- Plambeck, R. L., Wright, M. C. H. & Carlstrom, J. E. Velocity structure of the Orion-IRc2 SiO maser: evidence for an 80 AU diameter circumstellar disk. Astrophys. J. 348, L65–L68 (1990).
- Wright, M. C. H., Plambeck, R. L., Mundy, L. G. & Looney, L. W. SiO emission from a 1,000 AU disk in Orion KL. Astrophys. J. 455, L185–L188 (1995).
- Barvainis, R. The polarization of the SiO masers in Orion: maser emission from a rotating, expanding disk? Astrophys. J. 279, 358–362 (1984).
- Thronson, H. A. Jr et al. The Orion star-forming region: far-infrared and radio molecular observations. Astron. J. 91, 1350–1356 (1986).
- Wright, M. C. H., Plambeck, R. L. & Wilner, D. J. A multiline aperture synthesis study of Orion-KL. Astrophys. J. 469, 216–237 (1996).
- Genzel, R., Reid, M. J., Moran, J. M. & Downes, D. Proper motions and distances of H₂O maser sources. I. The outflow in Orion-KL. Astrophys. J. 244, 884–902 (1981).
- Allen, D. A. & Burton, M. G. Explosive ejection of matter associated with star formation in the Orion nebula. *Nature* 363, 54–56 (1993).
- Jones, B. F. & Walker, M. F. Proper motions of Herbig-Haro objects. VI. The M42 HH objects. Astron. J. 90, 1320–1323 (1985).
- Plambeck, R. L., Wright, M. C. H., Mundy, L. G. & Looney, L. W. Subarcsecond-resolution 86 GHz continuum maps of Orion KL. Astrophys. J. 455, L189–L192 (1995).
- Morita, K.-I., Hasegawa, T., Ukita, N., Okumura, S. K. & Ishiguro, M. Accurate positions of SiO masers in active star forming regions: Orion-KL, W51-IRS2, and Sgr-B2 MD5. *Publ. Astron. Soc. Jpn* 44, 373–380 (1992).
- Torrelles, J. M. et al. Radio continuum-H₂O maser systems in NGC 2071: H₂O masers tracing a jet (IRS 1) and a rotating proto-planetary disk of radius 20 AU (IRS 3). Astrophys. J. 505, 756–765 (1998).
 Zensus, J. A., Diamond, P. J. & Napier, P. J. (eds) Very Long Baseline Interferometry and the VLBA
- Lensus, J. A., Diantoni, F. J. & Napler, F. J. (eds) Very Long Baseline Interferometry and the VLDA (Astron. Soc. Pacific, San Francisco, 1995).
 Gaume, R. A., Wilson, T. L., Vrba, F. J., Johnston, K. J. & Schmid-Burgk, J. Water masers in Orion.
- Astrophys. J. 493, 940–949 (1998).
- Sugai, H. et al. Velocity field of the Orion-KL region in molecular hydrogen emission. Astrophys. J. 442, 674–678 (1995).
- Migenes, V., Johnston, K. J., Pauls, T. A. & Wilson, T. L. The distribution and kinematics of ammonia in the Orion-KL nebula: high-sensitivity VLA maps of the NH₃(3,2) line. *Astrophys. J.* 347, 294–301 (1989).
- Greenhill, L. J. et al. in The Impact of VLBI on Astrophysics and Geophysics (eds Reid, M. J. & Moran, J. M.) 253–254 (Kluwer, Dordrecht, 1988).
- Hu, X. Kinematic studies of Herbig-Haro objects in the Orion nebula. *Astron. J.* 112, 2712–2717 (1996).
 Chandler, C. J. & De Pree, C. G. Vibrational ground-state SiO J = 1–0 emission in Orion IRc2 imaged
- with the VLA. Astrophys. J. **455**, L67–L71 (1995).
- 25. Lucas, R. *et al.* Interferometric observations of SiO $\nu = 0$ thermal emission from evolved stars. *Astron. Astrophys.* **262**, 491–500 (1992).
- Danchi, W. C., Bester, M., Degiacomi, C. G., Greenhill, L. J. & Townes, C. H. Characteristics of dust shells around 13 late-type stars. *Astron. J.* 107, 1469–1513 (1994).
- Dougados, C., Lena, P., Ridgway, S. T., Christou, J. C. & Probst, R. G. Near-infrared imaging of the Becklin-Neugebauer-IRc2 region in Orion with subarcsecond resolution. *Astrophys. J.* 406, 112–121 (1993).
- Schleuning, D. A. Far-infrared and submillimeter polarization of OMC-1: evidence for magnetically regulated star formation. Astrophys. J. 493, 811–825 (1998).
- Rao, R., Crutcher, R. M., Plambeck, R. L. & Wright, M. C. H. High-resolution millimeter-wave mapping of linearly polarized dust emission: magnetic field structure in Orion. *Astrophys. J.* 502, L75– L78 (1998).
- Shu, F. H., Najita, J., Ostriker, E. C. & Shang, H. Magnetocentrifugally driven flows from young stars and disks. V. Asymptotic collimation into jets. Astrophys. J. 455, L155–L158 (1995).
- Wardle, M. & Königl, A. The structure of protostellar accretion disks and the origin of bipolar flows. Astrophys. J. 410, 218–238 (1993).
- Garay, G., Moran, J. M. & Haschick, A. D. The Orion-KL super water maser. Astrophys. J. 338, 244– 261 (1989).

