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flow line, reaching a minimum of $0.7 \, \rm cm \, yr^{-1}$ at Vostok station. Along the eastern shoreline a significant increase in accretion rate to $2.9 \, \rm cm \, yr^{-1}$ is necessary to produce the grounded accreted ice. It is possible that some of the accreted ice thickening along the eastern shoreline is the result of compressive flow and not accretion. The lower layers in the ice sheet $(3,300-2,800 \, \rm m)$ undergo a 7% thickening, in contrast to the shallower layers of the ice sheet $(2,400-1,900 \, \rm m)$ which thin by 5-13% when grounding. The accretion process is most evident along the lake shorelines.

In the grounded ice adjacent to the southeastern Lake Vostok shoreline, we observe an average of 295 m of accreted ice in 21 radar profiles (Fig. 1, inset). No accretion ice is imaged adjacent to the eastern grounding line in the northern 15 radar profiles. We estimate the ice flux out of the lake, projected onto a 165-km line downslope of the lake, parallel to local contours. Assuming an average thickness of 295 m of accretion ice and a mean velocity of 3 m yr⁻¹, we estimate an annual accreted ice volume flux of 0.146 km³ yr⁻¹ along this line. For a lake volume¹⁴ of 1,800 km³, the residence time is 13,300 years. Previous estimates of residence time range from 4,500 years (ref. 6) to 125,000 years (ref. 1). The lower estimate (4,500 years), derived from the He⁴/He³ ratio⁶, may be related to the formation of the accreted ice samples along the western shoreline in regions isolated from the main lake circulation—that is, the shallow embayment or the western grounding line. Deeper samples of accreted ice should be more representative of the open lake.

In the southern portion of Lake Vostok, the interaction between the lake and the overlying ice sheet is dominated by accretion. Earlier evidence for melting in this region was based on radar data oblique to the ice flow field, and is, we believe, incorrect. Ice flow over southern Lake Vostok has a strong along-lake component, and has had a consistent orientation for the past 16,000 years—assuming that the velocity has been constant over this time period. The accretion process dominates at the lake shorelines, although some accretion continues over the middle of the lake. The accretion ice is generally transported out of the lake along the southeastern lake margin. Most samples of accretion ice analysed to date are derived either from the shallow embayment or from the western grounding line, environments that are not representative of the open lake system.

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Development of anisotropic structure in the Earth's lower mantle by solid-state convection

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Seismological observations reveal highly anisotropic patches at the bottom of the Earth's lower mantle, whereas the bulk of the mantle has been observed to be largely isotropic¹⁻⁴. These patches have been interpreted to correspond to areas where subduction has taken place in the past or to areas where mantle plumes are upwelling, but the underlying cause for the anisotropy is unknown—both shape-preferred orientation of elastically heterogenous materials⁵ and lattice-preferred orientation of a homogeneous material⁶⁻⁸ have been proposed. Both of these mechanisms imply that large-strain deformation occurs within the anisotropic regions, but the geodynamic implications of the mechanisms differ. Shape-preferred orientation would imply the presence of large elastic (and hence chemical) heterogeneity whereas lattice-preferred orientation requires deformation at high stresses. Here we show, on the basis of numerical modelling incorporating mineral physics of elasticity and development of lattice-preferred orientation, that slab deformation in the deep lower mantle can account for the presence of strong anisotropy in the circum-Pacific region. In this model—where development of the mineral fabric (the alignment of mineral grains) is caused solely by solid-state deformation of chemically homogeneous mantle material—anisotropy is caused by large-strain deformation at high stresses, due to the collision of subducted slabs with the core-mantle boundary.

Lattice-preferred orientation (LPO) leads to the development of a mineral fabric in material that deforms primarily by dislocation creep. Previous work⁹ has shown that slabs cause high-stress regions in the lower mantle, which leads to localized regions of dislocation creep within a lower mantle dominated by diffusion creep under a wide range of rheological parameters. It has also been shown that hot, upwelling regions are low stress, and are therefore dominated by diffusion creep. The presence of dislocation creep is not in itself a

sufficient condition for the formation of LPO in slabs. In addition, the strain due to the dislocation creep must be greater than 100–200%. Here we investigate the circumstances under which this condition is met by a combined mineral-physics and dynamical-modelling approach. This requires the use of a composite rheology formulation involving a combination of diffusion and dislocation creep deformation mechanisms.

