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# **Geophysical Research Letters**

### **RESEARCH LETTER**

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#### **Key Points:**

- A narrow, high seismic velocity anomaly near the 660 km phase boundary splits a mantle-wide low seismic velocity anomaly beneath Samoa
- Harzburgite enrichment below the base of transition zone within a continuous hot thermochemical upwelling can explain the "Samoa gap'
- Anomaly strength can be matched by a harzburgite fraction of at least 0.925 and a temperature elevated by 125-175°C

#### **Supporting Information:**

- Data Set S1
- Supporting Information S1

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# Signals of 660 km topography and harzburgite enrichment in seismic images of whole-mantle upwellings

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Abstract Various changes in seismic structures across the mantle transition zone (MTZ) indicate that it may hamper thermal and chemical circulation. Here we show how thermal elevation of the postspinel phase transition at 660 km depth plus harzburgite segregation below this depth can project as narrow high-velocity anomalies in tomographic images of continuous thermochemical mantle upwellings. Model S40RTS features a narrow high-velocity anomaly of +0.8% near 660 km depth within the broad low-velocity structure beneath the Samoa hot spot. Our analyses indicate that elevation of the 660 phase boundary in a hot pyrolitic plume alone is insufficient to explain this anomaly. An additional effect of harzburgite enrichment is required and consistent with geodynamic simulations that predict compositional segregation in the MTZ, especially within thermochemical upwellings. The Samoa anomaly can be modeled with a 125-175°C excess temperature and a harzburgite enrichment below 660 of least 60% compared to a pyrolitic mantle.

### 1. Introduction

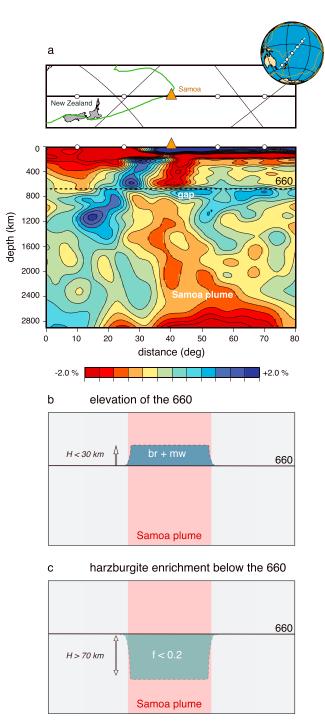
It is well established that the upper mantle transition zone has a profound influence on the structure of mantle flow. Seismic tomography has been successful in imaging the variable descent of subducting slabs into the deep mantle [e.g., Grand, 1994; Sigloch and Mihalynuk, 2013; Fukao and Obayashi, 2013], but tomographic constraints on the origin of mantle upwellings have remained ambiguous [e.g., Montelli et al., 2004; Wolfe et al., 2009; Styles et al., 2011; French and Romanowicz, 2015].

In this paper, we investigate the effects of compositional layering in the upper mantle transition zone on images of mantle upwellings. We consider a thermochemical upwelling from the lower mantle that has transported compositionally distinct material into the transition zone [e.g., Xie and Tackley, 2004; Brandenburg and van Keken, 2007; Nakagawa et al., 2010]. The layering originates from thermal perturbations of the 660 km phase boundary (i.e., the 660) and from the segregation of basaltic and harzburgitic components with intrinsically different densities [Irifune and Ringwood, 1993]. Using a forward modeling approach, we illustrate how anomalous layering near the 660 can project as a high-velocity seismic anomaly embedded within a whole-mantle low-velocity structure and be reminiscent of discontinuous flow across the transition zone.

Figure 1a is the working example of our analysis. It shows the shear velocity structure in the mantle beneath the southwestern Pacific according to S40RTS [Ritsema et al., 2011]. The dipping high-velocity anomaly between 10° and 35° is the subducting Pacific Plate. A broad low-velocity anomaly extends from the core-mantle boundary to the surface beneath the Samoa hot spot. We interpret this mantle-wide structure as a large-scale mantle upwelling related to hot spot volcanism on Samoa and call it the Samoa plume for simplicity. There is a gap in the Samoa plume near the base of the upper mantle transition zone, manifested as a positive wave speed anomaly with a maximum amplitude of  $\delta V_{\rm c} = 0.8\%$ . We will call this the Samoa gap from here on. The Samoa gap may imply that upward mantle flow is blocked near the 660. Here we hypothesize that the Samoa gap is due to the thermal elevation of the 660 and compositional heterogeneity around the base of the transition zone within a continuous thermochemical upwelling. Our modeling is informed by seismic estimates of 660 topography, geodynamic simulations of mantle mixing, and estimates of image resolution in tomographic model S40RTS.



