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Logarithmic corrections to the entropy of rotating black holes and black strings in AdS₅

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ABSTRACT: We investigate logarithmic corrections to the entropy of supersymmetric, rotating, asymptotically AdS_5 black holes and black strings. Within the framework of the AdS/CFT correspondence, the entropy of these black objects is determined, on the field theory side, by the superconformal index and the refined topologically twisted index of $\mathcal{N} = 4$ supersymmetric Yang-Mills theory, respectively. We read off the logarithmic correction from those field-theoretic partition functions. On the gravity side, we take the near-horizon limit and apply the Kerr/CFT correspondence whose associated charged Cardy formula describes the degeneracy of states at subleading order and determines the logarithmic correction to the entropy. We find perfect agreement between these two approaches. Our results provide a window into precision microstate counting and demonstrate the efficacy of low-energy, symmetry-based approaches such as the Kerr/CFT correspondence for asymptotically AdS black objects under certain conditions.

KEYWORDS: AdS-CFT Correspondence, Black Holes, Gauge-Gravity Correspondence

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1 Introduction

One of the fundamental results in the study of black holes during the 1970's was to establish that they can be attributed an entropy [1, 2], which further fits into a set of rules similar to the thermodynamic laws [3-5]. At the leading order in Newton's constant, the entropy was found to be universally given by a quarter of the area of the horizon. When further quantum effects are taken into consideration, logarithmic corrections to the entropy arise.

Entropy is an important bridge between the macroscopic and the microscopic worlds, and we expect any candidate quantum theory of gravity to provide an explanation for the macroscopic Bekenstein-Hawking entropy in terms of microscopic counting of degrees of freedom. As a leading candidate for a theory of quantum gravity, string theory has already provided, in the work of Strominger and Vafa, a microstate counting of the entropy for certain asymptotically flat black holes [6]. It has also provided, through work by Sen and collaborators, an explanation for the logarithmic corrections to the entropy [7].

The AdS/CFT correspondence, stating a mathematical equivalence between certain theories containing gravity in AdS and field theories on the boundary of AdS [8], potentially translates intricate questions of black hole dynamics into questions in unitary field theories. Using the AdS/CFT correspondence, Benini, Hristov and Zaffaroni have given a microscopic foundation to the entropy of certain magnetically charged AdS_4 black holes in terms of a field theory counting of microstates [9]. More recently, microscopic foundations in terms of field-theoretic partition functions have been provided for rotating, electrically charged, asymptotically AdS black holes [10–17]. For some of these black holes it has been shown that the field-theoretic, microscopic correction to the entropy matches precisely the macroscopic logarithmic corrections arising from one-loop supergravity [18–21].

The recent developments use the full UV complete description of gravity in AdS_{d+1} , the supersymmetric field theories in d = 3, 4, 5, 6. Guided by the principle of renormalization group (RG) flow, one naturally expects that a universal property such as the black hole entropy can be recovered from an effective theory without full knowledge of the UV complete description, as verified in previous cases [22]. For AdS black holes, applying the Kerr/CFT correspondence [23–25], a unified microscopic description of the entropy of asymptotically AdS_{4,5,6,7} black holes was given in [26] for extremal and extended in [27, 28] to near-extremal black holes.

The goal of this paper is to explore logarithmic corrections to the entropy of AdS_5 black holes and black strings from both the UV complete boundary $\mathcal{N} = 4$ SYM theory and the near-horizon 2-dimensional CFT. From the UV complete field theoretical point of view, the observables are the topologically twisted index and the superconformal index, both of which are initially given in the grand-canonical ensemble. Only after a Legendre transform to the microcanonical ensemble does the index probe the degeneracies of states giving rise to black hole entropy [10, 29]. We should also point out that the Kerr/CFT correspondence, being entirely based on the near horizon geometry, is not suitable to probe the entropy of the black hole in the grand-canonical ensemble since the values of the chemical potentials are to be fixed at the boundary of the full geometry. Hence, we probe the black hole entropy at fixed charge and angular momentum [30] in the microcanonical ensemble, and compare the results from the two approaches in this ensemble.¹

We review the logarithmic corrections as arising from the analysis of the large-N limit of the corresponding partition functions responsible for the UV complete counting of the entropy for the black holes and black strings in AdS₅. Namely, we consider the logarithmic corrections to the superconformal index and the topologically twisted index of $\mathcal{N} = 4$ SYM. We also exploit the fact that in the Kerr/CFT approach the degeneracy of states is given by the charged Cardy formula, which can and has been evaluated to the subleading order that contains logarithmic corrections [30–32]. Logarithmic corrections to the entropy can be computed macroscopically from one-loop contributions of the massless sector in the corresponding supergravity theory [33–35], and therefore provide a litmus test to any UV complete description. We find that both approaches yield the same logarithmic correction in the microcanonical ensemble.

This paper is organized as follows. In section 2 we review the AdS_5 black hole entropy from the superconformal index of the boundary $\mathcal{N} = 4$ SYM and from the nearhorizon Kerr/CFT, and then compute the logarithmic correction to the entropy using both approaches. In section 3 we apply similar techniques and use two approaches to compute

¹Further discussion about the entropy of black hole in different ensembles can be found in [7, 31].

the logarithmic correction to the entropy of rotating AdS_5 black strings. A brief summary and some proposals for the future research are presented in section 4. Some relevant facts about the special functions are collected in appendix A.

2 AdS₅ black holes

2.1 The superconformal index and black hole entropy

An efficient way to count $\frac{1}{16}$ -BPS states in $\mathcal{N} = 4$ SYM is to consider the theory on $S^1 \times S^3$ and evaluate the superconformal index (SCI) [36, 37]:

$$\mathcal{I}(\tau;\Delta) = \operatorname{Tr}\left[(-1)^F e^{-\beta \{\mathcal{Q}, \mathcal{Q}^{\dagger}\}} v_a^{Q_a} p^{J_1 + \frac{r}{2}} q^{J_2 + \frac{r}{2}} \right], \qquad (2.1)$$

where β is the circumference of S^1 , and F is the fermionic operator, while $Q_{a=1,2,3}$ are flavor charges with associated fugacities $v_a = e^{2\pi i \Delta_a}$. With r we denote the R-charge. The fugacities $p = e^{2\pi i \tau}$ and $q = e^{2\pi i \sigma}$ are associated to the angular momenta $J_{1,2}$ of S^3 , and the combinations $J_{1,2} + \frac{r}{2}$ commute with the supercharge Q. In what follows we set $\tau = \sigma$ for simplicity. Note that the counting of states that the SCI offers should be seen as performed in the grand-canonical ensemble, since we are keeping fixed the values of chemical potentials while summing over all possible charges.

According to the AdS/CFT correspondence, SU(N) $\mathcal{N} = 4$ SYM is dual to type IIB supergravity on AdS₅ × S⁵, in which one can find supersymmetric black hole solutions that are asymptotically AdS, rotating and electrically charged. Remarkably, in recent years plenty of evidences have been gathered indicating that $\mathcal{I}(\tau; \Delta)$ captures the entropy of such black holes [10–12] (see [38–44] for further developments and [45, 46] for a more complete list of references).

The SCI can be written as a contour integral over the holonomies of the gauge group [47, 48]:

$$\mathcal{I}(\tau; \Delta) = \kappa_N \int_0^1 \prod_{\mu=1}^{N-1} du_\mu \mathcal{Z}(u; \Delta, \tau) ,$$

$$\mathcal{Z}(u; \Delta, \tau) = \frac{\prod_{a=1}^3 \prod_{i \neq j} \widetilde{\Gamma}(u_{ij} + \Delta_a; \tau)}{\prod_{i \neq j} \widetilde{\Gamma}(u_{ij}; \tau)} ,$$

$$\kappa_N = \frac{(p; p)_{\infty}^{N-1}(q; q)_{\infty}^{N-1}}{N!} \prod_{a=1}^3 \left(\widetilde{\Gamma}(\Delta_a; \tau) \right)^{N-1} ,$$

(2.2)

where $(\cdot; \cdot)_{\infty}$ is the Pochhammer symbol, and $\widetilde{\Gamma}(u; \tau)$ is the elliptic Gamma function defined both in appendix A. There are two main approaches to evaluate the N-dimensional integral over the holonomies of the gauge group representing the SCI. The first approach relies on a direct application of the residue theorem, and yields what is known in the literature as the Bethe-Ansatz method. The second approach implements a saddle-point evaluation of the integral.

2.1.1 The Bethe-Ansatz approach

The location of the poles of (2.2) is given by the solutions to the set of equations:

$$Q_k(\hat{u}; \Delta, \tau) = 1, \quad \forall \ k = 1, \cdots, N,$$
(2.3)

where

$$Q_k(u;\Delta,\tau) = e^{2\pi i\lambda} \prod_{l=1(\neq k)}^N \prod_{a=1}^3 \frac{\theta_1(-u_{kl} + \Delta_a;\tau)}{\theta_1(u_{kl} + \Delta_a;\tau)}$$
(2.4)

are the Bethe-Ansatz operators and the values \hat{u} satisfying (2.3) are called Bethe-Ansatz solutions. We then define BA = { $\hat{u} \mid (2.3)$ is satisfied}. With λ we have denoted a Lagrange multiplier implementing the SU(N) constraint on the holonomies $\sum_{i=1}^{N} u_i \in \mathbb{Z}$, and $\theta_1(u;\tau)$ is the elliptic theta function defined in appendix A. Upon direct application of the residue theorem, $\mathcal{I}(\tau, \Delta)$ can be rewritten in terms of a discrete sum as:

$$\mathcal{I}(\tau; \Delta) = \kappa_N \sum_{\hat{u} \in BA} \mathcal{Z}(\hat{u}; \Delta, \tau) H(\hat{u}; \Delta, \tau)^{-1},$$

$$H(\hat{u}; \Delta, \tau) = \det \left[\frac{1}{2\pi i} \frac{\partial (Q_1, \cdots, Q_N)}{\partial (u_1, \cdots, u_{N-1}, \lambda)} \right].$$
(2.5)

Let us emphasize that (2.5) is not the full story, since the application of the residue theorem required the poles to be isolated, and there is enough evidence by now [49, 50] that this is not the case generically. We shall focus only on the contributions coming from isolated poles (see [51] for more detailed discussions on this point). A set of solutions to the equations (2.3) was found in [52] and it is given by:

$$u_{i} = u_{\hat{j},\hat{k}} = \bar{u} + \frac{\hat{j}}{m} + \frac{\hat{k}}{n} \left(\tau + \frac{r}{m}\right) ,$$

$$\hat{j} = 0, \cdots, m - 1 , \qquad \hat{k} = 0, \cdots, n - 1 ,$$

$$r = 0, \cdots, n - 1 ,$$
(2.6)

where N = mn, hence, each set $\{u_i\}$ in (2.6) can be labeled by the numbers $\{m, n, r\}$. These solutions to the Bethe-Ansatz equations for the SCI were, in fact, inspired by the set of solutions found in [52] for the Bethe-Ansatz equations associated to the topologically twisted index. We will discuss that case in section 3.1. However, in the large-N limit, it was possible to argue that the configuration corresponding to $\{1, N, 0\}$ contributed dominantly to the SCI. We shall refer to the $\{1, N, 0\}$ solution as the "basic" solution, namely

$$\hat{u}_{\text{basic}} = \left\{ u_i = \bar{u} + \frac{i}{N} \tau \, \Big| \, i = 1, 2, \cdots, N - 1 \right\} \bigcup \left\{ u_N = \bar{u} \right\} \,, \tag{2.7}$$

where \bar{u} is determined as

$$N\bar{u} + \frac{N(N-1)}{2N}\tau \in \mathbb{Z}.$$
(2.8)

The parameter \bar{u} enforces the SU(N) constraint $\sum_{i=1}^{N} u_i \in \mathbb{Z}$ on the holonomies, which, together with the periodicity properties of $\mathcal{I}(\tau; \Delta)$ allows us to obtain:

$$\bar{u} = \frac{k}{N} - \frac{N-1}{2N}\tau, \qquad k = 1, \cdots, N-1,$$
(2.9)

each of which contributes identically to the $\mathcal{I}(\tau; \Delta)$, we therefore have that, in the appropriate regime of chemical potentials:

$$\log \mathcal{I}(\tau; \Delta) \big|_{\text{Basic BA}} = -\frac{i\pi(N^2 - 1)}{\tau^2} \Delta_1 \Delta_2 \Delta_3 + \log N + \mathcal{O}(N^0).$$
(2.10)

We see that the log N in (2.10) has a purely combinatorial origin, whose precise form is quite insensitive to details about the theory in which $\mathcal{I}(\tau; \Delta)$ is being evaluated.