We choose rheological parameters for the upper and lower mantle based on mineral-physics observations^{10,11}, resulting in a viscosity profile that generally increases with depth and includes an increase at the base of the transition zone. We also use a yield stress approach¹² in the upper regions of our model to form subducting slabs. We place strain tracers in slab regions above the transition zone to track the evolution of deformation. Given the uncertainties related to first-principles calculations of fabric development, such as the critical shear stresses associated with specific slip systems and the nature of dynamic recrystallization, we assume that strain may be used as a proxy for the development of mineral fabric. During dislocation creep, material flows by the slipping of specific glide planes, resulting in an oriented array of crystal axes and promoting fabric development. On the other hand diffusion creep occurs by the migration of atoms involving grain-boundary sliding, resulting in a random orientation of crystal axes and tending to destroy any preexisting fabric. Therefore, we track strain only in regions dominated by dislocation creep. We compare the resulting strain field to mineral-physics deformation experiments in order to assess the expected seismic anisotropy.

The numerical calculations of mantle convection are solved using the conservation equations of mass, momentum and energy in the extended Boussinesq formulation¹³ using a finite-element approach. The modelling geometry is a two-dimensional quarter-cylinder with free-slip boundaries. We employ depth-dependent thermal expansivity and thermal conductivity^{14,15}. We allow our convection calculations to run for long enough to ensure that the initial condition does not affect the results. Details of the model setup and rheological formulation are given in previous work⁹. Parameters used for the calculations presented here are given in Table 1.

Lagrangian finite strain is calculated as a post-processing step,

using the velocity fields at each time step of the convection calculation. Strain is calculated for particles within regions dominated by dislocation creep by time-integrating the deformation gradient tensor for each strain tracer¹⁶. The strain calculation code was checked by comparison to analytical solutions of simple and pure shear as well as a combination of the two¹⁷. If a strain tracer leaves the regime of dislocation creep, its recorded strain magnitude is decreased as a function of further material strain. We assume that LPO is destroyed after the material stretches further than twice its original length.

The results of two calculations are given here (Figs 1 and 2) to illustrate the insensitivity of our results to the amount of internal heating and to the scaling geometry. The calculations vary in terms of heating mode and radial scaling. Figure 1 shows results from a calculation in which the mantle is entirely bottom-heated and scaled to preserve the volume ratio of the Earth's upper and lower mantle. Figure 2 shows results from a calculation in which the mantle is heated by both internal and bottom heating (\sim 50% each), and is scaled to preserve the surface area to volume ratio of the mantle in a spherical Earth. These scalings have been found to better approximate the heat and mass transfer of spherical models¹⁸. Both results yield heat flow and radial viscosity profiles that are consistent with observations^{19,20}.

Snapshots in time of the temperature and viscosity ratio fields are shown in Figs 1c-e and 2c-e. The viscosity ratio is defined as the ratio of the components corresponding to dislocation creep and diffusion creep. Regions in blue represent a positive ratio, indicating that they are dominated by diffusion creep. Red regions represent a negative ratio, indicating a domination of dislocation creep. Note that most of the upper mantle flows by dislocation creep. In contrast the interior of the upper-mantle slab flows primarily by diffusion creep, because the upper-mantle activation coefficients result in a transition stress that increases with decreasing temperature. Away from the slab region, lower-mantle flow is dominated by diffusion creep, particularly in the lowermost boundary region, where high temperatures lead to low viscosities and therefore low stresses.

Strain tracers are shown as points on the viscosity ratio fields. Superimposed on tracer points are vectors representing the finite

