## CALCULATIONS ON THE HELIUM ISOELECTRONIC SEQUENCE USING COULOMB GREEN'S FUNCTION VARIABLES

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Calculations on the ground states of the helium isoelectronic series are carried out using variational wavefunctions of the form  $\psi(x,y)$ , in which x and y are the combinations  $r_1 + r_2 \pm r_{12}$  occurring in the Coulomb Green's function. The results for helium are the most accurate to date using a two-variable wavefunction, accounting for 71.5% of the correlation energy.

The Coulomb Green's function is the solution under specified boundary conditions of the differential equations [1]:

$$(\frac{1}{2}k^{2} + \frac{1}{2}\nabla_{i}^{2} + Z/r_{i})G(r_{1}, r_{2}, k) = \delta(r_{12}), \quad i = 1, 2.$$
(1)

We use atomic units, whereby  $\hbar = e = \mu = 1$ . The sum of eqs. (1) over the two particle labels gives

$$(k^{2} + \frac{1}{2}\nabla_{1}^{2} + \frac{1}{2}\nabla_{2}^{2} + Z/r_{1} + Z/r_{2})G(r_{1}, r_{2}, k) = 2\delta(r_{12}),$$
(2)

which closely resembles the Schrödinger equation for a helium-like atom. As shown by Hostler [2], the Coulomb Green's function depends on just two variables:

$$x \equiv r_1 + r_2 + r_{12}, \quad y \equiv r_1 + r_2 - r_{12}, \tag{3}$$

rather than three  $- \sup r_1, r_2, r_{12}$  — which would be expected on the basis of rotational invariance alone. This reduction is a consequence of the SO(4) dynamical symmetry of the Coulomb problem. This higher symmetry is also connected with the existence of an additional constant of the motion, the Runge-Lenz vector.

Specifically the retarded Coulomb Green's function  $G^+(x, y, k)$ , is given by [3]:

$$G^{+}(x,y,k) = -(\pi i k)^{-1} (x - y)^{-1} (\partial/\partial x - \partial/\partial y)$$
  
×  $\Gamma(1 - i v) M_{iv}^{1/2}(-i k y) W_{iv}^{1/2}(-i k x),$   
 $v \equiv Z/k, \quad \text{Im} k > 0,$  (4)

in which M and W are Whittaker functions as defined by Buchholz [4].

In this note, we shall explore variational wavefunctions of the form  $\psi(x,y)$  to approximate <sup>1</sup>S states of the helium isoelectronic series. Rigorously, of course, three variables such as  $r_1, r_2, r_{12}$  are required to span the functional domain for S states. It is of interest, nevertheless, to see how far one can go with just x and y dependence.

The classical work of Hylleraas [5] was based on the independent variables:

$$s \equiv r_1 + r_2, \quad t \equiv r_1 - r_2, \quad u \equiv r_{12}.$$
 (5)

Since x = s + u and y = s - u, our computations will be equivalent to those of Hylleraas with the variable *t* absent.

For helium-like systems, the hamiltonian is given by

$$H = -\frac{1}{2}\nabla_1^2 - \frac{1}{2}\nabla_2^2 - Z/r_1 - Z/r_2 + 1/r_{12}.$$
 (6)

For  $\psi = \psi(x, y)$ , we write the variational integral

$$E = (\psi, H\psi)/(\psi, \psi) = (T + V)/N,$$
(7)

with

Volume 104, number 6

$$N = \frac{1}{48} \int_{0}^{\infty} \int_{0}^{x} (x^{2} + 4xy + y^{2})(x - y)^{2} |\psi(x, y)|^{2} dx dy,$$
(8)
$$T = \frac{1}{12} \int_{0}^{\infty} \int_{0}^{x} [(x^{2} + 2xy)(\partial \psi / \partial x)^{2} + (y^{2} + 2xy)(\partial \psi / \partial y)^{2}](x - y)^{2} dx dy,$$
(9)

$$V = -\frac{1}{4}Z \int_{0}^{\infty} \int_{0}^{x} (x+y)(x-y)^{2} |\psi(x,y)|^{2} dx dy$$
  
+  $\frac{1}{24} \int_{0}^{\infty} \int_{0}^{x} (x^{2} + 4xy + y^{2})(x-y) |\psi(x,y)|^{2} dx dy.$   
(10)

Specifically, we consider linear variational functions of the form

$$\psi(x,y) = e^{-\alpha(x+y)} \sum_{m,n} C_{m,n} x^m y^n.$$
 (11)