2.1.2 The saddle point approach and the Cardy-like expansion

Let us further reinforce the idea that the logarithmic correction to the SCI has a combinatorial origin. To do so we briefly reproduce here the saddle-point evaluation of (2.2)implemented in [45].

The strict Cardy-like limit. By the strict Cardy-like limit we mean that we keep only the most divergent term in a $\tau \to 0$ expansion.² The study of the strict Cardy-like limit was the subject of several works [11, 41, 54–57], and the main idea is to rewrite (2.2) in the following way:

$$\mathcal{I}(\tau; \Delta) = \kappa_N \int \prod_{\mu=1}^{N-1} du_\mu \exp\left(\frac{1}{\tau^2} S_{\text{eff}}(u; \Delta, \tau)\right), \qquad (2.11)$$

where $S_{\text{eff}}(u; \Delta, \tau)$ is appropriately defined such that $\mathcal{Z}(u; \Delta, \tau)$ in (2.2) is recovered. The $\frac{1}{\tau^2}$ factor can be used as a large control parameter to apply the saddle-point method in the strict Cardy-like limit. We are exploiting the fact that we already know the leading contribution in such limit is precisely of the order $\mathcal{O}\left(\frac{1}{\tau^2}\right)$. The saddle-point equations have the form:

$$\frac{\partial}{\partial u_{\mu}} S_{\text{eff}}(u; \Delta, \tau) \Big|_{u=\text{saddle}} = 0, \qquad (\mu = 1, \cdots, N-1).$$
(2.12)

The set with all identical holonomies, namely $u_i = u_j$ for all $i, j \in \{1, \dots, N\}$ [11, 55] is one of the most well-known solutions to (2.12). The effective action at this saddle point successfully counted the dual AdS₅ black hole microstates [11].

There are N distinct sets of identical holonomies satisfying the SU(N) constraint $\sum_{i=1}^{N} u_i \in \mathbb{Z}$, namely

$$u^{(m)} = \left\{ u_j^{(m)} = \frac{m}{N} \, \big| \, j = 1, \cdots, N \right\} \,, \qquad (m = 0, 1, \cdots, N - 1) \,. \tag{2.13}$$

Within the appropriate range of chemical potentials, the saddle points (2.13) yield the following effective action:

$$\frac{1}{\tau^2} \sum_{m=0}^{N-1} S_{\text{eff}}(u^{(m)}; \Delta, \tau) = \exp\left(-\frac{i\pi(N^2 - 1)}{\tau^2} \Delta_1 \Delta_2 \Delta_3 + \log N + \mathcal{O}(-1/|\tau|)\right).$$
(2.14)

²More refined limits were considered in [53]. The authors discussed the limit where $q = e^{2\pi i \tau}$ approaches roots of unit.

From (2.13) and (2.14) we see that the logarithmic correction has its origin in the multiplicity of the saddle points. This result remains true even for more generic $\mathcal{N} = 1$ toric quiver gauge theories, as emphasized in [45], which renders the log N correction a quite robust one. Note that we have not made use of the large-N limit here, therefore, provided that we remain at small values of τ , (2.14) holds for finite N (Evidence in favor of this has been given in [49]).

The Cardy-like expansion. With the Bethe-Ansatz approach we have learned that even for generic values of τ , in the large-N limit, the log N is the same and arises from degeneracies of Bethe-Ansatz solutions. At this point we have shown that also for finite N, in the strict Cardy-like limit, the log N has a combinatorial origin.

We now proceed to include subleading corrections in inverse powers of τ and show that, indeed, the log N remains unchanged. This is an important step, since it helps us build an intuition that we later import to a different situation, namely the refined topologically twisted index, where we have only access to the strict Cardy-like limit and argue about the possibility of the combinatorial nature of log N to remain true as we depart from this limit. We then focus on the effective action evaluated near the leading saddle-point solution (2.13). Following [45], we make the Ansatz for saddle-point solutions in the finite Cardy-like expansion,

$$u^{(m)} = \left\{ u_j^{(m)} = \frac{m}{N} + v_j \tau \mid v_j \sim \mathcal{O}(|\tau|^0), \sum_{j=1}^N v_j = 0 \right\}, \quad (m = 0, 1, \cdots, N-1), \quad (2.15)$$

and evaluate the effective action around this Ansatz. For a suitable choice of chemical potentials, the following expression was obtained:

$$\log \mathcal{I}(\tau; \Delta) = -\frac{i\pi(N^2 - 1)}{\tau^2} \Delta_1 \Delta_2 \Delta_3 + \log N + \mathcal{O}(e^{-1/|\tau|}).$$
(2.16)

The exponentially suppressed correction comes from the asymptotic expansion of the building blocks of the effective action, namely elliptic Gamma functions and Pochhammer symbols (see appendix A.1). An important aspect about (2.16) is that it includes all power-like corrections in τ , and rather remarkably, it is a series that truncates at the leading order. A prominent role in the technical evaluation of (2.16) was the cancellation of the $\mathcal{O}(\tau^0)$ contribution which was given in terms of the effective action of a matrix model of SU(N) level k = N Chern-Simons theory on S^3 (see [45] for a more detailed discussion).

What seems like a rather technical step when analyzed from the strictly mathematical perspective of looking at the asymptotic behavior of $S_{\text{eff}}(u; \Delta, \tau)$, becomes very natural when viewed from an effective field theory perspective. Such analysis was carried out in [58], where the Cardy-like (small τ) expansion was shown to geometrically correspond to shrinking the S^1 circle, thus leading to an effective field theory on S^3 organized in inverse powers of the circumference of S^1 .

In particular, a careful treatment of the Kaluza-Klein reduction on S^1 yields a result compatible with (2.16), where the log N is associated to degeneracies of vacua. This effective field theory approach clarifies the organization of the index in inverse powers of $|\tau|$ and further confirms the logarithmic term as certain degeneracy of vacua [53, 58]. Specifically, the effective field theory approach allows to establish the existence of a minimum of $S_{\text{eff}}(u; \Delta, \tau)$ at u = 0, which spontaneously breaks the one-form symmetry \mathbb{Z}_N of the 4d $\mathcal{N} = 4$ SYM theory. The fact that u = 0 spontaneously breaks \mathbb{Z}_N implies the existence of exactly N - 1 additional local minima which contribute equally to the index, hence the log N correction to the logarithm of the SCI.

Summarizing, the logarithmic correction to the logarithm of the SCI, which we refer to as $\Delta \log \mathcal{I}_{CFT_4}$, has been shown to be robust. In [45] it was originally obtained using two different approaches to evaluate the index: the saddle-point approximation and the Bethe-Ansatz approach. In the latter approach, the logarithmic term appears as the degeneracy of the Bethe-Ansatz solutions. The same logarithmic contribution was also shown to persist for a large class of $\mathcal{N} = 1$ superconformal field theories. The form of the logarithmic correction was further confirmed in [59], which provides an interpretation for certain exponentially suppressed terms. In [60], the logarithmic corrections were extended to other gauge groups and the results were shown to be compatible with the SU(N) analysis.

The black hole entropy is extracted from the SCI by implementing an inverse Laplace transformation, which yields the degeneracy of a state with given energy and charges. In the regime of large charges, we can reduce the inverse Laplace transformation to a Legendre transformation using the saddle point approximation. This is tantamount to changing from a grand-canonical ensemble to a microcanonical one. At the leading order in N, the two ways of approaching the entropy should be equivalent. However, when studying subleading structures we have to be more careful, since the very process of going from one ensemble to the other could modify the subleading corrections we are trying to probe. To be more specific, let us call ΔS_{CFT_4} the subleading logarithmic correction to the black hole entropy. Then we expect that in general $\Delta S_{\text{CFT}_4} = \Delta \log \mathcal{I}_{\text{CFT}_4} + (\text{corrections from changing ensemble})$. Let us now study more carefully the contribution coming from the change of ensemble.

2.2 The logarithmic correction associated to changing ensemble

We denote \mathcal{I}_{GC} as the index computeded in the grand-canonical ensemble and \mathcal{I}_{MC} as the index in the microcanonical ensemble, i.e., the index for fixed values of the charges. We consider D chemical potentials μ_I $(I = 1, \dots, D)$ satisfying the constraint,

$$\sum_{I=1}^{D} c_I \mu_I = n_0, \tag{2.17}$$

where $c_I = 1$ for μ_I associated to electric charges and $c_I = -1$ for μ_I associated to angular momenta. We implement the inverse Laplace transform which takes us from the grand-canonical ensemble to the microcanonical ensemble

$$\mathcal{I}_{\rm MC} = \int d^D \mu \, d\Lambda \exp\left\{\log \mathcal{I}_{\rm GC} - \sum_{I=1}^D Q_I \mu_I - \Lambda \left(\sum_{I=1}^D c_I \mu_I - n_0\right)\right\},\tag{2.18}$$

where Λ is the Lagrange multiplier associated to the constraint (2.17). Note that we have already considered the case of equal angular momenta when computing the index

in the grand-canonical ensemble. Otherwise, we would have also needed an additional Lagrange multiplier accounting for the constraint among rotations. We know the logarithmic corrections in the grand-canonical ensemble takes the form

$$\log \mathcal{I}_{\rm GC} = \log \mathcal{I}_{\rm GC}^{\rm (leading)} + \log N.$$
(2.19)

Imposing (2.19), we find that the index in the microcanonical ensemble takes the form

$$\mathcal{I}_{\rm MC} = N \int d^D \mu \, d\Lambda \exp\left[\log \mathcal{I}_{\rm GC}^{\rm (leading)} - \sum_{I=1}^D Q_I \mu_I - \Lambda \left(\sum_{I=1}^D c_I \mu_I - n_0\right)\right]. \tag{2.20}$$

We are now ready to implement the saddle point method, keeping the subleading logarithmic corrections associated to the one-loop determinant. The saddle point equations are given as

$$\frac{\partial}{\partial \mu_{I}} \left[\log \mathcal{I}_{\rm GC}^{\rm (leading)} - \sum_{I=1}^{D} Q_{I} \mu_{I} - \Lambda \left(\sum_{I=1}^{D} c_{I} \mu_{I} - n_{0} \right) \right] = 0,$$

$$\frac{\partial}{\partial \Lambda} \left[\log \mathcal{I}_{\rm GC}^{\rm (leading)} - \sum_{I=1}^{D} Q_{I} \mu_{I} - \Lambda \left(\sum_{I=1}^{D} c_{I} \mu_{I} - n_{0} \right) \right] = 0,$$
(2.21)

which leads to

$$\frac{\partial \log \mathcal{I}_{GC}^{(\text{leading})}}{\partial \mu_I} = Q_I + c_I \Lambda,$$

$$\sum_{I=1}^D c_I \mu_I = n_0.$$
(2.22)

A very important property of $\log \mathcal{I}_{GC}^{(\text{leading})}$ is its homogeneity of degree one in the chemical potentials. This implies the following crucial relation

$$\log \mathcal{I}_{\rm GC}^{\rm (leading)} = \sum_{I=1}^{D} \mu_I \frac{\partial \log \mathcal{I}_{\rm GC}^{\rm (leading)}}{\partial \mu_I}.$$
 (2.23)

Evaluating at the saddle point values, we obtain

$$\log \mathcal{I}_{\rm GC}^{\star(\text{leading})} = \sum_{I=1}^{D} \mu_I^{\star} \left(Q_I + c_I \Lambda \right), \qquad (2.24)$$

such that the saddle point imposed on (2.20) yields

$$\mathcal{I}_{\rm MC} \approx N \exp\left\{\log \mathcal{I}_{\rm GC}^{\star(\text{leading})} - \sum_{I=1}^{D} Q_{I} \mu_{I}^{\star} - \Lambda \left(\sum_{I=1}^{D} c_{I} \mu_{I}^{\star} - n_{0}\right) - \frac{1}{2} \log \det\left(H\right)\right\} \\ = N \exp\left\{\sum_{I=1}^{D} \mu_{I}^{\star} \left(Q_{I} + c_{I} \Lambda\right) - \sum_{I=1}^{D} Q_{I} \mu_{I}^{\star} - \Lambda \left(\sum_{I=1}^{D} c_{I} \mu_{I}^{\star} - n_{0}\right) - \frac{1}{2} \log \det\left(H\right)\right\} \\ = N e^{n_{0} \Lambda - \frac{1}{2} \log \det\left(H\right)}.$$
(2.25)