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**Figure 1.** (a) Vertical, SW-NE oriented cross section through the shear velocity model S40RTS [*Ritsema et al.*, 2011] centered on the Samoa hot spot. The Samoa plume is a broad low shear velocity anomaly from the core-mantle boundary to the surface and assumed to be a hot thermal upwelling. A high-velocity anomaly breaks the Samoa plume near the 660 km discontinuity (dashed line). This feature is called the Samoa gap in this paper. (b) Sketch of the expected 660 km phase boundary elevation due to the increased temperature in the upper mantle beneath Samoa. (c) Sketch of a layer in the uppermost lower mantle with a harzburgite-enriched composition. The 660 elevation (in Figure 1b) and the harzburgite-enriched layer (in Figure 1c) may be observed as high-velocity anomalies.

#### 2. Models of the Samoa Gap

#### 2.1. Temperature-Induced Phase Boundary Topography

The mineral phase transformation of ringwoodite (*ri*) to the postspinel phases bridgemanite (*br*) plus magnesiowüstite (*mw*) is responsible for the deepest global seismic discontinuity in the upper mantle at 660 km depth. Recent estimates indicate that this transition has a negative Clapeyron slope  $-2.9 \le \gamma_{660} \le -2.1$  MPa K<sup>-1</sup> [*Ye et al.*, 2014]. Hence, a temperature increase in the upper mantle of  $\Delta T = 250^{\circ}$ C would elevate the 660 by 13–18 km. Analyses of *P* and *S* wave reflections [e.g., *Flanagan and Shearer*, 1998; *Gu and Dziewonski*, 2002; *Deuss*, 2009] and conversions [e.g., *Schmandt et al.*, 2012; *Mulibo and Nyblade*, 2013; *Jenkins et al.*, 2016] indicate that topography on the 660 in the mantle is as large as 30 km.

In seismic images (Figure 1b), a locally elevated 660 would be visible as a thin high-velocity anomaly with respect to a mantle in which the 660 is unperturbed. Its vertical width is equivalent to the elevation of the 660, and the velocity contrast is determined by the shear velocity increase across the 660. The expected concurrent depression of the 410 due to the exothermic phase transition around that depth would produce a low-velocity anomaly. We ignore such an anomaly because it would likely not be observable within a large-scale low-velocity anomaly (i.e., the Samoa plume).

#### 2.2. Harzburgite Enrichment Below the 660

Melting of mantle peridotite generates a stratified oceanic lithosphere with layers of basalt and harzburgite. After plates subduct these two components are recycled back into the mantle to become a folded and stretched mechanical mixture (see *Xu et al.* [2008] for a discussion). Since harzburgite is denser than mid-ocean ridge basalt for a narrow depth range (~100 km) beneath the 660, the two components may segregate near the 660 [*Irifune and Ringwood*, 1993]. As a result the lower mantle will have a harzburgite-enriched composition between about 660 and 800 km depth and basalt enrichment directly above 660. Numerical simulations of thermochemical mantle convection [e.g., *Xie and Tackley*, 2004; *Brandenburg and van Keken*, 2007; *Nakagawa et al.*, 2010] demonstrate that a compositional gradient forms in the mantle and that basalt-harzburgite partitioning can be particularly strong within upwelling regions of the mantle and when  $\gamma_{660}$  or the density contrast between basalt and harzburgite are high [e.g., *van Summeren et al.*, 2009]. The shear velocity of a basalt-enriched composition above 660 will be lower than the background mantle and would contribute to the overall low velocities of a hot upwelling. By contrast, the shear speed in a harzburgite-enriched layer beneath the 660 is higher than in a mantle with a pyrolite composition, as illustrated in Figure 1c.

