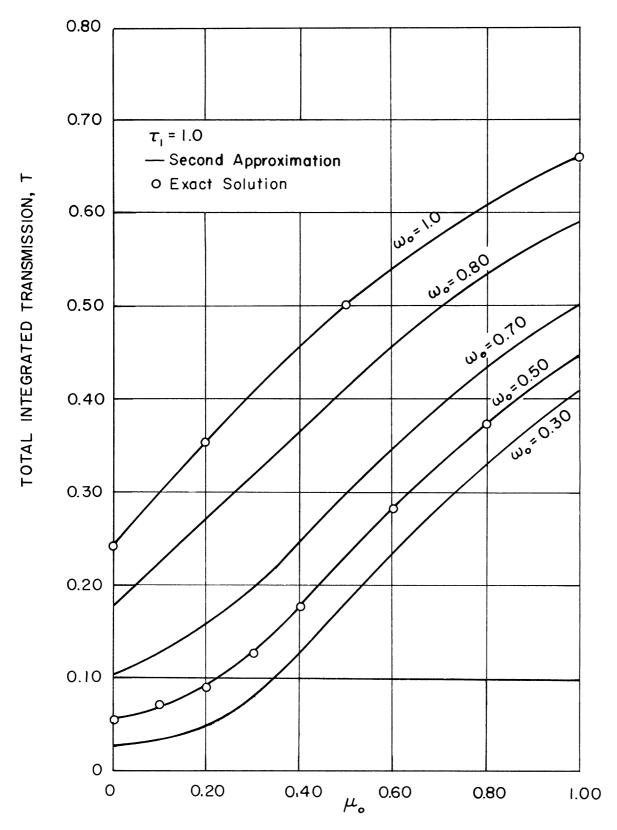


Figure III-8 Total integrated transmission as a function of incident angle, T vs  $\mu_{0},$  parameters of  $\omega_{0}$  , and  $\tau_{1}$  = l

#### H. Conclusions

From the excellent numerical check with the exact solutions of one dimensional problems with isotropic scattering, it is reasonable to believe that the diffusion-type approximation, if used with correct boundary conditions, can yield very good results which are adequate for all practical purposes. In the present work a systematic approach to the diffusion-type approximation has been developed. A method of determining the correct boundary conditions appropriate to each order of approximation has also been developed.

Since the exact solutions of the multiple scattering problem are generally difficult or impossible to obtain, the need for a comparatively simple, yet reasonably accurate approximate method of solution is obvious. It therefore seems worthwhile, to exploit further the possibilities of using the diffusion-type approximation. This extended investigation should proceed along the following two directions: (i) use of the suggested method to find approximate solutions for geometries where exact solutions do not exist, such as a point source in front of a parallel dispersion, (ii) extension of the method to anisotropic scattering, for which some exact solutions are now available for comparison.

#### SUMMARY AND CONCLUSIONS

- 1. A satisfactory method was developed for numerical integration of the transport equation for parallel plane radiation obliquely incident on a parallel-plane dispersion. The important advance is in the allowance for strongly anisotropic scattering. Previous results were limited to isotropic or Rayleigh (nearly isotropic) scattering.
- 2. The solutions are exact except for errors in the numerical procedure itself. These errors arise from numerical roundoff, from termination of the iterative procedure at an arbitrary point and from integration by quadrature. The error due to the use of quadrature arises from the use of a finite number of terms, i.e., from division of the integration into a finite number of intervals. The numerical results agree very well with previous results for special cases (isotropic and Rayleigh scattering, and anisotropic scattering by semi-infinite dispersions) indicating that the net numerical errors are not serious.
- 3. The numerical and graphical results illustrate the effects of
  - a) the phase function for single scattering
  - b) the optical thickness of the dispersion
  - c) the angle of incidence
  - d) the albedo for single scattering (ratio of scattered to scattered plus absorbed radiation)

- 4. The numerical results are limited to finite absorption since the iterative procedure did not converge satisfactorily for zero absorption. It appears feasible but unnecessary to develop a separate numerical procedure for zero absorption since the results for small finite absorption can readily be extrapolated to zero absorption.
- 5. The reflectance and transmission for strongly anisotropic scattering was found to differ decisively from that for isotropic and Rayleigh scattering. A similar result was previously found for semi-infinite dispersions. It is therefore concluded that solutions for isotropic scattering do not provide a reliable guide for the reflectance or transmission for dispersions of liquid or solid particles.
- 6. The transmission and reflectance were found to be dependent primarily on the fraction of the radiation scattered into the backward (or forward) hemisphere. Only slight dependence was found on the other characteristics of the phase function. This suggests that the complicated phase functions associated with large circumference to wavelength ratios can be approximated by simple functions without serious error if the back-scattered fraction is preserved. This result is important from a computational point-of-view since the extent and cost of the computations goes up rapidly with the number of terms used to represent the phase function.

- 7. Many approximate models have been proposed for multiple scattering. The solutions based on these models were compared with the "exact" values for both isotropic and anisotropic scattering. The six-flux model was found to be the best of these models but to be a fair approximation, Richard's modified diffusion model and the two-flux model were poorer in all cases and the simple diffusion model was seriously in error in all cases.
- 8. Despite their inability to provide good absolute values for the transmission and reflectance the above approximate models were successfully utilized to provide simple expressions for interpolation and extrapolation of the exact values.
- 9. A new variable-order diffusion-type model was developed which yields very accurate results for isotropic scattering. It is proposed to develop this model further for anisotropic scattering.
- 10. All accurate solutions for multiple scattering are currently limited to one-dimensional radiant transport. However both the numerical method and the variable order diffusion-type model hold promise for the computation of two-dimensional radiant transport.

#### APPENDIX A

# Derivation and List of Equations

l. Derivation of Explicit Relations for  $\textbf{b}_m(\tau_1,\,\mu,\,\mu_0)$  and  $\textbf{c}_m(\tau_1,\,\mu,\,\mu_0)\,.$ 

The diffuse reflected intensity is

$$\frac{I(0, \mu, \varphi)}{I_{o}(-\mu)} = \frac{\mu_{o}}{4\pi} \sum_{m=0}^{N} (2 - \delta_{o,m}) \cos m\varphi \sum_{l=m}^{N} (-1)^{l+m} F_{l}(\zeta_{l}, \mu, \mu_{o})$$
 I-14

substituting relation I-3 for  $F_{I}^{m}(z_{I},\mu,\mu_{\bullet})$  and letting

$$\omega_{L}^{m} = \omega_{o} A_{l} \frac{(l-m)!}{(l+m)!}$$

gives

$$\frac{I(0, \mu, \varphi)}{I_{\bullet}(-\mu)} = \frac{\mu_{\bullet}\omega_{\bullet}}{4\pi(\mu + \mu_{\bullet})} \left\{ \sum_{l=0}^{N} (-1)^{l} A_{l} \left[ \psi_{l}^{\circ}(\mu) \psi_{l}^{\circ}(\mu_{\bullet}) - \phi_{k}^{\circ}(\mu) \phi_{l}^{\circ}(\mu_{\bullet}) \right] + 2 \sum_{m=1}^{N} \cos m \varphi \sum_{l=m}^{N} (-1)^{l+m} A_{l} \frac{(l-m)!}{(l+m)!} \left[ \psi_{l}^{m}(\mu_{\bullet}) \psi_{k}^{m}(\mu) - \phi_{k}^{m}(\mu_{\bullet}) \phi_{k}^{m}(\mu) \right] \right\}$$
Al-2

and for N = 2

$$\cos 2\varphi = 2\cos^2 \varphi - 1$$

A. = 1

$$\frac{I(0,\mu,\varphi)}{I_{\circ}(-\mu)} = \frac{\mu_{\circ}\omega_{\circ}}{4\pi(\mu+\mu_{\circ})} \left\{ \left[ \psi_{\circ}^{\circ}(\mu_{\circ}) \psi_{\circ}^{\circ}(\mu) - \phi_{\circ}^{\circ}(\mu_{\circ}) \phi_{\circ}^{\circ}(\mu) \right] - A_{\circ} \left[ \psi_{\circ}^{\circ}(\mu_{\circ}) \psi_{\circ}^{\circ}(\mu) - \phi_{\circ}^{\circ}(\mu_{\circ}) \phi_{\circ}^{\circ}(\mu) \right] \right\}$$

$$+ A_{2} \left[ \psi_{2}^{\circ} (\mu_{\bullet}) \psi_{2}^{\circ} (\mu) - \phi_{2}^{\circ} (\mu_{\bullet}) \phi_{2}^{\circ} (\mu) \right] + A_{1} \cos \varphi \left[ \psi_{1}^{\prime} (\mu_{\bullet}) \psi_{1}^{\prime} (\mu) - \phi_{1}^{\prime} (\mu_{\bullet}) \phi_{1}^{\prime} (\mu) \right]$$

$$-A_{2} \frac{\cos \varphi}{3} \left[ \psi_{2}^{1} (\mu_{0}) \psi_{2}^{1} (\mu) - \phi_{2}^{1} (\mu_{0}) \phi_{2}^{1} (\mu) \right]$$

$$+A_{2} \frac{2\cos^{2} \varphi - 1}{6} \left[ \psi_{2}^{2} (\mu_{0}) \psi_{2}^{2} (\mu) - \phi_{2}^{2} (\mu_{0}) \phi_{2}^{2} (\mu) \right]$$
Al-3

Therefore, let

$$K = \frac{\mu_o \omega_o}{4\pi (\mu + \mu_o)}$$

$$Q_{\bullet} = \left[ \psi_{\bullet}^{\circ}(\mu_{\bullet}) \psi_{\bullet}^{\circ}(\mu) - \phi_{\bullet}^{\circ}(\mu_{\bullet}) \phi_{\bullet}^{\circ}(\mu) \right] - A_{\bullet} \left[ \psi_{\bullet}^{\circ}(\mu_{\bullet}) \psi_{\bullet}^{\circ}(\mu) - \phi_{\bullet}^{\circ}(\mu_{\bullet}) \phi_{\bullet}^{\circ}(\mu) \right]$$

$$+A_{2}[\psi_{1}^{\circ}(\mu_{0})\psi_{2}^{\circ}(\mu)-\phi_{1}^{\circ}(\mu_{0})\phi_{2}^{\circ}(\mu)]$$
 A1-5

$$Q_{1} = A_{1} \left[ \psi_{1}^{i} (\mu_{0}) \psi_{1}^{i} (\mu) - \phi_{1}^{i} (\mu_{0}) \phi_{1}^{i} (\mu) \right] - \frac{A_{2}}{3} \left[ \psi_{2}^{i} (\mu_{0}) \psi_{2}^{i} (\mu) - \phi_{2}^{i} (\mu_{0}) \phi_{2}^{i} (\mu) \right]$$

$$A1-6$$

$$Q_{2} = \frac{A_{2}}{6} \left[ \psi_{2}^{2}(\mu_{0}) \psi_{2}^{2}(\mu) - \phi_{2}^{2}(\mu_{0}) \phi_{2}^{2}(\mu) \right]$$
Al-7

and the expressions for  $b_m$  (7,  $\mu$ ,  $\mu$ ) become

$$b_0 = K(Q_0 - Q_2)$$
 Al-8a

$$b_1 = KQ_1$$
 Al-8b

$$b_2 = 2 kQ_2$$
 Al-8c

The diffuse reflected intensity may be written as

$$\frac{I(0, \mu, \mu_0)}{I_0(-\mu_0)} = b_0 + b_1 \cos \varphi + b_2 \cos \varphi$$
Al-9

Proceeding in the same manner to derive the relations for  $c_m(\tau_1,\;\mu,\;\mu_0), \; \text{the expression for the diffuse transmitted intensity}$  can be written as

$$\frac{I(\tau_{1},-\mu,\varphi)}{I_{o}(-\mu)} = \frac{\mu_{o}}{4\pi} \sum_{m=0}^{N} (2-\delta_{o,m}) \cos m\varphi \sum_{l=m}^{N} G_{l}^{m}(\tau_{l},\mu,\mu_{o})$$
 I-19

Equations I-14 and I-29 are of the same form with the term  $(-1)^{1+m} \, F_I^m \, (z_I,\mu,\mu_0) \quad \text{in I-14 replaced by} \quad G_I^m \, (z_I,\mu,\mu_0)$  in I-19. The only difference between  $F_I^m$  and  $G_I^m$  is that the positions of  $\Psi_I^m(\mu)$  and  $\Phi_I^m(\mu)$  are interchanged and the term  $\mu + \mu_0$  in  $F_I^m$  is replaced by  $\mu - \mu_0$  in  $G_I^m$ .

Therefore, if

- 1. All signs are made positive in front of the  $A_{l}^{\prime}$  s
- 2.  $\mu$  +  $\mu_{O}$  is replaced in K by  $\mu$   $\mu_{O}$
- 3.  $\psi_{\lambda}^{m}(\mu)$  and  $\phi_{\lambda}^{m}(\mu)$  are interchanged

The expression for the diffuse transmitted intensity may be written directly.

$$K' = \frac{\mu_{\circ} \omega_{\circ}}{4\pi (\mu - \mu_{\circ})}$$
 Al-10

$$Q_{o}' = \left[ \psi_{o}''(\mu_{o}) \, \phi_{o}''(\mu) - \phi_{o}''(\mu_{o}) \, \psi_{o}''(\mu) \right] + A_{1} \left[ \psi_{1}''(\mu_{o}) \, \phi_{1}''(\mu) - \phi_{1}''(\mu_{o}) \, \psi_{1}''(\mu) \right]$$

$$+ A_2 \left[ \psi_1^{\circ} (\mu_0) \phi_1^{\circ} (\mu) - \phi_1^{\circ} (\mu_0) \psi_1^{\circ} (\mu) \right]$$
 Al-11

$$Q_{1}' = A_{1} \left[ \psi_{1}' (\mu_{0}) \, \phi_{1}' (\mu) - \, \phi_{1}' (\mu_{0}) \, \psi_{1}' (\mu) \right] + \frac{A_{2}}{3} \left[ \psi_{2}' (\mu_{0}) \, \phi_{2}' (\mu) - \, \phi_{2}' (\mu_{0}) \, \psi_{2}' (\mu) \right] \quad \text{Al-l2}$$

$$Q_{2}' = \frac{A_{2}}{6} \left[ \psi_{1}^{2} (\mu_{0}) \phi_{1}^{2} (\mu) - \phi_{1}^{2} (\mu_{0}) \psi_{2}^{2} (\mu) \right]$$
Al-13

with the expressions for  $\mathbf{c}_{\mathrm{m}}$ 

$$C_{\bullet} = K' \left[ Q_{\bullet}' - Q_{2}' \right]$$
 Al-14a

$$c_i = \kappa' Q_i'$$

$$c_2 = 2K'Q_2'$$
 Al-14c

and the diffuse transmitted intensity may be written as

$$\frac{I(z_1,-\mu,\varphi)}{I_{\alpha}(-\mu_{\alpha})} = c_{\alpha} + c_{1}\cos\varphi + c_{2}\cos^{2}\varphi$$
 Al-15

## 2. Integrated Reflectance of Half Space for Rayleigh Scattering

There is a section on integral formulations in reference (4).

For Rayleigh scattering ( $\omega$ , =0), Equation I-151 in that report then reduces to, for k = 0 and 1,

$$(\frac{1}{2}\omega_0 A_{00} - 1)\mu_0 + U_1 + \frac{1}{2}\omega_2 A_{20}\mu_2 = 0$$
 A2-1

$$\left(1 + \frac{3}{2}\omega_{0}\mu A_{01}\right)U_{0} - 3\mu U_{1} + \left(2 + \frac{3}{2}\omega_{z} A_{z1}\mu\right)U_{z} = 0$$
 A2-2

where A and U are defined as

$$\Psi_{\bullet}(\mu) = U_{\bullet}(\mu) K(\mu)$$
 A2-3

$$A_{rk} = \int_{0}^{1} U_{r}(\mu') P(-\mu') K(\mu') d\mu'$$
A2-4

K is equivalent to the H function in Chandrasekhar's Radiative

Transfer (1) and P is the Legendre polynomial. According to

Equation I-157 in reference (5).

$$U_{4}(\mu) = P_{4}(\mu) \quad \text{when} \quad \mu = 0$$
 A2-5

By this additional condition Equations A2-1 and A2-2 can be solved for

 $\mu_0$ ,  $\mu_1$  and  $\mu_2$ .

$$U_{0} = 1 + \frac{3}{4} \omega_{2} A_{21} \mu + \frac{3}{4} \omega_{2} A_{20} \mu^{2}$$

$$U_{1} = \left(1 - \frac{1}{2} \omega_{0} A_{00} + \frac{1}{4} \omega_{2} A_{20}\right) \mu + \left[\frac{3}{4} \left(1 - \frac{1}{2} \omega_{0} A_{00}\right) \omega_{2} A_{21} + \frac{3}{8} \omega_{0} \omega_{2} A_{01} A_{20}\right] \mu^{2}$$

$$A2 - 7$$

$$U_{2} = -\frac{1}{2} - \frac{3}{4} \omega_{0} A_{01} \mu + \frac{3}{4} \left(2 - \omega_{0} A_{00}\right) \mu^{2}$$

$$A2 - 8$$

The K function is

$$\frac{1}{K(\mu)} = 1 - \frac{\mu}{2} \int_{-\mu + \mu'}^{1} \frac{\Psi(\mu') K(\mu')}{\mu + \mu'} d\mu',$$
 A2-9

and

$$\Psi(\mu') = \sum_{r=0}^{N} (-1)^{r} \omega_{r} U_{r} (\mu') U_{r} (-\mu').$$
 A2-10

The latter can be written in terms of A's using Equations A2-6, A2-7 and A2-8

$$\Psi(\mu') = \omega_0 + \frac{1}{4}\omega_2 + \left(-\frac{3}{2}\omega_2 + \frac{3}{4}\omega_0\omega_2 A_{\infty} - \frac{9}{16}\omega_0^2\omega_2 A_{01}^2 + \frac{3}{2}\omega_0\omega_2 A_{20} - \frac{9}{16}\omega_0\omega_2^2 A_{21}^2\right)\mu'^2 + \left(\frac{9}{4}\omega_2 - \frac{9}{4}\omega_0\omega_2 A_{\infty} + \frac{9}{16}\omega_0\omega_2^2 A_{20}^2\right)\mu'^4$$
A2-11

The A's can be cancelled out by the use of Equations (I-149) and (I-152) in the reference (5).

$$U_{k}(\mu) K(\mu) = P_{k}(\mu) + \frac{1}{2} \mu \sum_{r=0}^{N} (-1)^{r} \omega_{r} \Psi_{r}(\mu) \begin{cases} \frac{P_{k}(-\mu') \Psi_{r}(\mu') d\mu'}{\mu + \mu'}, & k = 0, 1, \dots, N \\ \frac{1}{2} \mu + \frac{1}{$$

For k = 0, multiplying A2-12 by  $P_{\text{O}}(\mu)$  then integrating from 0 to 1 gives

$$A_{\circ \circ} = 1 + \frac{1}{4} \omega_{\circ} A_{\circ \circ}^2 - \frac{1}{4} \omega_2 A_{21}^2$$
 A2-13