Acknowledgements. We thank M. Reid for assistance in recovering his VLA data, and C. Dougados and K. Menten for permission to use the infrared and radio images presented in Fig. 2. The National Radio Astronomy Observatory is a facility of the US NSF operated under cooperative agreement by Associated Universities, Inc. C.S. was supported by the NSF REU program. C.R.G. was supported by the NSF Stellar Astronomy and Astrophysics program.

Correspondence and requests for materials should be addressed to L.J.G. (e-mail: greenhill@cfa.harvard.edu)

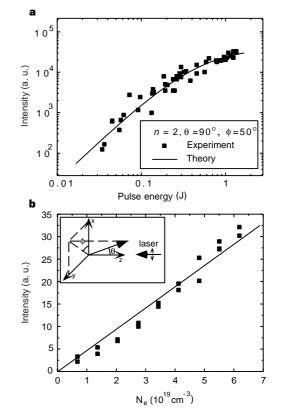
Experimental observation of relativistic nonlinear Thomson scattering

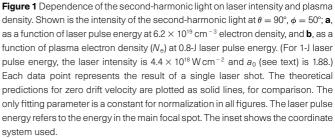
Szu-yuan Chen, Anatoly Maksimchuk & Donald Umstadter

Center for Ultrafast Optical Science, University of Michigan, Ann Arbor, Michigan 48109, USA

Classical Thomson scattering¹—the scattering of low-intensity light by electrons—is a linear process, in that it does not change the frequency of the radiation; moreover, the magnetic-field component of light is not involved. But if the light intensity is extremely high ($\sim 10^{18}$ W cm⁻²), the electrons oscillate during the scattering process with velocities approaching the speed of light. In this relativistic regime, the effect of the magnetic and electric fields on the electron motion should become comparable, and the effective electron mass will increase. Consequently, electrons in such high fields are predicted to quiver nonlinearly, moving in figure-of-eight patterns rather than in straight lines. Scattered photons should therefore be radiated at harmonics of the frequency of the incident light^{2–8}, with each harmonic having its own unique angular distribution^{4–6}. Ultrahigh-peak-power lasers⁹ offer a means of creating the huge photon densities required to study relativistic, or 'nonlinear' (ref. 6), Thomson scattering. Here we use such an approach to obtain direct experimental confirmation of the theoretical predictions of relativistic Thomson scattering. In the future, it may be possible to achieve coherent^{10,11} generation of the harmonics, a process that could be potentially utilized for 'table-top' X-ray sources.

We used a laser system that produces 400-fs-duration laser pulses at 1.053- μ m wavelength with a maximum peak power of 4×10^{12} W (4 TW). The laser beam, 50 mm in diameter, was focused with an *f*/3.3 parabolic mirror onto the front edge of a supersonic helium gas jet. The focal spot consisted of a gaussian spot with full-width at half-maximum (FWHM) of 7 μ m (containing 60% of the total energy) and a large (>100 μ m) dim spot. The helium gas was fully ionized by the foot of the laser pulse. A halfwave plate was used to rotate the axis of linear polarization of the laser beam in order to vary the azimuthal angle (ϕ) of observation. We define $\theta = 0^{\circ}$ as along the direction opposite to that of the laser propagation, and $\phi = 0^{\circ}$ as along the axis of linear polarization. In a linearly polarized laser field, electrons move in a figure-eight trajectory lying in the plane defined by the axis of linear polarization and the direction of beam propagation.