The integrals (8)-(10) are then given explicitly by

$$N = \frac{1}{48} \sum_{m,n,k,l} \{ [4,0] + 2[3,1] - 6[2,2] + 2[1,3] + [0,4] \},$$
(12)

$$T = \frac{1}{12} \alpha^2 N - \frac{1}{12} \sum_{m,n,k,l} C_{m,n} C_{k,l}$$

$$\times \{2\alpha [m([3,0] - 3[1,2] + 2[0,3]) + n(2[3,0] - 3[2,1] + [0,3])] - mk([2,0] - 3[0,2] + 2[-1,3]) - nl(2[3,-1] - 3[2,0] + [0,2])\}, \qquad (13)$$

$$V = \frac{1}{48} \sum_{m,n,k,l} C_{m,n} C_{k,l} \{(-12Z + 2)[3,0] + (12Z + 6)[2,1] + (12Z - 6)[1,2] \}^{\frac{1}{2}} + (-12Z - 2)[0,3] \}.$$
 (14)

We have introduced the notation

$$[I,J] \equiv (m+k+I, n+l+J),$$
(15)

in terms of the elementary integrals

Table 1	
Helium	calculations

Helium calculations with N-term basis

N	α	-E
1	0.84375	2.84765625
3	0.905	2.8912315
6	0.935	2.8916945
10	0.99	2.8917460
15	1.03	2.891750194

$$(m,n) \equiv \int_{0}^{\infty} \int_{0}^{x} e^{-2\alpha(x+y)} x^{m} y^{n} dx dy$$
$$= \sum_{r=0}^{n} \frac{(-1)^{r} n! (m-n-r)!}{(n-r)! (2\alpha)^{r+1} (4\alpha)^{m+n-r+1}} + \frac{n! m!}{(2\alpha)^{n+m+2}}.$$
(16)

The functional form (11) trivially includes the simple exponential  $\exp[-\alpha(x+y)]$ , equivalent to the scaled hydrogenic product  $\exp[-2\alpha(r_1+r_2)]$ . This is optimized for  $\alpha = 27/32$ , giving the approximate helium ground-state energy E = -2.84765625 hartree. With the inclusion of sets of linear terms containing higher powers of x and y, we obtain the results summarized in table 1. The exponential parameter has been reoptimized for each basis set. The three-term function (1, x, y) gives an optimal energy of -2.8912315H. The six-term function (adding  $x^2, xy, y^2$ ) gives -2.8916504 H. The ten-term function (adding  $x^3$ ,

Table 2 Coefficients for helium fifteen-term function,  $\alpha = 1.03$ 

C(0, 0)	1.00000000	
C(1, 0)	0.265456196	
C(0, 1)	-0.175206467	
C(2,0)	0.017240841	
C(1, 1)	0.080962524	
C(0, 2)	-0.041635911	
C(3,0)	0.002107589	
C(2,1)	-0.008768843	
C(1, 2)	0.001859248	
<i>C</i> (0, 3)	-0.000208398	
C(4, 0)	0.000080883	
C(3, 1)	0.001069403	
C(2, 2)	0.000208337	
C(1, 3)	-0.000588844	
<i>C</i> (0, 4)	0.000024063	

Table 3

	Z	α	- <i>E</i>	$-E(SCF)^{a}$	-E(exact) b)	% corr
	1	0.435	0.50975076	0.48793	0.52759152	55.0
	2	1.03	2.89175019	2.86168	2.90372433	71.5
	3	1.60	7.26902230	7.23641	7.27991339	75.0
	4	2.20	13.64514384	13.61130	13.65556622	76.5
	5	2.78	22.02081175	21.98623	22.03097156	77.3
	6	3.35	32.39625486	32.36119	32.40624658	77.8
	7	3.95	44.77157028	44.73616	44.78144513	78.2
•	8	4.58	59.14680624	59.11114	59.15659510	78.5
	9	5.05	75.52198941	75.48613	75.53171234	78.7 ·

Helium isoelectronic sequence ground-state energies using fifteen-term function

a) Ref. [7]. b) Ref. [8].

 $x^2y, xy^2, y^3$ ) gives -2.8917460 H. Finally, with fifteen terms (adding  $x^4, x^3y, x^2y^2, xy^3, y^4$ ), we obtain -2.891750194 H. The coefficients are shown in table 2. It is estimated that with additional terms the procedure would converge to an approximate energy of -2.89175065 H. Given the exact non-relativistic energy of -2.903724375 H [6] and the Hartree–Fock energy of -2.86168 H [7], it is seen that our calculation accounts for  $\approx 71.5\%$  of the correlation energy of helium.

In table 3, we present analogous calculations, using the fifteen-term function, on the  $1s^{2}$  <sup>1</sup>S ground states of the helium isoelectronic sequence from Z = 1 to Z = 9. The results are compared with the self-consistent-field calculations of Roothaan and Weiss [7] and the essentially exact 13th-order perturbation calculations of Scherr and Knight [8]. As shown in the last column of table 3, the fraction of correlation energy captured by  $\psi(x, y)$  increases to over 75% for twoelectron atoms beyond helium.

The calculations on helium-like systems described in this note are the most accurate to date based on wavefunctions dependent on just two variables.

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