Similarly for k = 1 and 2.

$$A_{11} = \frac{1}{2} - \frac{1}{4}\omega_{0} A_{01}^{2} - \frac{1}{4}\omega_{2} A_{21}^{2}$$
A2-14

$$A_{20} = -\frac{1}{8} \left( \omega_o A_{so}^2 + \omega_1 A_{2o}^2 \right) + \frac{3}{8} \left( \omega_o A_{si}^2 + \omega_2 A_{2i}^2 \right)$$
 A2-15

Then A2-11 can be simplified with A2-13, A2-14, and A2-15 to give

$$\Psi(\mu') = \omega_0 + \frac{1}{4}\omega_2 - \frac{3}{4}(2 - \omega_0)\omega_2 {\mu'}^2 + \frac{9}{4}(1 - \omega_0)\omega_2 {\mu'}^4$$
A2-16

The numerical solution of R was carried out as follows. The K function was first evaluated by the computer using an iterative procedure. Then the moments of K are computed.

$$\alpha_n = \int_a^b K(\mu) \mu^n d\mu$$
 A2-17

Then A's are computed from the values of a's and A2-6, A2-7, and A2-8. However it appears more convenient to use A2-6 and A2-8 to derive the following.

$$A_{00} = d_0 + \frac{3}{4}\omega_2 d_1 A_{21} + \frac{3}{4}\omega_2 d_2 A_{20}$$
A2-18

$$-A_{01} = \alpha_1 + \frac{3}{4}\omega_1\alpha_2A_{21} + \frac{3}{4}\omega_2\alpha_3A_{20}$$
 A2-19

$$A_{20} = -\frac{1}{2}\alpha_0 - \frac{3}{4}\omega_0\alpha_1A_{01} + \frac{3}{4}(2-\omega_0A_{00})\alpha_2$$
A2-20

$$-A_{21} = -\frac{1}{2}\alpha_{1} - \frac{3}{4}\omega_{0}\alpha_{2}A_{01} + \frac{3}{4}(2 - \omega_{0}A_{00})\alpha_{3}$$
 A2-21

Finally R can be found by

$$R = 1 - \frac{1}{\mu_{\bullet}} \Psi_{1}(\mu_{\bullet})$$

$$= 1 - \frac{1}{\mu_{\bullet}} U_{1}(\mu_{\bullet}) K(\mu_{\bullet})$$

$$= 1 - \left[1 - \frac{1}{2} \omega_{\bullet} A_{\bullet \bullet} + \frac{1}{4} \omega_{2} A_{2 \bullet} + \left(\frac{3}{4} \omega_{2} A_{2 1} + \frac{3}{8} \omega_{\bullet} \omega_{2} A_{2 1} A_{2 0}\right) - \frac{3}{8} \omega_{\bullet} \omega_{2} A_{\bullet \bullet} A_{2 1} \mu_{\bullet}\right] K(\mu_{\bullet})$$

$$A2-22$$

Sample values of the results for  $\omega_{\bullet}$  = 0.9, 0.6, and 0.3 are tabulated in Table A-1.

Table A-l The K-functions and Related Numerical Values for Computing the Integrated Reflectance of Half Space, Rayleigh Scattering

0.60 1.2442 0.6481

1.0969 0.5579

0.30

	ω.	= 0.90		ω₀ = 0.6	io .	ω <sub>0</sub> =	0.30	
$\mu_{O}$	$K(\mu_O)$	R	K	$(\mu_O)$	R	$K(\mu_O)$	R	
0.00	1.0000	0.687	1.	0000	0.369	1.0000	0.164	
0.10	1.1863	0.630	1.	0996	0.308	1.0436	0.128	
0.20	1.3144	0.591	1.	1578	0.273	1.0667	0.110	
0.30	1.4214	0.559	1.	2018	0.247	1.0832	0.098	
0.40	1.5147	0.531	1.	2371	0.227	1.0960	0.088	
0.50	1.5976	0.507	1.	2665	0.211	1.1063	0.081	
0.60	1.6722	0.485	1.	2915	0.197	1.1148	0.075	
0.70	1.7400	0.466	j l.	3131	0.186	1.1220	0.071	
0.80	1.8020	0.449	1.	3319	0.176	1.1282	0.067	
0.90	1.8591	0.433	1.	3486	0.168	1.1335	0.064	
1.00	1.9117	0.418	1.	3634	0.161	1.1382	0.061	
ω.	۹.	લ,	9 <sub>2</sub>	a <sub>3</sub>	A <sub>oo</sub>	-A.,	-A20	-A21
0.90	1.5582	0.8493	0.5868	0.4486	1.5198	0.8223	0.0295	0.1138

0.4391 0.3321

0.2817

0.3744

1.2252

1.0889

0.6351

0.5525

0.1165

0.1211

0.0203

0.0093

### 3. Solution for the Six-flux Formulation of the Slab Problem

The following solution of equations II-7a to II-8 subject to the boundary conditions  $I_1=1$  and  $I_3=0$  at Z=0 and  $I_2=I_4=0$  at  $Z=\tau_1$  is reproduced directly from a previous report to avoid errors in transcription. For this reason the symbolism is slightly different. In equations A3-1 to A3-31 T is used in place of  $\tau_1$  for the optical thickness of the dispersion,  $\Theta$  is used in place of  $\Theta_0$  for the angle the incident radiation makes with the normal to the slab, but as in this report Z refers to the normal optical distance through the slab.

All of the other symbols such as the A's, a's, b's, \( \Lambda \), p, etc. are defined as used: the definitions of these symbols apply only to equations A3-1 to A3-31 and have no relation to the definition of the symbols in the NOMENCLATURE or in the rest of this report.

$$I_{1} = \frac{A_{11}}{\Delta} e^{-p_{1}Z} + \frac{A_{12}}{\Delta} e^{-p_{1}(T-Z)} + \frac{A_{13}}{\Delta} e^{-p_{2}Z} + \frac{A_{14}}{\Delta} e^{-p_{2}(T-Z)}$$

$$A3-1$$

$$I_{2} = \frac{A_{21}}{\Delta} e^{-p_{1}Z} + \frac{A_{22}}{\Delta} e^{-p_{1}(T-Z)} + \frac{A_{23}}{\Delta} e^{-p_{2}Z} + \frac{A_{24}}{\Delta} e^{-p_{2}(T-Z)}$$
A3-2

$$I_{3} = \frac{A_{31}}{\Lambda} e^{-p_{1}Z} + \frac{A_{32}}{\Lambda} e^{-p_{1}(T-Z)} + \frac{A_{33}}{\Lambda} e^{-p_{2}Z} + \frac{A_{34}}{\Lambda} e^{-p_{2}(T-Z)}$$

and

$$I_4 = \frac{A_{41}}{\Delta} e^{-p_1 Z} + \frac{A_{42}}{\Delta} e^{-p_1 (T-Z)} + \frac{A_{43}}{\Delta} e^{-p_2 Z} + \frac{A_{44}}{\Delta} e^{-p_2 (T-Z)}$$

$$A_{3-4}$$

where:

$$\Delta = (a_1b_3 - a_3b_1)^2 - (a_2b_3 - a_3b_2)^2 e^{-2p_1T} - (a_1b_4 - a_4b_1)^2 e^{-2p_2T}$$

$$+2(a_1b_2 - a_2b_1)(a_3b_4 - a_4b_3)e^{-(p_1+p_2)T} + (a_2b_4 - a_4b_2)^2 e^{-2(p_1+p_2)T}$$

$$A3-5$$

$$A_{11} = a_1b_3(a_1b_3-a_3b_1)-a_1b_4(a_1b_4-a_4b_1)e^{-2p_2T}+a_1b_2(a_3b_4-a_4b_3)e^{-(p_1+p_2)T}$$

$$A_{21} = a_2b_3(a_1b_3-a_3b_1)-a_2b_4(a_1b_4-a_4b_1)e^{-2p_2T}+a_2b_2(a_3b_4-a_4b_3)e^{-(p_1+p_2)T}$$

$$A_{3-7}$$

$$A_{31} = b_1b_3(a_1b_3-a_3b_1)-b_1b_4(a_1b_4-a_4b_1)e^{-2p_3T}+b_1b_2(a_3b_4-a_4b_3)e^{-(p_1+p_2)T}$$

$$A_{41} = b_2b_3(a_1b_3-a_3b_1)-b_2b_4(a_1b_4-a_4b_1)e^{-2p_2T}+b_2^2(a_3b_4-a_4b_3)e^{-(p_1+p_2)T}$$

$$A_{12} = -a_2b_3(a_2b_3 - a_3b_2)e^{-p_1T} - a_2b_1(a_3b_4 - a_4b_3)e^{-p_2T} + a_2b_1(a_2b_4 - a_4b_2)e^{-(p_1 + 2p_2)T} \quad A5-10$$

$$A_{22} = -a_1b_3(a_2b_3 - a_3b_2)e^{-p_1T} - a_1b_1(a_3b_4 - a_4b_3)e^{-p_2T} + a_3b_4(a_2b_4 - a_4b_2)e^{-(p_1 + 2p_2)T} \quad A5-11$$

$$A_{32} = -b_2b_3(a_2b_3 - a_3b_2)e^{-p_1T} - b_2b_1(a_3b_4 - a_4b_3)e^{-p_2T} + b_2b_4(a_2b_4 - a_4b_2)e^{-(p_1 + 2p_2)T} \quad A5-12$$

$$A_{42} = -b_1b_3(a_2b_3 - a_3b_2)e^{-p_1T} - b_1^2(a_3b_4 - a_4b_3)e^{-p_2T} + b_1b_4(a_2b_4 - a_4b_2)e^{-(p_1 + 2p_2)T} \quad A5-15$$

$$A_{13} = -a_3b_1(a_1b_3 - a_3b_1) + a_3b_2(a_2b_3 - a_3b_2)e^{-2p_1T} + a_3b_4(a_1b_2 - a_2b_1)e^{-(p_1 + p_2)T} \quad A5-15$$

$$A_{23} = -a_4b_1(a_1b_3 - a_3b_1) + a_4b_2(a_2b_3 - a_3b_2)e^{-2p_1T} + a_4b_4(a_1b_2 - a_2b_1)e^{-(p_1 + p_2)T} \quad A5-16$$

$$A_{23} = -b_3b_1(a_1b_3 - a_3b_1) + b_3b_2(a_2b_3 - a_3b_2)e^{-2p_1T} + b_3b_4(a_1b_2 - a_2b_1)e^{-(p_1 + p_2)T} \quad A5-16$$

$$A_{43} = -b_4b_1(a_1b_3 - a_3b_1) + b_4b_2(a_2b_3 - a_3b_2)e^{-2p_1T} + b_4^2(a_1b_2 - a_2b_1)e^{-(p_1 + p_2)T} \quad A5-16$$

$$A_{43} = -b_4b_1(a_1b_3 - a_3b_1) + b_4b_2(a_2b_3 - a_3b_2)e^{-2p_1T} + b_4^2(a_1b_2 - a_2b_1)e^{-(p_1 + p_2)T} \quad A5-16$$

$$A_{44} = -a_4b_3(a_1b_2 - a_2b_1)e^{-p_1T} + a_4b_1(a_1b_4 - a_4b_1)e^{-p_2T} - a_4b_2(a_2b_4 - a_4b_2)e^{-(2p_1 + p_2)T} \quad A5-18$$

$$A_{24} = -a_4b_3(a_1b_2 - a_2b_1)e^{-p_1T} + a_4b_1(a_1b_4 - a_4b_1)e^{-p_2T} - a_4b_2(a_2b_4 - a_4b_2)e^{-(2p_1 + p_2)T} \quad A5-19$$

$$A_{34} = -b_4b_3(a_1b_2 - a_2b_1)e^{-p_1T} + a_4b_1(a_1b_4 - a_4b_1)e^{-p_2T} - a_4b_2(a_2b_4 - a_4b_2)e^{-(2p_1 + p_2)T} \quad A5-19$$

$$A_{34} = -b_4b_3(a_1b_2 - a_2b_1)e^{-p_1T} + b_4b_1(a_1b_4 - a_4b_1)e^{-p_2T} - b_4b_2(a_2b_4 - a_4b_2)e^{-(2p_1 + p_2)T} \quad A5-20$$

$$A_{34} = -b_4b_3(a_1b_2 - a_2b_1)e^{-p_1T} + b_4b_1(a_1b_4 - a_4b_1)e^{-p_2T} - b_4b_2(a_2b_4 - a_4b_2)e^{-(2p_1 + p_2)T} \quad A5-20$$

in which,

$$p_1 = \left\{ \frac{1}{2} \left[ 1 - \sqrt{1 - \sin^2 2 \, \Theta \, \left[ \frac{(C_1 - C_2)^2 - 4C_3^2}{(C_1 - C_2)^2} \right]} \right]^{1/2} \sqrt{C_1^2 - C_2^2} \, \sec \, \Theta \, \csc \, \Theta \right.$$

$$A3-22$$

 $A_{44} = -b_3^2(a_1b_2 - a_2b_1)e^{-p_1T} + b_3b_1(a_4b_4 - a_4b_1)e^{-p_2T} - b_3b_2(a_2b_4 - a_4b_2)e^{-(2p_1 + p_2)T}$ 

$$p_{2} = \left\{ \frac{1}{2} \left[ 1 + \sqrt{1 - \sin^{2} 2 \theta \left[ \frac{(C_{1} - C_{2})^{2} - 4C_{3}^{2}}{(C_{1} - C_{2})^{2}} \right]} \right]^{1/2} \sqrt{C_{1}^{2} - C_{2}^{2}} \text{ sec } \theta \text{ csc } \theta \text{ A3-23} \right]$$

$$a_1 = \frac{(C_1 + C_2) \sec \theta + p_1}{(C_1 + C_2) \sec \theta}$$

$$A3-24$$

$$a_2 = \frac{(C_1 + C_2) \sec \theta - p_1}{(C_1 + C_2) \sec \theta}$$

$$A3-25$$

$$a_3 = \frac{(C_1 + C_2) \sec \theta + p_2}{(C_1 + C_2) \sec \theta}$$
A3-26

$$a_4 = \frac{(C_1 + C_2) \sec \theta - p_2}{(C_1 + C_2) \sec \theta}$$

$$A3-27$$

$$b_1 = \frac{(C_1^2 - C_2^2) \sec^2 \theta - p_1^2}{2(C_1 + C_2) C_3 \sec^2 \theta} \frac{(C_1 + C_2) \csc \theta + p_1}{(C_1 + C_2) \csc \theta}$$
A3-28

$$b_2 = \frac{(C_1^2 - C_2^2) \sec^2 \theta - p_1^2}{2(C_1 + C_2) C_3 \sec^2 \theta} \frac{(C_1 + C_2) \csc \theta - p_1}{(C_1 + C_2) \csc \theta}$$
A3-29

$$b_3 = \frac{(C_1^2 - C_2^2) \sec^2\theta - p_2^2}{2(C_1 + C_2) C_3 \sec^2\theta} \frac{(C_1 + C_2) \csc\theta + p_2}{(C_1 + C_2) \csc\theta}$$
A3-30

and

$$b_4 = \frac{(C_1^2 - C_2^2) \sec^2 - \theta p_2^2}{2(C_1 + C_2) C_3 \sec^2 \theta} \frac{(C_1 + C_2) \csc \theta - p_2}{(C_1 + C_2) \csc \theta}$$
A3-31

#### APPENDIX B

## Tables of Computed Functions

The tabulated results obtained from the computer are presented in this appendix.

In the first portion of the appendix the complete set of  $\psi_{\boldsymbol{\ell}}^{\boldsymbol{m}}(\boldsymbol{\mu})$  and  $\phi_{\boldsymbol{\ell}}^{\boldsymbol{m}}(\boldsymbol{\mu})$  functions, the integrated reflectance R, the diffuse portion of the integrated transmission  $T_D$ , and the total integrated transmission T are given as a function of  $\boldsymbol{\mu}$  for each of the 36 sets of parameters used. The results are presented as one long table (Table B-O) which is subdivided into 36 problems. The values of the parameters corresponding to each problem are listed in Table I-2.

In the last portion of the appendix (Tables B-1 to B-5) tabular results are given which compare the exact solution with results obtained from approximate models.

Table B-O The  $\psi_{\ell}^{m}$  and  $\varphi_{\ell}^{m}$  Functions, Integrated Reflectances, Diffuse Portion of the Integrated Transmission, and Total Integrated Transmission for 36 sets of Parameters (Listed in Table I-2).