Parameter	Description	Units	Value, Fig. 1	Value, Fig. 2
ΔT	Temperature drop across mantle	K	3,000	3,000
α_{\circ}	Reference thermal expansivity	K^{-1}	3×10^{-5}	3×10^{-5}
00	Reference density	kg m ⁻³	4,500	4,500
\mathcal{L}_{p}	Specific heat	J kg ⁻¹ K ⁻¹	1,250	1,250
า	Mantle thickness	m	2.8×10^{6}	2.8×10^{6}
ro .	Reference thermal conductivity	$W m^{-1} K^{-1}$	5.6	5.6
7	Gravitational constant	m s ⁻²	9.8	9.8
60	Reference thermal diffusivity	$m^2 s^{-1}$	10 ⁻⁶	10 ⁻⁶
Di	Dissipation number		0.5	0.5
d _{um}	Upper mantle grain size	mm	2.0	2.0
d _{lm}	Lower mantle grain size	mm	1.0	1.0
n	Grain size index		2.5	2.5
1	Power law index		3.0	3.0
4 _{diff-um+}	Upper-mantle diffusion creep prefactor	$d_{um}^{m} Pa^{-1} s^{-1*} \\ d_{um}^{m} Pa^{-1} s^{-1*} \\ Pa^{-n} s^{-1*}$	1.3276×10^{-13}	2.6551×10^{-1}
A _{diff-lm+}	Lower-mantle diffusion creep prefactor	$d_{um}^{m} Pa^{-1} s^{-1}*$	6.2230×10^{-14}	1.3276×10^{-1}
A _{disl-um+}	Upper-mantle dislocation creep prefactor	Pa ⁻ⁿ s ⁻¹ *	4.7295×10^{-13}	9.4590×10^{-1}
A _{disl-um+}	Lower-mantle dislocation creep prefactor	Pa ⁻ⁿ s ⁻¹ *	1.2446×10^{-29}	2.6551×10^{-1}
diff-um	Upper-mantle diffusion creep activation coefficient		17	17
diff-lm	Lower-mantle diffusion creep activation coefficient		10	10
disl-um	Upper-mantle dislocation creep activation coefficient		31	31
disl-Im	Lower-mantle dislocation creep activation coefficient		10	10
r _d	Ductile yield stress	MPa	400	250
r'b	Brittle yield stress gradient	MPa km ⁻¹	5.33	3.33
surface	Non-dimensional surface radius		1.67813	1.4292
Bottom	Non-dimensional bottom radius		0.67813	0.4292
R _{interface}	Non-dimensional upper - lower mantle boundary radius		1.42	1.19
η_{max}	Non-dimensional maximum viscosity		1.0	1.0

 $^{^{\}star}$ Indices m and n are the grain size index and the power law index, respective

Non-dimensional minimum viscosity

Table 1 Calculation parameters

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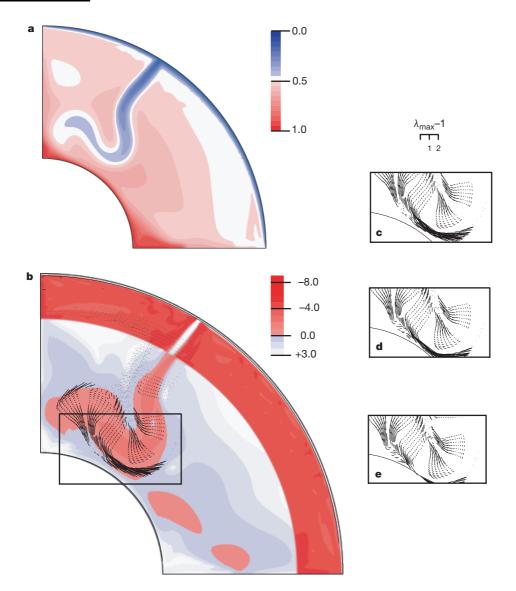


Figure 1 Snapshot of a slab impinging on the core-mantle boundary (CMB) in the bottomheated model. a, Non-dimensional temperature field and b, ratio of viscosity due to dislocation creep over that due to diffusion creep. Positive values indicate deformation is dominated by diffusion creep, and negative values indicate deformation is dominated by

dislocation creep. Superimposed are strain tracers and their associated strain vectors. Strain vectors represent maximum stretch, $\lambda_{\text{max}},$ and have a magnitude proportional to the stretch minus 1. The evolution of strain of the area of **b** shown in a box is given after successive intervals of 25 Myr in c, d, e.

strain. The vectors are in the direction of maximum principle strain, and have a length related to the stretch. The stretch is defined as the ratio of final length to the original length. Therefore, an undeformed particle has a stretch of unity. To differentiate better between strained and unstrained tracers the vector length is set to the stretch minus one, so only strained tracers have an associated vector.

