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Title: Assessing Margin-wide Rupture Behavior along the Cascadia Megathrust using 3-D Dynamic Rupture Simulations

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Assessing Margin-Wide Rupture Behaviors along the Cascadia Megathrust with 3-D Dynamic Rupture Simulations

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Key Points

- We design the first fully dynamic 3-D earthquake simulations based on geodetic coupling models for the Cascadia megathrust.
- Segmentation in the stress drop is needed to produce subsidence amplitudes consistent with observed megathrust earthquakes.
- Dynamic rupture simulations demonstrate physics-based controls on margin-wide rupture.

Abstract

From California to British Columbia, the Pacific Northwest coast bears an omnipresent earthquake and tsunami hazard from the Cascadia subduction zone. Multiple lines of evidence suggests that magnitude eight and greater megathrust earthquakes have occurred - the most recent being 321 years ago (i.e., 1700 A.D.). Outstanding questions for the next great megathrust event include where it will initiate, what conditions are favorable for rupture to span the convergent margin, and how much slip may be expected. We develop the first 3-D fully dynamic rupture simulations that are driven by fault stress, strength and friction to address these questions. The initial dynamic stress drop distribution in our simulations is constrained by geodetic coupling models, with segment locations taken from paleoseismic analyses. We document the sensitivity of nucleation location and stress drop to the final seismic moment and coseismic subsidence amplitudes. We find that the final earthquake size strongly depends on the amount of slip deficit in the central Cascadia region, which is inferred to be creeping

60 interseismically, for a given initiation location in southern or northern Cascadia. Several
61 simulations are also presented here that can closely approximate recorded coastal subsidence
62 from the 1700 A.D. event without invoking localized high-stress asperities along the down-dip
63 locked region of the megathrust. These results can be used to inform earthquake and tsunami
64 hazards for not only Cascadia, but other subduction zones that have limited seismic observations
65 but a wealth of geodetic inference.

66 **Plain Language Summary**

67 The largest earthquakes on Earth occur along faults that develop between two tectonic
68 plates that come into contact. Termed megathrust earthquakes, these catastrophic events are
69 responsible for generating both strong ground-shaking and tsunamis. The Cascadia megathrust
70 fault straddles the Pacific coastline of North America and from evidence in both the United
71 States and Japan, we know this fault last slipped in 1700 A.D. We have combined models of
72 strain buildup (geodetic coupling models) with state-of-the-art 3-D computer simulations to
73 understand the potential hazard of a future earthquake in Cascadia and show what factors might
74 lead to the fault slipping its entire length. We compare our simulations to geologic measurements
75 of permanent ground movement from 1700 A.D. Our results demonstrate that no matter where
76 the earthquake is allowed to start, coupling models showing strain accumulation to the top of the
77 fault easily leads to big earthquakes. We also look into what 1700 A.D. event may have looked
78 like and show several scenarios that fit the geologic data very closely. This work represents the
79 first set of 3-D simulations that use the laws of physics to see what may control the size of future
80 earthquakes in Cascadia.

81 **1. Introduction**

82 The Cascadia subduction zone megathrust dominates earthquake hazard in the United
83 States Pacific Northwest. It is oft-cited that the probability of a magnitude ~ 9 (M9) event
84 occurring in the coming decades is between 10 – 14 % (Peterson et al., 2014). The most recent
85 megathrust rupture in Cascadia occurred in 1700 A.D. and generated a transoceanic tsunami
86 (Heaton and Hartzell, 1987). Matching amplitudes of historical tsunami records from Japan
87 requires a magnitude between M8.7 - 9.2 for this earthquake (Satake et al., 1996, 2003). While
88 321 years have elapsed since this last event, the Holocene (<12 kya) earthquake record onshore
89 and offshore documents even older $M > 8$ megathrust events. (e.g., Atwater and Griggs, 2012;
90 Leonard et al., 2010; Kemp et al., 2018; Goldfinger et al., 2012, 2017).

91 *Geological and Geophysical Inferences on the State of Megathrust Segmentation*

92 Several geological and geophysical observations suggest the Cascadia megathrust
93 exhibits along-strike segmentation. For instance, there are systematic changes in the accretionary
94 wedge backstop geometry, seismicity, and interseismic slip patterns (e.g., Stone et al., 2018;
95 Bartlow, 2020; Watt and Brothers, 2020) that may indicate coseismic rupture patterns will also
96 be variable along-strike. The strongest observational constraints that may inform our
97 understanding of future great earthquakes come from paleoseismic and geodetic observations.
98 Underwater turbidite deposits, which can be generated from submarine landslides induced by
99 strong ground-shaking during megathrust earthquakes, have been extensively used to map along-
100 strike rupture extents (Goldfinger et al., 2003; 2012; 2017; Figure 1a). Analysis of the timing and
101 spatial extents of turbidite deposits suggests that the recurrence interval (RI) between megathrust
102 earthquakes could vary along the Cascadia margin. In particular, the RI estimated for northern
103 Cascadia ($> \sim 46^\circ$ latitude) exceeds 400 years whereas it is estimated to be less than 200 years for
104 the southern portions of Cascadia ($< \sim 43^\circ$ latitude; Goldfinger et al., 2017).

105 Decadal scale interseismic velocities measured at the Earth's surface by Global
106 Navigation Satellite System (GNSS) networks, tide gauge, and leveling data also find significant
107 variations in coupling (slip deficit) distribution along the margin. Regions in northern and
108 southern Cascadia have higher coupling suggesting they are accumulating strain that may be
109 released in a future great earthquake (Schmalzle et al., 2014; Li et al., 2018; Yousefi et al., 2020;
110 Figure 1b, c). However, the use of geodetic coupling inversions to place bounds on the future
111 down-dip or along-strike rupture extent is complicated by heterogeneous frictional properties
112 (Boulton et al., 2019) or the potential presence of stress shadows (Hetland and Simons, 2010;
113 Alemeida et al., 2018) and other factors (e.g., off-fault deformation). Therefore, the down-dip
114 extent of coupling and coseismic rupture may differ even though these inversion results are the
115 best available constraint on potential stress distributions for the Cascadia megathrust (Wang and
116 Trehu, 2016).

117 Another piece of evidence for segmentation comes from the behavior of episodic tremor
118 and slip (ETS) events along the megathrust. In GNSS displacement records, ETS manifests as
119 transient reversals in displacement indicative of slip on or near the megathrust at depths between
120 30 and 40 km. These slow slip episodes are often accompanied by a weak seismic signature
121 known as nonvolcanic tremor (Rodgers and Dragert, 2003). The character of ETS events varies

122 significantly along the margin. The northern (i.e., $> 47^\circ$) and southern ($< 43^\circ$) sections host more
123 frequent slip episodes with average recurrence intervals of 10 and 14 months and have higher
124 tremor density, whereas the central section of the megathrust hosts ETS approximately every 19
125 months (Brudzinski and Allen, 2007; Wech and Creager, 2008). Studies of the ETS source
126 region find that the phenomenon occurs in regions of significantly elevated V_p/V_s ratios (e.g.,
127 Audet et al. 2009; Delph et al. 2021) and that tremor, and constituent low-frequency earthquakes,
128 are extremely sensitive to small magnitude stress changes such as those from the solid Earth
129 tides (Royer et al. 2015). Collectively, these observations suggest that pore fluid pressures are
130 nearly lithostatic in the ETS source region. The Cascadia megathrust also features a transition
131 zone at depth that separates the ETS region from the region that is conventionally considered to
132 be locked (≤ 20 km; Hyndman, 2013). Known as the gap, this spatial disconnect in slip behavior
133 is also found in other subduction zones (Gao and Wang, 2017); the frictional behavior and shear
134 stress accumulation levels in the gap may or may not allow for deeper rupture (Ramos and
135 Huang, 2019).