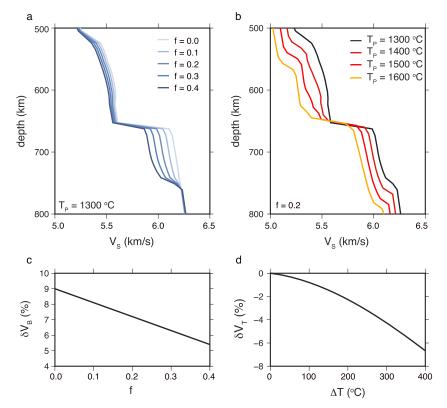
#### 3. Analysis

#### **3.1. Mineral Physics Constraints**

Experimental mineral physics results constrain our seismic models of the elevation of the 660 and harzburgite enrichment in the uppermost lower mantle. We compute theoretical profiles of shear velocity as described in *Cobden et al.* [2008], i.e., using PerPle\_X [*Connolly*, 2005], with the thermodynamic parameter database from *Stixrude and Lithgow-Bertelloni* [2011] and estimates of temperature and pressure-dependent anelasticity from *Goes et al.* [2004]. As in *Xu et al.* [2008], the mantle is composed of the Na-Ca-Fe-Mg-Al-Si (i.e, NCFMAS) oxides and regarded as a mechanical mixture of basalt and harzburgite in proportions *f* and 1–*f*, respectively. We calculate the reference profile for an adiabat with a potential temperature of 1300°C, suitable for the convective mid-ocean ridge basalt source mantle, and for a basalt fraction *f* = 0.2, roughly equivalent to the composition of pyrolite. We compute the elevation of the 660 due to a temperature increase using a Clapeyron slope of  $\gamma_{660} = -2.9$  MPa K<sup>-1</sup> and velocity anomalies due to changes in the composition of the uppermost lower mantle by changing *f*.

Figure 2a compares profiles of shear velocity for basalt fractions f between 0 and 0.4 and an adiabat for the reference potential temperature of 1300°C. In a purely harzburgitic mantle (i.e., f = 0), the 660 phase transition is entirely controlled by the  $ri \rightarrow bm + mw$  transition with a shear velocity jump of 9.0% at the 660. The shear velocity is up to 2% higher than in a pyrolitic mantle between 660 km and 760 km depth. In a mechanical mixture of basalt and harzburgite (i.e., f > 0), the phase transformation at the 660 is distributed over a finite pressure range because the  $gt \rightarrow mw$  in basalt occurs near 760 km depth. The shear velocity jump at the 660 decreases with increasing basalt fraction f from 7.2% for f = 0.2 to 5.4% for f = 0.4. Figure 2b compares profiles of shear velocity for adiabats with potential temperatures between 1300°C (our reference geotherm) and 1600°C (expected within a hot mantle upwelling). The basalt fraction f = 0.2. The elevation of the 660 increases with temperature, but the velocity increase across the 660 is not very sensitive to temperature.

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**Figure 2.** Shear velocity profiles calculated for mechanical mixtures of basalt and harzburgite in proportions f and 1 - f, respectively. (a) The basalt fraction f is varied from 0 to 0.4. The geotherm is an adiabat with a potential temperature of 1300°C. (b) The potential temperature is varied between 1300°C and 1600°C. The basalt fraction f = 0.2. (c) The shear velocity increase  $\delta V_B$  across the 660 as a function of basalt fraction f. (d) The shear velocity decrease  $\delta V_T$  in the uppermost lower mantle as a function of temperature increase  $\Delta T$ .

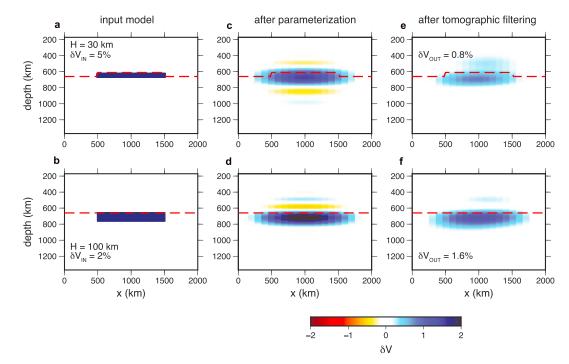
#### 3.2. Model Parameterization

We parameterize the elevation of the 660 and a layer below the 660 with a harzburgite-enriched composition as narrow rectangular blocks. The rectangular blocks have horizontal side lengths of 1000 km which correspond to the minimum horizontal scale that can be resolved by S40RTS. The vertical thickness *H* and the uniform velocity perturbation  $\delta V_{IN}$  are free parameters.