The Hessian H has the form

$$H = \begin{pmatrix} \frac{\partial^2 \log \mathcal{I}_{\rm GC}^{\rm (leading)}}{\partial \mu_1^2} & \cdots & \frac{\partial^2 \log \mathcal{I}_{\rm GC}^{\rm (leading)}}{\partial \mu_1 \partial \mu_D} & \frac{\partial^2 \log \mathcal{I}_{\rm GC}^{\rm (leading)}}{\partial \mu_1 \partial \Lambda} \\ & \ddots & \ddots & \ddots \\ & \ddots & \ddots & \ddots \\ \frac{\partial^2 \log \mathcal{I}_{\rm GC}^{\rm (leading)}}{\partial \mu_D \partial \mu_1} & \cdots & \frac{\partial^2 \log \mathcal{I}_{\rm GC}^{\rm (leading)}}{\partial \mu_D^2} & \frac{\partial^2 \log \mathcal{I}_{\rm GC}^{\rm (leading)}}{\partial \mu_D \partial \Lambda} \\ \frac{\partial^2 \log \mathcal{I}_{\rm GC}^{\rm (leading)}}{\partial \Lambda \partial \mu_1} & \cdots & \frac{\partial^2 \log \mathcal{I}_{\rm GC}^{\rm (leading)}}{\partial \Lambda \partial \mu_D} & \frac{\partial^2 \log \mathcal{I}_{\rm GC}^{\rm (leading)}}{\partial \Lambda^2} \end{pmatrix}.$$
(2.26)

Since $\log \mathcal{I}_{GC}^{(\text{leading})}$ is a homogeneous function of degree one, the chemical potentials can appear either in the numerator or the denominator in a way that the second derivative terms appearing along the diagonal of H vanish when μ_I appears in the denominator. To keep track of this we define a list of numbers $\{\delta_1, \dots, \delta_D\}$ such that δ_I vanishes when μ_I is in the numerator of $\log \mathcal{I}_{GC}^{(\text{leading})}$ and it is equal to one otherwise. This implies the following scaling of H

$$\det H \sim \det \begin{pmatrix} \mathcal{O}(N^2)\delta_1 & \cdots & \mathcal{O}(N^2) & c_1 \\ \cdot & \cdot & \cdot & \cdot \\ \cdot & \cdot & \cdot & \cdot \\ \mathcal{O}(N^2) & \cdots & \mathcal{O}(N^2)\delta_D & c_D \\ c_1 & \cdots & c_D & 0 \end{pmatrix} \sim \mathcal{O}(N^{2(D-1)}).$$
(2.27)

Defining D = d + 1, where d is the number of independent chemical potentials, the index computed in the microcanonical ensemble up to logarithmic corrections takes the form

$$\log \mathcal{I}_{\rm MC} \approx n_0 \Lambda + (1-d) \log N. \tag{2.28}$$

From (2.28) it is clear that from the CFT_4 point of view the origin of the logarithmic correction is two-fold: a part log N from the grand-canonical ensemble and another part $-d \log N$ from changing the grand-canonical ensemble to the microcanonical ensemble. Since the chemical potentials are constrained by having equal angular momenta as well as the BPS condition, we have d = 3 and therefore the logarithmic correction in the microcanonical ensemble is

$$\Delta S_{\rm CFT_4} = -2\log N. \tag{2.29}$$

We expect this 4-dimensional result to match with the subleading correction coming from the 2-dimensional Cardy formula.

2.3 Black hole, its entropy and near-horizon limit

The non-extremal asymptotically AdS₅ black hole background was found in [61]. In the Boyer-Lindquist coordinates $x^{\mu} = (t, r, \theta, \phi, \psi)$, the metric and the gauge field are given

by³

$$ds^{2} = -\frac{\left[(1+g^{2}r^{2})\rho^{2}dt + 2q\nu\right]dt}{\Xi\rho^{2}} + \frac{2q}{\rho^{2}\Xi}\nu^{2} + \frac{f}{\rho^{4}\Xi^{2}}(dt-\nu)^{2} + \frac{\rho^{2}dr^{2}}{\Delta_{r}} + \frac{\rho^{2}}{\Xi}\left(d\theta^{2} + \sin^{2}\theta\,d\phi^{2} + \cos^{2}\theta\,d\psi^{2}\right),$$
(2.30)

$$A = \frac{\sqrt{3} q}{\rho^2 \Xi} (dt - \nu) , \qquad (2.31)$$

where

$$\nu \equiv a \left(\sin^2 \theta \, d\phi + \cos^2 \theta \, d\psi \right), \qquad \Xi \equiv 1 - a^2 g^2,$$

$$\Delta_r \equiv \frac{(r^2 + a^2)^2 (1 + g^2 r^2) + q^2 + 2a^2 q}{r^2} - 2m,$$

$$\rho^2 \equiv r^2 + a^2, \qquad f \equiv 2m\rho^2 - q^2 + 2a^2 q g^2 \rho^2.$$
(2.32)

These black hole solutions are characterized by three independent parameters (m, a, q), and g is the inverse radius of AdS₅.

We are ultimately interested in exploring the black hole solution for the parameter space satisfying supersymmetry and extremality, i.e. BPS. The supersymmetric limit corresponds to

$$q = \frac{m}{1+2ag}.$$
(2.33)

However, this is not enough to ensure physical solutions and therefore we must also consider an additional constraint to prevent naked closed timelike curves, which in the BPS limit takes the form

$$m = \frac{2a(1+ag)^2(1+2ag)}{g}.$$
 (2.34)

Extremality occurs when the inner horizon and the outer horizon coincide, which for our solution gives the double root

$$r_0^2 = \frac{a(2+ag)}{g}.$$
 (2.35)

The macroscopic Bekenstein-Hawking entropy for the supersymmetric black hole, computed as a quarter of the area of the horizon (in units of $G_N = 1$), is

$$S_{\rm BH} = \frac{\pi^2 a^{3/2} \sqrt{2 + ag}}{g^{3/2} (1 - ag)^2} = 2\pi \sqrt{\frac{3Q^2}{g^2} - \frac{\pi}{2g^3}J}, \qquad (2.36)$$

where we have written it explicitly in terms of the electric charge, Q, and the angular momentum, J. The remarkable achievement of [10–12] was to obtain this expression for

³For simplicity, we consider the black hole with equal angular momenta $J_1 = J_2$ and equal electric charges $Q_1 = Q_2 = Q_3$.

the black hole entropy as the Legendre transform of the leading N^2 -part of the SCI (2.16), thus providing it with a microscopic explanation.

Given that the AdS/CFT correspondence geometrizes RG flow in the radial direction, it is convenient to consider zooming into a near-horizon region (IR), r_0 , while assuming a co-rotating frame:

$$r \to r_0 + \lambda \,\widetilde{r} \,, \quad t \to \frac{\widetilde{t}}{\lambda} \,, \quad \phi \to \widetilde{\phi} + g \frac{\widetilde{t}}{\lambda} \,, \quad \psi \to \widetilde{\psi} + g \frac{\widetilde{t}}{\lambda} \,,$$
 (2.37)

where we have also imposed both (2.33) and (2.34). Taking $\lambda \to 0$ brings us to a near-horizon region of the AdS₅ BPS black hole:

$$ds^{2} = \alpha_{1} \left[-\tilde{r}^{2} d\tau^{2} + \frac{d\tilde{r}^{2}}{\tilde{r}^{2}} \right] + \Lambda_{1}(\theta) \left[d\tilde{\phi} + \alpha_{2} \,\tilde{r} \,d\tau \right]^{2} + \Lambda_{2}(\theta) \left[d\tilde{\psi} + \beta_{1}(\theta) d\tilde{\phi} + \beta_{2}(\theta) \,\tilde{r} \,d\tau \right]^{2} + \alpha_{3} \,d\theta^{2} \,, \qquad (2.38)$$

where

$$\begin{aligned} \alpha_1 &= \frac{a}{2g(1+5ag)}, \\ \alpha_2 &= \frac{3a(1-ag)}{2(1+5ag)\sqrt{a\left(a+\frac{2}{g}\right)}}, \\ \alpha_3 &= \frac{2a}{g(1-ag)}, \\ \Lambda_1(\theta) &= \frac{4a(2+ag)\sin^2\theta}{g(1-ag)(4-ag+3ag\cos(2\theta))}, \\ \Lambda_2(\theta) &= \frac{a(4-ag+3ag\cos(2\theta))\cos^2\theta}{2g(1-ag)^2}, \\ \beta_1(\theta) &= \frac{6ag\sin^2\theta}{4-ag+3ag\cos(2\theta)}, \\ \beta_2(\theta) &= \frac{3g(1-ag)\sqrt{a\left(a+\frac{2}{g}\right)}}{(1+5ag)(4-ag+3ag\cos(2\theta))}. \end{aligned}$$
(2.39)

It is in the near-horizon limit at extremality where we find that the near-horizon geometry is locally a $U(1)^2$ -bundle over AdS₂. The asymptotic symmetries of this space can be studied via the Kerr/CFT correspondence, which associates to each U(1)-fiber in (2.38) a central charge and an effective temperature in the CFT₂. We can apply the Kerr/CFT correspondence to either U(1), and the results of the black hole entropy from the Cardy formula are the same [25, 26].

2.4 Kerr/CFT correspondence and charged Cardy formula

Let us briefly review the Cardy formula which determines the degeneracy of states in a CFT_2 . We are interested in its application up to and including the logarithmic corrections to the degeneracy of states, with constraints among the charges and chemical potentials.

We consider the partition function of a CFT_2 with n global U(1) symmetries expressed in the grand-canonical ensemble

$$Z(\tau, \bar{\tau}, \vec{\mu}) = \operatorname{Tr} e^{2\pi i \tau L_0 - 2\pi i \bar{\tau} \bar{L}_0 + 2\pi i \mu_i P^i}, \qquad (2.40)$$

where P^i are the conserved charges of the global U(1)'s, and μ_i are the corresponding chemical potentials. One particular property of the CFT₂ with conserved currents is that under modular transformations

$$\tau \to \tau' = \frac{a\tau + b}{c\tau + d}, \quad \mu_i \to \mu'_i = \frac{\mu_i}{c\tau + d}, \quad i = 1, \cdots, n,$$
 (2.41)

the partition function transforms as (in a special choice of normalization)

$$Z\left(\tau',\vec{\mu}'\right) = e^{-2\pi i \left(\frac{c\mu^2}{c\tau+d}\right)} Z(\tau,\vec{\mu}).$$
(2.42)

Therefore, the modular invariance of the partition function requires

$$Z(\tau, \bar{\tau}, \vec{\mu}) = e^{-\frac{2\pi i \mu^2}{\tau}} Z\left(-\frac{1}{\tau}, -\frac{1}{\bar{\tau}}, \frac{\vec{\mu}}{\tau}\right) , \qquad (2.43)$$

where $\mu^2 \equiv \mu_i \mu_j k^{ij}$ with k^{ij} denoting the matrix of the Kac-Moody levels of the U(1) currents. The modular invariance (2.43) implies that for small τ

$$Z(\tau, \bar{\tau}, \vec{\mu}) \approx e^{-\frac{2\pi i \mu^2}{\tau}} e^{-\frac{2\pi i E_L^v}{\tau} + \frac{2\pi i E_R^v}{\bar{\tau}} + \frac{2\pi i \mu_i p_v^i}{\tau}}, \qquad (2.44)$$

where E_L^v , E_R^v and p_v^i are the lowest eigenvalues of L_0 , \bar{L}_0 and P^i respectively. Moreover, we take E_L^v , E_R^v to be negative, and $p_v^i = 0$, corresponding to an electrically neutral vacuum. Note that (2.44) is the grand-canonical partition function and does not contain logarithmic corrections, as opposite to the analogous quantity (2.16) in CFT₄. This is already an indication that the Kerr/CFT computation is probing a different object in the grand-canonical ensemble. Nevertheless, Kerr/CFT still encodes information about the actual degeneracy of states by transforming to the microcanonical ensemble, i.e., the black hole sector with fixed charges and angular momentum.