To illustrate the temporal evolution of strain, Figs 1c-e and 2c-e show the strain configurations at successive times. Results from Fig. 1 reveal that although slab deformation in the lower mantle is dominated entirely by dislocation creep, the strain is not well developed because of low strain rates in the viscous slab. Directly above the core-mantle boundary (CMB), however, the magnitude of strain increases dramatically, resulting in a region of high strain that is directed laterally. Figure 1c-e reveals the time-dependent nature of the strain configuration. The overall strain magnitude in Fig. 1e is greatly reduced, and the directional sense is more random. Results are similar for the calculation shown in Fig. 2. Again, deformation in the slab is dominated by dislocation creep in a mantle otherwise dominated by diffusion creep. Figure 2b reveals a low magnitude lateral sense of strain directly above the CMB that increases with time (Fig. 2c-e) resulting in a strong laterally directed strain field. The strain markers also show that material is rotated as it approaches the upwellings that are constrained to the side boundaries of the model

Our results show that the details of the strain field are timedependent and can be quite complicated. Analysis of numerous convection results reveal that a lateral sense of high-magnitude extension (in excess of 100%) directly above the CMB is a general feature of the strain field. This feature tends to be long-lived but not permanent. As shown here, the strain pattern may decay into one that is characterized by a lower magnitude and inconsistent direction (that is, the progression from Fig. 1b-e). We also see that a coherent pattern of high-magnitude lateral strain tends to develop from a less developed state (that is, the progression

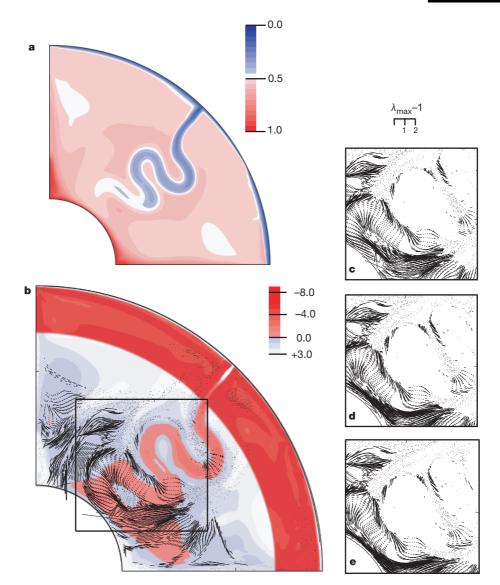


Figure 2 As Fig. 1, but for a model with bottom heating and internal heating.

from Fig. 2b-e). In general, we find that slab regions are characterized by a high degree of lateral strain directly above the CMB that occasionally reverts to a low-magnitude, more randomly oriented strain.

In order to assess the implications for seismic anisotropy, comparison to mineral-physics data is required. Experimental studies on LPO development show that the strength of LPO increases with strain, and reaches nearly steady-state values at a certain strain^{21–23}. This critical strain depends on the material as well as deformation conditions. A detailed experimental study of LPO development is now available for (Mg, Fe)O (ref. 23). Because of its large elastic anisotropy²⁴, LPO of (Mg Fe)O results in a detectable seismic anisotropy (\sim 1–2%)⁵, although its volume fraction is small (less than 20%). For (Mg, Fe)O, a shear strain of \sim 200% is needed to develop significant LPO and steady-state LPO is formed at strains of \sim 400–500% (ref. 23). At steady state, LPO of (Mg, Fe)O corresponding to horizontal shear results in $V_{\rm SH} > V_{\rm SV}$ anisotropy²³ where $V_{\rm SH}$ and $V_{\rm SV}$ correspond to horizontal and vertical components of the shear wave velocity, respectively.

Our numerical models show the development of strong subhorizontal shear strain near the base of the mantle. Therefore, LPO of (Mg, Fe)O provides a natural explanation for the spatial variation in anisotropy in the lower mantle. Contributions from another major component, (Mg, Fe)SiO₃ perovskite, are difficult to estimate because of the absence of experimental data on LPO. In the twophase mixture, (Mg, Fe)O is probably the weaker phase 10 and deforms more easily than (Mg, Fe)SiO₃. As a consequence, a higher degree of strain will occur in (Mg, Fe)O for a given deformation. In addition, if the results of LPO development in analogous material are used, the contribution from (Mg, Fe)SiO₃ perovskite is likely to be smaller than that of (Mg, Fe)O and the sense is opposite $(V_{\rm SV} > V_{\rm SH}$ for horizontal shear)^{6,23}. Once materials leave highstress regions, diffusion creep dominates and erases pre-existing LPO. Our models show highly time-dependent flow patterns, so the regions of strong anisotropy in these models correspond to regions with high-stress, large-strain deformation in the recent past. In contrast, if shape-preferred orientation due to laminated structures of palaeo-crust and ambient mantle is responsible for the anisotropy⁵, then anisotropic structures will have much longer lifetimes and would not represent recent dynamic regimes. Also, the origin of the high contrast in elastic properties necessary for shape-preferred orientation is not clear. (We note that the origin of anisotropy in the Central Pacific, probably related to plume upwell-

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ing, is not well constrained by the present work. It is possible that anisotropy in these regions is caused by shape-preferred orientation involving aligned melt pockets.)