PROBLEM 1 TAU = 0.050 ALBEDO FOR SINGLE SCATTERING = 0.90

A0 = 1.0000 A1 = 0.0000 A2 = 0.0000 F = 0.500 P = 0.3333

NUMBER OF ITERATIONS WAS 2

PROBLEM 2 TAU = 0.250 ALBEDO FOR SINGLE SCATTERING = 0.90

A0 = 1.0000 A1 = 0.0000 A2 = 0.0000 F = 0.500 P = 0.3333

NUMBER OF ITERATIONS WAS 4

μ	$\psi_{O}^{O}$	$\phi_{\circ}^{\circ}$	R	$\mathtt{T}_{\mathrm{D}}$	T
0.04691	1.0797	0.0467	0.498	0.360	0.365
0.23077	1.1993	0.4990	0.300	0.268	0.607
0.50000	1.2413	0.8317	0.173	0.161	0.768
0.76923	1.2611	0.9501	0.125	0.104	0.827
0.95309	1.2694	1.0001	0.105	0.085	0.854

PROBLEM 3 TAU = 0.500 ALBEDO FOR SINGLE SCATTERING = 0.90

A0 = 1.0000 A1 = 0.0000 A2 = 0.0000 F = 0.500 P = 0.3333

NUMBER OF ITERATIONS WAS 4

μ	$\psi_{O}^{O}$	$\phi_{o}^{o}$	R	$\mathtt{T}_{D}$	T
0.04691	1.0842	0.0275	0.548	0.286	0.286
0.23077	1.2530	0.2636	0.417	0.312	0.426
0.50000	1.3465	0.6497	0.279	0.239	0.607
0.76923	1.3896	0.8462	0.209	0.171	0.693
0.95309	1.4089	0.9249	0.179	0.136	0.728

μ	$\psi_{0}^{O}$	$\phi_{o}^{o}$	R	$\mathtt{T}_{\mathrm{D}}$	T
0.04691	1.0888	0.0165	0.598	0.197	0.197
0.23077	1.2915	0.1231	0.504	0.258	0.271
0.50000	1.4541	0.4012	0.394	0.263	0.398
0.76923	1.5427	0.6476	0.316	0.225	0 • 498
0.95309	1.5852	0.7553	0.279	0.186	0.536

PROBLEM 5 TAU = 1.500 ALBEDO FOR SINGLE SCATTERING = 0.90

A0 = 1.0000 A1 = 0.0000 A2 = 0.0000 F = 0.500 P = 0.3333

NUMBER OF ITERATIONS WAS 9

μ	$\psi_{o}^{o}$	$\phi_{\rm o}^{\rm o}$	R	$\mathtt{T}_{\mathrm{D}}$	T
0.04691	1.0910	0.0123	0.622	0.152	0.152
0.23077	1.3063	0.0845	0.537	0.205	0.207
0.50000	1.5009	0.2769	0.444	0.248	0.298
0.76923	1.6202	0.5158	0.370	0.253	0.395
0.95309	1.6774	0.6565	0.330	0.242	0.449

PROBLEM 6 TAU = 2.000 ALBEDO FOR SINGLE SCATTERING = 0.90

A0 = 1.0000 A1 = 0.0000 A2 = 0.0000 F = 0.500 P = 0.3333NUMBER OF ITERATIONS WAS 9

μ	ΨΟ	$\phi_{o}^{o}$	R	$\mathtt{T}_{\mathtt{D}}$	T
0.04691	1.0919	0.0098	0.632	0.129	0.129
0.23077	1.3122	0.0665	0.552	0.174	0.174
0.50000	1.5204	0.2145	0.465	0.233	0.251
0.76923	1.6567	0.4370	0.395	0.271	0.345
0.95309	1.7224	0.6056	0.355	0.290	0.413

```
PROBLEM 7 TAU = INFINITY ALBEDO FOR SINGLE SCATTERING = 0.90 A0 = 1.0000 A1 = 0.0000 A2 = 0.0000 F = 0.500 P = 0.3333 NUMBER OF ITERATIONS WAS 16
```

μ	$\psi_{O}^{O}$	ذ	R	$\mathtt{T}_{\mathrm{D}}$	${f T}$
0.04691	1.0936	0.0000	0.654	0.000	0.000
0.23077	1.3236	0.0000	0.581	0.000	0.000
0.50000	1.5554	0.0000	0.507	0.000	0.000
0.76923	1.7289	0.0000	0.452	0.000	0.000
0.95309	1.8259	0.0000	0.422	0.000	0.000

μ	$\psi_{O}^{O}$	$\phi_{o}^{o}$	R	$\mathtt{T}_{\mathrm{D}}$	T
0.04691	1.0355	0.3696	0.1931	0.1817	0.526
0.23077	1.0477	0.8694	0.0529	0.0665	0.871
0.50000	1.0516	0.9557	0.0273	0.0271	0.932
0.76923	1.0527	0.9877	0.0182	0.0173	0.954
0.95309	1.0531	0.9991	0.0148	0.0132	0.962

PROBLEM 9 TAU = 0.500 ALBEDO FOR SINGLE SCATTERING = 0.600

A0 = 1.0000 A1 = 0.0000 A2 = 0.0000 F = 0.500 P = 0.3333

NUMBER OF ITERATIONS WAS 5

μ	$\psi_{O}^{O}$	$\phi_{o}^{o}$	R	${ m T}_{ m D}$	T
0 • 04691	1.0507	0.0145	0.318	0.154	0.154
0.23077	1.1448	0.1939	0.232	0.168	0.283
0.50000	1.1958	0.5221	0.154	0.134	0.502
0.76923	1.2186	0.7046	0.114	0.101	0.623
0.95309	1.2283	0.7827	0.097	0.083	0.675

PROBLEM 10 TAU = INFINITY ALBEDO FOR SINGLE SCATTERING = 0.60

A0 = 1.0000 A1 = 0.0000 A2 = 0.0000 F = 0.500 P = 0.3333

NUMBER OF ITERATIONS WAS 7

μ	$\psi_{O}^{O}$	ø°o	R	$\mathtt{T}_{\mathrm{D}}$	T
0.04691	1.0520	0.0000	0.335	0.000	0.000
0.23077	1.1587	0.0000	0.267	0.000	0.000
0.50000	1.2456	0.0000	0.212	0.000	0.000
0.76923	1.3009	0.0000	0.177	0.000	0.000
0.95309	1.3288	0.0000	0.158	0.000	0.000

PROBLEM 12 TAU = 0.500 ALBEDO FOR SINGLE SCATTERING = 0.30 AO = 1.0000 A1 = 0.0000 A2 = 0.0000 F = 0.500 P = 0.3333 NUMBER OF ITERATIONS WAS 4

μ	$\psi_{\mathcal{O}}^{\mathcal{O}}$	ø°	R	$^{\mathrm{T}}\mathrm{D}$	T
0.04691	1.0232	0.0059	0.1416	0.0636	0.0636
0 • 23077	1.0632	0.1469	0.0999	0.0696	0.1842
0.50000	1.0844	0.4323	0.0658	0.0569	0.4248
0.76923	1.0936	0.6000	0.0484	0.0439	0.5659
0.95309	1.0974	0.6743	0.0410	0.0366	0.6284

PROBLEM 11 TAU = 0.050 ALBEDO FOR SINGLE SCATTERING = 0.30 AO = 1.0000 A1 = 0.0000 A2 = 0.0000 F = 0.500 P = 0.3333 NUMBER OF ITERATIONS WAS 2

μ	$\psi_{O}^{O}$	ذ	R	$\mathtt{T}_{\mathrm{D}}$	T
0.04691	1.0172	0.3564	0.0935	0.0874	0.431
0.23077	1.0237	0.8365	0.0265	0.0337	0.839
0.50000	1.0252	0.9307	0.0132	0.0139	0.918
0.76923	1.0256	0.9624	0.0087	0.0087	0.945
0.95309	1.0258	0.9740	0.0071	0.0067	0.956

PROBLEM 13 TAU = INFINITY ALBEDO FOR SINGLE SCATTERING = 0.30 A0 = 1.0000 A1 = 0.0000 A2 = 0.0000 F = 0.500 P = 0.3333 NUMBER OF ITERATIONS WAS 5

μ	$\psi_{O}^{O}$	$\phi_{\mathrm{o}}^{\mathrm{o}}$	R	$\mathtt{T}_{\mathrm{D}}$	T
0.04691	1.0233	0.0000	0.1438	0.0000	0.0000
0.23077	1.0663	0.0000	0.1079	0.0000	0.0000
0.50000	1.0975	0.0000	0.0817	0.0000	0.0000
0.76923	1.1159	0.0000	0.0663	0.0000	0.0000
0.95309	1.1248	0.0000	0.0588	0.0000	0.0000

PROBLEM 14 TAU = 0.050 ALBEDO FOR SINGLE SCATTERING = 0.90

A0 = 1.0000 A1 = 0.0000 A2 = 0.5000 F = 0.500 P = 0.400

NUMBER OF ITERATIONS FOR M = 0.1. AND 2 WERE 2.2. AND 2

μ	$\psi_{O}^{O}$	ψĺ	ψ <sup>0</sup> 2	ψ <u>1</u>	$\psi_2^1$	ψ2 2	$\phi_{o}^{o}$	$\phi_1^{\rm o}$
0.04691	1.0590	0.0000	-0.5158	0.0000	0.1409	3.0162	0.3864	0.0000
0.23077	1.0771	0.0000	-0.4452	0.0000	0.6760	2.8703	0.9132	0.0000
0.50000	1.0799	0.0000	-0.1493	0.0000	1.3040	2.2751	0.9835	0.0000
0.76923	1.0756	0.0000	0.3670	0.0000	1.4802	1.2387	1.0093	0.0000
0.95309	1.0705	0.0000	0.8457	0.0000	0.8688	0.2799	1.0155	0.0000

μ	ø2	$\phi_1^{\perp}$	$\phi_2^{\perp}$	φ <sup>2</sup>	R	$\mathtt{T}_{\mathrm{D}}$	T
0.04691	-0.1817	0.0000	0.0488	1.0452	0.299	0.283	0.627
0.23077	-0.3774	0.0000	0.5447	2.3280	0.077	0.100	0.905
0.50000	-0.1370	0.0000	1.1801	2.0618	0.042	0.038	0.943
0.76923	0.3433	0.0000	1.3871	1.1615	0.028	0.026	0.963
0.95309	0.8015	0.0000	0.8247	0.2639	0.023	0.020	0.969

μ	$\psi_{\mathrm{O}}^{\mathrm{O}}$	ψÎ	ψ <mark>0</mark>	$\psi_{\mathtt{l}}^{\mathtt{l}}$	1 ¥2	2 \\2	$\phi_{o}^{o}$	$\phi_1^{\rm o}$
0.04691	1.0857	0.0000	-0.5214	0.0000	0.1410	3.0222	0.0487	0.0000
0.23077	1.2102	0.0000	-0.4690	0.0000	0.6807	2.8997	0.5063	0.0000
0.50000	1.2461	0.0000	-0.1759	0.0000	1.3161	2.3058	0.8379	0.0000
0.76923	1.2525	0.0000	0.3446	0.0000	1.4954	1.2569	0.9439	0.0000
0.95309	1.2471	0.0000	0.8287	0.0000	0.8781	0.2822	0.9849	0.0000
μ	ø <sub>2</sub>	$\phi_1^1$	ø <sub>2</sub>	ø <sub>2</sub>	R	$T_{\mathrm{D}}$	T	
		0.0000	0.0010 0.2340	0.0257	0.500	0.356	0.361	
0.50000	-0.1737 -0.1237	0.0000	0.2340	1.0025 1.4149	0•299 0•173	0.267 0.163	0.606	
0.76923	0.2418	0.0000	1.0843	0.9141		0.103	0.770	
0.95309	0.6315	0.0000	0.6787	0.2182	0•125 0•105	0.107	0.830 0.859	
					0 -	· <del>-</del>	• • • •	

μ	ο Ψο	ο Ψ1	ψ2 0	l Vl	$\psi^1_2$	2 \\psi_2	$\phi_{\circ}^{\circ}$	$\phi_\mathtt{l}^\mathrm{o}$
0.04691	1.0908	0.0000	-0.5217	0.0000	0.1411	3.0222	0.0284	0.0000
0.23077	1.2680	0.0000	-0.4723	0.0000	0.6822	2.9054	0.2680	0.0000
0.50000	1.3570	0.0000	-0.1809	0.0000	1.3226	2.3157	0.6554	0.0000
0.76923	1.3863	0.0000	0.3414	0.0000	1.5049	1.2638	0.8365	0.0000
0.95309	1.3909	0.0000	0.8285	0.0000	0.8841	0.2839	0.9033	0.0000
	40	<sub>d</sub> l	<sub>4</sub> 1	ø <sub>2</sub>	_	_		
μ	$\phi_2^{\rm o}$	$\emptyset_{\mathtt{l}}$	$\phi_2^{\perp}$	φ <sub>2</sub>	R	$^{\mathrm{T}}\mathrm{D}$	${f T}$	
0.04691	-0.0018	0.0000	0.0002	0.0049	0.553	0.280	0.280	
0.23077	-0.0584	0.0000	0.0831	0.3492	0.419	0.308	0.422	
0.50000	-0.0838	0.0000	0.4960	0.8722	0.281	0.236	0.604	
0.76923	0.1671	0.0000	0.7948	0.6696	0.212	0.170	0.692	
0.95309	0.4789	0.0000	0.5304	0.1700	0.183	0.138	0.729	

```
PROBLEM 17 TAU = 1.000 ALBEDO FOR SINGLE SCATTERING = 0.90

AO = 1.0000 A1 = 0.0000 A2 = 0.5000 F = 0.500 P = 0.400

NUMBER OF ITERATIONS FOR M = 0, 1, AND 2 WERE 6, 3, AND 3
```

	0	ο Ψ1	φ2 Ψ2	1 Ψ1	1 \\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\	2 \\2	$\phi_{o}^{o}$	$\phi_1^{\rm o}$
μ <b>0•04691</b>	ψ <sub>0</sub> 1.0949		-0.5215	0.0000	0.1411	3.0223	0.0162	0.6000
0.23077	1.3044		-0.4711	0.0000	0.6826	2.9065	0.1226	0.0000
0.50000	1.4614		-0.1772	0.0000	1.3263	2.3196	0.4017	0.0000
0.76923	1.5334	0.0000	0.3485	0.0000	1.5116	1.2671	0.6412	0.0000
0.95309	1.5594	0.0000	0.8387	0.0000	0.8888	0.2848	0.7341	0.0000
μ	$\phi_2^{\rm o}$	$\phi_1^1$	$\phi_2^1$	$\phi_2^2$	R	$\mathtt{T}_{\mathrm{D}}$	T	
0.04691	0.0010	0.0000	0.0001	0.0010	0.601	0.190	0.190	
-	-0.0014	0.0000	0.0122	0.0459	0.504	0.253	0.266	
	-0.0205	0.0000	0.1900	0.3272	0.394	0.259	0.394	
0.76923	0.0917	0.0000	0.4247	0.3550	0.317	0.221	0.493	
0.95309	0.2876	0.0000	0.3220	0.1019	0.282	0.180	0.530	

μ 0.04691 0.23077 0.50000 0.76923 0.95309	1.0968 1.3178 1.5056 1.6080 1.6491	0.0000	ψ <sub>2</sub> -0.5213 -0.4699 -0.1731 0.3556 0.8476	1 0.0000 0.0000 0.0000 0.0000 0.0000	1 0.1411 0.6826 1.3270 1.5132 0.8901	2 \\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\	\$\phi_0^\circ\$ 0.0113 0.0800 0.2678 0.4952 0.6242	\$\phi_1^\circ\$ 0.0000 0.0000 0.0000 0.0000 0.0000
μ 0•04691 0•23077 0•50000 0•76923 0•95309	\$\psi_2^\circ\$ 0.0013 0.0066 0.0076 0.0656 0.1937	\$\psi_1\$ 0.0000 0.0000 0.0000 0.0000 0.0000	\$\psi_2^1\$ 0.0001 0.0027 0.0737 0.2260 0.1945	\$\psi_2^2\$ 0.0003 0.0076 0.1223 0.1870 0.0608	R 0.623 0.537 0.443 0.371 0.334	T <sub>D</sub> 0.142 0.194 0.238 0.241 0.226	T 0.142 0.196 0.288 0.383 0.433	

PROBLEM 19 TAU = 2.000 ALBEDO FOR SINGLE SCATTERING = 0.90 A0 = 1.0000 A1 = 0.0000 A2 = 0.5000 F = 0.500 P = 0.400 NUMBER OF ITERATIONS FOR M = 0.1, AND 2 WERE 11, 3, AND 3

μ	$\psi_{O}^{O}$	ο Ψ <u>1</u>	о ¥2	$\psi_{1}^{1}$	1 Ψ2	2 \\2	$\phi_{o}^{o}$	$\phi_1^\circ$
0.04691	1.0981	0.0000	-0.5212	0.0000	0.1411	3.0223	0.0092	0.0000
0.23077	1.3261	0.0000	-0.4692	0.0000	0.6826	2.9066	0.0632	0.0000
0.50000	1.5304	0.0000	-0.1709	0.0000	1.3271	2.3201	0.2048	0.0000
0.76923	1.6521	0.0000	0.3598	0.0000	1.5137	1.2678	0.4145	0.0000
0.95309	1.7034	0.0000	0.8528	0.0000	0.8905	0.2850	0.5620	0.0000
μ	ø <sub>2</sub> 0	$\phi_1^1$	ø <sub>2</sub>	ø <sub>2</sub>	R	$\mathtt{T}_{\mathtt{D}}$	Т	
0.04691	0.0012	0.0000	0.0000	0.0001	0.636	0.118	0.118	
0.23077	0.0077	0.0000	0.0009	0.0018	0.554	0.161	0.162	
0.50000	0.0177	0.0000	0.0291	0.0458	0.468	0.215	0.233	
0.76923	0.0562	0.0000	0.1200	0.0982	0.400	0.248	0.322	
0.95309	0.1438	0.0000	0.1170	0.0361	0.363	0.259	0.382	

PROBLEM 20 TAU = INFINITY ALBEDO FOR SINGLE SCATTERING = 0.90

A0 = 1.0000 A1 = 0.0000 A2 = 0.5000 F = 0.500 P = 0.400NUMBER OF ITERATIONS FOR M = 0.1, AND 2 WERE 16.3, AND 3

μ 0 • 04691 0 • 23077 0 • 50000 0 • 76923 0 • 95309	ψ <sub>0</sub> ° 1.0996 1.3362 1.5623 1.7192 1.7999		$\psi_{2}^{\circ}$ -0.5210 -0.4677 -0.1661 0.3701 0.8677	ψ <sub>1</sub> 0.0000 0.0000 0.0000 0.0000 0.0000	ψ <sub>2</sub> 0•1411 0•6826 1•3271 1•5138 0•8906	2 3.0223 2.9066 2.3207 1.2679 0.2850	\$\phi_0^\circ\$ 0.0000 0.0000 0.0000 0.0000 0.0000	\$\psi_1^\circ\$ 0.0000 0.0000 0.0000 0.0000 0.0000
μ 0•04691 0•23077 0•50000 0•76923 0•95309	\$\phi_2^0\$ 0.0000 0.0000 0.0000 0.0000 0.0000	\$\psi_1^1\$ 0.0000 0.0000 0.0000 0.0000 0.0000	\$\psi_2^1\$ 0.0000 0.0000 0.0000 0.0000 0.0000	\$\psi_2\$ 0.0000 0.0000 0.0000 0.0000 0.0000	R 0.656 0.580 0.507 0.453 0.424	T <sub>D</sub> 0.000 0.000 0.000 0.000 0.000	T 0.000 0.000 0.000 0.000 0.000	

PROBLEM 21 TAU = 1.000 ALBEDO FOR SINGLE SCATTERING = 0.90

A0 = 1.0000 A1 = 1.0000 A2 = 1.8165 F = 0.750 P = 0.5504

NUMBER OF ITERATIONS FOR M = 0, 1, AND 2 WERE 4, 11, AND 5

ψ2 Ψ2  $\phi_{o}^{o}$ 1  $\phi_1^{\circ}$ ψĺ Ψl Ψ2  $\psi_{o}$ 0.0164 0.0089 0.0211 -0.5329 3.1069 1.0266 0.1242 0.04691 1.1075 0.1332 0.0665 1.3101 0.1331 -0.5012 0.6458 3.1140 0.23077 1.0228 1.3954 0.3549 -0.2024 0.50000 0.8819 2.5446 0.4580 0.2319 1.3221 1.4059 0.7536 1.3537 0.6124 0.3586 0.6120 1.5497 0.4677 0.76923 0.7956 0.8892 0.2757 0.9220 0.3175 0.8755 0.6381 0.95309 1.2821

 $\phi_1^1$  $\phi_2^2$  $\phi_2^{o}$ Т R  $T_{D}$ μ 0.0006 0.0044 0.0052 0.203 0.203 0.0027 0.563 0.04691 0.23077 -0.0034 0.0788 0.290 0.303 0.0397 0.0494 0.438 0.4097 0.303 0.337 0.472 0.50000 -0.0162 0.2019 0.2957 0.76923 0.3034 0.4311 0.213 0.333 0.605 0.1277 0.5826 0.95309 0.3629 0.2116 0.4639 0.1220 0.173 0.307 0.657