Let us take a moment to understand the charges p_i of the theory, which include the angular momenta p_1, p_2 and the electric charges p_3, p_4, p_5 , originally coming from the AdS₅ black hole solution. Particularly, for the 5d BPS black hole of interest, such charges obey a linear constraint of the generic form

$$\sum_{i=1}^{5} b^{i} p_{i} = M, \qquad (2.45)$$

where b^i are some constant coefficients and M is related to the mass of the black hole. Therefore, (2.45) implements the BPS conditions (2.33) and (2.34). Since we are considering $p^i \sim \mathcal{O}(N^0)$, it can be seen from (2.45) that $M \sim N^0$. In the grand-canonical ensemble, we fix chemical potentials and admit all values of charges. We consider a linear constraint among chemical potentials

$$\sum_{i=1}^{5} s^{i} \mu_{i} = C , \qquad (2.46)$$

where C is a constant of the order $\mathcal{O}(N^0)$. The constraint (2.17) is a special case of (2.46). As we are going to see, this leads to the result that in terms of the scaling of N, $s^i \sim k_{ii}^{-1}$. Moreover, in order to compare to the CFT₄ with equal angular momenta, we consider an additional constraint of the form

$$\sum_{i=1}^{2} \alpha^{i} \mu_{i} = 0, \qquad (2.47)$$

of which the constraint $\mu_1 = \mu_2$ is a special case. To clarify how we use these constraints to derive the logarithmic corrections, we carefully change to the microcanonical ensemble by integrating over chemical potentials while respecting the constraints (2.46) and (2.47). The density of states $\rho(\tau, \bar{\tau}, \vec{\mu})$ can be expressed as the inverse Fourier transform of $Z(\tau, \bar{\tau}, \vec{\mu})$

$$\rho(E_L, E_R, \vec{p}) = \int d\tau \, d\bar{\tau} \, d^n \mu \, d\lambda_1 \, d\lambda_2 \exp\left[2\pi i S(\mu, \tau, \bar{\tau}) + 2\pi i \lambda_1 \left(\sum_{i=1}^5 s^i \mu_i - C\right) + 2\pi i \lambda_2 \sum_{i=1}^2 \alpha^i \mu_i\right],\tag{2.48}$$

$$S(\mu,\tau,\bar{\tau}) = -\frac{\mu^2}{\tau} - \frac{E_L^v}{\tau} + \frac{E_R^v}{\bar{\tau}} + \frac{\mu_i p_v^i}{\tau} - \tau E_L + \bar{\tau} E_R - \mu_i p^i , \qquad (2.49)$$

where E_L , E_R and p^i are the eigenvalues of L_0 , \bar{L}_0 and P^i , respectively, and n denotes the number of independent chemical potentials. Before we proceed, we take a moment to discuss the scaling of the various expressions and parameters involved. This is a crucial step in understanding which terms contribute to the subleading corrections of the entropy. The modular parameters are order-1 parameters: $\tau, \bar{\tau} \sim \mathcal{O}(N^0)$. Similarly, we take $p^i \sim N^0$ and $C \sim N^0$. From (2.49), we find that $\mu^2 \sim \mu_i p^i \tau$, which solving for the scaling of μ_i gives

$$\mu_i \sim p^j k_{ij} \tau \quad \Rightarrow \quad \mu_i \sim \sum_j p^j k_{ij} \sim (s^i)^{-1}, \tag{2.50}$$

where we have made the summation over the indices explicit, to make it clear that the highest order in the summation should be the scaling of μ_i and s^i . Likewise, $E_L \sim E_R \sim E_L^v \sim E_R^v \sim \mu^2 \sim \sum_{i,j} p^i p^j k_{ij}$.

Therefore, we have related the different parameters to the matrix k_{ij} , where the scaling can be found via the Kac-Moody levels. There are two types of levels that we are interested in. The Kac-Moody level from the SU(2) rotation, i.e. k^{11} or k^{22} , is proportional to the central charge c [62], which is of the order of Newton's constant $G^{-1} \sim N^2$. The Kac-Moody levels from U(1) gauge symmetries, i.e. k^{ii} (i > 2), are proportional to N^{-2} [30, 63]. For the BPS AdS₅ black hole, the N-dependences of various factors are

$$E_L^v, E_R^v \sim N^2, \quad 4E_L - \mathcal{P}^2, E_R \sim N^2, \quad k^{11}, k^{22} \sim N^2, \quad k^{33}, k^{44}, k^{55} \sim N^{-2}, \quad (2.51)$$

which implies that $k_{11}, k_{22} \sim N^{-2}$ and $k_{33}, k_{44}, k_{55} \sim N^2$. Moreover, this also implies that

$$s^1 = s^2 \sim N^2, \quad s^3 = s^4 = s^5 \sim N^{-2},$$
 (2.52)

or likewise $s_1 = s_2 \sim N^0$, $s_3 = s_4 = s_5 \sim N^0$. Due to the definition of μ_i and p_i as $\mu^2 = \mu_i \mu_j k^{ij}$ and $\mathcal{P}^2 \equiv p_i p_j k^{ij}$, we lower and raise the indices of μ_i, p_i and s_i with k_{ij} . However, for α_i we do not need to raise indices with k_{ij} , as α_1 and α_2 are the net scalings due to that the right-hand side of (2.47) is zero.

With these scalings in mind, we can now proceed to compute the saddle point. From (2.48) and (2.49) let us define

$$\widetilde{S}(\mu, \tau, \bar{\tau}) \equiv S(\mu, \tau, \bar{\tau}) + \lambda_1 \left(\sum_{i=1}^5 s^i \mu_i - C \right) + \lambda_2 \sum_{i=1}^2 \alpha^i \mu_i \,. \tag{2.53}$$

The equations for the fixed points have the form

$$i = 1, 2: \qquad \frac{\partial \widetilde{S}}{\partial \mu^i} = -2\frac{k_{ij}\mu^j}{\tau} + \frac{(p_v)_i}{\tau} - p_i + \lambda_1 s_i + \lambda_2 \alpha_i = 0, \qquad (2.54a)$$

$$i = 3, 4, 5: \quad \frac{\partial \tilde{S}}{\partial \mu^i} = -2\frac{k_{ij}\mu^j}{\tau} + \frac{(p_v)_i}{\tau} - p_i + \lambda_1 s_i = 0,$$
 (2.54b)

$$\frac{\partial \tilde{S}}{\partial \tau} = \frac{\mu^2}{\tau^2} + \frac{E_L^v}{\tau^2} - \frac{\mu_i p_v^i}{\tau^2} - E_L = 0, \qquad (2.54c)$$

$$\frac{\partial S}{\partial \bar{\tau}} = -\frac{E_R^v}{\bar{\tau}^2} + E_R = 0, \qquad (2.54d)$$

$$\frac{\partial \widetilde{S}}{\partial \lambda_1} = \sum_{i=1}^5 s^i \mu_i - C = 0, \qquad (2.54e)$$

$$\frac{\partial \widetilde{S}}{\partial \lambda_2} = \sum_{i=1}^2 \alpha^i \mu_i = 0.$$
(2.54f)

We define the values of the saddle to be $(\mu_i)_0$, τ_0 and $\overline{\tau}_0$, such that (2.54a) gives

$$(\mu_i)_0 = \begin{cases} \frac{1}{2} k_{ij} \left(p_v^j - p^j \tau_0 + \lambda_1 s^j \tau_0 + \lambda_2 \alpha^j \tau_0 \right), & i = 1, 2, \\ \frac{1}{2} k_{ij} \left(p_v^j - p^j \tau_0 + \lambda_1 s^j \tau_0 \right), & i = 3, 4, 5, \end{cases}$$
(2.55)

where we can redefine p^i by shifting it as follows

$$\tilde{p}^{i} \equiv \begin{cases} p^{i} - \lambda_{1}s^{i} - \lambda_{2}\alpha^{i}, & i = 1, 2, \\ p^{i} - \lambda_{1}s^{i}, & i = 3, 4, 5. \end{cases}$$
(2.56)

Therefore, we rewrite

$$(\mu_i)_0 = \frac{1}{2} k_{ij} \left(p_v^j - \tilde{p}^j \tau_0 \right) \,, \tag{2.57}$$

with τ_0 satisfying

$$\tau_0^2 E_L = \mu_0^2 + E_L^v - (\mu_i)_0 p_v^i \,. \tag{2.58}$$

Using (2.57), we find that

$$\mu_0^2 = \frac{1}{4} \left[k_{ij} p_v^i p_v^j + \tau_0^2 k_{ij} \tilde{p}^i \tilde{p}^j \right] - \frac{1}{2} k_{ij} \tilde{p}^i p_v^j \tau_0 = \frac{1}{4} k_{ij} \tau_0^2 \tilde{p}^i \tilde{p}^j , \qquad (2.59)$$

where $k_{im}k^{il} = \delta_m^l$, and in the second equality we have assumed that the vacuum is electrically neutral, i.e. $p_v^i = 0$. Inserting (2.59) in (2.58), we obtain

$$\tau_0 = \pm i \sqrt{\frac{4(-E_L^v)}{4E_L - k_{ij}p^i p^j}}.$$
(2.60)

The saddle point for $\bar{\tau}$ trivially is $\bar{\tau}_0 = \pm i \sqrt{\frac{-E_R^o}{E_R}}$. Consequently, $(\mu_i)_0$ given by (2.57) now has the form

$$(\mu_i)_0 = -\frac{1}{2}k_{ij}\tilde{p}^j\tau_0 = \mp ik_{ij}\tilde{p}^j\sqrt{\frac{-E_L^v}{4E_L - k_{ij}\tilde{p}^i\tilde{p}^j}}.$$
(2.61)

Imposing (2.54f), we find

$$\sum_{i=1}^{2} \alpha^{i}(\mu_{i})_{0} = -k_{ij} \alpha^{i} \tilde{p}^{j} \sqrt{\frac{E_{L}^{v}}{4E_{L} - k_{ij} \tilde{p}^{i} \tilde{p}^{j}}} = 0 \quad \Rightarrow \quad \alpha^{1} k_{11} \tilde{p}^{1} + \alpha^{2} k_{22} \tilde{p}^{2} = 0.$$
 (2.62)

Choosing the normalization of the Kac-Moody levels such that $k_{11} = k_{22}$, we can solve for the Lagrange multiplier λ_2

$$\alpha^{1}k_{11}\tilde{p}^{1} + \alpha^{2}k_{22}\tilde{p}^{2} = 0,$$

$$\Rightarrow k_{11}\left(\alpha^{1}(p^{1} - \lambda_{1}s^{1} - \lambda_{2}\alpha^{1}) + \alpha^{2}(p^{2} - \lambda_{1}s^{2} - \lambda_{2}\alpha^{1})\right) = 0,$$

$$\Rightarrow \alpha^{1}p^{1} + \alpha^{2}p^{2} - \lambda_{1}(\alpha^{1}s^{2} + \alpha^{2}s^{2}) - \lambda_{2}((\alpha^{1})^{2} + (\alpha^{2})^{2}) = 0.$$
(2.63)

We now set $\alpha^1 = -\alpha^2$, since both p^1 and p^2 should have the same scaling. Therefore, we obtain

$$\alpha_1, \alpha_2 \sim 1, \tag{2.64}$$

and

$$\alpha^1 \sim \alpha^2 \sim 1\,,\tag{2.65}$$

where we do not need to raise indices with k_{ij} here, as α_1 and α_2 are the net scalings due to that the right-hand side of (2.47) is zero. Moreover, $s^1 = s^2$ as they correspond to the equal angular momenta. We then find from (2.63) that

$$\alpha^{1}(p^{1} - p^{2}) = 2(\alpha^{1})^{2}\lambda_{2},
\Rightarrow \lambda_{2} = \frac{p^{1} - p^{2}}{2\alpha^{1}},$$
(2.66)

which vanish for $p^1 = p^2$. This implies that λ_2 does not affect the logarithmic corrections to the entropy for the case of equal angular momenta, as its contribution to the determinant of the Hessian matrix is of the order $\mathcal{O}(N^0)$. We are also interested in the scaling of λ_1 . From (2.54e), we find

$$\sum_{i}^{5} s^{i}(\mu_{i})_{0} = -k_{ij}s^{i}\tilde{p}^{j}\sqrt{\frac{E_{L}^{v}}{4E_{L} - k_{ij}\tilde{p}^{i}\tilde{p}^{j}}} = C.$$
(2.67)

As we expect the scaling to remain the same for any arbitrary values of the charges, we consider a special case $p^i = p_v^i = 0$ and find from (2.56) and (2.67) that

$$\lambda_{1} \sum_{i=1}^{5} k_{ii} s^{i} s^{i} = C \sqrt{\frac{4E_{L} - \lambda_{1}^{2} \sum_{i=1}^{5} k_{ii} s^{i} s^{i}}{E_{L}^{v}}},$$

$$\Rightarrow \qquad \lambda_{1}^{2} \left(\sum_{i=1}^{5} k_{ii} s^{i} s^{i} \right)^{2} = C^{2} \left(\frac{4E_{L}}{E_{L}^{v}} - \frac{\lambda_{1}^{2} \sum_{i=1}^{5} k_{ii} s^{i} s^{i}}{E_{L}^{v}} \right), \qquad (2.68)$$

$$\Rightarrow \qquad \lambda_{1}^{2} \left[\left(\sum_{i=1}^{5} k_{ii} s^{i} s^{i} \right)^{2} + C^{2} \frac{\sum_{i=1}^{5} k_{ii} s^{i} s^{i}}{E_{L}^{v}} \right] = 4C^{2} \frac{E_{L}}{E_{L}^{v}}.$$

Let us now discuss the scalings of each of these terms. Given (2.51) and (2.52), we have at the leading order $E_L \sim E_L^v \sim N^2$, $C \sim N^0$ and $k_{ii}s^is^i \sim N^2$ and therefore

$$\lambda_1 \sim N^{-2}.\tag{2.69}$$

The leading order value of the degeneracy is obtained by evaluating the action at the saddle point values, which gives

$$\log \rho_0 = \sqrt{E_L^v \left(4E_L - k_{ij}\tilde{p}^i\tilde{p}^j\right)} + 2\sqrt{E_R E_R^v}.$$
(2.70)

We would like to comment on the scaling with respect to N in (2.70). At the leading order, $\lambda_1 s^i \sim \mathcal{O}(N^{-2})$ which implies that $\mathcal{O}(\tilde{p}^i) \sim \mathcal{O}(p^i)$. Therefore, (2.70) coincides with the leading order of the degeneracy with \tilde{p}^i replaced by p^i . This is important as we can see that the constraint we imposed only affects the subleading order of the entropy.