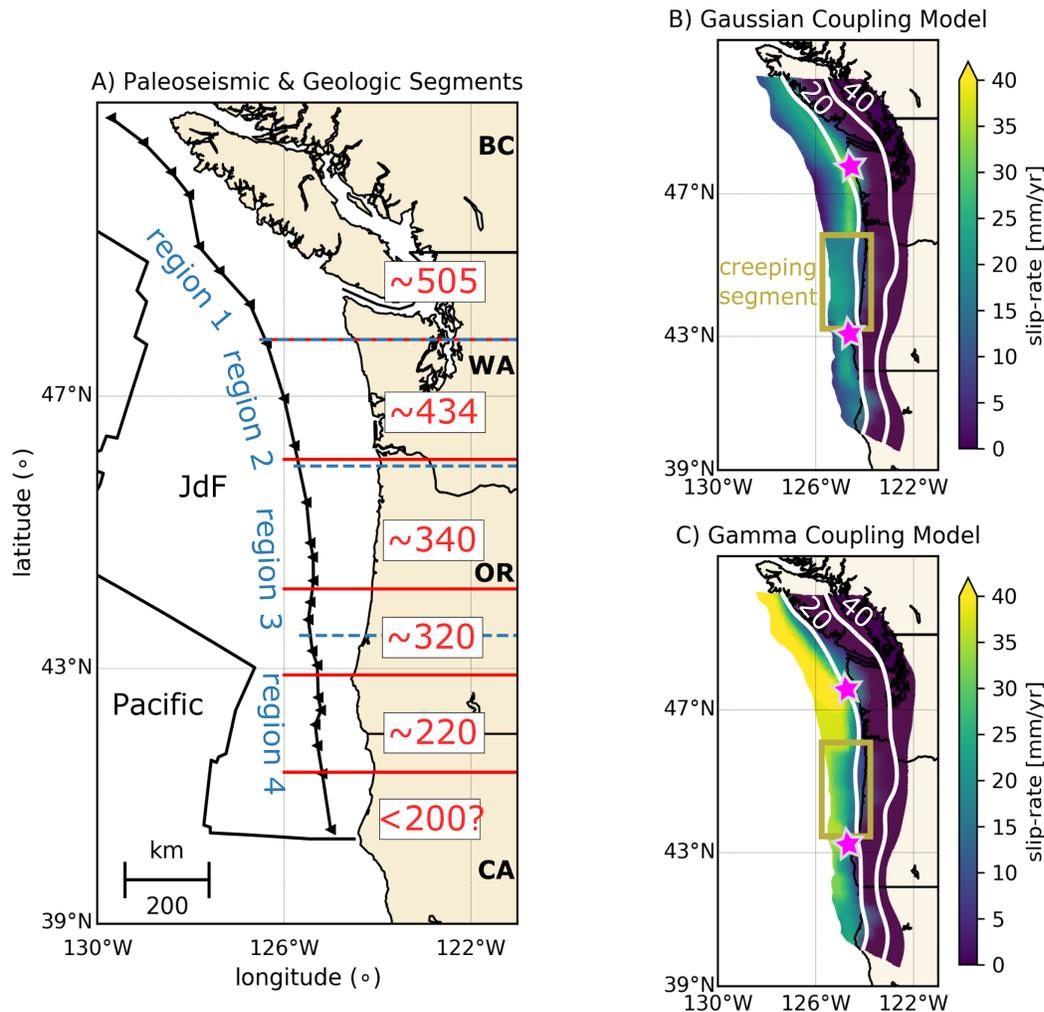
136 *Cascadia Earthquake Source Models*

137 What are ways to anticipate how a future Cascadia megathrust earthquake may behave?
138 One way to assess the hazard posed by large seismic events in Cascadia is to use kinematic
139 rupture simulations. Kinematic rupture simulations are commonplace due to the straightforward
140 relationship between fault slip and the recorded elastic displacement field once the Green's
141 functions are known, allowing these types of models to be run at lower computational cost.
142 Using the kinematic framework, potential locations of strong-ground motion sources along the
143 fault, sedimentary basin amplification or tsunami generation have been assessed (Olsen et al.,
144 2008; Delorey et al., 2014; Wirth et al., 2018; Frankel et al., 2018; Melgar et al., 2016; Roten et
145 al., 2019; Wirth et al., 2019 a, b). Most of these kinematic rupture models calibrate first-order
146 rupture parameters (slip, slip-rate, rise-time or rupture speed) from the few large megathrust
147 earthquakes observed in other subduction zones (e.g., M8.2 2003 Tokachi-Oki, M9.2 2004
148 Sumatra-Andaman, M8.8 2010 Maule, M9.1 2011 Tohoku-Oki). Kinematic simulations provide
149 important constraints on strong ground motions felt onshore, but because they must assume a slip
150 distribution before computing the elastic wavefield they cannot answer *what* controls the final
151 rupture size.

152 To account for source physics, fully dynamic rupture simulations can be used to
153 investigate what controls the final rupture size and kinematic rupture properties like rupture
154 speed. Dynamic rupture simulations are self-consistent, physics-based numerical models that
155 describe the entire earthquake rupture process (nucleation, propagation, and arrest) that is
156 coupled a constitutive fault friction law (e.g., Madariaga and Olson, 2008). To date, 2-D dynamic
157 rupture models in Cascadia have focused on tsunami generation (e.g., Lotto et al., 2018) or how
158 frictional and stress conditions in the transition zone may influence down-dip rupture extent
159 (Ramos and Huang, 2019).

160 Here we develop 3-D dynamic rupture simulations to explore how variable strain
161 accumulation rates, frictional behavior and hypocenter location influence megathrust rupture
162 dynamics. We will use geodetic coupling results from Schmalzle et al. (2014), who utilized
163 GNSS time series information spanning several decades for their coupling inversions. These
164 coupling distributions represent two possible end-member scenarios for strain accumulation near
165 the deformation front: either there is interseismic creep at shallow megathrust depths (hereafter
166 referred to as the Gaussian coupling model; Figure 1b) or it is fully coupled (hereafter referred to
167 as the Gamma coupling model; Figure 1c). Specifically, these coupling models will be used to
168 estimate the dynamic stress drop, which is defined as the difference between the initial shear
169 stress and dynamic fault strength. Dynamic stress drop is a key parameter determining how much
170 energy is available for rupture propagation (Kanamori and Rivera, 2004). Our dynamic stress
171 drop levels are further constrained by strain accumulation times and segment locations adopted
172 from paleoseismic studies (i.e., Goldfinger et al., 2017; Figure 1a). We compare the resulting
173 coseismic uplift and subsidence patterns to available paleoseismic measurements and discuss
174 which classes of models allow margin-wide ruptures to develop. We find the final earthquake
175 size is sensitive to earthquake nucleation location (e.g., northern vs. southern Cascadia) and the
176 distribution of relative dynamic stress drop. The principal control on margin-wide rupture, when
177 using these particular end member geodetic coupling models, is the relative dynamic stress drop
178 amplitude in the central Cascadia region ($\sim 43 - 47^\circ$ latitude). The results also suggest that Gamma
179 coupling models tend to produce larger earthquakes, even if shallow subducted sediment has a
180 slip-strengthening or velocity-strengthening frictional behavior. Another intriguing question is if
181 geodetic coupling models can inform our understanding about the 1700 A.D. earthquake when
182 incorporated into a dynamic rupture simulation. To that end, we also present several rupture

183 simulations that provide a close fit to the 1700 A.D. event.



184
185 **Figure 1.**

186 Cascadia subduction zone study area. A) Megathrust segmentation (segments are separated by red lines) suggested from offshore
187 turbidite deposits (Goldfinger et al., 2017) with corresponding estimated segment recurrence intervals in years. Primary
188 morphotectonic regions identified by Watts and Brothers (2020) are superposed (blue dashed lines). B) Gaussian and C) Gamma
189 coupling models from Schmalzle et al., (2014) projected onto the Slab2 megathrust geometry. The Gaussian coupling model
190 assumes interseismic creep at shallow megathrust depths whereas the Gamma model assumes high strain accumulation. The
191 inferred region of the creeping segment is denoted by a yellow box. Magenta stars denote rupture initiation locations in our dynamic
192 rupture models. Thick white lines are megathrust depth contours (kilometers). JdF = Juan de Fuca plate.

193
194 **2. Methodology**

195 We solve for 3-D elastodynamic earthquake rupture using SeisSol, a powerful open-
196 source software package that implements the Arbitrary high-order DERivative-Discontinuous

197 Galerkin (ADER-DG) approach to simulate wave propagation coupled to spontaneous dynamic
198 rupture (de la Puente et al., 2009; Pelties et al., 2012; Heinecke et al., 2014; Uphoff et al.,
199 2017). The capability of SeisSol to solve for complex source dynamics and incorporate realistic
200 geometric features, such as bathymetry, topography and fault zone structure (e.g., Ulrich et al.,
201 2019a, b; Wollherr et al., 2019) nicely lends itself to our purposes of investigating how
202 heterogeneous megathrust stresses influence rupture behavior.

203 We generate an unstructured 3-D tetrahedral mesh for the Cascadia subduction zone that
204 spans over 1100 km along-strike (39.0 to 51.0 degrees latitude, -127.5 to -121.0 degrees
205 longitude) and we use static refinement to increase resolution locally. The average on-fault
206 element edge size (h) is 2.5 km, and the maximum depth of the fault mesh is 50 km (Hayes et al.,
207 2018) and includes over 440,000 unstructured triangular elements. We account for the large-
208 scale variations in the free-surface geometry by meshing the ETOPO1 topography and
209 bathymetry dataset to ~ 1 km average element size near the coastline. In all of our simulations,
210 we use ADER-DG with fifth order accuracy (polynomial order $p = 4$) in time and space.

211 We ensure simulation results are sufficiently resolved by following the procedure
212 established in Wollherr et al. (2018) to estimate the process zone, the region behind the rupture
213 front where the fault strength drops from its static to dynamic level. For the 2.5-km fault mesh,
214 the median process zone width (Λ_m) is ~ 1.1 km. The recommended number of elements needed
215 to resolve Λ_m in a purely elastic setup with depth-dependent initial conditions is 2 - 3 ($p = 4$). The
216 quadrature points approach utilized in SeisSol (Pelties et al., 2014) ensures each element edge
217 length is sampled $p + 2$ times. Given our setup, Λ_m is sampled by ~ 2.7 elements which is within
218 the recommended range. The expected relative percent error in the rupture arrival time, peak
219 slip-rate, and final slip are 0.09, 8.32, and 0.71, respectively (Wollherr et al., 2018). While the
220 peak slip-rate relative error is slightly larger than the 7% recommended by Day et al., (2005) for
221 elastic rupture problems, we compare our model-predicted slip and rupture size to higher
222 resolution meshes with $h = 1$ km and $h = 0.5$ km and observe negligible changes, which gives us
223 confidence that these first order rupture features are correctly resolved. The highest resolution
224 mesh has more than 50 million elements and requires 22 hours on 40 nodes of the supercomputer
225 SuperMUC-NG at the Leibniz Supercomputing Centre, Germany.

226

227

228 *2.1 Constraining dynamic rupture with geodetic coupling models*

229 In our simulations, potential shear stress distributions are informed by geodetic coupling
230 models. The Schmalzle et al., (2014) inversion for slip rate deficit was performed with respect to
231 a Cascadia megathrust geometry predating Slab2 (McCrory et al., 2012) and as such, we first
232 map the geodetic coupling models to our megathrust geometry through a bilinear interpolation
233 using the cartesian horizontal plane coordinates. But the effect of this transformation does not
234 distort the main features of the coupling models (Figure 1b, c).