For a model that represents a 660 perturbation, *H* corresponds to the elevation of the 660. The shear velocity jump  $\delta V_{\rm B}$  at the phase boundary (see Figure 2c) and the shear velocity reduction  $\delta V_{\rm T}$  due to the increased temperature determine  $\delta V_{\rm IN}$  (see Figure 2d).  $\delta V_{\rm IN} = \delta V_{\rm B} + \delta V_{\rm T}$ . For a model that represents a harzburgite-enriched layer beneath the 660, *H* is the layer thickness and the shear velocity anomaly  $\delta V_{\rm IN}$  depends on the composition ( $\delta V_{\rm C}$ ) and temperature  $\delta V_{\rm T}$ . Thus,  $\delta V_{\rm IN} = \delta V_{\rm C} + \delta V_{\rm T}$ , where  $\delta V_{\rm C}$  and  $\delta V_{\rm T}$  have opposite signs. If the layer is pure harzburgite (f = 0) and there is no temperature anomaly, then  $\delta V_{\rm IN}$  is about 2%. If  $\Delta T$  exceeds about 200°C, the shear velocity in the harzburgite-enriched layer is lower than that in the reference model and the layer may not be visible.

#### 3.3. Tomographic Filtering

To estimate how phase boundary topography and harzburgite enrichment in the uppermost lower mantle would be imaged tomographically, we use the model resolution matrix  $\mathcal{R}$  of S40RTS. We first project the test structure (i.e., the rectangular block) into the model parameterization of S40RTS, which consists of spherical harmonics up to degree 40, and 21 vertical spline basis functions. After projection into S40RTS parameterization the linear operator  $\mathcal{R}$  is applied to produce a tomographically filtered version of the input model. Application of  $\mathcal{R}$  distorts and dampens the input model due to incomplete and heterogeneous data coverage, and model regularization, but it does not include the effects of inaccurate forward modeling [*Ritsema et al.*, 2007].



**Figure 3.** Resolution test showing how a rectangular block-shaped velocity anomaly (a) above the 660 (the red dashed line) and (b) below the 660 would be imaged in S40RTS. The anomaly has horizontal side lengths of 1000 km. In Figure 3a the thickness H = 30 km and  $\delta V_{IN} = 5\%$ . In Figure 3b the thickness H = 100 km and  $\delta V_{IN} = 2\%$ . The anomalies are drawn with vertical exaggeration for clarity. (c and d) The anomalies after projection into S40RTS parameterization. (e and f) The anomalies after reparameterization and filtering by  $\mathcal{R}$ . The highest recovered anomaly in tomographically filtered model is  $\delta V_{OUT}$ .

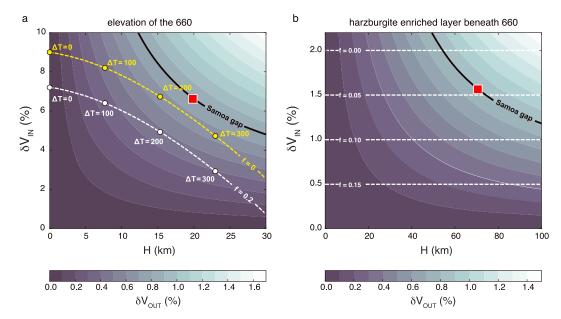
Figure 3 shows how a 30 km elevation of the 660 (in Figure 3a) and a 100 km thick layer of compositional heterogeneity below the 660 (in Figure 3b) would be imaged in S40RTS. After projection into S40RTS parameterization (Figures 3c and 3d) the high-velocity rectangular blocks are thicker and the velocity anomalies are weaker because the spacing of the vertical splines in S40RTS is large compared to *H*. The amplitude reduction is strongest for the thinnest layer. After filtering (Figures 3e and 3f), the velocity perturbations have been reduced further by a factor of about 2.