Note that we have more than one saddle points, namely one for each choice of signs in the values of $\tau_0, \bar{\tau}_0, (\mu_i)_0$. However, one saddle dominates over the others as $S(\mu, \tau, \bar{\tau})$ is exponentially suppressed. To see this explicitly, let us take

$$\tau_{0} = i\epsilon_{\tau} \sqrt{\frac{4(-E_{L}^{v})}{4E_{L} - k_{ij}p^{i}p^{j}}}, \quad \bar{\tau}_{0} = i\epsilon_{\bar{\tau}} \sqrt{\frac{-E_{R}^{v}}{E_{R}}}, \quad \mu_{i,0} = -i\epsilon_{\tau}k_{ij}p^{j} \sqrt{\frac{-E_{L}^{v}}{4E_{L} - k_{ij}p^{i}p^{j}}}, \quad (2.71)$$

where ϵ_{τ} and $\epsilon_{\bar{\tau}}$ take on values of ± 1 . Then, under the constraints imposed by the Lagrange multipliers, the density of states (2.48) can be approximated by the saddle points

$$\rho_{0} = \sum_{\epsilon_{\tau},\epsilon_{\bar{\tau}}=\pm 1} \exp\left(2\pi i S(\mu,\tau,\bar{\tau})\right)$$

$$= \sum_{\epsilon_{\tau},\epsilon_{\bar{\tau}}=\pm 1} \exp\left\{2\pi i \left[-\frac{i\epsilon_{\tau}}{2} \frac{\mathcal{P}^{2}\sqrt{-E_{L}^{v}}}{\sqrt{4E_{L}-\mathcal{P}^{2}}} - i\epsilon_{\tau}(-E_{L}^{v})\sqrt{\frac{4E_{L}-\mathcal{P}^{2}}{4(-E_{L}^{v})}} + i\epsilon_{\bar{\tau}}(-E_{R}^{v})\sqrt{\frac{E_{R}}{-E_{R}^{v}}}\right] - i\epsilon_{\tau}E_{L}\sqrt{\frac{-4E_{L}^{v}}{4E_{L}-\mathcal{P}^{2}}} + i\epsilon_{\bar{\tau}}E_{R}\sqrt{\frac{-E_{R}^{v}}{E_{R}}} - (-i\epsilon_{\tau})\mathcal{P}^{2}\sqrt{\frac{-E_{L}^{v}}{4E_{L}-\mathcal{P}^{2}}}\right]\right\}$$

$$= \sum_{\epsilon_{\tau},\epsilon_{\bar{\tau}}=\pm 1} \exp\left[2\pi \left(\epsilon_{\tau}\sqrt{-E_{L}^{v}(4E_{L}-\mathcal{P}^{2})} - \epsilon_{\bar{\tau}}\sqrt{-4E_{R}^{v}E_{R}}\right)\right]. \qquad (2.72)$$

If we now select the combination of $\epsilon_{\tau} = 1$ and $\epsilon_{\bar{\tau}} = -1$, which maximizes the exponent in (2.72), we can write

$$\rho_0 \sim \exp\left[2\pi \left(\sqrt{-E_L^v(4E_L - \mathcal{P}^2)} + \sqrt{-4E_R^v E_R}\right)\right] + \dots, \qquad (2.73)$$

where the dots denote the exponentially suppressed terms of subleading non-logarithmic order. To summarize, given the behavior of Z, the degeneracy ρ can be determined using the saddle-point approximation with the dominant saddle at

$$\tau_0 = \sqrt{\frac{4E_L^v}{4E_L - \mathcal{P}^2}}, \quad \bar{\tau}_0 = -\sqrt{\frac{E_R^v}{E_R}}, \quad \mu_{i,0} = -k_{ij} p^j \sqrt{\frac{E_L^v}{4E_L - \mathcal{P}^2}}, \quad (2.74)$$

where k_{ij} is the inverse matrix of k^{ij} , and $\mathcal{P}^2 \equiv p^i p^j k_{ij}$. Note that the saddle-point values τ_0 and $(\mu_i)_0$ are parametrically small, as $\mathcal{P}^2 \gg |E_L|$, which is reminiscent of the 4d Cardy limit originally used in [10, 11] and recently clarified in [53, 58]. Moreover, the Cardy limit in 2d, which assumes that the levels of the theory is much larger than the Casimir energy, is compatible with the Cardy limit in 4d, which focuses on small chemical potentials and large charges, as they both address the high energy states of the theory and in our particular case address the entropy of extremal black holes.

At the saddle (2.74), the density of states ρ reaches its extremum ρ_0 , and the corresponding entropy is

$$S(\mu_0, \tau_0, \bar{\tau}_0) = \log \rho_0 \approx 2\pi \sqrt{-E_L^v(4E_L - \mathcal{P}^2)} + 2\pi \sqrt{-E_R^v(4E_R)} \,. \tag{2.75}$$

This expression is also called the charged Cardy formula in [64], which implies a microcanonical ensemble of black hole microstates. If we apply $E_L^v = E_R^v = -c/24$, and define the temperatures $T_{L,R}$ through

$$E_L - \frac{\mathcal{P}^2}{4} = \frac{\pi^2}{6} c T_L^2, \quad E_R = \frac{\pi^2}{6} c T_R^2, \quad (2.76)$$

we can rewrite the entropy (2.75) as

$$S = \frac{\pi^2}{3}cT_L + \frac{\pi^2}{3}cT_R, \qquad (2.77)$$

where T_R is proportional to the physical Hawking temperature T_H . This formula coincides, at the leading order, with the canonical ensemble version of the charged Cardy formula, and has been successfully used in a variety of cases [23, 65–68]. However, we emphasize that obtaining (2.77) did not involve a change of ensemble, as we merely re-identified certain combinations.

From the near-horizon CFT_2 and the Kerr/CFT correspondence we know that for the BPS AdS_5 black hole

$$c_L = \frac{9\pi a^2}{G_N g (1 - ag)(1 + 5ag)} = \frac{18N^2 (ag)^2}{(1 - ag)(1 + 5ag)},$$
(2.78)

$$T_L = \frac{1+5ag}{3a(1-ag)\pi} \sqrt{a\left(a+\frac{2}{g}\right)},$$
(2.79)

where we have used the AdS_5/CFT_4 dictionary $\frac{1}{2}N^2 = \frac{\pi}{4G_N}\ell_5^3 = \frac{\pi}{4G_Ng^3}$. Note that both c_L and T_L are dimensionless. Consequently, the BPS AdS₅ black hole entropy at the leading order in N is given by the Cardy formula

$$S_{\rm CFT_2} = \frac{\pi^2}{3} c_L T_L = \frac{2N^2 \pi (ag)^{3/2} \sqrt{2+ag}}{(1-ag)^2} \,. \tag{2.80}$$

This near-horizon CFT_2 result matches the macroscopic Bekenstein-Hawking entropy of the black hole (2.36), as shown in [24–26].

2.5 Logarithmic corrections from near-horizon CFT₂

To derive the logarithmic corrections to the black hole entropy from the near-horizon CFT₂, we evaluate the Cardy formula beyond its leading saddle-point value by including its Gaussian correction. Namely, we consider a logarithmic correction ΔS_{CFT_2} obtained from expanding τ , $\bar{\tau}$ and $\vec{\mu}$ to the quadratic order around the saddle point given by (2.74). The result is

$$\Delta S_{\rm CFT_2} = -\frac{1}{2} \log \frac{\det \mathcal{A}}{(2\pi)^{n+2}}, \qquad (2.81)$$

where \mathcal{A} is the Hessian of the exponent in the integrand of (2.48) around the saddle point (2.74), and has the form $\mathcal{A}_{\mu\nu} = \frac{\partial^2 \widetilde{S}}{\partial x^{\mu} \partial x^{\nu}}$, where $x^{\mu} = \{\tau, \overline{\tau}, \lambda, \mu^{i=1,\dots,n}\}$, whose only non-trivial elements in the presence of constraints are

$$\frac{\partial^2 \widetilde{S}}{\partial \tau \partial \mu^i} = 2 \frac{k_{ij}(\mu^j)_0}{\tau_0},
\frac{\partial^2 \widetilde{S}}{\partial \tau^2} = -\frac{2}{\tau_0^3} \left(k_{ij}(\mu^i)_0(\mu^j)_0 + E_L^v \right),
\frac{\partial^2 \widetilde{S}}{\partial \overline{\tau}^2} = 2 \frac{E_R^v}{\overline{\tau}_0^3},
\frac{\partial^2 \widetilde{S}}{\partial \lambda_1 \partial \mu^i} = s_i, \quad (i = 1, \cdots, 5)
\frac{\partial^2 \widetilde{S}}{\partial \lambda_2 \partial \mu^i} = \alpha_i, \quad (i = 1, 2)
\frac{\partial^2 \widetilde{S}}{\partial \mu^i \partial \mu^j} = -2 \frac{k_{ij}}{\tau_0}.$$
(2.82)

We see that k_{11} and k_{22} come from the SU(2) rotation, which corresponds to the angular momenta, while k^{ii} (i > 2) come from the U(1) gauge symmetries. At the subleading order the Hessian takes the form

$$\det \mathcal{A} = \frac{(2\pi)^{n+2}}{16} \left(-E_L^v\right)^{-\frac{n+1}{2}} \left(4E_L - \mathcal{P}^2\right)^{\frac{n+3}{2}} \left(-E_R^v\right)^{-\frac{1}{2}} \left(4E_R\right)^{\frac{3}{2}} \det(H), \qquad (2.83)$$

where

$$H \sim \left(\frac{k_{ij} |\mathcal{S}^T|}{|\mathcal{S}|| 0}\right),\tag{2.84}$$

and

$$S = \begin{pmatrix} s_1 & s_2 & s_3 & s_4 & s_5 \\ \alpha_1 & \alpha_2 & 0 & 0 & 0 \end{pmatrix}.$$
 (2.85)

Note that this result is different than in [30], as we have considered two linear constraints on the chemical potentials.

For supersymmetric extremal (BPS) black holes, one of the Frolov-Thorne temperatures T_R vanishes, as it is proportional to the Hawking temperature, and only the left sector contributes to the black hole entropy. Consequently, (2.83) for BPS black holes becomes

$$\left(\det \mathcal{A}\right)_{\rm BPS} = \frac{(2\pi)^{n+2}}{16} \left(-E_L^v\right)^{-\frac{n+1}{2}} \left(4E_L - \mathcal{P}^2\right)^{\frac{n+3}{2}} \det(H), \qquad (2.86)$$

where

$$-E_L^v = \frac{c}{24}, \quad 4E_L - \mathcal{P}^2 = \frac{2\pi^2}{3}cT_L^2.$$
 (2.87)

With the scalings in (2.51), (2.52) and (2.64), the Hessian takes on the N-dependence

$$\det H \sim N^2, \tag{2.88}$$

such that

$$(\det \mathcal{A})_{AdS_5 Black Hole} \sim (N^2)^{-\frac{n+1}{2}} (N^2)^{\frac{n+3}{2}} (N^2) = N^4.$$
 (2.89)

Note that the result is independent of n since the scaling of E_L^v and $4E_L - \mathcal{P}^2$ are equal. Therefore, the logarithmic correction to the leading-order BPS AdS₅ black hole entropy (2.80) is

$$\Delta S_{\rm CFT_2} = -\frac{1}{2} \log \frac{\det \mathcal{A}}{(2\pi)^{n+2}} = -2 \log N + \mathcal{O}(1) \,, \tag{2.90}$$

which precisely agrees with ΔS_{CFT_4} in (2.29). However, we also observe that in contrast to the CFT₄ approach, where the logarithmic correction originates from both the grandcanonical ensemble and the process of changing ensemble, the logarithmic correction in the CFT₂ approach comes purely from changing the ensemble. This is because the Kerr/CFT is intrinsically a near-horizon field theoretic approach, and it does not probe the boundary of the full geometry, where the values of the chemical potentials in the grand-canonical ensemble are fixed. Hence, in Kerr/CFT it is more natural to consider the entropy in the microcanonical ensemble, and we only expect the results of the black hole entropy with logarithmic correction from the CFT₄ and the CFT₂ approaches match in this ensemble.