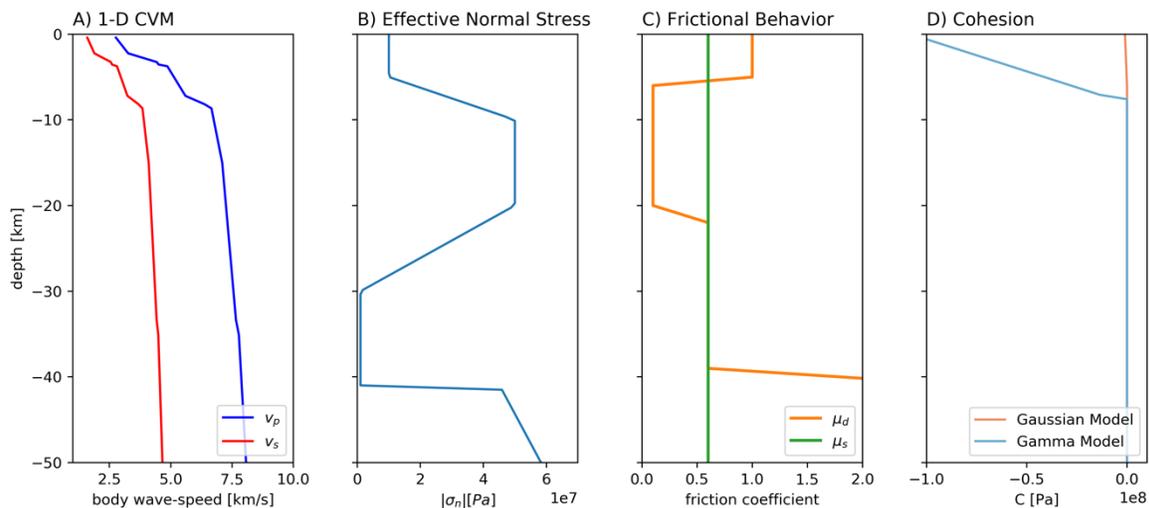
235 We define the parameter T as the time needed for a certain level of slip deficit to
236 accumulate on a section of the megathrust. The product of slip rate deficit (coupling) and T is
237 slip deficit. T should not be interpreted as the RI, but rather as another way to quantify relative
238 dynamic stress drop along the megathrust. From these slip distributions, we estimate the static
239 stress drop using Poly3D, a three-dimensional, polygonal element, displacement discontinuity
240 boundary element method, which accounts for nonplanar megathrust geometry and the free-
241 surface effect due to buried slip (Thomas, 1993).

242 Initial shear stress is then estimated by adding the static stress drop to the dynamic fault
243 strength. Calculating the initial shear stress in this manner is known as the complete stress drop
244 assumption and assumes that slip deficit is accumulated linearly in the along-dip fault dimension
245 and will be entirely released during coseismic rupture (Yang et al., 2019a, b; Hok et al., 2011).
246 This shear stress distribution is first resampled to an average grid spacing of ~ 3 km and then
247 linearly interpolated onto the fault mesh. We note that we initialize stress values and friction
248 parameters with a high order sub-element resolution (e.g., Pelties et al., 2014). For all dynamic
249 rupture simulations considered, we compare the results to the 1700 A.D. subsidence
250 measurements along the coast where available (Wang et al, 2013), and to recorded subsidence
251 amplitudes from other $\sim M9$ earthquakes (e.g., 2011 Tohoku, 1960 Chile, 1964 Alaska).

252 *2.2 Material Properties and Fault Strength*

253 Wave propagation is simulated within a heterogenous, linearly elastic medium where the
254 elastic moduli (lame parameters) vary as a function of depth. The average 1-D velocity structure
255 is taken from the Cascadia 3-D Community Velocity Model (3D-CVM) for P and S waves
256 (Figure 2a; Stephenson et al., 2017). Since the goal of this study is to calculate upper plate
257 deformation and rupture extent (along-dip and along-strike) for a given dynamic stress drop
258 distribution, we believe this is a satisfactory simplification to make. We estimate that we can

259 resolve a cutoff seismic frequency up to ~ 0.4 Hz in the near fault region. High frequency (>1 Hz)
 260 broadband ground motions can be calculated at a higher computational cost if an appropriate 3-D
 261 velocity model is utilized. The current 3D-CVM was developed with respect to an older
 262 Cascadia subduction zone geometry (i.e., McCrory et al., 2012) and thus, we leave direct
 263 extrapolation of this 3-D velocity model to our model geometry for future work.



264
 265 **Figure 2.** Material properties, strength, and frictional conditions for dynamic rupture simulations. A) Smoothed 1-D CVM velocity
 266 model for Cascadia (Stephenson et al., 2017). B) Effective normal stress extended beyond 40 km depth from Ramos & Huang
 267 (2019). C) Dynamic and static frictional coefficients with depth. D) Frictional cohesion for the Gaussian and Gamma coupling
 268 models.

269 Effective normal stress accounts for pore pressure counteracting vertical lithostatic stress
 270 on the fault. We use the depth-dependent effective normal stress distribution for Cascadia
 271 presented in Ramos & Huang (2019) that includes low strength levels (1 MPa) in the ETS region
 272 (Figure 2b). These incredibly low effective stress conditions in the ETS region are supported by
 273 observations on the sensitivity of tremor and low-frequency earthquakes to small magnitude
 274 stress changes (e.g., Rubinstein et al., 2007; Royer et al., 2015), stress orientations in the ETS
 275 region (e.g., Newton and Thomas, 2020), and low stress drops of ETS events (e.g., Gao et al.,
 276 2012). For lack of in-situ fault stress information, we assume a linear stress gradient above and
 277 below the locked region (10 – 20 km depth) that are consistent with other Cascadia megathrust
 278 simulations (Liu and Rice, 2009; Li and Liu, 2016). Such assumptions are simple but allow us to
 279 focus on how heterogeneous shear stresses on the megathrust contribute to first order rupture
 280 characteristics.

281 *2.3 Fault friction law*

282 The physics controlling the inelastic breakdown process in our dynamic simulations is
 283 given by a nonsingular linear slip-weakening friction law (Palmer and Rice, 1973). This
 284 constitutive friction law allows us to idealize rupture as a propagating shear-crack. It is described
 285 by the static (μ_s) and dynamic (μ_d) friction coefficients and a critical slip-weakening distance
 286 (D_c).

287 We set $\mu_s = 0.6$ and $\mu_d = 0.1$ within the locked region of the megathrust ($5 \text{ km} \leq \text{depth} \leq$
 288 20 km) [Figure 2c]. Because Ramos & Huang (2019) showed that rupture can penetrate the gap
 289 or generate strong free-surface reflections if its frictional behavior is slip-weakening at depths $<$
 290 5 km and at depths $> 25 \text{ km}$ (together with a highly negative stress drop), we set μ_d equal to or
 291 above μ_s in these regions (Figure 2c). D_c is set to a constant level of 1 m or 2 m. $D_c = 2 \text{ m}$ is
 292 selected in the dynamic rupture model in which the stress and strength conditions of Ramos &
 293 Huang (2019) are extrapolated along strike, for consistency with the 2-D dynamic rupture
 294 simulations. $D_c = 1 \text{ m}$ is used for the dynamic rupture models based on the heterogeneous
 295 geodetic coupling prestress distributions. Our range of D_c values are consistent with those used
 296 in slip-weakening simulations of the Tohoku-Oki earthquake, which constrained D_c using the
 297 frequency range of back-projection results (Huang et al., 2014). We make minimalistic
 298 assumptions for cohesion in the upper 5 km of the megathrust (Figure 2d). Due to the nearly zero
 299 dynamic stress drop amplitudes near the deformation front for Gaussian coupling models, the
 300 cohesion gradient can be low (Figure 2d). But in the case of the Gamma coupling models,
 301 relatively higher cohesion levels ($\sim 5 \text{ MPa}$ average) are locally needed at shallow fault depths to
 302 prevent fault failure at the start of the simulation (Figure 2d).

303 *2.4 Rupture Initiation*

304 Fault pre-stress conditions influence the estimated critical nucleation size when using a
 305 linear slip-weakening friction law. The theoretical critical nucleation radius that permits
 306 spontaneous dynamic rupture to initiate in a 3-D linearly elastic and homogeneous media has
 307 been derived by Day (1982) and is given by,

$$308 \quad r_c^t = \frac{7\pi G(S+1)D_c}{24 \Delta\tau_d} \quad (1)$$

309 where G is the shear modulus, S is the relative fault strength defined as the ratio between
310 strength excess (static fault strength minus initial shear stress) and dynamic stress drop ($\Delta\tau_d$).
311 Expression (1) provides a sufficient means to initiate and sustain dynamic rupture propagation
312 for the 3-D dynamic rupture model that is adapted from 2-D dynamic rupture simulations
313 presented in Ramos & Huang (2019). For the prestress distributions derived from the
314 heterogeneous coupling models, we determine the best numerical nucleation size through a trial-
315 and-error approach. We find that critical nucleation radii are within $\sim 10\%$ of the theoretically
316 predicted value calculated from equation (1). Rupture initiation is prescribed by a smooth space
317 and time dependent function, leading to an imposed rupture velocity that decreases away from
318 the hypocenter and allows a gradual transition from forced to spontaneous rupture (Harris et al.,
319 2018). Rupture nucleation locations are chosen within the areas containing the highest dynamic
320 stress drop distribution (see Figure 1b and c). Each dynamic rupture simulation is run for 420
321 seconds (7 minutes) to allow seismic waves to propagate to the edge of the model domain.