PROBLEM 22 TAU = 1.000 ALBEDO FOR SINGLE SCATTERING = 0.90

A0 = 1.0000 A1 = 1.0000 A2 = 1.0490 F = 0.750 P = 0.4821

NUMBER OF ITERATIONS FOR M = 0, 1, AND 2 WERE 5, 5, AND 4

٦ 1 ψ<sub>1</sub> 0 Ψ2 Ψ2 Ψl u 0.0207 -0.5263 1.0276 3.0559 1.0984 0.0185 0.0102 0.1228 0.04691 1.2944 2.9865 0.1422 0.0727 0.1295 -0.4847 1.0351 0.6288 0.23077 1.3958 0.9181 1.2706 0.4702 0.2396 0.3481 -0.1923 2.4050 0.50000 1.3865 1.3195 0.7712 0.4682 0.6065 0.3426 0.6611 1.4793 0.76923 0.7926 0.8450 0.3063 0.8782 0.2971 0.9051 0.6317 0.95309 1.3458

μ	ø°2	$\phi_1^{\perp}$	$\phi_{\perp}^{2}$	$\phi_2^2$	${\tt R}$	$^{\mathrm{T}}\mathrm{D}$	T
0.04691	0.0011	0.0023	0.0036	0∙0်24	0.564	0.222	0.222
0.23077	-0.0002	0.0334	0.0392	0.0574	0.445	0.305	0.318
0.50000	-0.0134	0.1777	0.2572	0.3569	0.310	0.341	0.476
0.76923	0.1145	0.2583	0.5119	0.3825	0.218	0.325	0.598
0.95309	0.3290	0.1647	0.3853	0.1092	0.175	0.295	0.645

PROBLEM 23 TAU = 1.000 ALBEDO FOR SINGLE SCATTERING = 0.90A0 = 1.0000 A1 = 1.0000 A2 = 0.0000 F = 0.750 P = 0.3889NUMBER OF ITERATIONS FOR M = 0 AND 1 WERE 6 AND 4

μ	ο Ψ <sub>Ο</sub>	ψl	о \$2	1 \\psi_1	$\psi_{\mathcal{O}}^{1}$	2 \\psi_2	$\phi_{\circ}^{\circ}$	ø <sub>1</sub> o
0.04691 0.23077 0.50000 0.76923 0.95309	1.0856 1.2689 1.3856 1.4151 1.4126	0.0208 0.1282 0.3445 0.6033 0.7913	0.0000 0.0000 0.0000 0.0000 0.0000	1.0288 1.0498 0.9607 0.7185 0.3424	0.0000 0.0000 0.0000 0.0000 0.0000	0.0000 0.0000 0.0000 0.0000 0.0000	0.0198 0.1464 0.4720 0.7817	0.0112 0.0767 0.2433 0.4675 0.6300
	0	٦	-	•				

μ	$\phi_2^{\circ}$	$\phi_1^\perp$	$\phi_2^1$	$\phi_2^2$	R	$\mathtt{T}_{\mathrm{D}}$	T
0.04691	0.0000	0.0018	0.0000	0.0000	0.559	0.238	0.238
0.23077	0.0000	0.0269	0.0000	0.0000	0.448	0.316	0.329
0.50000	0.0000	0.1540	0.0000	0.0000	0.315	0.345	0.480
0.76923	0.0000	0.2180	0.0000	0.0000	0.220	0.324	0.596
0.95309	0.0000	0.1312	0.0000	0.0000	0.174	0.296	0.646

0.23077 0.50000 0.76923	1.0745 1.0643 1.0438	0.2166 0.4890 0.7628	-0.5187 -0.4485 -0.1510 0.3689	1.0385 1.0213 0.9086 0.6676	0.6486 1.2790 1.4635	3.0397 2.9005 2.3066 1.2528	0.9285 1.0007 1.0309	0.4807 0.7542
0.95309	1.0262	0.9499	0.8511	0.3151	0.8916	0.2812	1.0390	

μ	$\phi_2^{\rm o}$	$\phi_1^1$	$\phi_2^{\perp}$	$\phi_2^2$	R	$\mathtt{T}_{\mathrm{D}}$	T
0.23077 0.50000 0.76923	-0.1837 -0.3821 -0.1373 0.3473 0.8097	0.8605 0.8387 0.6440	1.2127 1.4139	1.0601 2.3706 2.0874 1.1753 0.2670	0.287 0.060 0.023 0.009 0.004	0.295 0.117 0.055 0.045 0.040	0.639 0.922 0.960 0.982 0.989

PROBLEM 25 TAU = 0.250 ALBEDO FOR SINGLE SCATTERING = 0.90

A0 = 1.0000 A1 = 1.7321 A2 = 1.0000 F = 0.9330 P = 0.4821

NUMBER OF ITERATIONS FOR M = 0, 1, AND 2 WERE 6, 10, AND 3

ш	ψ <sub>O</sub>	ο Ψ1	ο Ψ2	1 Ψ1	ΨŽ	2 \\\\2	$\phi_{0}^{\circ}$	$\phi_{\mathtt{l}}^{\mathrm{o}}$
0.04691								
0.23077								
0.50000								
0.76923	1.1337	0.7363	0.3424	0.7104	1.4325	1.2914	1.0700	0.7177
0.95309	1.0806	0.9348	0.8344	0.3328	0.8507	0.2901	1.1597	0.9336

μ	$\phi_{2}^{\circ}$	$\phi_1^1$	$\phi_2^1$	$\phi_2^2$	R	$\mathtt{T}_{\mathrm{D}}$	Т
0.04691	-0.0116	0.0303	0.0248	0.0378	0.466	0.394	0.396
0.23077	-0.1771	0.4320	0.3258	1.0475	0.238	0.335	0.669
0.50000	-0.1199	0.6730	0.9280	1.4696	0.097	0.256	0.853
0.76923	0.2664	0.5879	1.1875	0.9455	0.043	0.226	0.933
0.95309	0.6743	0.3072	0.7458	0.2255	0.018	0.230	0.979

μ	ψO	ο Ψ1	o V2	1 Ψ1	1 \\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\	ψ <sub>2</sub>	$\phi_{o}^{o}$	$\phi_1^{\circ}$
0.04691	1.0926	0.0235	-0.5262	1.0532	0.1057	3.0527	0.0337	0.0156
0.23077	1.2466	0.1545	-0.4854	1.1044	0.5701	2.9763	0.3067	0.1173
0.50000	1.2615	0.4199	-0.1968	1.0005	1.1953	2.3878	0.7653	0.3632
0.76923	1.1958	0.7115	0.3311	0.7317	1.4104	1.3067	1.0156	0.6422
0.95309	1.1267	0.9164	0.8262	0.3421	0.8419	0.2939	1.1318	0.8434

и	Ø2	$\phi_1^1$	$\phi_2^1$	$\phi_2^2$	R	${f T}_{f D}$	T
0.04691	-0.0023	0.0014	0.0160	0.0098	0.500	0.331	0.331
0.23077	-0.0588	0.1901	0.1720	0.3764	0.333	0.392	0.507
0.50000	-0.0779	0.4782	0.6520	0.9228	0.163	0.355	0.723
0.76923	0.1919	0.4925	0.9454	0.7034	0.078	0.311	0.833
0.95309	0.5234	0.2768	0.6291	0.1781	0.041	0.293	0.885

PROBLEM 27 TAU = 1.000 ALBEDO FOR SINGLE SCATTERING = 0.90 A0 = 1.0000 A1 = 1.7321 A2 = 1.0000 F = 0.9330 P = 0.4821 NUMBER OF ITERATIONS FOR M = 0.1, AND 2 WERE 5, 10, AND 4

μ	Ψο	ψĺ	ψS	ψ <u>1</u>	ψ2 -	ψ2 Ψ2	$\phi_{\circ}^{\circ}$	$\phi_1^{\rm o}$
0.04691	1.0951	0.0223	-0.5262	1.0537	0 • 1051	3.0528	0.0212	0.0119
0.23077	1.2736	0.1408	-0.4859	1.1107	0.5618	2.9790	0.1630	0.0852
0.50000	1.3347	0.3832	-0.1988	1.0178	1.1727	2.3969	0.5354	0.2784
0.76923	1.2749	0.6723	0.3281	0.7470	1.3913	1.3145	0.8972	0.5418
0.95309	1.1966	0.8816	0.8239	0.3493	0.8334	0.2959	1.0802	0.7334
μ	$\phi_{2}^{\circ}$	$\phi_1^1$	$\phi_{\mathbf{o}}^{\mathbf{l}}$	ø2	R	$\mathtt{T}_{D}$	T	

0.04691  $0.0\overline{0}16$   $0.0\overline{0}53$   $0.0\overline{0}80$   $0.0\overline{0}23$  0.5300.259 0.259 0.23077 0.0028 0.0544 0.0677 0.0563 0.359 0.396 0.372 0.50000 -0.0043 0.2286 0.3185 0.3540 0.239 0.418 0.553 0.76923 0.1298 0.3207 0.5855 0.3758 0.131 0.419 0.692 0.95309 0.3496 0.2072 0.4387 0.1085 0.079 0.400 0.750

μ	о ψо	ψÔ	ψ <sub>2</sub>	$\psi_{\mathtt{l}}^{\mathtt{l}}$	1 \\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\	2 \\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\	$\phi_{o}^{o}$	$\phi_1^\circ$
			-0.5262					0.0100
0.23077 0.50000			-0.1968		0.5606 1.1674			0.0702 0.2253
0.76923	1.3398	0.6369	0.3314	0.7513	1.3850	1.3159	0.7702	0.4588
0.95309	1.2661	0.8436	0.8278	0.3516	0.8302	0.2963	0.9945	0.6437

u	ø <sub>2</sub> o	$\phi_\mathtt{l}^{\scriptscriptstyle \perp}$	$\phi_{\perp}^{2}$	$\phi_2^2$	R	$\mathtt{T}_{\mathrm{D}}$	T
0.04691	0.0020	0.0027	0.0043	0.0007	0.551	0.215	0.215
0.23077	0.0118	0.0234	0.0341	0.0118	0.426	0.303	0.305
0.50000						0.397	0.447
0.76923						0.446	0.588
0.95309	0.2542	0.1496	0.3044	0.0653	0.118	0.456	0.663

PROBLEM 29 TAU = INFINITY ALBEDO FOR SINGLE SCATTERING = 0.90 A0 = 1.0000 A1 = 1.7321 A2 = 1.0000 F = 0.9330 P = 0.4821 NUMBER OF ITERATIONS FOR M = 0, 1, AND 2 WERE 21, 6, AND 4

μ 0•04691 2•23077 0•50000 0•76923 0•95309	1.1010 1.3142 1.4608 1.4964 1.4740	0.1180	$\psi_{2}^{\circ}$ $-0.5257$ $-0.4825$ $-0.1869$ $0.3511$ $0.8547$	1.0538 1.1116 1.0222 0.7530 0.3527	$\begin{matrix} 1 \\ \psi_2 \\ 0 \cdot 1049 \\ 0 \cdot 5604 \\ 1 \cdot 1659 \\ 1 \cdot 3823 \\ 0 \cdot 8283 \end{matrix}$	2 \$\psi_2\$ 3.0528 2.9792 2.3983 1.3163 0.2965	0.0000 0.0000 0.0000 0.0000 0.0000	0.0000 0.0000 0.0000 0.0000 0.0000
0.04691 0.23077 0.50000 0.76923 0.95309	\$\phi_2^\circ\$\$ 0.0000 0.0000 0.0000 0.0000 0.0000	\$\psi_1^1\$ 0.0000 0.0000 0.0000 0.0000 0.0000	\$\psi_2^1\$ 0.0000 0.0000 0.0000 0.0000 0.0000	\$\psi_2\$ 0.0000 0.0000 0.0000 0.0000 0.0000	R 0.595 0.489 0.377 0.292 0.244	T <sub>D</sub> 0.000 0.000 0.000 0.000 0.000	T 0.000 0.000 0.000 0.000 0.000	

PROBLEM 30 TAU = 0.050 ALBEDO FOR SINGLE SCATTERING = 0.60

A0 = 1.0000 A1 = 1.7321 A2 = 1.0000 F = 0.9330 P = 0.4821

NUMBER OF ITERATIONS FOR M = 0, 1, AND 2 WERE 2, 3, AND 3

μ 0.04691 0.23077 0.50000 0.76923 0.95309	1.0402 1.0492 1.0421 1.0287 1.0172	0.2212	-0.5109 -0.4387 -0.1419 0.3755 0.8552	1 1.0247 1.0051 0.8937 0.6575 0.3108	1 0.1261 0.6568 1.2862 1.4676 0.8630	2 42 3.0240 2.8803 2.2836 1.2134 0.2791	\$\phi_0^0\$ 0.3734 0.8846 0.9689 0.9992 1.0083	\$\phi_1^\circ\$ 0.0249 0.2046 0.4716 0.7431 0.9290
μ 0•04691	φ <sup>0</sup> <sub>2</sub> -0.1790	$\phi_{1}^{1}$ 0.3618	$\phi_{2}^{1}$	ø <sub>2</sub> 2 1•0 <b>5</b> 01	R <b>0 • 184</b>	т <sub>D</sub> 0.188	T 0.531	

μ	о ¥о	ψĺ	ο Ψ2	1 Ψ1	1 \\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\	2 \\psi_2	$\phi_{\circ}^{\circ}$	$\phi_1^{\circ}$
0.04691	1.0558	0.0332	-0.5145	1.0331	0.1189	3.0322	0.0166	0.0077
0.23077	1.1385	0.1894	-0.4584	1.0523	0.6123	2.9283	0.2128	0.0737
0.50000	1.1435	0.4574	-0.1660	0.9453	1.2398	2.3389	0.5810	0.2827
0.76923	1.1049	0.7390	0.3567	0.6924	1 • 4399	1.2776	0.7966	0.5394
0.95309	1.0662	0.9338	0.8444	0.3252	0.8533	0.2871	0.8951	0.7262
μ	ø <sub>2</sub>	$\phi_1^1$	$\phi_2^1$	ø <sub>2</sub>	R	$\mathtt{T}_{\mathrm{D}}$	Т	
0.04691	-0.0010	0.0077	0.0091	0.0063	0.293	0.165	0.165	
0.23077	-0.0526	0.1566	0.1321	0.3578	0.180	0.204	0.319	
0.50000	-0.0593	0.4118	0.5804	0.8883	0.085	0.197	0.565	
0.76923	0.2020	0.4266	0.8720	0.6804	0.040	0.178	0.701	
0.95309	0.5251	0.2327	0.5743	0.1726	0.021	0.169	0.761	

μ	o ∜o	ψ <u>0</u>	ο Ψ2	1 Ψ <u>1</u>	1 \(\psi_2\)	2 ¥2	$\phi_{o}^{o}$	ø <sub>1</sub>
0.04691	1.0568	0.0327		1.0332	0.1187	3.0322	0.0000	0.0000
0.23077	1.1499	0.1833	-0.4583	1.0551	0.6084	2.9300	0.0000	0.0000
0.50000	1.1795	0.4381	-0.1652	0.9540	1.2279	2.3452	0.0000	0.0000
0.76923	1.1517	0.7134	0.3589	0.7003	1.4298	1.2833	0.0000	0.0000
0.95309	1.1128	0.9074	0.8483	0.3286	0.8493	2.8867	0.0000	0.0000
μ	ø°2	$\phi_1^1$	$\phi_2^1$	ø <sub>2</sub>	R	$\mathtt{T}_{\mathtt{D}}$	Т	
0.04691	0.0000	0.0000	0.0000	0.0000	0.303	0.000	0.000	
0.23077	0.0000	0.0000	0.0000	0.0000	0 • 206	0.000	0.000	

0.000

0.000

0.000

0.000

0.000

0.000

0.50000 0.0000 0.0000 0.0000 0.0000 0.124

0.76923 0.0000 0.0000 0.0000 0.0000 0.073

0.95309 0.0000 0.0000 0.0000 0.0000 0.048

μ	ψ <mark>Ο</mark>	o \\[ \psi_1	ψ <sub>2</sub>	1 Ψ <u>1</u>	1 Ψ2	2 \\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\	$\phi_{o}^{o}$	$\phi_1^{\circ}$
0.04691	1.0195	0.0427	-0.5036	1.0115	0.1335	3.0086	0.3581	0.0204
0.23077	1.0243	0.2260	-0.4293	0.9889	0.6652	2.8602	0.8435	0.1952
0.50000	1.0206	0.4964	-0.1333	0.8796	1.2928	2.2667	0.9371	0.4622
0.76923	1.0146	0.7671	0.3817	0.6480	1.4712	1.2341	0.9681	0.7321
0.95309	1.0086	0.9520	0.8590	0.3066	0.8643	0.2769	0.9785	0.9167
μ	ø°2	$\phi_1^1$	ø2	φ <sub>2</sub> 2	R	T <sub>D</sub>	T	
0.04691	-0.1748	0.3526	0.0553	1.0404	0.090	0.089 0.040	0.435 0.845	
0.23077	-0.3513	0.8080	0.5581	2.3141	0.021			
0.50000	-0.1216	0.8028	1.1889	2.0531	0.008	0.019	0.924	
0.76923	0.3581	0.6137	1.3923	1.1569	0.003	0.015	0.952	
0.95309	0.8157	0.2948	0.8268	0.2629	0.001	0.013	0.961	

PROBLEM 34 TAU = 0.500 ALBEDO FOR SINGLE SCATTERING = 0.30

A0 = 1.0000 A1 = 1.7321 A2 = 1.0000 F = 0.9330 P = 0.4821

NUMBER OF ITERATIONS FOR M = 0, 1, AND 2 WERE 4, 3, AND 3

μ 0.04691 0.23077 0.50000 0.76923 0.95309	1.0257 1.0598 1.0600 1.0421 1.0249	0.2135	ψ <sub>2</sub> -0.5052 -0.4374 -0.1432 0.3743 0.8553	1 \\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\	1 \$\psi_2\$ 0.1304 0.6460 1.2733 1.4604 0.8609	$\begin{array}{c} 2\\ \psi_2 \\ 3.0125 \\ 2.8831 \\ 2.2931 \\ 1.2504 \\ 0.2808 \\ \end{array}$	\$\phi_0^\circ\$ 0.0064 0.1539 0.4567 0.6399 0.7238	\$\phi_1^\circ\$ 0.0030 0.0456 0.2259 0.4624 0.6369
	d <sup>o</sup>	$\phi_{-}^{1}$	$\phi_{\alpha}^{1}$	ø <sup>2</sup>	R	$\mathtt{T}_{\mathcal{D}}$	Т	

и	$\phi_{\mathbf{o}}^{\mathbf{o}}$	$\phi_{\mathfrak{I}}^{\perp}$	$\phi_{\perp}^{2}$	$\phi_2^2$	R	$\mathtt{T}_{\mathrm{D}}$	Т
0.04691	-0.0004	0.0033	0.0040	0.0030	0.131	0.064	0.063
0.23077	-0.0496	0.1312	0.1015	0.3409	0.075	0.083	0.198
0.50000	-0.0502	0.3602	0.5241	0.8568	0.034	0.084	0.452
0.76923	0.2046	0.3757	0.8164	0.6591	0.015	0.079	0.601
0.95309	0.5202	0.2030	0.5400	0.1674	0.008	0.076	0.668

TAU = INFINITY ALBEDO FOR SINGLE SCATTERING = 0.30 PROBLEM 35 A0 = 1.0000 A1 = 1.7321 A2 = 1.0000 F = 0.9330 P = 0.4821NUMBER OF ITERATIONS FOR M = 0, 1, AND 2 WERE 4, 3, AND 3