3 AdS₅ black strings

3.1 AdS₅ black string entropy from boundary $\mathcal{N} = 4$ SYM

A rotating AdS_5 black string solution in gauged supergravity has been discussed in [64, 69], where it was shown that its leading-order entropy can be obtained from the refined topologically twisted index of $\mathcal{N} = 4$ SYM on $S^2 \times T^2$. The topologically twisted index of $\mathcal{N} = 4$ SYM with gauge group SU(N) is defined as the supersymmetric index of the theory on $T^2 \times S^2$ with a topological twist on S^2 [70, 71], its Hamiltonian interpretation being

$$Z(p_a, \Delta_a) = (-1)^F e^{2\pi i \tau \{\mathcal{Q}, \mathcal{Q}^\dagger\}} e^{i\Delta_a J_a} .$$

$$(3.1)$$

The topologically twisted index depends on a set of chemical potentials, Δ_a , for the generators of flavor symmetries (a = 1, 2, 3), a modular parameter of the torus τ and magnetic fluxes p_a . The topologically twisted index of $\mathcal{N} = 4$ SYM with gauge group SU(N) admits a presentation as an integral over the space of holonomies in the following way:

$$Z(p_a, \Delta_a) = \frac{1}{N!} \sum_m \oint_{\mathcal{C}} \prod_{\mu=1}^{N-1} \left(du_\mu \eta(q)^2 \right) \mathcal{Z}_{TT}(u, \Delta_a, \tau, p_a) ,$$

$$\mathcal{Z}_{TT}(u, \Delta_a, \tau, p_a) = \prod_{i,j=1}^N \left[\frac{\theta_1(u_{ij}; \tau)}{i\eta(q)} \prod_{a=1}^3 \left(\frac{i\eta(q)}{\theta_1(u_{ij} + \Delta_a; \tau)} \right)^{m_{ij} - p_a + 1} \right] ,$$
(3.2)

where $\eta(q)$ is the Dedekind eta function that we define in appendix A. We can evaluate (3.2) as the sum over residues [71] which takes the following explicit form:

$$Z(p_a, \Delta_a) = \eta(q)^{2(N-1)} \sum_{\hat{u} \in BA} \prod_{i,j=1}^{N} \left[\prod_{a=1}^{3} \left(\frac{\theta_1(u_{ij}; \tau)}{\theta_1(u_{ij} + \Delta_a; \tau)} \right)^{1-p_a} \right] H^{-1}(\hat{u}, \Delta, \tau), \quad (3.3)$$

where, analogously to the SCI discussed in section 2.1, BA stands for the set of solutions to the Bethe-Ansatz equations (2.3), and $H(\hat{u}, \Delta, \tau)$ is the Jacobian defined in (2.5). The location of a set of such residues was found in [52] and have the form given by (2.6) labeled by $\{u_i\}$ with integers $\{m, n, r\}$.

In fact, the set of solutions found in [52] inspired the evaluation of the SCI carried in [12], where the $\{u_i\}$ are also organized according to equation (2.6), however, in the large-N limit, it was possible to argue that the configuration corresponding to $\{1, N, 0\}$ was dominant. For fixed $\{m, n, r\}$, it is possible to count how many values of \bar{u} give non-equivalent contributions to topologically twisted index (by non-equivalent we mean, those which are not identified by periodicity $u \sim u + 1$ or $u \sim u + \tau$). Once again, imposing the SU(N) constraint we find that:

$$\bar{u} = \frac{k}{N} - \frac{1}{2N} \left[n(m-1) + m(n-1) \left(\tau + \frac{r}{m} \right) \right],$$

$$k = 0, 1, \cdots, N-1,$$
(3.4)

which reduces to (2.9) for $\{m, n, r\} = \{1, N, 0\}$. We then conclude that there is a degeneracy factor of N for each $\{m, n, r\}$ configuration contributing to the topologically twisted index. To argue that there is no other contribution of the same order that spoils the value of the coefficient of log N would require a more detailed study of the large-N behavior of the topologically twisted index, which has been studied recently in [51] at the leading order in N. A systematic study of subleading corrections to the topologically twisted index still remains an open problem. It is, however, very tempting to conjecture that indeed, there is no contribution other than the one originated from degeneracy of Bethe-Ansatz solutions and, consequently, the coefficient of $\log N$ is 1 also for the topologically twisted index in the grand-canonical ensemble.

One can further refine the topologically twisted index by adding a rotation on S^2 [70]. This will modify the integral expression (3.2) through the appropriate fugacities associated to the rotation on S^2 , namely $\xi = e^{2\pi i \omega}$. To be concrete, we would have:

$$Z(p_a, \Delta_a)_{\text{refined}} = \frac{1}{N!} \sum_m \oint_{\mathcal{C}} \prod_{\mu=1}^{N-1} \left(du_\mu \eta(q)^2 \right) \mathcal{Z}_{TTref}(u, \Delta_a, \tau, p_a) ,$$

$$\mathcal{Z}_{TTref}(u, \Delta_a, \tau, p_a) = \prod_{i,j=1}^N \left[\frac{\theta_1 \left(u_{ij} + 2\omega j; \tau \right)}{i\eta(q)} \prod_{a=1}^3 \left(\frac{i\eta(q)}{\theta_1 \left(u_{ij} + \Delta_a + 2\omega j; \tau \right)} \right)^{m_{ij} - p_a + 1} \right] .$$
(3.5)

The refined topologically twisted index has been studied, in the strict Cardy-like limit, in [69], where the correction due to the refinement could be factored out in the following way:

$$Z(p_a, \Delta_a)_{\text{refined}}\big|_{\tau \to 0} = Z(p_a, \Delta_a) Z_\omega \,, \tag{3.6}$$

where $Z(p_a, \Delta_a)$ is the unrefined topologically twisted index, and Z_{ω} is the correction associated to the refinement. The explicit form of Z_{ω} is irrelevant to us, while only the fact that it is independent on u, p_a and Δ_a will be important. To the best of our knowledge, the direct application of the Bethe-Ansatz approach to the refined topologically twisted index has not been performed yet. However, we can exploit the fact that in the Cardy-like limit there is a simple connection to the unrefined index, namely (3.6), and based on the intuition we have gained by studying the SCI, to argue that the combinatorial origin of log N corrections is still there at small τ , therefore we do not expect it to go away as we depart from the Cardy-like limit.

As we have discussed in the AdS₅ black hole case, the logarithmic correction to the entropy can be seen as essentially arising from the degeneracy of dominant Bethe-Ansatz solutions to the appropriate partition function of the boundary $\mathcal{N} = 4$ SYM. As in the case of the SCI, the logarithmic correction we compute for the topologically twisted index is in the grand-canonical ensemble. However, since we find that the result matches that of the microcanonical ensemble, we conjecture that there are no additional logarithmic contributions associated to the change of ensembles. Therefore, for the BPS rotating AdS₅ black string considered in [64, 69], the logarithmic correction to $\log Z^{(\text{leading})}(p_a, \Delta_a)$ can be obtained from the degeneracy of dominant residues contributing to the topologically twisted index of $\mathcal{N} = 4$ SYM, i.e.

$$\Delta \log Z(p_a, \Delta_a) = \log N.$$
(3.7)

This result has the same origin (in the Bethe-Ansatz treatment [12]) as in the SCI, and we expect a similar robustness as the logarithmic correction to the AdS_5 black hole.

Since $\log Z^{(\text{leading})}(p_a, \Delta_a) \sim N^2$ and it is homogeneous of degree one in the chemical potentials, it is possible to apply the result of section 2.2 to conclude that the logarithmic correction has an additional contribution from the change of ensemble which again takes

the form $-d \log N$, where d = 3 is the number of independent chemical potentials for the rotating AdS₅ black string. We then conclude that

$$\Delta S_{\rm CFT_4} = (1 - d) \log N + \mathcal{O}(1) = -2 \log N + \mathcal{O}(1).$$
(3.8)

3.2 AdS₅ black string entropy from near-horizon CFT₂

The near-horizon geometry of the rotating AdS_5 black string solution is [64]

$$ds^{2} = \frac{(-\mathcal{M}\Pi)^{2/3}}{\Theta^{2}} \left[-r^{2} d\tau^{2} + \frac{dr^{2}}{r^{2}} + \frac{\mathcal{W}\Theta^{2}}{\mathcal{M}^{2}} \left(dy - \frac{\mathcal{M}}{\Theta\sqrt{\mathcal{W}}} r d\tau \right)^{2} \right] + \frac{(-\mathcal{M})^{2/3}}{\Pi^{1/3}} \left[d\theta^{2} + \sin^{2}\theta \left(d\varphi + \frac{\mathcal{J}}{\mathcal{M}} dy \right)^{2} \right], \qquad (3.9)$$

where $\mathcal{M} \equiv -p_1 p_2 p_3$ is the product of magnetic charges, and \mathcal{J} is the angular momentum, while

$$\Theta \equiv (p^{1})^{2} + (p^{2})^{2} + (p^{3})^{2} - 2(p^{1}p^{2} + p^{1}p^{3} + p^{2}p^{3}),$$

$$\Pi \equiv (-p^{1} + p^{2} + p^{3})(p^{1} - p^{2} + p^{3})(p^{1} + p^{2} - p^{3}),$$

$$\mathcal{W} \equiv \frac{-4q_{0}p^{1}p^{2}p^{3} - \mathcal{J}^{2}}{\Theta},$$
(3.10)

with q_0 denoting the momentum added along the black string direction. Using the standard Kerr/CFT correspondence, we obtain the central charge of the near-horizon CFT₂

$$c_L = \frac{6\mathcal{M}}{G_4\Theta} \,. \tag{3.11}$$

This central charge was found in [64] as a Brown-Henneaux central charge [72].

To compute the black string entropy using the Cardy formula, we still need the Frolov-Thorne temperature, which can be computed from the standard formalism for the Kerr/CFT correspondence [73]

$$T_L = \frac{\sqrt{\mathcal{W}\Theta}}{2\pi\mathcal{M}} \,. \tag{3.12}$$

Therefore, the Cardy formula leads to the rotating AdS_5 black string (BS) entropy

$$S_{\rm BS} = \frac{\pi^2}{3} c_L T_L = \frac{\pi \sqrt{\mathcal{W}}}{G_4},$$
 (3.13)

which is the same as the leading-order rotating AdS_5 black string entropy [64, 69].

For the logarithmic correction to the AdS_5 black string entropy from the near-horizon CFT_2 , we apply the same technique as the AdS_5 black hole case. As mentioned in [64, 69], the rotating AdS_5 black string solution has one angular momentum and three electric charges. Similar to the BPS AdS_5 black hole case,

$$c \sim N^2$$
, $k^{11} \sim N^2$, $k^{ii} \sim N^{-2}$ $(i = 2, 3, 4)$, (3.14)

where we take n = 4 in the general formula (2.86) due to the following reason. Three U(1) electric charges have three corresponding chemical potentials Δ_a subject to a constraint,

hence there are only two indepedent U(1) electric charges. The angular momentum \mathcal{J} appearing in the second line of (3.9) can be viewed as an additional U(1), which can be treated in the same way as a U(1) electric charge [74], while in the first line of (3.9) there is actually another angular momentum hidden in the BTZ part of the metric. Hence, from the near-horizon region of the rotating AdS₅ black string there are still one angular momentum and three U(1) charges (including \mathcal{J}), which are independent of each other. Unlike the AdS₅ black hole case where we can choose one of the two angular momenta, for the rotating AdS₅ black string the way of counting near-horizon symmetries is unique.