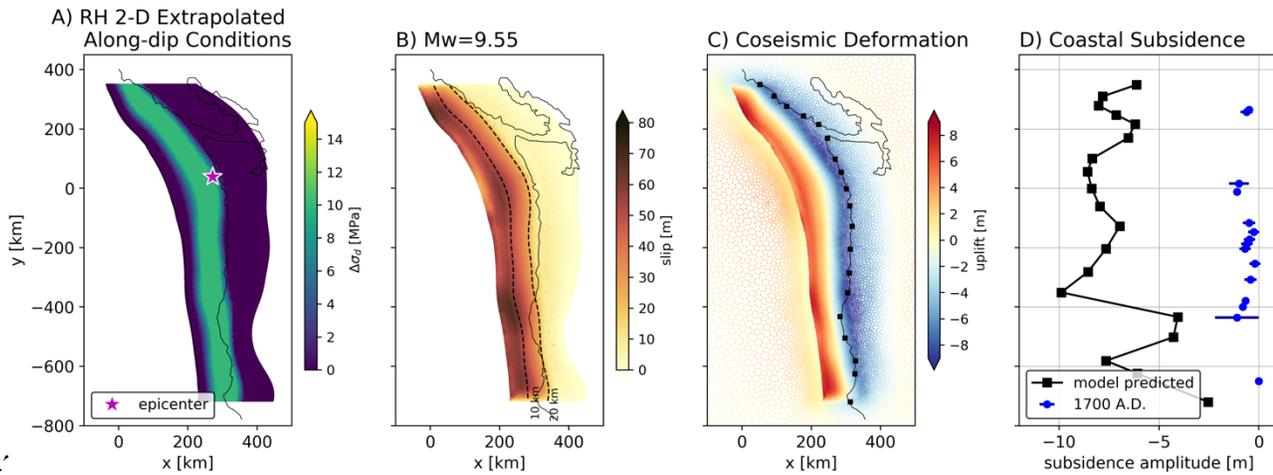
322 **3. Results**

323 *3.1 Translating 2-D rupture simulations to 3-D*

324 A 3-D dynamic rupture model that assumes a relatively homogeneous dynamic stress drop
325 profile along the locked region of the megathrust is shown in Figure 3. Previously developed 2-D
326 dynamic rupture simulations (Ramos and Huang, 2019) were relative to a specific location in
327 northern Cascadia, which is where we initiate rupture (Figure 3a). Such a laterally uniform
328 dynamic stress drop distribution is unlikely given observations of geophysical and geological
329 megathrust segmentation (e.g., Watt and Brothers, 2020). However, we develop such a simulation
330 to demonstrate 1) what a megathrust event would appear as if there was a strong gradient in shear
331 stress-rate from the locked to gap regions (20 - 30 km depth) across the margin and 2) how this
332 scenario would influence coastal subsidence amplitudes.

333 In spite of the low dynamic stress drop (< 5 MPa) at depths shallower than 10 km and
334 slip-strengthening friction, coseismic slip is able to reach the deformation front with amplitudes
335 exceeding 60 m in most locations along-strike (Figure 3b). The along-strike variation of slip at
336 the deformation front exhibits two peaks north and south of the hypocenter - even though the
337 initial dynamic stress drop distribution is laterally invariant, the final coseismic slip pattern is not
338 (Figure S1). This might be attributed to changes in the along-strike megathrust dip angle. There
339 are also small amounts of slip (< 5 m) in the gap region. The coseismic hinge-line, separating

340 regions of subsidence from regions of uplift, is entirely offshore (Figure 3c). Subsidence levels
 341 exceeding 5 m are observed along most of the coastline (Figure 3 c, d). This exceeds subsidence
 342 measurements from the 1700 A.D. event (Wang et al., 2013) by at least a factor of two because
 343 the earthquake is much larger than an M9 (Figure 3d). Such subsidence amplitudes are also much
 344 larger than the maximum levels observed for the 2011 M9.0 Tohoku (~1.1 m, Hashima et al.,
 345 2016), 1964 M 9.4 Alaska (~ 2.4 m, Plafker et al., 1969) or the largest ever recorded event, the
 346 1960 M9.5 Chile Earthquake (~2.7 m, Plafker and Savage, 1970).



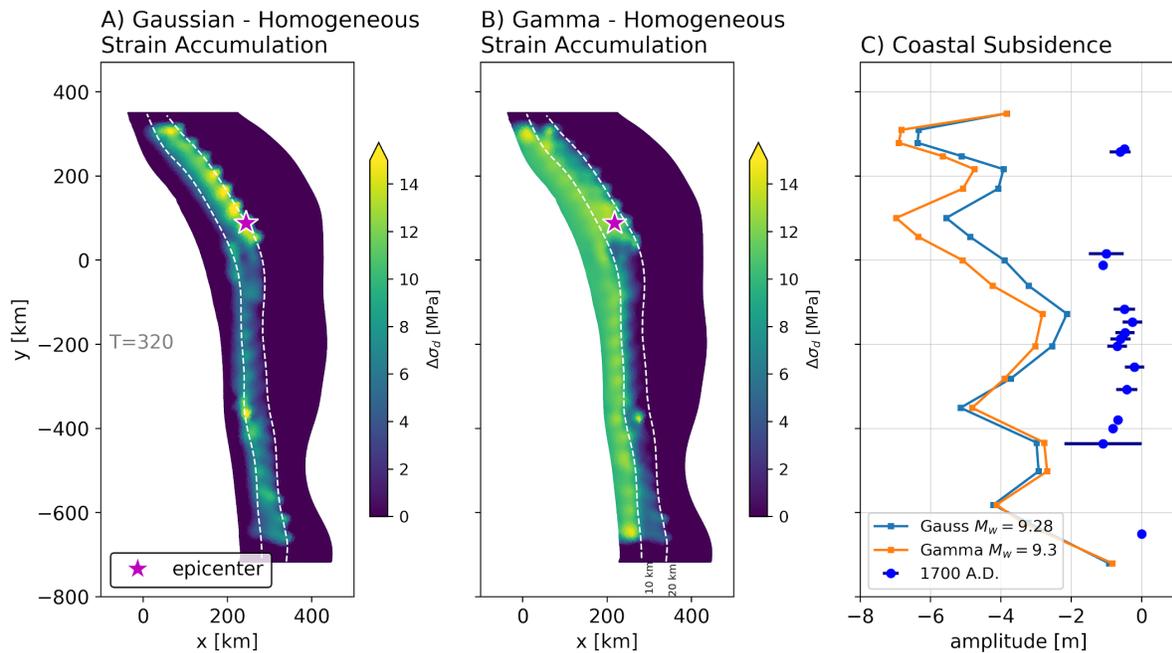
347 **Figure 3.** Results of the dynamic rupture model in which the stress and strength conditions of Ramos & Huang (2019) are
 348 extrapolated along-strike (RH). A) Along-strike dynamic stress drop and considered epicenter location. B) Final megathrust slip-
 349 distribution and moment-magnitude. The black dashed lines indicate the 10 and 20 km depth whereas the solid black line denotes
 350 the coastline. C) Coseismic uplift (red) and subsidence (blue) along the Cascadia margin. Squares signify the coastline. D) Model
 351 predicted (black squares; same as panel C) and paleoseismic observations of estimated subsidence during the 1700 A.D. rupture
 352 (Wang et al., 2013).
 353

354 Interestingly, this model generates a down-dip rupture front that can reach and ‘jump’
 355 across the gap region, despite the negative stress drop in the gap combined with a slip-neutral
 356 frictional behavior (Figure S2). Due to dynamic stress perturbations carried by seismic waves
 357 (‘dynamic unclamping’, Oglesby and Mai, 2012; Figure S2), this down-dip rupture front is most
 358 likely triggered by temporal stress changes and made possible by the incredibly low static fault
 359 strength here (i.e., 0.6 MPa).

360 3.2 Uniform Gaussian and Gamma coupling models

361 We now explore dynamic rupture scenarios based on the Gaussian and Gamma coupling
 362 models. We start with simulations that assume uniform T level (Figures 4 a and b). T is set to its
 363 320 years, the time elapsed since the most recent event (Goldfinger et al, 2017). In this

364 parameterization, the highest dynamic stress drop amplitude is located in the northern Cascadia
 365 region for both Gaussian and Gamma distributions (Figure 4a, b), which is where spontaneous
 366 rupture is initiated. The location of highest dynamic stress drop is not coincident between the
 367 Gaussian and Gamma coupling models, and hence the hypocenter locations are slightly different.
 368 Uniform T for both coupling models generates margin-wide rupture with coastal subsidence
 369 amplitudes that again exceed 1700 A.D. (Figure 4c). The Gamma coupling model has higher
 370 dynamic stress drop than the Gaussian model near the deformation front, which leads to a 1 to 2-
 371 meter difference in subsidence amplitude for the northern ($0 < Y < 200$ km) region of the
 372 megathrust (Figure 4c). These subsidence amplitudes, while lower than the 2-D extrapolated
 373 model, still surpass the estimated subsidence amplitudes of the largest recorded global
 374 megathrust earthquakes (i.e., 1960 Chile, 1964 Alaska). This result demonstrates that the
 375 uniform T coupling model overestimates the amount of slip deficit accumulated since 1700 A.D.



376
 377 **Figure 4.** Comparison between Gaussian and Gamma dynamic stress drop distributions and the resultant subsidence patterns
 378 assuming the maximum strain accumulation time (T) of 320 years (i.e., time since the last great earthquake in 1700 A.D.) A)
 379 Dynamic stress drop distribution for the Gaussian coupling model. B) Dynamic stress drop distribution for the Gamma coupling
 380 model. Both ruptures are nucleated in northern Cascadia (magenta star). C) Model predicted subsidence along the coastline
 381 compared to 1700 A.D. measurements.

382 When comparing the along-dip gradient of dynamic stress drop between the simple
 383 (Figure 3) and heterogeneous 3-D models (Figure 4), we note that the smoother model extends

384 slightly deeper (Figure S2). The amplitude of coseismic subsidence is probably more strongly
385 controlled by the dynamic stress drop gradient towards the coastline (Figure S3). A point to note
386 is that the region of higher relative dynamic stress drop in the northern Cascadia region ($0 \leq Y$
387 ≤ 200 km) is also where there are limited paleoseismic measurements from 1700 A.D. Thus,
388 while geodetic coupling models are well constrained here, the few along-strike subsidence
389 measurements limit rigorous comparison to physics-based model predictions.