	0	0	0	1	1	.2	d <sup>0</sup>	d <sup>O</sup>
μ	$\Psi_{\mathbf{O}}$	$\Psi_{\mathtt{l}}$	$\psi_{\mathbf{Z}}$	$\Psi_\mathtt{l}$	$\Psi_{2}$	2 \\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\	$\varphi_{_{\bigcirc}}$	$\varphi_{\mathtt{l}}$
0.04691	1.0258	0.0407	-0.5052	1.0152	0.1304	3.0125	0.0000	0.0000
0.23077	1.0620	0.2124	-0.4374	1.0102	0.6448	2.8838	0.0000	0.0000
0.50000	1.0670	0.4793	-0.1433	0.9042	1.2696	2.2959	0.0000	0.0000
0.76923	1.0499	0.7538	0.3744	0.6646	1.4578	1.2530	0.0000	0.0000
0.95309	1.0313	0.9427	0.8559	0.3132	0.8602	2.8149	0.0000	0.0000
		7	7	0				
μ	$\phi_2^{\rm o}$	$\phi_1^\perp$	$\phi_2^{\perp}$	$\phi_2^2$	R	$\mathtt{T}_{\mathtt{D}}$	${f T}$	
0.04691	0.0000	0.0000	0.0000	0.0000	0.133	0.000	0.000	
0.23077	0.0000	0.0000	0.0000	0.0000	0.080	0.000	0.000	
0.50000	0.0000	0.0000	0.0000	0.0000	0.042	0.000	0.000	
0.76923	0.0000	0.0000	0.0000	0.0000	0.020	0.000	0.000	
0 4 1 0 7 2 3	0.000	3 \$ 3 0 0 0	5 5 5 5 5 5	2 4 2 0 0 0	00020		0.000	

0.000 0.000

TAU = INFINITY ALBEDO FOR SINGLE SCATTERING = 0.90 PROBLEM 36  $A2 = 2 \cdot 2415$  F =  $0 \cdot 9336$ A0 = 1.0000A1 = 1.7346P = 0.5821NUMBER OF ITERATIONS FOR M = 0, 1, AND 2 WERE 2, 5, AND 5 (NOTE--FOR M = 0 INITIAL FUNCTIONS WERE THOSE OF REFERENCE 4)

0.95309 0.0000 0.0000 0.0000 0.0000 0.011

μ	ο Ψ <sub>Ο</sub>	ο Ψ1	ο Ψ2	l Ψl	1 ψ2	2 \(\psi_2\)	$\phi_{0}^{\circ}$	$\phi_1^{\circ}$
0.04691	1.1185	0.0185	-0.5370	1.0516	0.1080	3.1376	0.0000	0.0000
0.23077	1.3550	0.1179	-0.5124	1.0896	0.5911	3.1955	0.0000	0.0000
0.50000	1.4868	0.3121	-0.2079	0.9561	1.2596	2.6406	0.0000	0.0000
0.76923	1.4699	0.5412	0.3771	0.6583	1.5179	1.4695	0.0000	0.0000
0.95309	1.3948	0.7094	0.9349	0.2909	0.9173	0.3331	0.0000	0.0000
μ	$\phi_2^{\circ}$	$\phi_{\underline{1}}^{\underline{1}}$	$\phi_2^1$	ø <sub>2</sub>	R	$^{\mathrm{T}}$ D	Т	
0.04691	0.0000	0.0000	0.0000	0.0000	0.604	0.000	0.000	
0 • 23077	0.0000	0.0000	0.0000	0.0000	0.486	0.000	0.000	
0.50000	0.0000	0.0000	0.0000	0.0000	0.376	0.000	0.000	
0.76923	0.0000	0.0000	0.0000	0.0000	0.296	0.000	0.000	
0.95309	0.0000	0.0000	0.0000	0.0000	0.256	0.000	0.000	

Table B-1 Comparison of Exact and Approximate Integrated Reflectances

Isotropic and Rayleigh Scattering

	$\mu_{O}$	Exact. iso*	Six-flux, iso.	Six-flux,Ray.	Two-flux,iso.	Richards, iso.
$\omega_{\bullet}$ = 0.90, $\tau_{1}$ = $\infty$ :	0.00000	0 (88	0.560	0.7/0		
	0.00000	0.677	0.569	0.569	0.519	0.500
	0.04691	0.654	0.560	0.560	0.519	0.489
	0.23077	0.581	0.527	0.527	0.519	0.448
	0.50000	0.507	0.484	0.485	0.519	0.400
	0.76923	0.452	0.440	0.441	0.519	0.361
	0.95309	0.422	0.389	0.391	0.519	0.338
	1.00000	0.416	0.333	0.337	0.519	0.333
$\omega_{\bullet} = 0.90, \tau_{1} = 1.0$		a (a)		1		- 1
	0.00000	0.624	0.533	0.534	0.519	0.457
	0.04691	0.598	0.520	0.521	0.519	0.446
	0.23077	0.504	0.466	0.467	0.496	0.393
	0.50000	0.394	0.375	0.376	0.421	0.316
	0.76923	0.316	0.303	0.304	0.330	0.257
	0.95309	0.279	0.259	0.260	0.293	0.227
	1.00000	0.271	0.137	0.138	0.284	0.220
$\omega_{\bullet} = 0.90, \tau_{1} = 0.50$		01				
	0.00000	0.584	0.507	0.509	0.519	0.428
	0.04691	0.548	0.492	0.494	0.519	0.413
	0.23077	0.417	0.390	0.391	0.420	0.333
	0.50000	0.279	0.267	0.267	0.285	0.231
	0.76923	0.209	0.199	0.199	0.211	0.171
	0.95309	0.179	0.166	0.167	0.182	0.147
	1.00000	0.170	0.079	0.079	0.175	0.141
$\omega_{\bullet}=0.90, \tau_{1}=0.25$						
	0.00000	0.571	0.488	0.491	0.519	0.405
	0.04691	0.498	0.463	0.467	0.505	0.388
	0.23077	0.300	0.283	0.283	0.302	0.243
	0.50000	0.173	0.166	0.166	0.175	0.146
	0.76923	0.125	0.117	0.117	0.124	0.103
	0.95309	0.105	0.096	0.096	0.103	0.084
	1.00000	0.101	0.043	0.043	0.099	0.081
$\omega_{\bullet} = 0.60, \tau_{1} = 0.50$						
	0.00000	0.345	0.235	0.234	0.225	0.176
	0.04691	0.318	0.228	0.227	0.225	0.168
	0.23077	0.232	0.190	0.193	0.211	0.135
	0.50000	0.154	0.137	0.138	0.164	0.092
	0.76923	0.114	0.103	0.104	0.129	0.069
	0.95309	0.097	0.082	0.084	0.112	0.059
	1.00000	0.093	0.034	0.035	0.108	0.057
$\omega_{\bullet}=0.30, \tau_{1}=0.50$ :		_				
	0.00000	0.156	0.091	0.091	0.089	0.064
	0.04691	0.142	0.088	0.088	0.089	0.061
	0.23077	0.100	0.075	0.076	0.089	0.048
	0.50000	0.066	0.056	0.057	0.086	0.033
	0.76923	0.048	0.042	0.043	0.079	0.025
	0.95309	0.041	0.032	0.034	0.074	0.021
	1.00000	0.040	0.012	0.014	0.072	0.020

<sup>\*</sup> Extrapolated by Lagrange's formula for  $\mu_{0}$  = 0 and  $\mu_{0}$  = 1.0.

Table B-2 Comparison of Exact and Approximate Total Transmissions

Isotropic and Rayleigh Scattering

ω <sub>•</sub> =0.90,τ <sub>1</sub> =1.0:	$\mu_{O}$	Exact. iso.*	Six-flux, iso.	Six-flux,Ray.	Two-flux, iso.	Richards, iso.
<b>6.</b> 20. 90, 11 - 110.	0.00000	0.183	0.181	0.173	0.000	0.211
	0.04691	0.197	0.194	0.185	0.001	0.221
	0.23077	0.271	0.265	0.256	0.189	0.269
	0.50000	0.398	0.413	0.412	0.420	0.385
	0.76923	0.498	0.523	0.523	0.549	0.492
	0.95309	0.536	0.567	0.568	0.609	0.549
	1.00000	0.541	0.562	0.564	0.621	0.562
$\omega_{\bullet}=0.90, \tau_{1}=0.50$			•			
••,·· <u>·</u>	0.00000	0.257	0.240	0.228	0.000	0.278
	0.04691	0.286	0.256	0.244	0.025	0.292
	0.23077	0.426	0.415	0.412	0.395	0.403
	0.50000	0.607	0.611	0.610	0.621	0.582
	0.76923	0.693	0.710	0.710	0.724	0.686
	0.95309	0.728	0.749	0.749	0.767	0.733
	1,00000	0.742	0.742	0.743	0.776	0.742
$\omega_{\bullet} = 0.90, \tau_{1} = 0.25$	:					
•,.,.	0.00000	0.277	0.277	0.264	0.000	0.322
	0.04691	0.365	0.308	0.294	0.137	0.340
	0.23077	0.607	0.597	0.594	0.600	0.574
	0.50000	0.768	0.768	0.768	0.776	0.748
	0.76923	0.827	0.837	0.837	0.844	0.823
	0.95309	0.854	0.864	0.864	0.871	0.853
	1.00000	0.861	0.859	0.859	0.877	0.859
$\omega_{\bullet} = 0.60, \tau_{1} = 0.50$	:					
, , ,	0.00000	0.139	0.102	0.095	0.000	0.104
	0.04691	0.154	0.109	0.101	0.000	0.108
	0.23077	0.283	0.255	0.252	0.061	0.217
	0.50000	0.502	0.491	0.491	0.269	0.454
	0.76923	0.623	0.618	0.620	0.421	0.588
	0.95309	0.675	0.668	0.670	0.496	0.648
	1.00000	0.694	0.661	0.664	0.512	0.661
$\omega_{\bullet}=0.30, \tau_{1}=0.50$	:					
• ±	0.00000	0.057	0.038	0.035	0.000	0.039
	0.04691	0.064	0.040	0.037	0.000	0.041'
	0.23077	0.184	0.168	0.167	0.026	0.156
	0.50000	0.425	0.417	0.418	0.186	0.399
	0.76923	0.566	0.561	0.562	0.335	0.546
	0.95309	o.628	0.622	0.622	0.411	0.612
	1.00000	0.651	0.626	0.629	0.430	0.626

<sup>\*</sup> Extrapolated by Lagrange's formula for  $\mu_{0}$  = 0 and  $\mu_{0}$  = 1.0.

Table B-3 Comparison of Exact and Approximate Integrated Reflectances Anisotropic Scattering, F = 0.9330, P = 0.4821