#### 4. Results

The contours in Figure 4 show how the peak recovered velocity anomaly, which we refer to as  $\delta V_{OUT}$ , varies as a function of H and  $\delta V_{IN}$ . The  $\delta V_{OUT}$  depends linearly on  $\delta V_{IN}$  and nonlinearly on H. It is highest for the thickest layers when reparameterization affects the amplitude reduction the least.

In Figure 4a, the highest value for  $\delta V_{OUT}$  of 1.7% is obtained when the 660 is elevated by 30 km and the shear velocity jump across the 660 is 10%. The Samoa gap of 0.8% (see Figure 1) can be explained if the 660 is elevated by at least 15 km. For a shear velocity jump as small as 5%, the 660 elevation must be 25 km or more. The smallest shear velocity perturbation  $\delta V_{\rm IN}$  in combination with the smallest *H* for which  $\delta V_{\rm OUT} = 0.8\%$  are 7% and 18 km, respectively. The combinations of *H* and  $\delta V_{\rm IN}$  consistent with a temperature-induced elevation of the 660, as discussed in section 3.1, are indicated by dashed lines in Figure 4a. We consider a mantle mixture with a pyrolitic composition (f = 0.2) and a harzburgitic mantle (f = 0) and assume that  $\gamma_{660} = -2.9$  MPa K<sup>-1</sup>. The highest values of  $\delta V_{OUT}$  are obtained when  $\Delta T$  is about 200–250°C, depending on composition. Within this temperature range, the 660 elevation is about 15–20 km.  $\delta V_{OUT}$  approaches 0.8% if the mantle is composed of pure harzburgite, but it is smaller than 0.5% for a pyrolitic mantle.

Figure 4b shows that the Samoa gap of 0.8% can be better explained by a harzburgite-enriched layer with a thickness of at least 50 km. The highest value for  $\delta V_{OUT}$  of 1.7% is obtained when shear velocity jump is 2.1% higher than in the ambient mantle within 100 km thick layer below the 660. The smallest values for  $\delta V_{IN}$ 



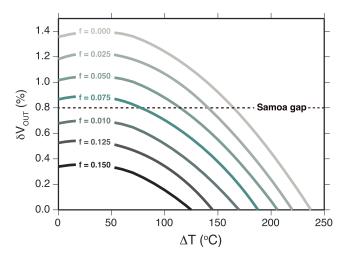
**Figure 4.** Contours of the peak recovered velocity anomaly  $\delta V_{OUT}$  obtained by tomographic filtering of input models using  $\mathcal{R}$ . An input model is defined by the assumed layer thickness H (along the x axis) and velocity anomaly  $\delta V_{IN}$  (along the y axis) and represents (a) an elevation of the 660 or (b) a layer in the uppermost lower mantle with a harzburgite-enriched composition. The 0.8% contour corresponds to the Samoa gap near the 660 within the Samoa plume (see Figure 1). The red square is a corner point where  $\delta V_{OUT} = 0.8\%$  for the smallest values of  $\delta V_{IN}$  and H. In Figure 4a, the dashed lines show the combinations of  $\delta V_{IN}$  and H consistent with an elevation of the 660 due to the presence of a temperature  $\Delta T$ , indicated with solid circles. The yellow and white lines correspond to assumed basalt fractions of f = 0 and f = 0.2, respectively. In Figure 4b, the dashed lines show the values of  $\delta V_{IN}$  consistent with harzburgite enrichment below the 660 for f = 0, f = 0.05, f = 0.10, and f = 0.15.

and *H* for which  $\delta V_{OUT} = 0.8\%$  are 1.5% and 70 km, respectively. The shear velocity increase  $\delta V_{IN}$  in this layer decreases with increasing basalt fraction *f*. If the layer has a thickness of 100 km, the harzburgite fraction must be higher than 0.925 to match the Samoa gap of 0.8%.