390 *3.3 Segmented Gaussian and Gamma coupling models*

391 We find that in order to produce coseismic uplift and subsidence amplitudes more
392 consistent with the paleoseismic Cascadia measurements and data from other megathrust
393 earthquakes (i.e., ± 2 m), we must prescribe along-strike variations of T, with T amplitudes lower
394 than 320 years for a particular segment. This is especially needed for the northern and southern
395 regions of the Cascadia megathrust, where both the Gaussian and Gamma coupling models
396 predict higher subsidence amplitudes than observed if T is set to 320 years. We refer readers to
397 the discussion section on the possible meaning of these lower T values.

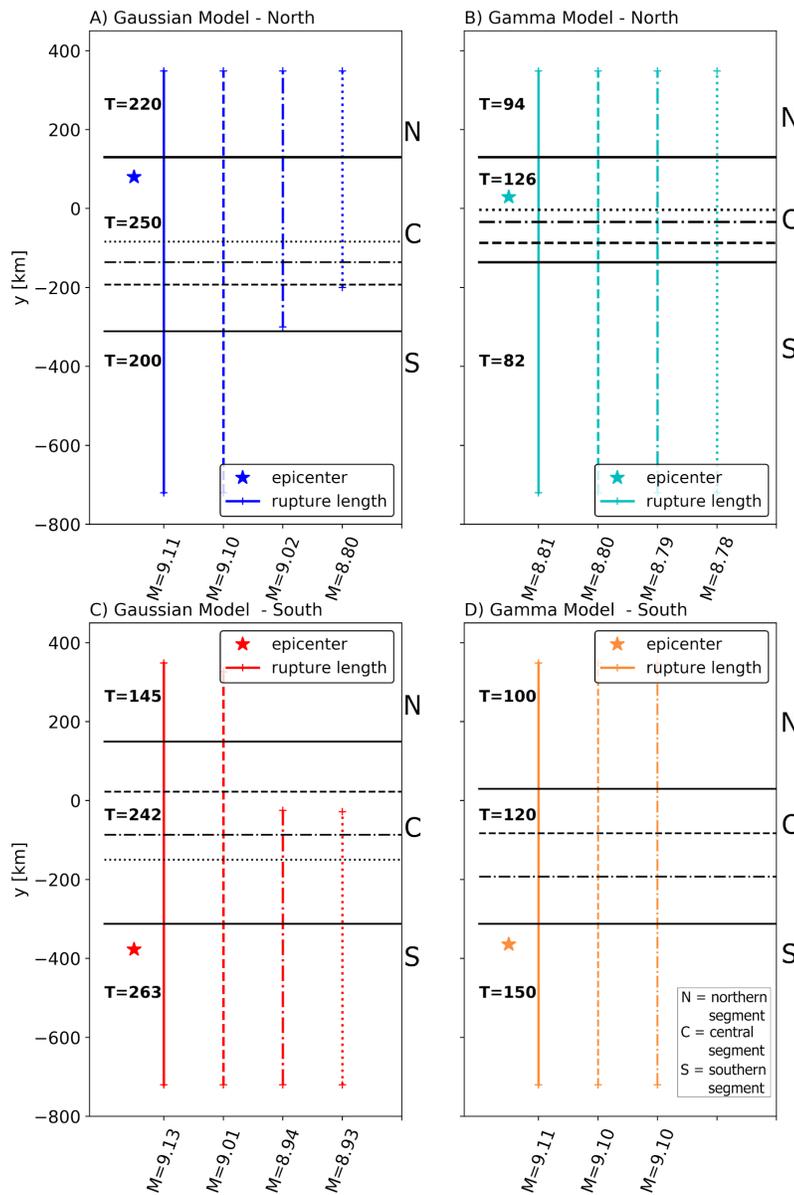
398 Our partitioning of the margin is informed by paleoseismic (Goldfinger et al., 2012,
399 2017), ETS (Brudzinski and Allen, 2007), and morphotectonic studies (Watt and Brothers, 2020)
400 in Cascadia. The following dynamic rupture models are parameterized using at least three
401 segments. This choice is conservative - we found through trial-and-error that two segment
402 models cannot match first-order 1700 A.D. subsidence patterns as well as three-segment
403 models. We note that some geologic models may suggest up to five segments (e.g., Goldfinger
404 et al., 2017; Wang et al. 2013) and thus there may be multiple ways to partition T levels along-
405 strike.

406 We first study three segment rupture models that are nucleated in the northern Cascadia
407 region to see how our choice of T and segment width affect final rupture length (Figure 6a). We
408 find that placing segment limits near ~ 46 and 43 degrees latitude (Y ranges from 180 to -350 km;
409 Figure 5a), together with T levels between 200 – 250 years (~ 8 – 10-m slip deficit), leads to
410 margin-wide rupture. The position of these segment limits corresponds to changes in estimated
411 RI level, tremor patterns and forearc morphology (Figure 1; Goldfinger et al., 2017; Watt and
412 Brothers, 2020). The middle segment encompasses most of the creeping region offshore Oregon.
413 Holding T levels constant, we systematically move the location of the southern segment
414 boundary southward until margin-wide rupture is no longer observed. An average slip deficit of

415 nearly 2 m over a width of ~80 km is needed to drive rupture through the creeping section and
416 into the southern end of Cascadia (Figure 5a). The higher coupling in the northernmost segment
417 ($Y > 200$ km) allows for rupture to propagate north of the epicenter in all cases. In contrast, if we
418 use the Gamma coupling model, margin-wide rupture is much easier to attain even with lower
419 relative stress drop (lower T values) (Figure 5b). Lower T levels are used in the Gamma rupture
420 simulations as higher values are not required to achieve margin-wide rupture with the Gamma
421 distribution. This result demonstrates the sensitivity of margin-wide rupture to the stress level in
422 the shallow portions of the fault. As the length of the central segment becomes shorter,
423 moment-magnitude only weakly decreases (by ~0.01) for Gamma ruptures. Gamma ruptures
424 nucleated in northern Cascadia can feature shallow, narrow slip distributions and low rupture
425 speeds ranging from ~1 to 2 km/s in the central region of the megathrust (Figure S4).

426 For dynamic ruptures initiated in southern Cascadia, we found that slightly higher T
427 levels (relative to ruptures nucleated in the north) are a necessary condition to sustain rupture
428 propagation, particularly through the central Cascadia region (Figure 5c, d). Gaussian models
429 that lead to a margin-wide rupture required an additional slip deficit of 3 m over a length of ~60
430 km in the central segment (i.e., Figure 5c, dashed line) compared to non-margin-wide rupture
431 event (i.e., Figure 5c, dot-dashed line). Similar to what was observed for ruptures initiated in
432 northern Cascadia, Gamma coupling models tend to generate margin-wide ruptures at much
433 lower slip deficit (i.e., Figure 5d). Higher relative T levels in the southernmost segment is
434 required in order for rupture to initiate and propagate outside the region of spontaneous rupture
435 initiation, given our slip-weakening friction parameters (i.e., μ_d , μ_s , D_c) and effective normal
436 stress that bound the fracture energy. The Cascadia megathrust dips more steeply below Oregon
437 and this probably influences the initial stages of ruptures that propagate from south to north more
438 than those that rupture north to south. In general, Gamma model results suggest that only a
439 narrow region of concentrated higher dynamic stress drop is sufficient for promoting margin-
440 wide rupture, even if slip-strengthening friction or higher sediment cohesion levels are
441 present. We will now discuss our assumptions about sediment friction in the shallow most
442 portions of the megathrust, and its effect on rupture size.

443

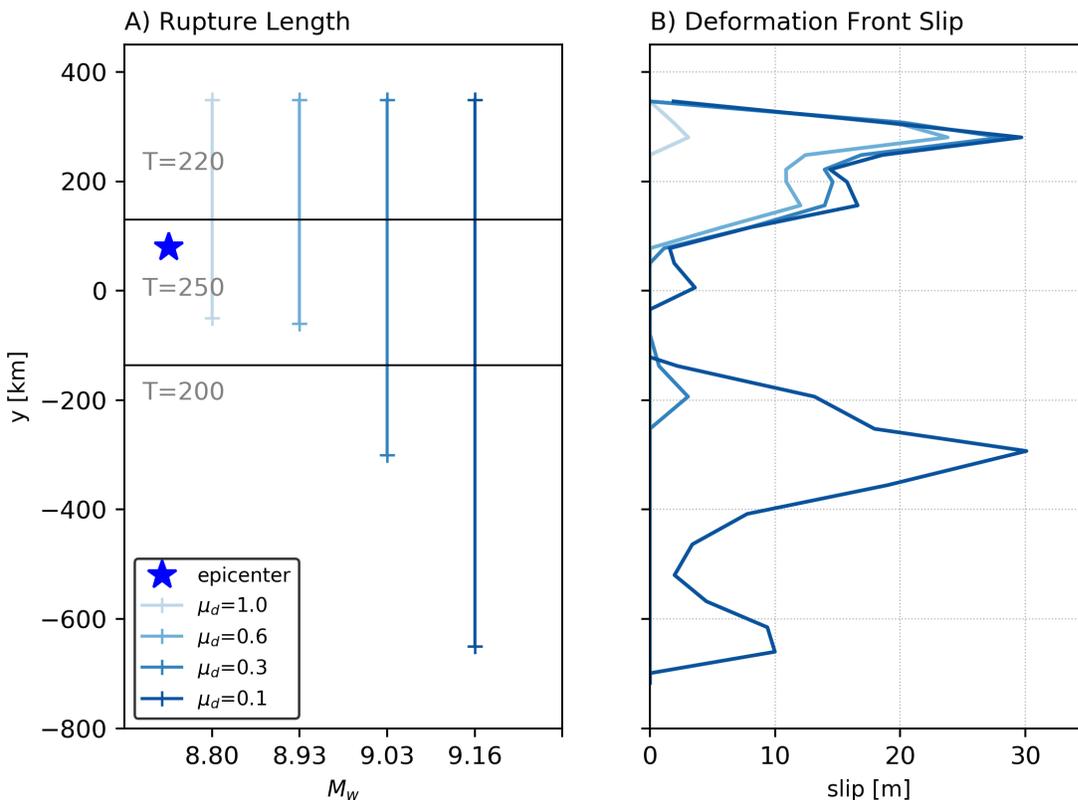


444
 445 **Figure 5.** Gaussian and Gamma dynamic rupture simulations nucleated in northern or southern Cascadia (colored star). In each
 446 plot, the line length corresponds to the along-strike rupture extent. The line style corresponds to the variable segment location. A)
 447 Gaussian ruptures where the width of the central segment, containing the nucleation asperity, is varied until margin-wide slip no
 448 longer occurs. The T levels (relative dynamic stress drop) remains constant for each simulation. B) Same idea as A but for Gamma
 449 rupture simulations. The higher stress drop at shallower depths ($5 \leq \text{km}$) provides enough energy for ruptures to span the entire
 450 Cascadia margin. Moment-magnitude is plotted along the x-axis on all plots. C) and D) show Gaussian and Gamma ruptures
 451 nucleated in southern Cascadia, respectively. For each subfigure, the Northern (N), central (C), or southern (S) segment regions
 452 are denoted.