$ \begin{array}{c} \mathbf{r}_1 = \mathbf{oo}: \\ \mathbf{r}_1 = \mathbf{oo}: \\ \mathbf{r}_1 = \mathbf{oo}: \\ \hline \\ 0.00000 & 0.628 & 0.467 & 0.195 & 0.567 & 0.433 & 0.195 \\ 0.04691 & 0.595 & 0.457 & 0.195 & 0.530 & 0.421 & 0.195 \\ 0.23077 & 0.489 & 0.425 & 0.195 & 0.396 & 0.362 & 0.143 \\ 0.50000 & 0.377 & 0.383 & 0.195 & 0.239 & 0.259 & 0.089 \\ 0.76923 & 0.292 & 0.338 & 0.195 & 0.0131 & 0.196 & 0.064 \\ 0.995309 & 0.244 & 0.286 & 0.195 & 0.079 & 0.161 & 0.054 \\ 1.00000 & 0.233 & 0.228 & 0.195 & 0.067 & 0.074 & 0.052 \\ \hline \\ \boldsymbol{\omega}_{\bullet} = 0.90 \\ \boldsymbol{\tau}_{1} = 0.50: \\ \hline \\ \begin{array}{c} 0.00000 & 0.547 & 0.417 & 0.195 & 0.552 & 0.405 & 0.195 \\ 0.04691 & 0.500 & 0.400 & 0.186 & 0.466 & 0.371 & 0.155 \\ 0.23077 & 0.333 & 0.277 & 0.098 & 0.238 & 0.183 & 0.056 \\ 0.50000 & 0.163 & 0.170 & 0.052 & 0.097 & 0.099 & 0.028 \\ 0.76923 & 0.078 & 0.120 & 0.035 & 0.043 & 0.068 & 0.019 \\ 0.995309 & 0.041 & 0.098 & 0.029 & 0.018 & 0.055 & 0.015 \\ 1 = \mathbf{oo}: \\ \hline \\ \boldsymbol{\omega}_{\bullet} = 0.60 \\ \boldsymbol{\tau}_{1} = \mathbf{oo}: \\ \hline \\ \boldsymbol{\omega}_{\bullet} = 0.60 \\ \boldsymbol{\tau}_{1} = \mathbf{oo}: \\ \hline \\ \boldsymbol{\omega}_{\bullet} = 0.30 \\ \boldsymbol{\tau}_{1} = \mathbf{oo}: \\ \hline \\ \boldsymbol{\omega}_{\bullet} = 0.30 \\ \boldsymbol{\tau}_{1} = \mathbf{oo}: \\ \hline \\ \boldsymbol{\omega}_{\bullet} = 0.30 \\ \boldsymbol{\tau}_{1} = \mathbf{oo}: \\ \hline \\ \boldsymbol{\omega}_{\bullet} = 0.30 \\ \boldsymbol{\tau}_{1} = \mathbf{oo}: \\ \hline \\ \boldsymbol{\sigma}_{\bullet} = 0.30 \\ \boldsymbol{\tau}_{1} = \mathbf{oo}: \\ \hline \\ \boldsymbol{\sigma}_{\bullet} = 0.30 \\ \boldsymbol{\tau}_{0.95309} & 0.042 & 0.047 & 0.014 & 0.034 & 0.029 \\ 0.04691 & 0.133 & 0.054 & 0.014 & 0.034 & 0.029 \\ 0.995309 & 0.048 & 0.071 & 0.046 & 0.085 & 0.077 & 0.027 \\ 0.76923 & 0.048 & 0.071 & 0.046 & 0.085 & 0.077 & 0.027 \\ 0.04691 & 0.133 & 0.057 & 0.014 & 0.131 & 0.056 & 0.014 \\ 0.023077 & 0.080 & 0.047 & 0.014 & 0.131 & 0.056 & 0.014 \\ 0.023077 & 0.080 & 0.047 & 0.014 & 0.034 & 0.029 & 0.013 \\ 0.23077 & 0.080 & 0.047 & 0.014 & 0.034 & 0.029 & 0.014 \\ 0.023077 & 0.080 & 0.047 & 0.014 & 0.034 & 0.029 & 0.014 \\ 0.023077 & 0.080 & 0.047 & 0.014 & 0.034 & 0.029 & 0.011 \\ 0.76923 & 0.020 & 0.030 & 0.014 & 0.034 & 0.029 & 0.011 \\ 0.76923 & 0.020 & 0.030 & 0.014 & 0.034 & 0.029 & 0.011 \\ 0.76923 & 0.020 & 0.030 & 0.014 & 0.034 & 0.029 & 0.011 \\ 0.76923 & 0.020 & 0.030 & 0.014 & 0.008 & 0.014 & 0.008 \\ 0.095309 & 0.011 & 0.020 & 0.014 & 0$	0.00	$\mu_{O}$	Exact*	Six-flux	Two-flux	<b>w</b> .=0.90	Exact*	Six-flux	Two-flux
1 0.00000 0.628 0.467 0.195 0.567 0.433 0.195 0.00691 0.595 0.457 0.195 0.530 0.421 0.195 0.23077 0.489 0.425 0.195 0.396 0.362 0.143 0.5000 0.5000 0.377 0.383 0.195 0.239 0.259 0.089 0.76923 0.292 0.338 0.195 0.131 0.196 0.064 0.95309 0.244 0.286 0.195 0.079 0.161 0.054 1.0000 0.233 0.228 0.195 0.067 0.079 0.161 0.054 1.0000 0.233 0.228 0.195 0.067 0.079 0.161 0.054 0.066 0.066 0.0661 0.500 0.400 0.186 0.466 0.371 0.155 0.23077 0.333 0.277 0.098 0.238 0.133 0.056 0.23077 0.333 0.277 0.098 0.238 0.183 0.056 0.23077 0.333 0.277 0.098 0.238 0.183 0.056 0.056 0.05500 0.163 0.170 0.052 0.097 0.099 0.028 0.76923 0.078 0.120 0.035 0.043 0.068 0.019 0.95309 0.041 0.098 0.029 0.018 0.055 0.015 1.00000 0.030 0.041 0.028 0.030 0.013 0.022 0.014 0.230 0.020 0.030 0.041 0.028 0.030 0.022 0.014 0.230 0.056 0.23077 0.206 0.142 0.046 0.293 0.157 0.046 0.23077 0.206 0.124 0.120 0.046 0.293 0.157 0.046 0.23077 0.206 0.124 0.120 0.046 0.085 0.079 0.099 0.029 0.0500 0.124 0.120 0.046 0.085 0.077 0.006 0.124 0.120 0.046 0.085 0.077 0.007 0.097 0.090 0.0	•								
0.04691   0.595   0.457   0.195   0.530   0.421   0.195     0.23077   0.489   0.425   0.195   0.396   0.362   0.143     0.5000   0.377   0.383   0.195   0.239   0.259   0.089     0.76923   0.292   0.338   0.195   0.131   0.196   0.064     0.95309   0.244   0.286   0.195   0.079   0.161   0.054     1.0000   0.233   0.228   0.195   0.067   0.074   0.052	$\tau_1 = \infty$ :	0.00000	0.628	0.467	0.195	.1-2.00	0.567	0.433	0.195
0.23077 0.489 0.425 0.195 0.396 0.362 0.143 0.50000 0.377 0.383 0.195 0.239 0.259 0.089 0.76923 0.292 0.338 0.195 0.131 0.196 0.064 0.95309 0.244 0.286 0.195 0.079 0.161 0.054 1.00000 0.233 0.228 0.195 0.067 0.074 0.052   ■ ■ 0.90  T = 0.50:								0.421	0.195
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$								0.362	0.143
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$		•					0.239	0.259	0.089
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$		•			0.195		0.131	0.196	0.064
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$			-	0.286	0.195		0.079	0.161	0.054
$ \begin{array}{c} {\bf \tau_1 = 0.50;} \\ {\bf \tau_1 = 0.50;} \\ & & & & & & & & & & & & & & & & & & $			0.233	0.228	0.195		0.067	0.074	0.052
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	$\omega_{\bullet} = 0.90$								
0.00000 0.547 0.417 0.195 0.552 0.405 0.195 0.0466 0.0466 0.371 0.155 0.23077 0.333 0.277 0.098 0.238 0.183 0.056 0.50000 0.163 0.170 0.052 0.097 0.099 0.028 0.76923 0.078 0.120 0.035 0.043 0.068 0.019 0.95309 0.041 0.098 0.029 0.018 0.055 0.015 1.0000 0.336 0.165 0.046 0.330 0.062 0.014 0.028 0.013 0.022 0.014 0.046 0.04691 0.333 0.160 0.046 0.23077 0.206 0.124 0.120 0.046 0.085 0.077 0.027 0.76923 0.073 0.097 0.046 0.085 0.019 0.330 0.160 0.046 0.23077 0.206 0.124 0.120 0.046 0.085 0.077 0.027 0.76923 0.073 0.097 0.046 0.085 0.077 0.027 0.76923 0.073 0.097 0.046 0.085 0.077 0.027 0.76923 0.073 0.097 0.046 0.085 0.077 0.027 0.76923 0.073 0.097 0.046 0.008 0.014 0.015 0.016	$\tau_1 = 0.50$ :								
0.04691 0.500 0.400 0.186 0.466 0.371 0.155 0.23077 0.333 0.277 0.098 0.238 0.183 0.056 0.50000 0.163 0.170 0.052 0.097 0.099 0.028 0.76923 0.078 0.120 0.035 0.043 0.068 0.019 0.95309 0.041 0.098 0.029 0.018 0.055 0.015 1.00000 0.030 0.041 0.028 0.013 0.022 0.014  ■ ■ 0.60  ▼1 = ∞:  0.00000 0.336 0.165 0.046 0.330 0.162 0.046 0.04691 0.3303 0.160 0.046 0.293 0.157 0.046 0.23077 0.206 0.142 0.046 0.293 0.157 0.046 0.23077 0.206 0.142 0.046 0.085 0.077 0.027 0.76923 0.073 0.097 0.046 0.085 0.077 0.027 0.76923 0.073 0.097 0.046 0.040 0.055 0.020 0.95309 0.048 0.071 0.046 0.021 0.041 0.017 1.00000 0.043 0.041 0.046 0.021 0.041 0.017 1 = ∞:  ■ ■ 0.30  ▼1 = ∞:  ■ 0.00000 0.153 0.057 0.014 0.036 0.015 0.054 0.014 0.23077 0.080 0.047 0.014 0.131 0.056 0.014 0.23077 0.080 0.047 0.014 0.131 0.054 0.014 0.23077 0.080 0.047 0.014 0.131 0.054 0.014 0.23077 0.080 0.047 0.014 0.034 0.029 0.011 0.50000 0.042 0.039 0.014 0.035 0.020 0.008 0.95309 0.011 0.020 0.014 0.015 0.020 0.008	1					$\tau_1 = 0.25$			
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$		0.00000	0.547					•	
0.50000 0.163 0.170 0.052 0.097 0.099 0.028 0.76923 0.078 0.120 0.035 0.043 0.068 0.019 0.95309 0.041 0.098 0.029 0.018 0.055 0.015 1.00000 0.030 0.041 0.028 0.013 0.022 0.014 0.008 0.060 0.0000 0.336 0.165 0.046 0.330 0.162 0.046 0.04691 0.303 0.160 0.046 0.293 0.157 0.046 0.23077 0.206 0.142 0.046 0.085 0.077 0.027 0.76923 0.073 0.097 0.046 0.085 0.077 0.027 0.76923 0.073 0.097 0.046 0.040 0.055 0.020 0.95309 0.048 0.071 0.046 0.021 0.041 0.017 1.00000 0.043 0.041 0.046 0.021 0.041 0.017 0.016 0.04691 0.133 0.054 0.041 0.046 0.016 0.012 0.016 0.020 0.04691 0.133 0.054 0.041 0.046 0.016 0.012 0.016 0.020 0.04691 0.133 0.054 0.041 0.046 0.016 0.012 0.016 0.016 0.012 0.016 0.055 0.020 0.95309 0.048 0.071 0.046 0.016 0.012 0.016 0.012 0.016 0.055 0.020 0.056 0.047 0.014 0.034 0.029 0.014 0.23077 0.080 0.047 0.014 0.034 0.075 0.043 0.013 0.50000 0.042 0.039 0.014 0.034 0.029 0.011 0.76923 0.020 0.030 0.014 0.034 0.029 0.011 0.76923 0.020 0.030 0.014 0.034 0.029 0.011 0.76923 0.020 0.030 0.014 0.034 0.029 0.011 0.76923 0.020 0.030 0.014 0.015 0.020 0.008 0.95309 0.011 0.020 0.014 0.005 0.008 0.014 0.007		0.04691	0.500						
$\begin{array}{c} 0.76923 & 0.078 & 0.120 & 0.035 & 0.043 & 0.068 & 0.019 \\ 0.95309 & 0.041 & 0.098 & 0.029 & 0.018 & 0.055 & 0.015 \\ 1.00000 & 0.030 & 0.041 & 0.028 & 0.013 & 0.022 & 0.014 \\ \hline \textbf{$\omega$} = 0.60 \\ \hline \textbf{$\tau_1$} = \textbf{$\infty$}; \\ \hline & 0.00000 & 0.336 & 0.165 & 0.046 & 0.330 & 0.162 & 0.046 \\ 0.04691 & 0.303 & 0.160 & 0.046 & 0.293 & 0.157 & 0.046 \\ 0.23077 & 0.206 & 0.142 & 0.046 & 0.180 & 0.119 & 0.039 \\ 0.50000 & 0.124 & 0.120 & 0.046 & 0.085 & 0.077 & 0.027 \\ 0.76923 & 0.073 & 0.097 & 0.046 & 0.040 & 0.055 & 0.020 \\ 0.95309 & 0.048 & 0.071 & 0.046 & 0.021 & 0.041 & 0.017 \\ 1.00000 & 0.043 & 0.041 & 0.046 & 0.021 & 0.041 & 0.017 \\ 1.00000 & 0.043 & 0.041 & 0.046 & 0.018 & 0.012 & 0.016 \\ \hline \textbf{$\omega$} = 0.30 \\ \hline \textbf{$\tau_1$} = \textbf{$\infty$}; \\ \hline \textbf{$0.04691} & 0.133 & 0.054 & 0.014 & 0.131 & 0.056 & 0.014 \\ 0.04691 & 0.133 & 0.054 & 0.014 & 0.131 & 0.056 & 0.014 \\ 0.23077 & 0.080 & 0.047 & 0.014 & 0.131 & 0.054 & 0.013 \\ 0.59000 & 0.042 & 0.039 & 0.014 & 0.034 & 0.029 & 0.011 \\ 0.76923 & 0.020 & 0.030 & 0.014 & 0.035 & 0.020 & 0.008 \\ 0.95309 & \textbf{$0.011} & 0.020 & 0.014 & 0.008 & 0.014 & 0.007 \\ \hline \end{array}$					-		_	_	•
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$			_	· ·	•		- 1		
$\begin{array}{c} \mathbf{u}_{\bullet} = 0.60 \\ \mathbf{v}_{1} = \infty; \\ 1 = \infty; \\ 0.00000 & 0.336 & 0.165 & 0.046 \\ 0.04691 & 0.303 & 0.160 & 0.046 & 0.293 & 0.157 & 0.046 \\ 0.23077 & 0.206 & 0.142 & 0.046 & 0.180 & 0.119 & 0.039 \\ 0.50000 & 0.124 & 0.120 & 0.046 & 0.085 & 0.077 & 0.027 \\ 0.76923 & 0.073 & 0.097 & 0.046 & 0.040 & 0.055 & 0.020 \\ 0.95309 & 0.048 & 0.071 & 0.046 & 0.021 & 0.041 & 0.017 \\ 1.00000 & 0.043 & 0.041 & 0.046 & 0.016 & 0.012 & 0.016 \\ \hline \mathbf{v}_{1} = \infty; \\ \mathbf{v}_{2} = 0.30 \\ \mathbf{v}_{3} = 0.30 \\ \mathbf{v}_{1} = 0.50; \\ \mathbf{v}_{3} = 0.30 \\ \mathbf{v}_{1} = 0.50; \\ \mathbf{v}_{2} = 0.30 \\ \mathbf{v}_{3} = 0.0000 & 0.153 & 0.057 & 0.014 & 0.151 & 0.056 & 0.014 \\ 0.04691 & 0.133 & 0.054 & 0.014 & 0.131 & 0.054 & 0.014 \\ 0.23077 & 0.080 & 0.047 & 0.014 & 0.131 & 0.054 & 0.014 \\ 0.23077 & 0.080 & 0.047 & 0.014 & 0.034 & 0.029 & 0.011 \\ 0.76923 & 0.020 & 0.030 & 0.014 & 0.034 & 0.029 & 0.011 \\ 0.76923 & 0.020 & 0.030 & 0.014 & 0.008 & 0.014 & 0.007 \\ 0.95309 & 0.011 & 0.020 & 0.014 & 0.008 & 0.014 & 0.007 \\ \hline \end{array}$					•		_		
$\begin{array}{c} \pmb{\omega}_{\bullet} = 0.60 \\ \pmb{\tau}_{1} = \infty; \\ \hline \\ 0.00000 & 0.336 & 0.165 & 0.046 & 0.330 & 0.162 & 0.046 \\ 0.04691 & 0.303 & 0.160 & 0.046 & 0.293 & 0.157 & 0.046 \\ 0.23077 & 0.206 & 0.142 & 0.046 & 0.180 & 0.119 & 0.039 \\ 0.50000 & 0.124 & 0.120 & 0.046 & 0.085 & 0.077 & 0.027 \\ 0.76923 & 0.073 & 0.097 & 0.046 & 0.040 & 0.055 & 0.020 \\ 0.95309 & 0.048 & 0.071 & 0.046 & 0.021 & 0.041 & 0.017 \\ 1.00000 & 0.043 & 0.041 & 0.046 & 0.016 & 0.012 & 0.016 \\ \hline \pmb{\omega}_{\bullet} = 0.30 \\ \pmb{\tau}_{1} = \infty; \\ \hline \hline \\ \hline & 0.00000 & 0.153 & 0.057 & 0.014 & 0.151 & 0.056 & 0.014 \\ 0.04691 & 0.133 & 0.054 & 0.014 & 0.131 & 0.054 & 0.014 \\ 0.23077 & 0.080 & 0.047 & 0.014 & 0.131 & 0.054 & 0.014 \\ 0.23077 & 0.080 & 0.047 & 0.014 & 0.075 & 0.043 & 0.013 \\ 0.50000 & 0.042 & 0.039 & 0.014 & 0.034 & 0.029 & 0.011 \\ 0.76923 & 0.020 & 0.030 & 0.014 & 0.035 & 0.020 & 0.008 \\ 0.95309 & \textbf{0.011} & 0.020 & 0.014 & 0.008 & 0.014 & 0.007 \\ \hline \end{array}$		,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,			-				
		1.00000	0.030	0.041	0.028		-	0.022	0.014
0.00000 0.336 0.165 0.046 0.330 0.162 0.046 0.04691 0.303 0.160 0.046 0.293 0.157 0.046 0.23077 0.206 0.142 0.046 0.180 0.119 0.039 0.50000 0.124 0.120 0.046 0.085 0.077 0.027 0.76923 0.073 0.097 0.046 0.040 0.055 0.020 0.95309 0.048 0.071 0.046 0.021 0.041 0.017 1.00000 0.043 0.041 0.046 0.016 0.012 0.016  w. = 0.30  τ <sub>1</sub> = ∞:  0.00000 0.153 0.057 0.014 0.151 0.056 0.014 0.04691 0.133 0.054 0.014 0.131 0.054 0.014 0.23077 0.080 0.047 0.014 0.131 0.054 0.014 0.23077 0.080 0.047 0.014 0.075 0.043 0.013 0.50000 0.042 0.039 0.014 0.034 0.029 0.011 0.76923 0.020 0.030 0.014 0.005 0.008 0.95309 0.011 0.020 0.014 0.008 0.014 0.007						•			
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\tau_1 = \infty$ :			2.765	0.01.6	τ <sub>1</sub> = 0.50		0 160	0 0)16
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$ au_1 = \infty$ : $ 0.00000                              $	0 30	1.00000	0.043	0.041	0.040	w = 030		0,0111	
0.00000 0.153 0.057 0.014 0.151 0.056 0.014 0.04691 0.133 0.054 0.014 0.131 0.054 0.014 0.23077 0.080 0.047 0.014 0.075 0.043 0.013 0.50000 0.042 0.039 0.014 0.034 0.029 0.011 0.76923 0.020 0.030 0.014 0.015 0.020 0.008 0.95309 0.011 0.020 0.014 0.008 0.014 0.007	•								
0.04691 0.133 0.054 0.014 0.131 0.054 0.014 0.23077 0.080 0.047 0.014 0.075 0.043 0.013 0.50000 0.042 0.039 0.014 0.034 0.029 0.011 0.76923 0.020 0.030 0.014 0.015 0.020 0.008 0.95309 <b>0.</b> 011 0.020 0.014 0.008 0.014 0.007	1 = 00:	0.00000	0.153	0.057	0.014	11- 0.70	0.151	0.056	0.014
0.23077       0.080       0.047       0.014       0.075       0.043       0.013         0.50000       0.042       0.039       0.014       0.034       0.029       0.011         0.76923       0.020       0.030       0.014       0.015       0.020       0.008         0.95309       0.011       0.020       0.014       0.008       0.014       0.007				- 1				0.054	0.014
0.50000       0.042       0.039       0.014       0.034       0.029       0.011         0.76923       0.020       0.030       0.014       0.015       0.020       0.008         0.95309       0.011       0.020       0.014       0.008       0.014       0.007				•				•	0.013
0.76923 0.020 0.030 0.014 0.015 0.020 0.008 0.95309 <b>0.</b> 011 0.020 0.014 0.008 0.014 0.007		•		•				_	0.011
0.95309 0.011 0.020 0.014 0.008 0.014 0.007		•		<b>.</b>			_	0.020	0.008
				_	0.014		0.008	0.014	
		1.00000		0.008	0.014		0.007	0.003	0.007

<sup>\*</sup> Extrapolated by Lagrange's formula for  $\mu_0^{}$  = 0 and 1.0.

Table B $^{\perp}$  Comparison of Exact and Approximate Total Transmissions Anisotropic Scattering, F = 0.9330, P = 0.4821

0.00	$\mu_{O}$	Exact*	Six-flux	Two-flux		Exact*	Six-flux	Two-flux
ω <sub>0</sub> = 0.90					ω <sub>•</sub> =0.90			
$\tau_1 = 1.0:$	0.00000	0.236	0.220	0.000	$\tau_1 = 0.50$ :		2 262	0.000
	0.04691	•		0.000		0.291	0.269	0.000
	-	0.259	0.233	0.041		0.331	0.288	0.198
	0.23077	0.372	0.339	0.511		0.507	0.515	0.711
	0.50000	0.553	0.523	0.730		0.723	0.707	0.853
	0.76923	0.692	0.632	0.814		0.833	0.790	0.902
	0.95309	0.750	0.673	0.847		0.885	0.821	0.920
	1.00000	0.761	0.662	0.853		0.904	0.811	0.923
$\omega_{\bullet} = 0.90$								
$\tau_1 = 0.25$ :								
	0.00000	0.297	0.299	0.000				
	0.04691	0.396	0.354	0.439				
	0.23077	0.669	0.698	0.842				
	0.50000	0.853	0.835	0.923				
	0.76923	0.933	0.889	0.949				
	0.95309	0.979	0.905	0.959				
	1.00000	0.991	0.900	0.961				
$\omega_{\bullet} = 0.60$					ω <sub>•</sub> =0.30			
$\tau_1 = 0.50$ :					$\tau_1 = 0.50$ :			
	0.00000	0.143	0.100	0.000	<b>-</b>	0.041	0.033	0.000
	0.04691	0.165	0.106	0.009		0.063	0.035	0.000
	0.23077	0.319	0.303	0.386		0.198	0.175	0.210
	0.50000	0.565	0.550	0.644		0.452	0.445	0.487
	0.76923	0.701	0.671	0.751		0.601	0.587	0.626
	0.95309	0.761	0.716	0.794		0.668	0.645	0.686
	1.00000	0.783	0.710	0.803		0.692	0.652	0.698
		, ,		5		- • - / -	0.075	0.000

<sup>\*</sup> Extrapolated by Lagrange's formula for  $\mu_{\text{O}}$  = 0 and 1.0

Table B-5 Comparison of Exact and Approximate Integrated Reflectances and Total Transmissions

Anisotropic Scattering,  $F_1$  = 0.750,  $\omega_{\bullet}$ = 0.90,  $\tau_1$  = 1.0

			-	· 1	, •	_	
	Integrate	d Reflects	ances			al Transmiss:	
	$\mu_{o}$	Exact*	Six-flux	Two-flux	Exact*	Six-flux	Two-flux
P = 0.5821:	0.00000	0.600	0.484	0.402	0.184	0.171	0.000
	0.04691	0.563	0.471	0.402	0.203	0.183	0.006
	0.23077	0.438	0.411	0.357	0.303	0.289	0.310
	0.50000	0.303	0.315	0.261	0.472	0.467	0.560
	0.76923	0.213	0.248	0.201	0.605	0.583	0.678
	0.95309	0.173	0.209	0.173	0.657	0.629	0.727
	1.00000	0.165	0.105	0.168	0.665	0.621	0.738
P = 0.4821:							
	0.00000	0.598	0.483	0.402	0.203	0.191	0.000
	0.04691	0.564	0.470	0.402	0.222	0.203	0.006
	0.23077	0.445	0.410	0.357	0.318	0.295	0.310
	0.50000	0.310	0.314	0.261	0.476	0.469	0.560
	0.76923	0.218	0.247	0.201	0.598	0.582	0.678
	0.95 <b>3</b> 09	0.175	0.207	0.173	0.645	0.626	0.727
	1.00000	0.166	0.103	0.168	0.652	0.618	0.738
P = 0.3889:							
	0.00000	0.589	0.480	0.402	0.221	0.213	0.000
	0.04691	0.559	0.467	0.402	0.238	0.225	0.006
	0.23077	0.448	0.409	0.357	0.329	0.308	0.310
	0.50000	0.315	0.314	0.261	0.480	0.471	0.560
	0.76923	0.220	0.246	0.201	0.596	0.580	0.678
	0.95309	0.174	0.205	0.173	0.646	0.622	0.727
	1.00000	0.164	0.101	0.168	0.656	0.614	0.738

<sup>\*</sup> Extrapolated by Lagrange's formula for  $\mu_{\text{O}}$  = 0 and 1.0

### APPENDIX C

## Computer Programs

All computer programs were written in MAD (Michigan Algorithm Decoder) compiler language and were compiled and executed on the IBM-704. A description of the language is given in reference (10). Programs written in this compiler language may be compiled and run on either the IBM-704 or 709 and in the near future provision will be made for processing MAD programs on the IBM-7090.

The programs are presented in the following pages with brief explanatory notes.

### 1. Solution of the Integral Equations

# a) Input Data

One general program was written to solve any set of integral equations for an arbitrary value of N and m. In order to specify the problem and its method of solution the following information is required. (The MAD symbols are placed in parenthesis for use in reading the listing of the programs).