The reasoning of section 2.5 can be followed in its entirety except that from the start there are only 4 chemical potentials, one conjugate to angular momentum and three conjugate to electric charges, obeying one constraint. This is in contrast with the AdS₅ black hole with 5 chemical potentials, two conjugate to angular momenta and three conjugate to electric charges. The scalings of the Kac-Moody levels and other parameters are the same. Moreover, only one Lagrange multiplier is needed, λ_1 , and we find that det $H \sim N^2$, as in the case of the AdS₅ black hole with the same final result as in (2.90). Consequently,

$$(\det \mathcal{A})_{AdS_5 Black String} \sim (N^2)^{-\frac{n+1}{2}} (N^2)^{\frac{n+3}{2}} (N^2) = N^4,$$
 (3.15)

and the logarithmic correction to the leading-order AdS_5 black string entropy (3.13) is

$$\Delta S_{\rm CFT_2} = -\frac{1}{2} \log \frac{\det \mathcal{A}}{(2\pi)^{n+2}} = -2 \log N + \mathcal{O}(1) \,. \tag{3.16}$$

4 Discussion

In this paper we have explored logarithmic corrections to asymptotically AdS_5 supersymmetric extremal, rotating, electrically charged black holes and black strings. For each case we examined the microstate counting in the context of $\mathcal{N} = 4$ SYM whereby it reduces to a combinatorial contribution from the space of solutions. We also approached the logarithmic corrections to the entropy by considering the microstate counting in the near-horizon geometry and its dual CFT_2 , where the logarithmic corrections arise as subleading contributions in the Cardy formula for the degeneracy of states. We found that the results from both approaches precisely match for both AdS_5 black holes and rotating black strings. It is instructive to write our boundary CFT₄ result as $(1 - d) \log N$ to note that the logarithmic correction has two contributions, one that has a completely combinatorial origin and is rather universal, namely, $\log N$, while the other contribution from the change of ensemble depending on the number of independent chemical potentials of the theory, $-d \log N$. Since we have 3 independent chemical potentials, we obtain $-2 \log N$ as a correction to the microscopic entropy. The calculation using the Cardy formula cannot resolve such splitting of the logarithmic correction, and only after being in microcanonical ensemble, an agreement can be found. The reason is that the Kerr/CFT correspondence is a field-theoretic approach intrinsic to the near-horizon geometry. It is, therefore, insensitive to certain details of the UV region. Our agreement in using the Cardy formula to its logarithmic precision should come more as a surprise than as a foregone conclusion. There is precedent where the Cardy formula leads to the wrong answer for logarithmic corrections [74]. Although the subtleties in applying the Cardy formula beyond its intrinsic regime are numerous, we expect that our positive results indicate the existence of resolutions which take into account particular properties of the spectrum [75, 76].

It would be interesting to derive the logarithmic corrections directly from the macroscopic one-loop contribution in type IIB supergravity. It is also natural to extend our near-horizon analysis to asymptotically AdS black holes in other dimensions. This route is certain to encounter obstructions in the form of zero modes, as is the case for asymptotically AdS₄ and AdS₆ black holes. Indeed, it has been shown that the one-loop supergravity contribution to the logarithmic corrections for asymptotically AdS₄ black holes [18] is different from the one obtained in the near-horizon approach [77, 78]. Our work indicates that given the absence of obstructions (zero modes) in odd-dimensional AdS spacetimes the counting can be performed at the near-horizon level, paving the way for a quantum entropy formula à la Sen [79]. It will also be interesting to explore the implications of our near-horizon results within supergravity localization along the lines of [80, 81].

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A Special functions

Here we summarize the definitions of special functions used in the paper. The Dedekind eta function is defined as

$$\eta(q) = q^{\frac{1}{24}} \prod_{k=1}^{\infty} \left(1 - q^k \right), \quad \text{Im}(\tau) > 0, \qquad (A.1)$$

with $q = e^{2\pi i \tau}$. The Pochhammer symbol is defined as

$$(z;q)_{\infty} = \prod_{k=0}^{\infty} (1 - zq^k).$$
 (A.2)

The elliptic theta functions which are relevant to us have the following product form:

$$\theta_0(u;\tau) = \prod_{k=0}^{\infty} (1 - e^{2\pi i (u+k\tau)}) (1 - e^{2\pi i (-u+(k+1)\tau)}), \qquad (A.3a)$$

$$\theta_1(u;\tau) = -ie^{\frac{\pi i\tau}{4}} (e^{\pi i\tau} - e^{-\pi i\tau}) \prod_{k=1}^{\infty} (1 - e^{2\pi i k\tau}) (1 - e^{2\pi i (k\tau+u)}) (1 - e^{2\pi i (k\tau-u)})$$
$$= ie^{\frac{\pi i\tau}{4}} e^{-\pi i u} \theta_0(u;\tau) \prod_{k=1}^{\infty} (1 - e^{2\pi i k\tau}).$$
(A.3b)

The elliptic gamma function and the "tilde" elliptic gamma function are defined as

$$\Gamma(z; p, q) = \prod_{j,k=0}^{\infty} \frac{1 - p^{j+1} q^{k+1} z^{-1}}{1 - p^j q^k z}, \qquad (A.4a)$$

$$\widetilde{\Gamma}(u;\sigma,\tau) = \prod_{j,k=0}^{\infty} \frac{1 - e^{2\pi i [(j+1)\sigma + (k+1)\tau - u]}}{1 - e^{2\pi i [j\sigma + k\tau + u]}} \,. \tag{A.4b}$$

A.1 Asymptotic behavior

For a small $|\tau|$ with fixed $0 < \arg \tau < \pi$, the Pochhammer symbol can be approximated as

$$\log(q;q)_{\infty} = -\frac{\pi i}{12} \left(\tau + \frac{1}{\tau}\right) - \frac{1}{2} \log(-i\tau) + \mathcal{O}\left(e^{\frac{2\pi \sin(\arg\tau)}{|\tau|}}\right). \tag{A.5}$$

To study asymptotic behaviors of elliptic functions, it is useful to introduce the function $\{u\}_{\tau}$, as

$$\{u\}_{\tau} \equiv u - \lfloor \operatorname{Re}(u) - \cot(\arg\tau)\operatorname{Im}(u) \rfloor \quad (u \in \mathbb{C}), \qquad (A.6)$$

which satisfies

$$\{u\}_{\tau} = \{\tilde{u}\}_{\tau} + \check{u}\tau, \qquad \{-u\}_{\tau} = \begin{cases} 1 - \{u\}_{\tau} & (\tilde{u} \notin \mathbb{Z}), \\ -\{u\}_{\tau} & (\tilde{u} \in \mathbb{Z}), \end{cases}$$
(A.7)

where we have defined $\tilde{u}, \check{u} \in \mathbb{R}$ as

$$u = \tilde{u} + \check{u}\tau \,. \tag{A.8}$$

The elliptic theta function $\theta_0(u; \tau)$ can be approximated for a small $|\tau|$ with fixed $0 < \arg \tau < \pi$ as

$$\log \theta_0(u;\tau) = \frac{\pi i}{\tau} \{u\}_{\tau} (1 - \{u\}_{\tau}) + \pi i \{u\}_{\tau} - \frac{\pi i}{6\tau} (1 + 3\tau + \tau^2) + \log \left(1 - e^{-\frac{2\pi i}{\tau} (1 - \{u\}_{\tau})}\right) \left(1 - e^{-\frac{2\pi i}{\tau} \{u\}_{\tau}}\right) + \mathcal{O}\left(e^{\frac{2\pi \sin(\arg \tau)}{|\tau|}}\right).$$
(A.9)

The elliptic theta function $\theta_1(u;\tau)$ is approximated for a small $|\tau|$ with fixed $0 < \arg \tau < \pi$ as

$$\log \theta_{1}(u;\tau) = \frac{\pi i}{\tau} \{u\}_{\tau} (1 - \{u\}_{\tau}) - \frac{\pi i}{4\tau} (1 + \tau) + \pi i \lfloor \operatorname{Re}(u) - \cot(\arg \tau) \operatorname{Im}(u) \rfloor + \frac{1}{2} \log \tau + \log \left(1 - e^{-\frac{2\pi i}{\tau} (1 - \{u\}_{\tau})}\right) \left(1 - e^{-\frac{2\pi i}{\tau} \{u\}_{\tau}}\right) + \mathcal{O}\left(e^{\frac{2\pi \sin(\arg \tau)}{|\tau|}}\right).$$
(A.10)

For a small $|\tau|$ with fixed $0 < \arg \tau < \pi$, the elliptic gamma function can be approximated as

$$\log \widetilde{\Gamma}(u;\tau) = 2\pi i \, Q(\{u\}_{\tau};\tau) + \mathcal{O}\left(|\tau|^{-1} e^{\frac{2\pi \sin(\arg\tau)}{|\tau|} \min(\{\widetilde{u}\},1-\{\widetilde{u}\})}\right),\tag{A.11}$$

provided $\tilde{u} \not\rightarrow \mathbb{Z}$ (see [56] for example), and the function $Q(\cdot; \cdot)$ is defined as:

$$Q(u;\tau) \equiv -\frac{B_3(u)}{6\tau^2} + \frac{B_2(u)}{2\tau} - \frac{5}{12}B_1(u) + \frac{\tau}{12}, \qquad (A.12)$$

with $B_n(u)$ being the *n*-th Bernoulli polynomial.

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References

- [1] J.D. Bekenstein, Black holes and entropy, Phys. Rev. D 7 (1973) 2333 [INSPIRE].
- [2] S.W. Hawking, Particle Creation by Black Holes, Commun. Math. Phys. 43 (1975) 199 [Erratum ibid. 46 (1976) 206] [INSPIRE].
- [3] J.M. Bardeen, B. Carter and S.W. Hawking, The Four laws of black hole mechanics, Commun. Math. Phys. 31 (1973) 161 [INSPIRE].
- [4] J.D. Bekenstein, Generalized second law of thermodynamics in black hole physics, Phys. Rev. D 9 (1974) 3292 [INSPIRE].
- [5] S.W. Hawking, Black Holes and Thermodynamics, Phys. Rev. D 13 (1976) 191 [INSPIRE].
- [6] A. Strominger and C. Vafa, Microscopic origin of the Bekenstein-Hawking entropy, Phys. Lett. B 379 (1996) 99 [hep-th/9601029] [INSPIRE].
- [7] A. Sen, Microscopic and Macroscopic Entropy of Extremal Black Holes in String Theory, Gen. Rel. Grav. 46 (2014) 1711 [arXiv:1402.0109] [INSPIRE].
- [8] J.M. Maldacena, The Large N limit of superconformal field theories and supergravity, Int. J. Theor. Phys. 38 (1999) 1113 [hep-th/9711200] [INSPIRE].
- F. Benini, K. Hristov and A. Zaffaroni, Black hole microstates in AdS₄ from supersymmetric localization, JHEP 05 (2016) 054 [arXiv:1511.04085] [INSPIRE].
- [10] A. Cabo-Bizet, D. Cassani, D. Martelli and S. Murthy, Microscopic origin of the Bekenstein-Hawking entropy of supersymmetric AdS₅ black holes, JHEP 10 (2019) 062
 [arXiv:1810.11442] [INSPIRE].
- [11] S. Choi, J. Kim, S. Kim and J. Nahmgoong, Large AdS black holes from QFT, arXiv:1810.12067 [INSPIRE].
- [12] F. Benini and P. Milan, Black Holes in 4D $\mathcal{N} = 4$ Super-Yang-Mills Field Theory, Phys. Rev. X 10 (2020) 021037 [arXiv:1812.09613] [INSPIRE].
- [13] S. Choi and S. Kim, Large AdS₆ black holes from CFT_5 , arXiv:1904.01164 [INSPIRE].
- [14] G. Kántor, C. Papageorgakis and P. Richmond, AdS_7 black-hole entropy and $5D \mathcal{N} = 2$ Yang-Mills, JHEP **01** (2020) 017 [arXiv:1907.02923] [INSPIRE].