453

454 *3.4 Effect of up-dip frictional behavior*

455 In all simulations presented so far, we have assumed the influence of subducting
 456 sediments will lead to slip-strengthening frictional behavior in the upper 5-km of the megathrust
 457 along-strike. We now relax this assumption and let the dynamic friction level vary from slip-
 458 strengthening to slip-weakening conditions (Figure 6) using the reference Gaussian model of
 459 Figure 5a (solid line). In all simulations, we fix T levels and segment locations, while testing
 460 varying dynamic friction coefficients in the near-margin region. Neither slip-strengthening ($\mu_d >$
 461 0.6) nor slip-neutral ($\mu_d = \mu_s = 0.6$) friction leads to margin-wide rupture for this particular
 462 parameterization (Figure 6a); only a slip-weakening behavior at shallow depths allows rupture to
 463 spontaneously grow into a margin-wide event. The effect of dynamic friction level on slip at the
 464 deformation front is shown in Figure 6b. We observe high slip amplitudes (>25 m) in northern
 465 and southern Cascadia and reduced slip in central Cascadia (Figure 6b). In the margin-wide
 466 rupture case (e.g., $\mu_d = 0.1$), this slip pattern is similar to other Gaussian coupling models.



467
 468 **Figure 6.** Gaussian coupling models with variable sediment frictional behavior in the upper 5 km of the megathrust. A) Along-
 469 strike rupture lengths (colored lines) as function of dynamic friction coefficient. B) Slip at the deformation-front for each scenario
 470 shown in A.

471 *3.5 Effect of down-dip locking depth*

472 Estimating the seismogenic zone from the available geodetic data and paleoseismic
473 measurements (Hyndman, 2013; Wang and Trehu, 2016) is fraught with uncertainty because of
474 their lack of offshore resolution. In both the Gaussian and Gamma coupling models, the down-
475 dip limit of coupling (positive stress drop) is near 20 km depth (Figure 1; Schmalzle et al., 2014),
476 broadly consistent with thermal models proposed for this subduction zone (Wang et al. 1995;
477 Hyndman, 2013; Cozzens and Spinelli, 2012). To assess how locking depth influences rupture
478 width, length and subsidence amplitudes, we now relax this assumption and let locking depth
479 vary. Note that the locking depth is meant as the maximum depth where slip-weakening
480 frictional behavior exists (i.e., $\mu_d < \mu_s$). Again, we start with the three-segment Gaussian
481 simulation (Fig 5a, solid line), which does not break through the central Cascadia region (Figure
482 7a). Slip-weakening behavior with $\mu_d = 0.1$ is initially set to end at 20 km depth and we
483 systematically extend locking depth by two kilometers until 30 km (Figure 7a). A dynamic
484 rupture simulation assuming a 15-km locking depth is also shown for sake of comparison. We
485 observe that moment magnitude increases ($8.8 < M_w < 9.2$) due to the propagation of rupture
486 into the gap and slow-slip regions of the fault. Ruptures progressively extend further south for
487 greater locking depth, but do not become margin-wide (Figure 7a).

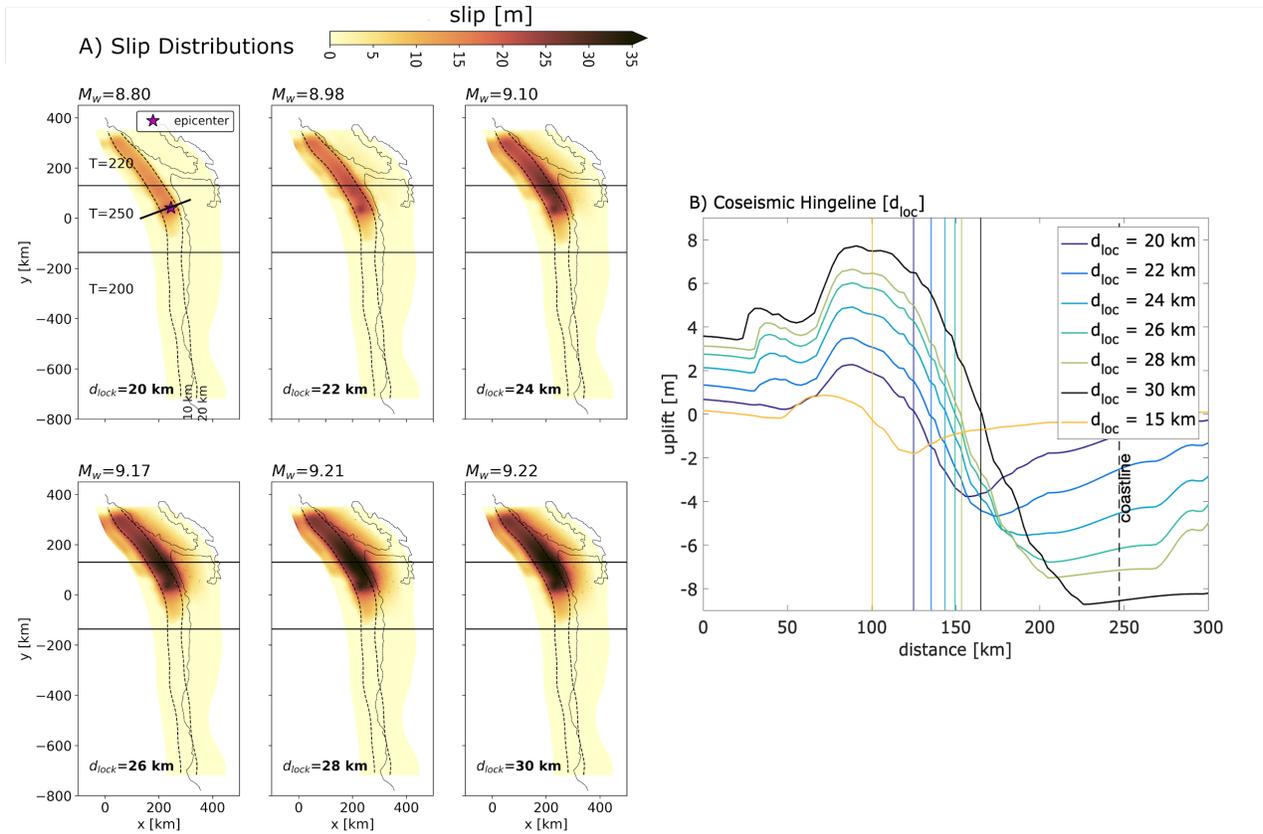
488 We select a 2-D profile near the hypocenter along-strike to assess how the model predicts
489 coseismic subsidence and amplitude patterns change in the margin-perpendicular direction
490 (Figure 7b). The maximum uplift and subsidence amplitudes increases by ~ 1 m for every 2-km
491 increase in locking depth. For deeper locking depths, the coseismic hinge-line moves closer to
492 shore, although all hinge-lines remain at least 100 km offshore for the profile selected in northern
493 Cascadia (Figure 7b).

494 *3.6 Fitting 1700 A.D. subsidence measurements*

495 Previous elastic rupture models have shown that coastal subsidence measurements from
496 1700 A.D. can be well fit with high slip-patches positioned along-strike. Wang et al., (2013) used
497 static models with four distinct asperities with T levels ranging from 450 – 550 years (18 – 22 m
498 slip deficit) to reproduce the subsidence amplitudes. In these static models, the greatest locking
499 depth was taken to coincide near the 350 C isotherm as this is where silica-rich
500 lithologies would be expected to transition from velocity-weakening to velocity-strengthening
501 frictional behavior (Wang et al., 2003). 3-D kinematic simulations used a range of locking depths

502 (~10 - 30 km) and determined that, in the presence of subevents, a locking depth near ~15 km
 503 provided the strongest fit to the subsidence data (Wirth et al., 2019b).