1)	Run Number	Α	five-character	number	for	identification
		of	output (RUN)			

Number of terms in expansion of phase function (NTERM)

	τ <sub>l</sub>	Optical thickness of dispersion (TAU)
	TOL	Iteration will be stopped whenever the maximum absolute value of the change in any function is less than TOL (TOL)
	ω.	Albedo for single scattering (ALBEDO)
2)	$a_0 \dots a_N$	Coefficients of an angular distribution function, A(0)A(NTERMS)
3)	$\mu_{ m O}, \ldots, \mu_{ m N}$	Set of values of the independent variable for use in numerical integration (See equation I-35), U(0)U(NPOINT)
4)	$\mathbf{w}_{\mathrm{O}}, \ldots, \mathbf{w}_{\mathrm{N}}$	Weight coefficients for numerical integration (See equation I-35), W(O)W(NPOINT)
5)	Ψ''' (μ;) (0)	Initial value of Psi Functions, PSI(0,L,I)
	i = 0,, p	
	λ = m,, N	
6)	$\phi_{I}^{m}\left(\mu_{i}\right)^{(0)}$	Initial values of PHi Functions, PHI(0,L,I)
	i = 0,, p	
	1 = m,, N	

The preceeding six groups of information are referred to as the "basic data package." They completely define a solution to the integral equations or any trial solution. Thus, the computer starts with a basic data package as an initial approximation; it then performs a certain number of iterations; and, finally, produces as output a new basic data package. The basic data package is preceded by a header card

giving the maximum number of iterations allowable and information as to which of the special options (described next) are desired.

## b) Special Options

The following options are available in the routine:

1) Items 5 and 6 (i.e. the initial values of the functions) of the data above may be deleted in which case the following expressions will be used for the initial functions.

$$\Psi_{I}^{m}(\mu_{i})^{(o)} = P_{I}^{m}(\mu_{i})$$

$$\Phi_{I}^{m}(\mu_{i})^{(o)} = e^{-\tau_{i}/\mu_{i}} P_{I}^{m}(\mu_{i})$$

- 2) The results of a computation may be punched on cards in the form of a basic data package for use in performing more iterations later or for use in another program if desired.
- 3) A reversal of direction of the change in  $\psi_{l}^{\bullet}$  or  $\phi_{l}^{\bullet}$  may be interpreted as convergence and the iteration stopped when this occurs if desired.

In addition the following features are available for speeding up the solution under certain special circumstances.

- a) if  $\tau_i > 80$ , calculation of the  $\phi_i(\mu_i)$  will be deleted and  $\phi_{\perp}^{m}(\mu_i)$  will be set to zero
- b) if any angular distribution coefficient is zero (for example if  $a_{\bf k}$  = 0) then calculation of the corresponding functions will not be calculated but set to zero

$$\Psi_{R}^{m}(\mu) \equiv 0$$

$$\Phi_{L}^{m}(\mu) \equiv 0$$

The number of equations in a set for use in estimating the computer time or cost in I-56 or I-58 is equal to the number of non-zero terms in the phase function.

## c) Restrictions

At the time of compilation of the MAD program an upper limit must be placed on the following so that the computer may properly allocate storage to the various arrays.

- 1) Maximum value of N, i.e. the maximum number of terms in the phase function (MAXL)
- 2) Maximum number of consecutive approximations which may be stored in the computer at one time (NMAX). Iteration will proceed computing f\*(0), f\*(1),...,f\*(NMAX) after which f\*(0) will be set to f\*(NMAX) and the process repeated.
- 3) Maximum number of points to be used in numerical integration (MAXPT)

Large values of NMAX result in a slightly more efficient program, however NMAX may be any value greater than zero. The total storage required including the approximately 5000 locations required by the program and its subroutines is given by

STORAGE = 
$$2(NMAX + 1)(MAXL + 1)(MAXPT + 1)$$
  
+  $4(MAXL + 1)(MAXPT + 1) + 4(MAXPT) + 5000$ 

In practice, the large computer time required for large values of N and/or large values of p is a more serious restriction than the limitation on storage.

An additional restriction is that all numbers expressed in floating point form must be of the following range.

Thus, if  $\mu_{\text{min}}$  is the smallest non-zero value of  $\mu_{\text{i}}^{},$  then

$$e^{-\tau_i/\mu_{min}}$$
 must be greater than  $10^{-18}$ 

in order to evaluate  $\mathbf{F}_{\mathbf{K}}^{\mathbf{m}}(\boldsymbol{\tau}_{1},\boldsymbol{\mu}_{\mathbf{min}},\boldsymbol{\mu}_{\mathbf{min}})$  which is needed in performing some of the integrations. (Note: This does not apply if  $\boldsymbol{\tau}_{1} \geqslant 80$  as the  $\boldsymbol{\varnothing}_{1}^{\mathbf{m}}$  are not computed in this case.)

The only restriction on the integration procedure is that

$$\sum_{i=0}^{p} W_i = 1.0 \pm 10^{-7}, \qquad \sum_{i=0}^{p} W_i \mu_i = 0.5 \pm 10^{-7},$$

or in other words, the integrals

$$\int_a^1 dx = 1, \qquad \int_a^1 x dx = 0.5$$

must be done correctly.

d) General Outline of the Programs

The general outline of the programs is as follows:

- 1) A few initializing calculations are performed to give the computer information as to the maximum size of various arrays.
- 2) The header card is read followed by the basic data package.
- 3) Iterations are performed until convergence is indicated or until the allowable number of iterations is exceeded.
- 4) The results are punched and/or printed.

A printed copy of the MAD Program follows.

```
5TH10001
SCOMPILE MAD, PUNCH OBJECT, PRINT OBJECT
               GENERALIZED PROGRAM FOR SOLUTION OF LIGHT SCATTERING
          R
               INTEGRAL EQUATIONS
          R
           BOOLEAN
                     C1, C2, C4, C5, PCH, BL,BJ,BK,CONVG,THICK,OMIT,SIGN,
               CONVG1
           1
            INTEGER I,J,K,L,M,N,NPOINT,NTERMS,MAXIT,ITCNT,INT1,INT2,
              FIXARG, RUN, NMAX1, TYPE, TIME, , 10, 11, 12
           DIMENSION BLOOP (100)
           VECTOR VALUES NMAX = 3
           VECTOR VALUES MAXL = 3
           VECTOR VALUES MAXPT = 20
          R FOLLOWING DIMENSIONED (NMAX+1)*(MAXL+1)*(MAXPT+1) - 1
           DIMENSION PSI(335, PSIDIM), PHI(335, PHIDIM)
              FOLLOWING DIMENSIONED (MAXL+1)*(MAXPT+1) + 1
           DIMENSION S(83,SDIM),POL(83,POLDIM),CHPSI(83,CHPSID),
               CHPHI(83, CHPHID)
           R FOLLOWING DIMENSIONED MAXPT
           DIMENSION SS(20), W(20), U(20), BJ(20)
              FOLLOWING DIMENSIONED MAXL
           DIMENSION PSIDEV(3), PHIDEV(3), OMEGA(3),
                                                           A(3).BL(3).BK(3).
               OMIT(3)
           DIMENSION PSIDIM(3), PHIDIM(3), SDIM(2), POLDIM(2), CHPSID(2),
               CHPHID(2)
            VECTOR VALUES PSIDIM = 3
            VECTOR VALUES PHIDIM = 3
            VECTOR VALUES SDIM = 2
            VECTOR VALUES POLDIM = 2
            VECTOR VALUES CHPSID = 2
            VECTOR VALUES CHPHID = 2
           PSIDIm(3) = MAXPT + 1
            PSIDIm(2) = MAXL + 1
            PSIDIm(1) = PSIDIm(2)*PSIDIm(3) + PSIDIm(3) + 1
            PHIDIm(3) = MAXPT + 1
            PHIDIm(2) = MAXL + 1
            PHIDIm(1) = PHIDIm(2)*PHIDIm(3) + PHIDIm(3) + 1
            SDIM(2) = MAXPT + 1
            SDIM(1) = SDIM(2) + 1
            POLDIM(2) = MAXPT + 1
            POLDIm(1) = POLDIm(2) + 1
            CHPSID(2) = MAXPT + 1
            CHPSID(1) = CHPSID(2) + 1
            CHPHID(2) = MAXPT + 1
            CHPHID(1) = CHPHID(2) + 1
            EXECUTE ZERO (BL ... BL (MAXL) , BJ ... BJ (MAXPT) , BK ... BK (MAXL) )
START
            READ FORMAT INBOOL, I, BL...BL(I)
            READ FORMAT INBOOL, I,BJ...BJ(I)
            READ FORMAT INBOOL, I, BK...BK(I)
READ FORMAT IN1, MAXIT, TYPE, PCH, SIGN
            READ FORMAT INA, RUN, M, NTERMS, NPOINT, TAU, TOL, ALBEDO
            PRINT FORMAT FORM1, RUN, M, NTERMS, NPOINT, MAXIT, TOL, TAU, ALBEDO
            THICK = OB
           WHENEVER TAU •GE• 80•, THICK = 1B WHENEVER TYPE •E• 0
               PRINT FORMAT FORM9, $READ IN$
            OR WHENEVER TYPE .E. 1
               PRINT FORMAT FORM9, $CALCULATED FROM LEGENDRE POLYNOMIALS$
            END OF CONDITIONAL
```

```
READ FORMAT IN2, A...A(NTERMS)
           PRINT FORMAT FORM2
           PRINT FORMAT FF1, A...A(NTERMS)
           EXECUTE ZERO . (OMEGA . . . OMEGA (NTERMS) . OMIT . . . OMIT (NTERMS))
           THROUGH ST6, FOR L = M,1,L.G.NTERMS
                            A(L)*FACT.(L-M)/FACT.(L+M)*ALBEDO
           OMEGA(L) =
ST<sub>6</sub>
           WHENEVER OMEGA(L) \bulletE \bullet 0 \bullet 0 OMIT(L) = 1B
           PRINT FORMAT FORM3
           PRINT FORMAT FF1, OMEGA...OMEGA(NTERMS)
           READ FORMAT IN2, U...U(NPOINT)
           PRINT FORMAT FORM4
           PRINT FORMAT FF1, U...U(NPOINT)
           READ FORMAT IN2, W...W(NPOINT)
           PRINT FORMAT FORM5
           TEST = -0.5
           WEIGHT = -1.0
           THROUGH SL6, FOR I = 0,1,I.G.NPOINT
           TEST = TEST + W(I)*U(I)
           WEIGHT = WEIGHT + W(I)
SL6
           PRINT FORMAT FF1, W...W(NPOINT), WEIGHT, TEST
           WHENEVER .ABS.WEIGHT.G.0.0000001
              PRINT FORMAT BCD, $1SUM OF WEIGHTS NOT 1.0$
           OR WHENEVER .ABS.TEST.G.0.0000001
              PRINT FORMAT BCD, $1SUM OF W(I) *U(I) NOT 0.5$
           OR WHENEVER M.G. NTERMS
              PRINT FORMAT BCD, $1M IS GREATER THAN NTERMS$
           OTHERWISE
              TRANSFER TO AWAY
           END OF CONDITIONAL
           EXECUTE ERROR.
AWAY
           CONTINUE
           THROUGH S1, FOR L = M,1,L.G.NTERMS
           THROUGH SI, FOR J = 0,1,J.G.NPOINT
           POL(L,J) = LEG \cdot (U(J),M,L)
51
           PRINT FORMAT FORM6
           THROUGH ST3, FOR L =M,1,L.G.NTERMS
           PRINT FORMAT FF4,L,M,POL(L,0)...POL(L,NPOINT)
ST3
           WHENEVER TYPE .E. 0
           THROUGH SL1, FOR L = M,1,L.G.NTERMS
           READ FORMAT IN2, PSI(0,L,0)...PSI(0,L,NPOINT),PSIDEV(L)
SL1
           THROUGH SL2, FOR L = M,1,L.G.NTERMS
SL2
           READ FORMAT IN2, PHI(0,L,0)...PHI(0,L,NPOINT),PHIDEV(L)
           OR WHENEVER TYPE .E. 1
           THROUGH SL7A, FOR L = M,1,L.G.NTERMS
           PSIDEV(L) = 0.
           PHIDEV(L) = 0.
           WHENEVER OMIT(L)
              EXECUTE ZERO. (PSI(0,L,0)...PSI(0,L,NPOINT))
              EXECUTE ZERO.(PHI(0,L,0)...PHI(0,L,NPOINT))
               TRANSFER TO SL7A
           END OF CONDITIONAL
           THROUGH SL7, FOR J = 0,1,J.G.NPOINT
           PSI(0,L,J) = POL(L,J)
           WHENEVER U(J) .G. O.
                                  .AND. TAU/U(J).L.80.
              PHI(0,L,J) = EXP \cdot (-TAU/U(J)) *POL(L,J)
           OTHERWISE
              PHI(0,L,J) = 0.
SL7
           END OF CONDITIONAL
```

```
CONTINUE
SL7A
            OTHERWISE
            EXECUTE ERROR.
            END OF CONDITIONAL
            PRINT FORMAT FORM7
            THROUGH ST4, FOR L = M,1,L.G.NTERMS
            PRINT FORMAT FF2, L, M, PSI(0, L, O) ... PSI(0, L, NPOINT)
ST4
            PRINT FORMAT FORMS
            THROUGH ST5, FOR L = M,1,L.G.NTERMS
            PRINT FORMAT FF3,L,M,PHI(0,L,0)...PHI(0,L,NPOINT)
ST5
            CONVG = OB
            ITCNT = 1
            PRINT FORMAT BCD, $1BEGINNING OF ITERATIVE SOLUTION$
            TIME1 = TIME \cdot (0)
            THROUGH S6, FOR N = 1,1,N.G.NMAX.OR.CONVG.OR.ITCNT.G.MAXIT
LOOP
            EXECUTE ZERO · (PSIDEV(M) · · · PSIDEV(NTERMS))
            THROUGH A1, FOR L = M,1,L.G.NTERMS
            WHENEVER OMIT(L)
               EXECUTE ZERO.(PSI(N.L.O)...PSI(N.L.NPOINT))
               TRANSFER TO A1
            END OF CONDITIONAL
            THROUGH A2, FOR J = 0,1,J.G.NPOINT
            WHENEVER U(J) .E. O.
               PSI(N_1L_1J) = POL(L_1J)
               TRANSFER TO A5
            END OF CONDITIONAL
            EXECUTE SETCHK.
            SUMX = 0.
            THROUGH A3, FOR I = 0,1,1.G.NPOINT
            XXX = 0.
            THROUGH A4, FOR K = M,1,K.G.NTERMS
            WHENEVER OMIT(K), TRANSFER TO A4
            TS1 = (-1.0) \cdot P \cdot (K+L) *F \cdot (0)
            WHENEVER C1 \cdot S(K \cdot I) = TS1
            XXX = XXX + TS1
A4
            CONTINUE
            WHENEVER C2.SS(I) = XXX
            SUMX = SUMX + W(I)*XXX*POL(L*I)
A 3
            PSI(N_{\bullet}L_{\bullet}J) = POL(L_{\bullet}J) + 0.5*U(J)*SUMX
            TS1 = PSI(N+L+J) - PSI(N-1+L+J)
Α5
            EXECUTE CHECK.
            WHENEVER .ABS.TS1.G.PSIDEV(L), PSIDEV(L) = .ABS. TS1
A 2
            CONTINUE
A1
           R . . . . . . . .
            EXECUTE ZERO.(PHIDEV(M)...PHIDEV(NTERMS))
            WHENEVER THICK
            THROUGH SL10, FOR L = M,1,L.G.NTERMS
            EXECUTE ZERO (PHI(N,L,0) ... PHI(N,L, NPOINT))
SL10
            TRANSFER TO ON
            END OF CONDITIONAL
            THROUGH B1, FOR L = M,1,L.G.NTERMS
            WHENEVER OMIT(L)
                EXECUTE ZERO.(PHI(N,L,0)...PHI(N,L,NPOINT))
                TRANSFER TO B1
            END OF CONDITIONAL
            THROUGH B2, FOR J = 0,1,J.G.NPOINT
            WHENEVER U(J) .E. 0.
```

```
PHI(N,L,J) = 0.
           TRANSFER TO B5
OR WHENEVER TAU/U(J) .G. 80.
              EXPON = 0.
           OTHERWISE
              EXPON = EXP \cdot (-TAU/U(J))
           END OF CONDITIONAL
           EXECUTE SETCHK.
           SUMY = 0.
           THROUGH B3, FOR I = 0,1,1.G.NPOINT
           YYY = 0.
           THROUGH B4, FOR K = M,1,K.G.NTERMS
           WHENEVER OMIT(K), TRANSFER TO B4
           TS1 = G_{\bullet}(0)
           WHENEVER C4, S(K,I) = TS1
           YYY = YYY + TS1
B4
           CONTINUE
           WHENEVER C5, SS(I) = YYY
           SUMY = SUMY + W(I)*YYY*POL(L,I)
В3
           PHI(N,L,J) = EXPON*POL(L,J) + 0.5*U(J)*SUMY
           TS1 = PHI(N,L,J) - PHI(N-1,L,J)
B5
           EXECUTE CHECK.
R2
           WHENEVER .ABS.TS1 .G. PHIDEV(L), PHIDEV(L) = .ABS. TS1
В1
           CONTINUE
          R . . . . . . . .
                                       •••••••••••
           PRINT FORMAT HEAD1, ITCNT
ON
           PRINT FORMAT FF1, U...U(NPOINT)
           CONVG = 1B
           CONVG1 = OB
           THROUGH S7A, FOR L = M,1,L.G.NTERMS
           WHENEVER OMIT(L), TRANSFER TO STA
           THROUGH S11, FOR J = 0,1,J.G.NPOINT
           TS1 = PSI(N,L,J) - PSI(N-1,L,J)
           WHENEVER L.NE.O, TRANSFER TO S11
           WHENEVER TS1*CHPSI(L,J).L.O. .AND. ITCNT.G.1 .AND. SIGN,
              CONVG1 = 1B
511
           CHPSI(L,J) = TS1
           PRINT FORMAT FF2,L,M,CHPSI(L,0)...CHPSI(L,NPOINT),PSIDEV(L)
S7A
           WHENEVER PSIDEV(L).G.TOL, CONVG = OB
           THROUGH S7B, FOR L = M,1,L.G.NTERMS
           WHENEVER OMIT(L), TRANSFER TO S7B
           THROUGH S12, FOR J = 0,1,J.G.NPOINT
           TS1 = PHI(N_1,J) - PHI(N_1,J)
           WHENEVER L.NE.O, TRANSFER TO S12
           WHENEVER TS1*CHPHI(L,J).L.O. .AND. ITCNT.G.1 .AND. SIGN,
              CONVG1 = 1B
512
           CHPHI(L_J) = TS1
           PRINT FORMAT FF3, L, M, CHPHI(L, O)... CHPHI(L, NPOINT), PHIDEV(L)
S7B
           WHENEVER PHIDEV(L).G.TOL, CONVG = OB
           CONVG = CONVG .OR. CONVG1
56
           ITCNT = ITCNT + 1
           NMAX1 = N-1
           WHENEVER CONVG
              PRINT FORMAT BCD, $4 SUCCESSFUL SOLUTION$
           OR WHENEVER ITCHT .G. MAXIT
              PRINT FORMAT BCD, $4DESIRED ACCURACY COULD NOT BE OBTAINED
          1 IN SPECIFIED NUMBER OF ITERATIONS$
           OTHERWISE
```

```
THROUGH ALPHA5, FOR L = M,1,L.G.NTERMS
               THROUGH ALPHA5, FOR I = 0,1,1.G.NPOINT
               PSI(0,L,I) = PSI(NMAX1,L,I)
               PHI(0,L,I) = PHI(NMAX1,L,I)
ALPHA5
               TRANSFER TO LOOP
           END OF CONDITIONAL
           TIME2 = TIME \cdot (0)
            PRINT FORMAT FTIME, (TIME2-TIME1) *3.6
            PRINT FORMAT BCD, $0 VALUES OF FUNCTIONS AFTER LAST ITERATIONS
            N = NMAX1
            THROUGH S13, FOR L = M,1,L.G.NTERMS
           PRINT FORMAT FF2, L, M, PSI(N, L, O) ... PSI(N, L, NPOINT), PSIDEV(L)
513
            THROUGH S14, FOR L = M,1,L.G.NTERMS
            PRINT FORMAT FF3,L,M,PHI(N,L,0)...PHI(N,L,NPOINT),PHIDEV(L)
514
            WHENEVER PCH
            PRINT FORMAT BCD, $0RESULTS PUNCHED$
            INA(4) = RUN
            IN2(3) = RUN
            PUNCH FORMAT INA, RUN, M, NTERMS, NPOINT, TAU, TOL, ALBEDO
            PUNCH FORMAT IN2, A...A(NTERMS)
PUNCH FORMAT IN2, U...U(NPOINT)
            PUNCH FORMAT IN2, W...W(NPOINT)
            THROUGH SL3, FOR L = M,1,L.G.NTERMS
            PUNCH FORMAT IN2, PSI(NMAX1,L,0)...PSI(NMAX1,L,NPOINT),
SL3
               PSIDEV(L)
           1
            THROUGH SL4, FOR L = M,1,L.G.NTERMS
            PUNCH FORMAT IN2, PHI(NMAX1,L,0)...PHI(NMAX1,L,NPOINT),
SL4
              PHIDEV(L)
            END OF CONDITIONAL
            TRANSFER TO START
           R
               INTERNAL FUNCTIONS
           R
           R
            INTERNAL FUNCTION
            ENTRY TO F.
            FUNCTION RETURN OMEGA(K)*(PSI(N-1,K,J)*PSI(N-1,K,I)
                 - PHI(N-1,K,J)*PHI(N-1,K,I))/(U(I)+U(J))
            ENTRY TO G.
            WHENEVER I.NE.J
            FUNCTION RETURN OMEGA(K)*(PHI(N-1,K,J)*PSI(N-1,K,I)
                  - PSI(N-1,K,J)*PHI(N-1,K,I))/(U(J)-U(I))
           1
            OTHERWISE
               WHENEVER I.E.O
                   IO = 0
                   I1 = 1
                   I2 = 2
               OR WHENEVER I.E.NPOINT
                   IO = NPOINT - 2
                   I1 = NPOINT - 1
                   I2 = NPOINT
               OTHERWISE
                   IO = I-1
                   I1 = I
                   I2 = I+1
               END OF CONDITIONAL
            X = U(I)
            \chi_0 = \chi(10)
```

```
X1 = \psi(I1)
           X2 = U(I2)
           FO = PHI(N-1,K,IO)
           F1 = PHI(N-1,K,I1)
           F2 = PHI(N-1,K,I2)
           PHPR = (F2-F0)/(X2-X0) + (2.*X-X0-X2)*
              ((x2-x1)*F0 - (x2-x0)*F1 + (x1-x0)*F2)/
              ((x1-x0)*(x2-x0)*(x2-x1))
           FO = PSI(N-1,K,IO)
           F1 = PSI(N-1,K,I1)
           F2 = PSI(N-1,K,I2)
           PSPR = (F2-F0)/(X2-X0) + (2*X-X0-X2)*
              ((x2-x1)*F0 - (x2-x0)*F1 + (x1-x0)*F2)/
              ((x1-x0)*(x2-x0)*(x2-x1))
           FUNCTION RETURN OMEGA(K)*(PSI(N-1,K,J)*PHPR
              - PHI(N-1,K,J)*PSPR)
           END OF CONDITIONAL
           END OF FUNCTION
           INTERNAL FUNCTION (FIXARG)
           INTEGER FIXARG
           ENTRY TO FACT.
           TS1 = 1.
           THROUGH ALPHA2, FOR INT1 = 1,1,INT1.G.FIXARG
ALPHA2
           TS1 = TS1*INT1
           FUNCTION RETURN TS1
           END OF FUNCTION
          R
           INTERNAL FUNCTION
           ENTRY TO CHECK.
           WHENEVER BJ(J) . AND . BL(L)
              PRINT FORMAT NGFO, $J$, J, U(J)
              THROUGH SL5, FOR K = M,1,K.G.NTERMS
SL5
              WHENEVER BK(K), PRINT FORMAT NGEO, $K$, K,
                 S(K,0)...S(K,NPOINT)
              PRINT FORMAT GEO,
                                    J.SS...SS(NPOINT),TS1,
                 PSI(N,L,J),PHI(N,L,J)
           END OF CONDITIONAL
           FUNCTION RETURN
           ENTRY TO SETCHK.
           WHENEVER BL(L) . AND . BJ(J)
              C1 = 1B
              C2 = 1B
              C4 = 1B
              C5 = 1B
           OTHERWISE
              C1 = 0B
              C2 = 0B
              C4 = 0B
              C5 = OB
           END OF CONDITIONAL
           FUNCTION RETURN
           END OF FUNCTION
          R
          R
              FORMAT SPECIFICATIONS AND PRE-SET CONSTANTS
          R
          VECTOR VALUES INA = $C5,315$,$,3F10.$,$5,$22,$,$
                                                                 5H$,$ $,
```

```
1 $*$
 VECTOR VALUES IN1 = $415,5F10.5*$
VECTOR VALUES IN2 = $1P5E14$,$.7,$2,$,$
VECTOR VALUES INBOOL = $15,$5,6011/(7011)*$
                                               5H$.$ $,$*$
 VECTOR VALUES HEAD1 = $37HOCHANGE IN FUNCTION DURING ITERATIO
       13*$
 VECTOR VALUES GIO = $1H0I9,11110/(S10,11110)*$
 VECTOR VALUES GFO = $1H0,15,84,10F11.6/($10,10F11.6)*$
VECTOR VALUES GEO = $1H0,15,84,1P10E11.3/($10,1P10E11.3)*$
 VECTOR VALUES BCD = $20C6*$
 VECTOR VALUES NGFO = $1H0,C1,2H =,13,S3,10F11.6/
    (S10,10F11.6)*$
 VECTOR VALUES NGEO = $1H0,C1,2H =,13,S3,1P10E11.3/
    (S<sub>1</sub>0,1P10E11.3)*$
 VECTOR VALUES FORM1 = $40H1SOLUTION OF INTEGRAL EQUATIONS FOR
        C5//6H M = I5,S5,18HNUMBER OF TERMS =
                                                       15,55,
219HNUMBER OF POINTS = 15/32H MAXIMUM NUMBER OF ITERATIONS
3= I5,55,33HDESIRED DEGREE OF CONVERGENCE = 1PE10.1/
4 7H TAU =
               1PE11.3,S5,32H ALBEDO FOR SINGLE SCATTERING =
   1PE11.3*$
 VECTOR VALUES FTIME = $19HOTIME REQUIRED WAS
                                                      F6.1,
1 8H SECONDS *$
 VECTOR VALUES FORM2 = $30H0PHASE FUNCTION--A...A(NTERMS)
 VECTOR VALUES FORM3 =
1 $47HOMODIFIED PHASE FUNCTION--OMEGA...OMEGA(NTERMS)
 VECTOR VALUES FORM4 = $28HOVALUES OF MU--U...U(NPOINT)
                                                               *5
 VECTOR VALUES FORM5 = $30HOWEIGHT FACTORS--W...W(NPOINT)
 VECTOR VALUES FORM6 =
1 $47HOLEGENDRE POLYNOMIALS--P(L,M) FOR U...U(NPOINT)
 VECTOR VALUES FORM7 =
1$47HOINITIAL FUNCTIONS--PSI(L,M) FOR U...U(NPOINT)
 VECTOR VALUES FORM8 =
1$47H0INITIAL FUNCTIONS--PHI(L.M) FOR U...U(NPOINT)
                                                          *$
 VECTOR VALUES FORM9 = $19HOINITIAL FUNCTIONS 8C6*$
 VECTOR VALUES FF1 = $1H0, $9, 1P9E12.4/($10, 1P9E12.4) *$
 VECTOR VALUES FF2 = $5HOPSI( II, 1H, II, 2H) 1P 9E12.4/
1 (S10,1P 9E12.4)*$
 VECTOR VALUES FF3 = $5H0PHI( II, 1H, II, 2H) 1P 9E12.4/
   (S10,1P 9E12.4)*$
 VECTOR VALUES FF4 = $5H0 P( II, 1H, II, 2H) 1P 9E12.4/
1 (S10+1P 9E12+4)*$
 END OF PROGRAM
```