Our results show that slabs are characterized by high stress, resulting in deformation dominated by dislocation creep within a mantle otherwise dominated by diffusion creep. We find complicated strain fields associated with deformation due to dislocation creep, but one consistent feature appears to be a large degree of laterally directed strain directly above the CMB. When examined in the context of mineral-physics experiments, we predict that this strain field results in significant seismic anisotropy, with $V_{\rm SH} > V_{\rm SV}$. This work shows that LPO of (Mg, Fe)O is a likely candidate for the seismic anisotropy observed near the CMB in slab regions. Although other processes may contribute to the formation of anisotropy⁵, they are not required, and solid-state processes within a homogeneous material may suffice.

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A ceratopsian dinosaur from China and the early evolution of Ceratopsia

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Ceratopsians (horned dinosaurs) represent one of the last and the most diverse radiations of non-avian dinosaurs¹⁻⁴. Although recent systematic work unanimously supports a basal division of Ceratopsia into parrot-like psittacosaurids and frilled neoceratopsians, the early evolution of the group remains poorly understood, mainly owing to its incomplete early fossil record. Here we describe a primitive ceratopsian from China. Cladistic analysis posits this new species as the most basal neoceratopsian. This new taxon demonstrates that some neoceratopsian characters evolved in a more incremental fashion than previously known and also implies mosaic evolution of characters early in ceratopsian history.

The lacustrine lower part of the Yixian Formation of western Liaoning, China, has produced many spectacular fossil remains, including feathered dinosaurs⁵. Recently a number of important three-dimensionally preserved vertebrate fossils, including specimens of a new ceratopsian, have been collected from the fluvial facies that form the lowest part of the Yixian Formation⁶. Some radiometric analyses suggest these deposits are older than 128 Myr and younger than 139 Myr (ref. 5) (Early Cretaceous period) but others imply they are older than 145 Myr (ref. 7) (Late Jurassic period).

Ceratopsia Marsh, 1890 Neoceratopsia Sereno, 1986

Liaoceratops yanzigouensis gen. et sp. nov.

Etymology. The generic name is derived from the provincial name 'Liaoning', and the suffix 'ceratops', commonly used for horned dinosaur names. The specific name 'yanzigou' refers to the village near which the holotype was found.

Diagnosis. Neoceratopsian characterized by sutures between premaxilla, maxilla, nasal and prefrontal intersecting at a common point high on the side of the snout, possession of several tubercles on the ventra margin of the angular, a foramen on the posterior face of the quadrate near the articulation with the quadratojugal, a small tubercle on the dorsal border of the foramen magnum, and a thick posterior border of the parietal frill.

Holotype. IVPP (Institute of Vertebrate Paleontology and Paleoanthropology, Beijing) V12738, an almost complete skull (Fig. 1).

Referred specimen. IVPP V12633, a juvenile skull (Fig. 2).

Locality and horizon. Yanzigou and Lujiatun, Shangyuan, western Liaoning, China; the lowest part of the Yixian Formation, probably Early Cretaceous.

Description. The holotype skull of *Liaoceratops* is comparable in size to other basal ceratopsians such as Psittacosaurus and Chaoyangsaurus. Progressive, but incomplete, sutural closures between skull elements suggest that it derives from a subadult individual. The rostral appears unkeeled as in Psittacosaurus. As in many neoceratopsians, the rostral bears lateral processes along the buccal margin. The premaxilla forms much of the side of the snout and extends posterodorsally to reach the prefrontal, as in psittacosaurids and basal ornithopods. Three cylindrical premaxillary teeth are present as in the primitive neoceratopsian Archaeoceratops8. The snout is relatively wide in dorsal view, and tapers abruptly, being intermediate in form between that of psitta-