504



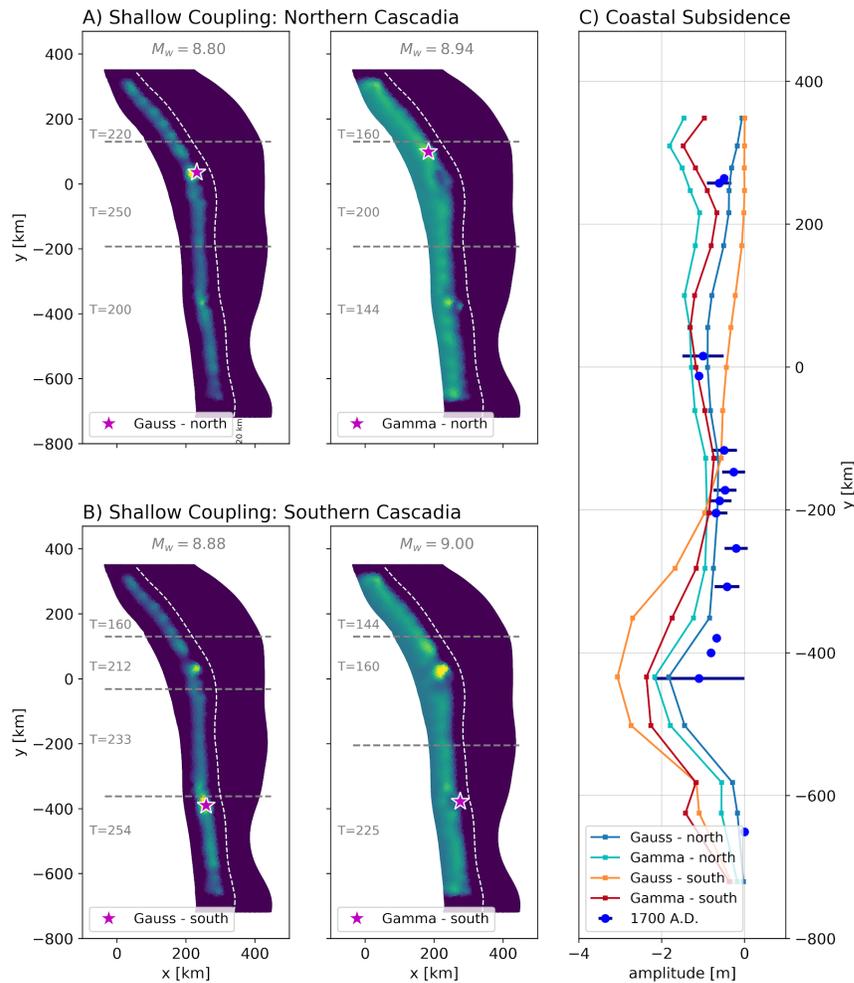
505

506 **Figure 7.** Effect of down-dip locking depth on coseismic slip distribution and uplift/subsidence patterns. A) Final slip distributions
 507 for the Gaussian coupling model nucleated in northern Cascadia. The earthquake is nucleated at the red star and a profile of
 508 uplift/subsidence at the free surface is plotted in figure B (black line through red star). In each panel, the locking depth is
 509 systematically deepened by 2 km. B) Model-predicted coseismic subsidence and uplift for the range of locking depths studied. The
 510 coastline is plotted for reference. Each solid line represents the coseismic hinge-line and is colored by its respective locking depth.
 511 We also show a shallower locking depth (15 km) in yellow for comparison.

512

We present four 3-D dynamic rupture scenarios derived from Gaussian and Gamma
 513 coupling distributions with a shallower locking depth at 15 km, but we also test a deeper locking
 514 depth (see Discussion). The T levels and segment locations were selected through a trial-and-
 515 error approach. These dynamic source models show 1700 A.D. subsidence data can be
 516 reasonably fit without invoking high amplitude slip deficit or subevents (Figure 8). Ruptures
 517 initiated in northern Cascadia with modest T levels (≤ 250 years) can match subsidence data
 518 with three segments (Figure 8a) whereas we find that four segments are required for the

519 Gaussian rupture model initiated in southern Cascadia (Figure 8b). The Gaussian-type simulation
 520 initiated in southern Cascadia has a final rupture length ~ 100 km shorter than the other ruptures.
 521 We note that the Gaussian simulation nucleated in northern Cascadia provides the best fit to the
 522 1700 A.D. subsidence data (Figure 8b, blue line).



523
 524 **Figure 8.** Gaussian and Gamma coupling models with shallow locking depth (15 km) that match coastal 1700 A.D. subsidence
 525 measurements to first order. A) Ruptures nucleated in northern Cascadia. B) Ruptures nucleated in southern Cascadia. C)
 526 Comparison of predicted coastal subsidence from simulations shown in A and B.

527 4. Discussion

528 4.1 What allows large earthquakes to develop along the Cascadia megathrust?

529 We observe margin-wide ruptures under conditions of higher relative dynamic stress drop
 530 amplitudes in the inferred creeping region of the central Cascadia megathrust (i.e., higher T

531 levels relative to the other segments). Alternatively, we also show that margin-wide ruptures are
532 promoted by slip weakening frictional behavior at shallow depth (Figure 6).

533 When the Gaussian or Gamma coupling models are used to generate heterogeneous shear
534 stress distributions, there are two natural locations to initiate spontaneous rupture: in northern or
535 southern Cascadia. Our results suggest that if rupture initiates in southern Cascadia, higher T
536 levels are required to sustain rupture through the central creeping region for Gaussian stress
537 distributions (Figure 5c). This is due to the combined effects from a lower slip-rate deficit
538 (inherent to both geodetic coupling models) and the generally narrower seismogenic rupture area
539 offshore Oregon caused by an increasing megathrust dip angle in this region.

540 A notable feature of our dynamic rupture simulations is that large earthquakes (M_w 8.8
541 and above) can be generated at much lower T levels than previously suggested from static
542 models (i.e., Wang et al., 2013). An explanation for this comes from dynamic effects within the
543 wedge. For instance, even though Gaussian simulations have little to no slip deficit extending to
544 the deformation front, reflections within the accretionary wedge appear to drive rupture
545 propagation along-strike. While a more realistic rheology within the wedge would certainly
546 affect wave propagation, our models suggest that wavefield inference at shallow depths could be
547 a viable mechanism to sustain rupture (Huang et al., 2014). Velocity-weakening, or in our case,
548 slip-strengthening, friction is a common assumption in dynamic rupture simulations of
549 megathrust earthquakes to represent the frictional behavior of sediments near the trench (Kozdon
550 and Dunham, 2013). One may also explicitly incorporate a subducting sediment channel
551 structure with depth-varying rigidity using slip-weakening friction (i.e., Ulrich et al., 2020). The
552 presence of clays or fluids within the megathrust fault zone can complicate the frictional
553 behavior, however (Saffer and Tobin, 2011). While Cascadia is well-known to have significant
554 sediment blanketing the trench along-strike with variable state of consolidation (Han et al.,
555 2017), there are no studies that directly sampled Cascadia megathrust fault gouge and subject
556 them to high-velocity friction experiments (Seyler et al., 2020). The assumption of slip-
557 strengthening friction in the upper 5 km in our dynamic rupture simulations is therefore modest
558 and will greatly benefit from offshore drilling data. We do not repeat the exercise of lowering the
559 dynamic friction level for ruptures in southern Cascadia, but we expect that a similar behavior
560 would occur.

561 The bulk of this study explored T levels that are below 320 years and exhibit
562 segmentation along-strike. It is geologically reasonable to presume the convergence rates of the
563 Juan de Fuca and Gorda plates have been stable for ~ 0.78 Ma (Demets et al., 2010). When
564 megathrusts ruptures occur, they can relieve the accumulated slip deficit partially or completely
565 (Hardebeck et al., 2012). Given this assumption, our results could be interpreted to mean that the
566 slip-rate deficit is not accumulated the same everywhere along the Cascadia megathrust. This is
567 consistent with suggested RI segmentation (e.g., Goldfinger et al., 2017), some of which may be
568 expressed as along-strike changes in fault vergence patterns and outer wedge geometry (Gulick
569 et al., 1998; Watt and Brothers, 2020). There is also the possibility that the coupling models we
570 used in our study overestimated the slip-rate deficit due to inelastic off-fault deformation (e.g.,
571 Baker et al., 2013). On the other hand, if the slip-rate deficit from plate convergence is being
572 accommodated along Cascadia at a constant rate, then in order to achieve reasonable coastal
573 subsidence amplitudes (> -2 m) our models imply a partial stress drop during the 1700 A.D.
574 event. We also lack precise constraints on the absolute shear stress and fault strength levels, for
575 which we had to make reasonable assumptions in order to estimate the relative dynamic stress
576 drop. Because there are no geophysical recordings of the 1700 A.D event, a detailed slip
577 inversion will remain out of reach. However, studies have inferred potential far-field tsunami
578 patterns of this earthquake using simple kinematic rupture models and Japanese written records
579 (e.g., Satake et al., 2003; 2020). Seafloor deformation predicted from dynamic rupture models
580 may be readily incorporated into kinematic tsunami propagation simulations - differences in
581 nucleation or peak slip locations (Figure 8) may alter our understanding of historic tsunamigenic
582 earthquakes in Cascadia.

583 *4.2 Explaining 1700 A.D. subsidence patterns with Dynamic Rupture Simulations*

584 The geodetic coupling models we use show positive stress drop down to ~ 20 km depth
585 (Figure 1). On the other hand, 3-D kinematic simulations were able to match 1700 A.D.
586 subsidence data assuming positive stress drop does not extend deeper than a fixed coupling level
587 closer to 15 km depth (i.e., 1 cm/yr contour from Burgette et al., 2009; McCaffrey et al., 2013;
588 Wirth et al., 2019b). We show that dynamic rupture simulations which taper stress drop to 0 MPa
589 below 16 km depth can agree well with the 1700 A.D subsidence data, particularly for ruptures
590 initiated in northern Cascadia (Figure 8). While shallower locking depths generally provide a
591 stronger fit to the paleoseismic data, we were also able to construct a dynamic rupture simulation

592 with a 20-km locking depth that fits the data just as well (Figure S5). This result suggests that if
593 T levels are sufficiently low along the fault, the subsidence patterns can probably be fit by an
594 even deeper locking depth (> 20 km depth). The influence of a deeper locking depth is to move
595 the coseismic hinge-line landwards and increase the amplitude of subsidence and uplift (i.e.,
596 Figure 7b). As discussed in Kanda and Simons (2012), either the location of peak interseismic
597 uplift-rate or greatest coseismic subsidence can provide a stronger constraint on the extent of
598 coupling as opposed to the hinge-line. Unfortunately, both the interseismic uplift (i.e., Krogstad
599 et al., 2016) and paleoseismic subsidence data (i.e., Wang et al., 2013) are limited in the along-
600 dip direction for this subduction zone. We thus caution using only paleoseismic subsidence data
601 to uniquely constrain the down-dip rupture limit in Cascadia.