2. Subroutine to Calculate Associated Legendre Polynomials  $A \text{ subroutine was written to calculate } P_1^m(x) \text{.}$  The program should be self explanatory.

```
LEG 001
*COMPILE MAD, PUNCH OBJECT
            EXTERNAL FUNCTION (X+M+N)
            INTEGER N.M.K.FIXARG, INT1, TWOLOG.
            ENTRY TO LEG.
            WHENEVER X.G.1. .OR. X .L.O. .OR. M.G. N
               PRINT FORMAT BCD, $4BAD ARGUMENT FOR LEGENDRE POLYNOMIAL SU
           1BROUTINE$
               PRINT FORMAT FORMA, X, M, N
               ERROR RETURN
            END OF CONDITIONAL
            U = (1 \cdot -X)/2 \cdot
            T = F_ACT \cdot (N+M)/(FACT \cdot (M)*FACT \cdot (N-M))
            SUM = T
            THROUGH S1, FOR K = 1,1,K,G,N-M
            T = -T*(N+M+K)*(N-M-K+1) / (K*(M+K))
            WHENEVER TWOLOG. (T) + TWOLOG. (U) .L. -115,
           1 TRANSFER TO OUT
            T = T*U
            SUM = SUM + T
S1
OUT
            WHENEVER M.E.O. FUNCTION RETURN SUM
            WHENEVER X.E.1., FUNCTION RETURN O.
            V = SQRT \cdot (1 \cdot -X \cdot X)/2 \cdot
            FUNCTION RETURN SUM*V.P.M
            INTERNAL FUNCTION (FIXARG)
            ENTRY TO FACT.
            TS1 = 1.
            THROUGH BETA2, FOR INT1 = 1,1,INT1.G.FIXARG
BETA2
            TS1 = TS1*INT1
            FUNCTION RETURN TS1
            END OF FUNCTION
            VECTOR VALUES BCD = $20C6*$
            VECTOR VALUES FORMA = $1H0 1PE15.6.2I15*$
            END OF FUNCTION
*ASSEMBLE, PUNCH OBJECT
                                                                      2LOG 001
        ORG 0
        PGM
        PZE SIZE
        PZE
        BCD 1TWOLOG
        PZE START
        REL
        ORG 0
START
        CAL 1,4
        STA *+1
        CLA **
        SSP
        SUB TWO
        ARS 27
        TRA 2,4
OCT 2000000000000
TWO
        SYN *
SIZE
        END
```

```
REF 001
*COMPILE MAD, PUNCH OBJECT , PRINT OBJECT
               INTEGRATED REFLECTANCE PROGRAM
          R
          R
           INTEGER I, J, K, L, M, N, NPOINT, NTERMS, INT1, INT2, FIXARG, RUN
           BOOLEAN PCH
           VECTOR VALUES NMAX = 0
           VECTOR VALUES MAXL = 6
           VECTOR VALUES MAXPT = 50
          R FOLLOWING DIMENSIONED (NMAX+1)*(MAXL+1)*(MAXPT+1) - 1
           DIMENSION PSI(356, PSIDIM), PHI(356, PHIDIM)
          R FOLLOWING DIMENSIONED MAXPT
           DIMENSION W(50), U(50), T(50), R(50)
          R FOLLOWING DIMENSIONED MAXL
           DIMENSION OMEGA(6), A(6), PSIDEV(6), PHIDEV(6)
           DIMENSION PSIDIM(3), PHIDIM(3)
           VECTOR VALUES PSIDIM = 3
           VECTOR VALUES PHIDIM = 3
           PSIDIm(3) = MAXPT + 1
           PSIDIM(2) = MAXL + 1
           PSIDIM(1) = PSIDIM(2)*PSIDIM(3) + PSIDIM(3) + 1
           PHIDIm(3) = MAXPT + 1
           PHIDIm(2) = MAXL + 1
           PHIDIm(1) = PHIDIm(2)*PHIDIm(3) + PHIDIm(3) + 1
            READ FORMAT IN1, PCH
START
           READ FORMAT IN1, RUN, M, NTERMS, NPOINT, TAU, TOL, ALBEDO
            WHENEVER M.NE. 0
               PRINT FORMAT BCD, $1M NOT ZERO$
               EXECUTE ERROR.
            END OF CONDITIONAL
            READ FORMAT IN2, A...A(NTERMS)
            EXECUTE ZERO . (OMEGA . . . OMEGA (NTERMS))
            THROUGH ST6, FOR L = M,1,L.G.NTERMS
            OMEGA(L) = A(L)*FACT*(L-M)/FACT*(L+M)*ALBEDO
ST6
            READ FORMAT IN2, U...U(NPOINT)
            READ FORMAT IN2, W. . . W(NPOINT)
            THROUGH SL1, FOR L = M,1,L.G.NTERMS
            READ FORMAT IN2, PSI(0,L,0)...PSI(0,L,NPOINT),PSIDEV(L)
SL1
            THROUGH SL2, FOR L = M,1,L.G.NTERMS
            READ FORMAT IN2, PHI(0,L,0)...PHI(0,L,NPOINT),PHIDEV(L)
SL2
            PRINT FORMAT BCD, $11NPUT DATA$
            PRINT FORMAT GIO, NPOINT, M, NTERMS, RUN
            PRINT FORMAT GFO, NTERMS, OMEGA...OMEGA(NTERMS)
PRINT FORMAT GFO, 0, U...U(NPOINT)
            PRINT FORMAT GFO,0, W... W(NPOINT)
            PRINT FORMAT GEO, 0, TAU, TOL
            THROUGH ST3, FOR L = M.1.L.G.NTERMS
ST3
            PRINT FORMAT GEO, L, PSI(0, L, 0)... PSI(0, L, NPOINT), PSIDEV(L)
            THROUGH ST4, FOR L = M,1,L.G.NTERMS
            PRINT FORMAT GEO, L, PHI(0, L, 0)... PHI(0, L, NPOINT), PHIDEV(L)
ST4
            THROUGH A2, FOR J = 0,1,J.G.NPOINT
            SUMX = 0.
            SUMY = 0.
            THROUGH A3, FOR I = 0,1,I,G,NPOINT
            XXX = 0.
            YYY = 0.
            THROUGH A4, FOR K = M,1,K.G.NTERMS
            XXX = XXX + (-1.0).P.K * F.(0)
            YYY = YYY + G_{\bullet}(0)
A4
```

```
SUMX = SUMX + W(I)*U(I)*XXX
            SUMY = SUMY + W(I)*U(I)*YYY
A3
           R(J) = 0.5*SUMX
            T(J) = 0.5*SUMY
A 2
           PRINT FORMAT BCD, SOREFLECTION AND TRANSMISSION COEFFICIENTSS
           PRINT FORMAT GEO, 1, R. . . . R(NPOINT)
           PRINT FORMAT GEO, 2, T. . . . T(NPOINT)
           WHENEVER PCH
               PUNCH FORMAT IN1, RUN, M, NTERMS, NPOINT, TAU, TOL, ALBEDO
               PUNCH FORMAT IN2, R...R(NPOINT)
               PUNCH FORMAT IN2, T...T(NPOINT)
           END OF CONDITIONAL
           TRANSFER TO START
               INTERNAL FUNCTIONS
          R
           INTERNAL FUNCTION
           INTEGER IO+I1+I2
           ENTRY TO F.
           FUNCTION RETURN OMEGA(K)*(PSI(N-1,K,J)*PSI(N-1,K,I)
                 - PHI(N-1*K*J)*PHI(N-1*K*I))/(U(I)+U(J))
           1
          R
           ENTRY TO G.
           WHENEVER I.NE.J
           FUNCTION RETURN OMEGA(K)*(PHI(N-1,K,J)*PSI(N-1,K,I)
                 - PSI(N-1,K,J)*PHI(N-1,K,I))/(U(J)-U(I))
           OTHERWISE
               WHENEVER I.E.O
                  IO = 0
                  I1 = 1
                  I2 = 2
               OR WHENEVER I.E.NPOINT
                  IO = NPOINT - 2
                  I1 = NPOINT - 1
                  I2 = NPOINT
               OTHERWISE
                  IO = I-1
                  I1 = I
                  I2 = I+1
              END OF CONDITIONAL
           X = U(I)
           xo = v(10)
           X1 = \psi(I1)
           X2 = U(12)
           FO = PHI(N-1,K,IO)
           F1 = pHI(N-1,K,I1)
           F2 = PHI(N-1,K,I2)
           PHPR = (F2-F0)/(X2-X0) + (2*X-X0-X2)*
              ((x2-x1)*F0 - (x2-x0)*F1 + (x1-x0)*F2)/
              ((\chi 1-\chi 0)*(\chi 2-\chi 0)*(\chi 2-\chi 1))
           FO = PSI(N-1,K,IO)
           F1 = psi(N-1,K,I1)
           F2 = PSI(N-1,K,I2)
           PSPR = (F2-F0)/(X2-X0) + (2*X-X0-X2)*
              ((\chi 2-\chi 1)*F0 - (\chi 2-\chi 0)*F1 + (\chi 1-\chi 0)*F2)/
               ((x1-x0)*(x2-x0)*(x2-x1))
           FUNCTION RETURN OMEGA(K)*(PSI(N-1,K,J)*PHPR
```

- PHI(N-1,K,J)\*PSPR)

```
END OF CONDITIONAL
             END OF FUNCTION
            R
             INTERNAL FUNCTION (FIXARG)
             INTEGER FIXARG
             ENTRY TO FACT.
             TS1 = 1.
             THROUGH ALPHA2, FOR INT1 = 1,1,INT1.G.FIXARG
ALPHA2
             TS1 = TS1*INT1
             FUNCTION RETURN TS1
             END OF FUNCTION
                 FORMAT SPECIFICATIONS AND PRE-SET CONSTANTS
            R
             VECTOR VALUES IN1 = $415,5F10.5*$
             VECTOR VALUES IN2 = $1P5E14.7,52*$
             VECTOR VALUES GIO = $1H019,11110/(S10,11110)*$
             VECTOR VALUES GFO = $1H0, I5, S4, 10F11.6/(S10, 10F11.6) *$

VECTOR VALUES GEO = $1H0, I5, S4, 1P10E11.3/(S10, 1P10E11.3) *$

VECTOR VALUES BCD = $20C6*$
             VECTOR VALUES N = 1
             END OF PROGRAM
```

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