- [15] J. Nahmgoong, 6d superconformal Cardy formulas, JHEP 02 (2021) 092 [arXiv:1907.12582]
 [INSPIRE].
- [16] S. Choi, C. Hwang and S. Kim, Quantum vortices, M2-branes and black holes, arXiv:1908.02470 [INSPIRE].
- [17] J. Nian and L.A. Pando Zayas, Microscopic entropy of rotating electrically charged AdS₄ black holes from field theory localization, JHEP 03 (2020) 081 [arXiv:1909.07943] [INSPIRE].
- [18] J.T. Liu, L.A. Pando Zayas, V. Rathee and W. Zhao, One-Loop Test of Quantum Black Holes in anti-de Sitter Space, Phys. Rev. Lett. 120 (2018) 221602 [arXiv:1711.01076] [INSPIRE].
- [19] D. Gang, N. Kim and L.A. Pando Zayas, Precision Microstate Counting for the Entropy of Wrapped M5-branes, JHEP 03 (2020) 164 [arXiv:1905.01559] [INSPIRE].
- [20] F. Benini, D. Gang and L.A. Pando Zayas, Rotating Black Hole Entropy from M5 Branes, JHEP 03 (2020) 057 [arXiv:1909.11612] [INSPIRE].
- [21] L.A. Pando Zayas and Y. Xin, Universal logarithmic behavior in microstate counting and the dual one-loop entropy of AdS₄ black holes, Phys. Rev. D 103 (2021) 026003
 [arXiv:2008.03239] [INSPIRE].
- [22] A. Strominger, Black hole entropy from near horizon microstates, JHEP 02 (1998) 009 [hep-th/9712251] [INSPIRE].
- [23] M. Guica, T. Hartman, W. Song and A. Strominger, The Kerr/CFT Correspondence, Phys. Rev. D 80 (2009) 124008 [arXiv:0809.4266] [INSPIRE].
- [24] H. Lü, J. Mei and C.N. Pope, Kerr/CFT Correspondence in Diverse Dimensions, JHEP 04 (2009) 054 [arXiv:0811.2225] [INSPIRE].
- [25] D.D.K. Chow, M. Cvetič, H. Lü and C.N. Pope, Extremal Black Hole/CFT Correspondence in (Gauged) Supergravities, Phys. Rev. D 79 (2009) 084018 [arXiv:0812.2918] [INSPIRE].
- [26] M. David, J. Nian and L.A. Pando Zayas, Gravitational Cardy Limit and AdS Black Hole Entropy, JHEP 11 (2020) 041 [arXiv:2005.10251] [INSPIRE].
- [27] J. Nian and L.A. Pando Zayas, Toward an Effective CFT_2 from $\mathcal{N} = 4$ Super Yang-Mills and Aspects of Hawking Radiation, JHEP **07** (2020) 120 [arXiv:2003.02770] [INSPIRE].
- [28] M. David and J. Nian, Universal entropy and Hawking radiation of near-extremal AdS₄ black holes, JHEP 04 (2021) 256 [arXiv:2009.12370] [INSPIRE].
- [29] S. Murthy, The growth of the $\frac{1}{16}$ -BPS index in 4d $\mathcal{N} = 4$ SYM, arXiv:2005.10843 [INSPIRE].
- [30] A. Pathak, A.P. Porfyriadis, A. Strominger and O. Varela, Logarithmic corrections to black hole entropy from Kerr/CFT, JHEP 04 (2017) 090 [arXiv:1612.04833] [INSPIRE].
- [31] A. Sen, Logarithmic Corrections to Schwarzschild and Other Non-extremal Black Hole Entropy in Different Dimensions, JHEP 04 (2013) 156 [arXiv:1205.0971] [INSPIRE].
- [32] S. Carlip, Logarithmic corrections to black hole entropy from the Cardy formula, Class. Quant. Grav. 17 (2000) 4175 [gr-qc/0005017] [INSPIRE].
- [33] S. Banerjee, R.K. Gupta and A. Sen, Logarithmic Corrections to Extremal Black Hole Entropy from Quantum Entropy Function, JHEP **03** (2011) 147 [arXiv:1005.3044] [INSPIRE].
- [34] S. Banerjee, R.K. Gupta, I. Mandal and A. Sen, Logarithmic Corrections to N = 4 and N = 8 Black Hole Entropy: A One Loop Test of Quantum Gravity, JHEP 11 (2011) 143
 [arXiv:1106.0080] [INSPIRE].

- [35] A. Sen, Logarithmic Corrections to N = 2 Black Hole Entropy: An Infrared Window into the Microstates, Gen. Rel. Grav. 44 (2012) 1207 [arXiv:1108.3842] [INSPIRE].
- [36] C. Romelsberger, Counting chiral primaries in N = 1, d = 4 superconformal field theories, Nucl. Phys. B 747 (2006) 329 [hep-th/0510060] [INSPIRE].
- [37] J. Kinney, J.M. Maldacena, S. Minwalla and S. Raju, An Index for 4 dimensional super conformal theories, Commun. Math. Phys. 275 (2007) 209 [hep-th/0510251] [INSPIRE].
- [38] A. González Lezcano and L.A. Pando Zayas, Microstate counting via Bethe Ansätze in the 4d $\mathcal{N} = 1$ superconformal index, JHEP **03** (2020) 088 [arXiv:1907.12841] [INSPIRE].
- [39] A. Lanir, A. Nedelin and O. Sela, *Black hole entropy function for toric theories via Bethe* Ansatz, JHEP 04 (2020) 091 [arXiv:1908.01737] [INSPIRE].
- [40] A. Arabi Ardehali, J.T. Liu and P. Szepietowski, c a from the $\mathcal{N} = 1$ superconformal index, JHEP 12 (2014) 145 [arXiv:1407.6024] [INSPIRE].
- [41] A. Amariti, I. Garozzo and G. Lo Monaco, Entropy function from toric geometry, Nucl. Phys. B 973 (2021) 115571 [arXiv:1904.10009] [INSPIRE].
- [42] A. Arabi Ardehali, High-temperature asymptotics of supersymmetric partition functions, JHEP
 07 (2016) 025 [arXiv:1512.03376] [INSPIRE].
- [43] L. Di Pietro and M. Honda, Cardy Formula for 4d SUSY Theories and Localization, JHEP 04 (2017) 055 [arXiv:1611.00380] [INSPIRE].
- [44] F. Benini, E. Colombo, S. Soltani, A. Zaffaroni and Z. Zhang, Superconformal indices at large N and the entropy of AdS₅ × SE₅ black holes, Class. Quant. Grav. **37** (2020) 215021
 [arXiv:2005.12308] [INSPIRE].
- [45] A. González Lezcano, J. Hong, J.T. Liu and L.A. Pando Zayas, Sub-leading Structures in Superconformal Indices: Subdominant Saddles and Logarithmic Contributions, JHEP 01 (2021) 001 [arXiv:2007.12604] [INSPIRE].
- [46] A. Gonzalez Lezcano, Precision microstate counting of black hole entropy from N = 1 toric quiver gauge theories, Ph.D. Thesis, SISSA, Trieste, Italy (2021) [INSPIRE].
- [47] C. Romelsberger, Calculating the Superconformal Index and Seiberg Duality, arXiv:0707.3702 [INSPIRE].
- [48] F.A. Dolan and H. Osborn, Applications of the Superconformal Index for Protected Operators and q-Hypergeometric Identities to N = 1 Dual Theories, Nucl. Phys. B 818 (2009) 137 [arXiv:0801.4947] [INSPIRE].
- [49] A.G. Lezcano, J. Hong, J.T. Liu and L.A. Pando Zayas, The Bethe-Ansatz approach to the $\mathcal{N} = 4$ superconformal index at finite rank, JHEP **06** (2021) 126 [arXiv:2101.12233] [INSPIRE].
- [50] F. Benini and G. Rizi, Superconformal index of low-rank gauge theories via the Bethe Ansatz, JHEP 05 (2021) 061 [arXiv:2102.03638] [INSPIRE].
- [51] J. Hong, The topologically twisted index of $\mathcal{N} = 4$ SU(N) Super-Yang-Mills theory and a black hole Farey tail, JHEP 10 (2021) 145 [arXiv:2108.02355] [INSPIRE].
- [52] J. Hong and J.T. Liu, The topologically twisted index of $\mathcal{N} = 4$ super-Yang-Mills on $T^2 \times S^2$ and the elliptic genus, JHEP **07** (2018) 018 [arXiv:1804.04592] [INSPIRE].
- [53] A. Arabi Ardehali and S. Murthy, The 4d superconformal index near roots of unity and 3d Chern-Simons theory, JHEP 10 (2021) 207 [arXiv:2104.02051] [INSPIRE].

- [54] A. Cabo-Bizet, D. Cassani, D. Martelli and S. Murthy, The asymptotic growth of states of the $4d \mathcal{N} = 1$ superconformal index, JHEP 08 (2019) 120 [arXiv:1904.05865] [INSPIRE].
- [55] M. Honda, Quantum Black Hole Entropy from 4d Supersymmetric Cardy formula, Phys. Rev. D 100 (2019) 026008 [arXiv:1901.08091] [INSPIRE].
- [56] A. Arabi Ardehali, Cardy-like asymptotics of the 4d $\mathcal{N} = 4$ index and AdS₅ blackholes, JHEP 06 (2019) 134 [arXiv:1902.06619] [INSPIRE].
- [57] J. Kim, S. Kim and J. Song, A 4d N = 1 Cardy Formula, JHEP 01 (2021) 025
 [arXiv:1904.03455] [INSPIRE].
- [58] D. Cassani and Z. Komargodski, EFT and the SUSY Index on the 2nd Sheet, SciPost Phys. 11 (2021) 004 [arXiv:2104.01464] [INSPIRE].
- [59] O. Aharony, F. Benini, O. Mamroud and P. Milan, A gravity interpretation for the Bethe Ansatz expansion of the N = 4 SYM index, Phys. Rev. D 104 (2021) 086026
 [arXiv:2104.13932] [INSPIRE].
- [60] A. Amariti, M. Fazzi and A. Segati, The SCI of $\mathcal{N} = 4 USp(2N_c)$ and $SO(N_c)$ SYM as a matrix integral, JHEP 06 (2021) 132 [arXiv:2012.15208] [INSPIRE].
- [61] Z.W. Chong, M. Cvetič, H. Lü and C.N. Pope, General non-extremal rotating black holes in minimal five-dimensional gauged supergravity, Phys. Rev. Lett. 95 (2005) 161301
 [hep-th/0506029] [INSPIRE].
- [62] J.M. Maldacena and A. Strominger, Universal low-energy dynamics for rotating black holes, Phys. Rev. D 56 (1997) 4975 [hep-th/9702015] [INSPIRE].
- [63] G. Compere, K. Murata and T. Nishioka, Central Charges in Extreme Black Hole/CFT Correspondence, JHEP 05 (2009) 077 [arXiv:0902.1001] [INSPIRE].
- [64] S.M. Hosseini, K. Hristov, Y. Tachikawa and A. Zaffaroni, Anomalies, Black strings and the charged Cardy formula, JHEP 09 (2020) 167 [arXiv:2006.08629] [INSPIRE].
- [65] A. Castro, A. Maloney and A. Strominger, Hidden Conformal Symmetry of the Kerr Black Hole, Phys. Rev. D 82 (2010) 024008 [arXiv:1004.0996] [INSPIRE].
- [66] S. Haco, S.W. Hawking, M.J. Perry and A. Strominger, Black Hole Entropy and Soft Hair, JHEP 12 (2018) 098 [arXiv:1810.01847] [INSPIRE].
- [67] S. Haco, M.J. Perry and A. Strominger, *Kerr-Newman Black Hole Entropy and Soft Hair*, arXiv:1902.02247 [INSPIRE].
- [68] M. Perry and M.J. Rodriguez, Central charges for AdS black holes, Class. Quant. Grav. 39 (2022) 045009 [arXiv:2007.03709] [INSPIRE].
- [69] S.M. Hosseini, K. Hristov and A. Zaffaroni, Microstates of rotating AdS₅ strings, JHEP 11 (2019) 090 [arXiv:1909.08000] [INSPIRE].
- [70] F. Benini and A. Zaffaroni, A topologically twisted index for three-dimensional supersymmetric theories, JHEP 07 (2015) 127 [arXiv:1504.03698] [INSPIRE].
- [71] S.M. Hosseini, A. Nedelin and A. Zaffaroni, The Cardy limit of the topologically twisted index and black strings in AdS₅, JHEP 04 (2017) 014 [arXiv:1611.09374] [INSPIRE].
- [72] J.D. Brown and M. Henneaux, Central Charges in the Canonical Realization of Asymptotic Symmetries: An Example from Three-Dimensional Gravity, Commun. Math. Phys. 104 (1986) 207 [INSPIRE].

- [73] G. Compère, The Kerr/CFT correspondence and its extensions, Living Rev. Rel. 15 (2012) 11 [arXiv:1203.3561] [INSPIRE].
- [74] A. Sen, Logarithmic Corrections to Rotating Extremal Black Hole Entropy in Four and Five Dimensions, Gen. Rel. Grav. 44 (2012) 1947 [arXiv:1109.3706] [INSPIRE].
- [75] T. Hartman, C.A. Keller and B. Stoica, Universal Spectrum of 2d Conformal Field Theory in the Large c Limit, JHEP 09 (2014) 118 [arXiv:1405.5137] [INSPIRE].
- [76] B. Mukhametzhanov and A. Zhiboedov, Modular invariance, tauberian theorems and microcanonical entropy, JHEP 10 (2019) 261 [arXiv:1904.06359] [INSPIRE].
- [77] J.T. Liu, L.A. Pando Zayas, V. Rathee and W. Zhao, Toward Microstate Counting Beyond Large N in Localization and the Dual One-loop Quantum Supergravity, JHEP 01 (2018) 026 [arXiv:1707.04197] [INSPIRE].
- [78] I. Jeon and S. Lal, Logarithmic Corrections to Entropy of Magnetically Charged AdS4 Black Holes, Phys. Lett. B 774 (2017) 41 [arXiv:1707.04208] [INSPIRE].
- [79] A. Sen, Quantum Entropy Function from AdS₂/CFT₁ Correspondence, Int. J. Mod. Phys. A 24 (2009) 4225 [arXiv:0809.3304] [INSPIRE].
- [80] A. Dabholkar, N. Drukker and J. Gomes, Localization in supergravity and quantum AdS₄/CFT₃ holography, JHEP 10 (2014) 090 [arXiv:1406.0505] [INSPIRE].
- [81] J. Nian and X. Zhang, Entanglement Entropy of ABJM Theory and Entropy of Topological Black Hole, JHEP 07 (2017) 096 [arXiv:1705.01896] [INSPIRE].