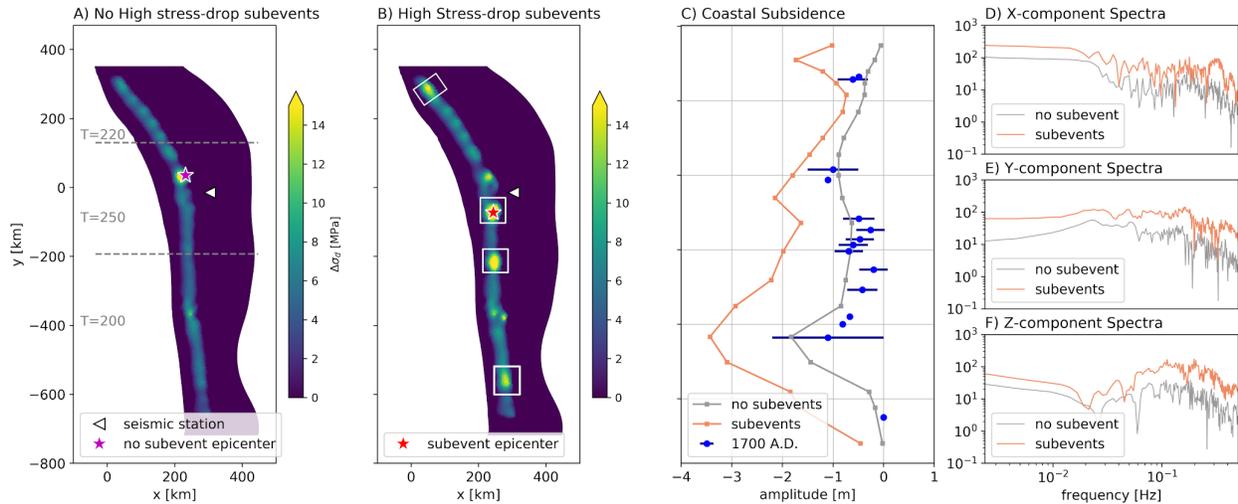
602 *4.3 The Potential of Heterogeneous Down-dip Frictional Properties*

603 The next Cascadia megathrust rupture may or may not include high-frequency seismic
604 energy radiated near the down-dip limit of slip (e.g., Lay, 2015), but one way to accomplish this
605 is to superimpose high stress drop subevents (>15 MPa) at several locations along-strike (Figure
606 9). For these simulations, we assume a locking depth of 15 km. The influence of subevents,
607 compared to a coupling model with no subevents, is to increase the subsidence amplitude and
608 generate higher relative seismic frequencies.

609 To conceptually demonstrate that heterogeneous D_c can also generate relatively higher
610 seismic frequencies in the specific case of the Cascadian margin, we also design a dynamic
611 rupture simulation containing several 16 km^2 asperities near the down-dip edge of the locked
612 megathrust that have lower $D_c = 1$ m with $D_c = 2$ m everywhere else (Figure S6). These D_c levels
613 are chosen to be consistent with already presented 3-D rupture simulations that can resolve the
614 cohesive zone widths. In this particular model, the effect of a heterogeneous D_c distribution is to
615 increase waveform amplitudes and high-frequency energy, with stations further away from the
616 hypocenter showing this more clearly (Figure S6).

617 What is unclear are properties most conducive to generating high frequency seismic
618 radiation down-dip, but dynamic rupture simulations for the 2011 Tohoku-Oki earthquake
619 showed that heterogeneous frictional properties or strength distributions (Huang et al., 2014;
620 Galvez et al, 2014) might account for these observations. We note that Cascadia is remarkably
621 different from the Japanese or Chilean subduction zones. In particular, the subducting interface
622 of the Juan de Fuca plate is relatively smooth along most of the margin compared to in the

623 aforementioned regions (van Rijsingen et al., 2018) and consequently, the interface topography
 624 may not provide an obvious explanation for future high seismic frequencies radiated down-dip.



625
 626 **Figure 9.** Comparison between dynamic rupture simulations with and without high dynamic stress drop subevents positioned near
 627 the down-dip edge of locking (~15 km) along-strike. A) Gaussian dynamic stress drop distribution without subevents. White
 628 triangle denotes synthetic seismogram receiver location. B) Gaussian dynamic stress drop distribution with superposed ~15 – 20
 629 MPa subevents (white boxes). C) Coastal subsidence for both models with 1700 A.D. observations for comparison. D), E), and F)
 630 show the raw spectral amplitudes of the x-, y-, and z-component velocity seismograms, respectively. The influence of the subevents
 631 is to increase the high-frequency amplitudes recorded (bold colored lines).

632 4.4 Limitations and Future Directions

633 Our study incorporates a physically consistent source model that emphasizes the
 634 importance of frictional and stress conditions necessary to generate M9-type ruptures. For lack of
 635 detailed information on the velocity structure in the accretionary prism and the highly simplified
 636 1-D CVM used, our 3-D dynamic rupture simulations do not capture accurate wave propagation
 637 effects along the Cascadia margin. Forecasting accurate ground motions during megathrust
 638 earthquakes is important, especially for subduction zones with limited or no seismic recordings
 639 (Frankel et al., 2018; Wirth et al., 2018). Developing dynamic rupture simulations that account
 640 for 3-D source, site, and path effects is one future direction that would, for instance, lead to more
 641 physically informed hazard estimates (e.g., Wirth et al., 2020).

642 Another limitation of our study is the rheology assumed: a linearly elastic body. There is
 643 potential for off-fault inelastic deformation in the wedge where there is a significant sediment
 644 volume (Ma, 2012). We note further that our choice of a linear slip-weakening friction law

645 allows us to assess first order along-strike and along-dip rupture limits, similar to Ramos &
646 Huang (2019). Modifying the friction law (and adjusting the finite element mesh resolution
647 accordingly) to account for strong rate-weakening would permit us to test a wider range of
648 rupture styles. Understanding what fault zone lithologies are present along the Cascadia
649 megathrust would also be helpful in assigning realistic dynamic friction levels during coseismic
650 rupture.

651 To improve the predictive capability of dynamic rupture simulations, offshore (e.g. near-
652 trench) geodetic measurements are needed. It would be particularly valuable if information about
653 the interseismic uplift-rate could be constrained offshore, to extend existing leveling data
654 onshore (Krogstad et al., 2016). This would reduce the ambiguity in geodetic coupling models
655 and improve our understanding of the spatial relationships between upper plate deformation and
656 intra-plate slip behavior (Bruhat and Segall, 2017; Watts and Brothers, 2020; Malatesta et al.,
657 2021). Our study stresses the importance of the spatial variation in coupling, especially in
658 the central Cascadia region where confirming the presence of lower coupling offshore Oregon is
659 critical for both kinematic and dynamic rupture simulation predictions.

660 Other geophysical measurements that have not been incorporated in this suite of dynamic
661 rupture simulations include inferences made about the seismogenic width from the arguably
662 highest resolution geophysical dataset available: the free-air gravity field. Basset et al., (2015)
663 observed that trench-parallel ridges in the free-air gravity anomaly field correlate well to the top
664 of slow-slip and tremor across the Cascadia forearc. If such trench-parallel features in the gravity
665 field are a proxy for down-dip rupture extent, then the transition from slip-weakening to slip-
666 strengthening frictional behavior may extend to depths greater than 20 km in some regions of
667 Cascadia. More work is needed to identify what geologic or geophysical features are most
668 indicative of future coseismic rupture limits, especially in subduction zones like Cascadia that
669 have not experienced megathrust events during the modern era of instrumentation.

670 **5. Conclusions**

671 Developing realistic seismic source models for the Cascadia megathrust is of paramount
672 importance to assist with seismic and tsunami risk mitigation. We present 3-D dynamic rupture
673 simulations that incorporate different hypotheses for megathrust strain accumulation based on
674 geodetic coupling models. We show that in order for margin-wide, ‘M9’ type ruptures to
675 develop, there must be a sufficiently high relative dynamic stress drop in the central Cascadia

676 region. Moreover, a slip weakening behavior or moderate slip deficit close to the deformation-
677 front can greatly facilitate margin-wide ruptures. Along-strike variations in the slip deficit
678 pattern relative to the geodetic coupling models are required to match available paleoseismic
679 data in our dynamic rupture models. We note that strain accumulation times lower than those
680 suggested from paleoseismic studies provide a better fit to the subsidence data, which might
681 suggest coupling models are overpredicting the slip-rate deficit or there was incomplete stress
682 drop from the last megathrust rupture. A close fit to 1700 A.D. subsidence data can be achieved
683 using Gaussian or Gamma coupling distributions with locking depths of 15 or 20-km depth,
684 obviating the need to call upon localized, high amplitude slip asperities along the down-dip
685 region of the seismogenic zone.

686 This work is a step forward in using fully dynamic rupture simulations for seismic hazard
687 analysis where there have been no instrumentally recorded ruptures. Kinematic rupture
688 properties (e.g., rise-time, slip-rate and rupture speed) and static seafloor displacement from our
689 dynamic simulations can be readily incorporated into existing 3-D kinematic rupture simulations
690 or inform tsunami propagation and coastal inundation models.

691

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700 results are available on UM Deep Blue (<https://deepblue.lib.umich.edu/>). The open-source SeisSol
701 software is available for download and installation at <https://github.com/SeisSol/SeisSol>

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