#### 5. Discussion and Conclusions

In this paper, we have demonstrated that thermal elevations of the postspinel phase transition at 660 km depth and basalt segregation at the top of the lower mantle can project as narrow high-velocity anomalies in tomographic images of continuous thermochemical mantle upwellings. Even though our analysis used postspinel Clapeyron slopes on the high end of those determined from mineral physics [*Ye et al.*, 2014; *Hirose*, 2002; *Weidner and Wang*, 1998], the elevation of the 660 alone is not sufficient to explain the relatively high shear velocity in the Samoa gap. This is because of the competing effects of the phase transition and temperature on shear velocity in a pyrolytic mantle. A harzburgite-enriched layer within the uppermost lower mantle is an essential feature of the model. It is consistent with geodynamic simulations of mantle mixing which predict strong compositional layering around the 660 [e.g., *Xie and Tackley*, 2004; *Nakagawa et al.*, 2010; *van Summeren et al.*, 2009]. It may also help explain the change in the pattern of seismic velocity heterogeneity across the 660 [e.g., *Gu et al.*, 2001; *Ritsema et al.*, 2004].

In Figure 5 we show which combinations of mantle temperature anomaly  $\Delta T$  and composition below 660 can explain the high-velocity anomaly of 0.8% in the gap within the thermochemical Samoa plume (see Figure 1). The uppermost lower mantle is enriched in harzburgite due to compositional segregation across the 660. Assuming a Clapeyron slope for the 660 phase boundary of  $\gamma_{660} = -2.9$  MPa K<sup>-1</sup>, we find that if the average harzburgite fraction is >0.925 (i.e., a basalt fraction of f < 0.075) between 660 and 760 km depth, a temperature anomaly of  $\Delta T = 125 - 175^{\circ}$ C in the transition zone can explain the Samoa anomaly. Although phase boundary topography plus high-velocity material below or around 660 km depth are required to match the Samoa gap, estimates of composition and thermal anomaly are subject to substantial uncertainties associated with the mineral physics constraints on transition zone shear velocities [*Stixrude and Lithgow-Bertelloni*, 2011; *Cammarano et al.*, 2003].



**Figure 5.** The expected values of  $\delta V_{\text{OUT}}$  in the Samoa gap for a model of the Samoa plume as a continuous thermochemical upwelling across the transition zone that has elevated the 660 and includes a 100 km thick zone below the 660 with a harzburgite-enriched (basalt-depleted) composition.  $\delta V_{\text{OUT}}$  is determined as a function of the temperature anomaly  $\Delta T$  and for variable basalt fraction in the uppermost lower mantle.

Our analysis has focused on the mantle beneath the Samoa hot spot where image resolution in the transition zone is relatively high and where the effects are most obvious. The Samoa plume and gap are also apparent in tomographic models based on different data sets and modeling strategies (supporting information Figure S1). However, a quantitative comparison of image resolution for all models is necessary to determine whether these tomographic images are consistent with S40RTS and our analysis.

While this study has concentrated on anomalies caused by phase boundary effects in regions of mantle upwelling, we expect that within a cold slab the  $ri \rightarrow bm + mw$  transition will occur at a greater depth and thus introduce a thin low wave speed anomaly. The Pacific slab anomaly (Figure 1a) is diminished by about 0.5% near 660, which is consistent with this interpretation. Figure 4a indicates that if the average slab composition is close to pyrolite, a velocity anomaly of 0.5% can be explained by a thermally induced phase boundary deflection of 15 km, corresponding to a temperature decrease within the slab of  $\Delta T = 200^{\circ}$ C.

If harzburgite enrichment is important in the mantle, we should expect to see high-velocity anomalies in other regions of mantle upwelling. Anomalies similar to the gap in the Samoa plume are indeed apparent beneath the Azores, Canary, Galapagos, and Hawaii hot spots (supporting information Figure S2), although they are much weaker and not as obviously layered, most likely due to the relatively poor tomographic image resolution in the transition zone beneath regions far from the western Pacific subduction zones [e.g., *Ritsema et al.*, 2004; *Houser et al.*, 2008] (supporting information Figure S3).

While we suggest that high-velocity layering within broad low seismic velocity anomalies is consistent with dynamically predicted basalt-harzburgite segregation, this prediction depends on the postspinel Clapeyron slope [*Weinstein*, 1992], mantle viscosity [*Brandenburg and van Keken*, 2007], and the relative densities of basalt, harzburgite, and pyrolite as a function of pressure and temperature, each with significant uncertainties. These parameters determine whether or not global compositional stratification near 660 develops over the convective timescale of Earth [*Nakagawa et al.*, 2010]. Further work should test our observation and interpretation of the compositional segregation within mantle plumes.

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