letters to nature

letters to nature

Although the observation of harmonics in laser–plasma (or laser–electron-beam) interactions has been made by several groups^{12–16}, that alone is insufficient to identify unambiguously nonlinear Thomson scattering and its underlying dynamics. Several other mechanisms might generate continuum or harmonics under our experimental conditions, and, therefore, need to be isolated and discriminated from the signal generated by nonlinear Thomson scattering: (1) continuum generated from self-phase-modulation of the laser beam in the gas, (2) harmonics generated from atomic nonlinear susceptibility of the gas or, especially, from the ionization process¹⁷, (3) continuum generated either from (relativistic) self-phase-modulation of laser pulses in the plasma, or from electron–electron bremsstrahlung and electron–ion bremsstrahlung, and (4) harmonics generated from the interaction of laser pulses with a transverse electron-density gradient¹⁴.

The main focal spot of the laser pulse undergoes relativisticponderomotive self-channelling when high laser power and high gas density are used¹⁸. (The ponderomotive force is generated by the gradient of the optical pressure of the laser beam). Side-imaging $(\theta = 90^\circ)$ of the first-harmonic light (at the laser frequency) from nonlinear Thomson scattering shows that the laser channel has a diameter of $<10 \,\mu$ m FWHM. However, interferograms¹⁸ show that the diameter of the plasma column is about $100-200 \,\mu$ m, which is created by the wings with intensities $>10^{15} \,\text{W cm}^{-2}$ (the ionization threshold). Therefore, the light generated from laser–gas interactions should be observed to originate from the entire region of plasma, rather than from the narrow laser channel. Results of side-imaging ($\theta = 90^\circ$ and arbitrary ϕ) of the second and third harmonics using a matching interference filter (10-nm bandwidth) show that the signal is emitted only from the narrower laser channel.

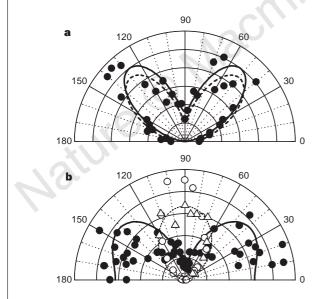
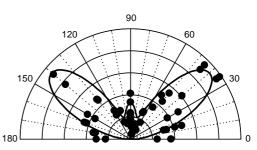


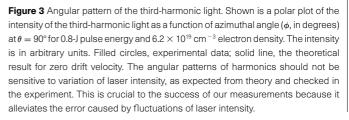
Figure 2 Angular pattern of the second-harmonic light. Shown are polar plots of the intensity of the second-harmonic light as a function of azimuthal angle (ϕ , in degrees) for 0.8-J pulse energy and 6.2×10^{19} cm⁻³ electron density at $\theta = 90^{\circ}$ (**a**) and $\theta = 51^{\circ}$ (**b**). The intensity is in arbitrary units. Filled circles, experimental data; solid and dashed lines, theoretical results for zero and nonzero drift velocity (v = 0.2c) in the laser propagation direction, respectively. Open circles and triangles, experimental data for the first-harmonic signal taken at two different runs under the same conditions; dotted line, the theoretical result. Its dipole radiation pattern (peaked at $\phi = 90^{\circ}$) confirms that there is no depolarization effect in the plasma and that the collective effect of plasmas on the angular pattern is not significant (at least outside a narrow cone along the axis of laser propagation). Although the data for $\phi = 180^{\circ}$ to 360° are not plotted, they should be a mirror image of the data for $\phi = 0^{\circ}$ to 180° because of the intrinsic symmetry of the laser field, as predicted theoretically. Such symmetry has been verified in the experiment.

In addition, the images of the harmonics have spatial distributions similar to the images of the first-harmonic light, and their profiles vary in the same way as the laser power and gas density are changed. This rules out the possibility that the harmonic signal observed in the side images is a result of laser–gas interaction (mechanisms (1) and (2) above).

According to theory⁴⁻⁶, the harmonic signal generated from nonlinear Thomson scattering should have two important features. First, it is linearly proportional to the electron density because it is an incoherent single-electron process (the harmonics generated from a collection of electrons interfere with each other destructively, leaving only an incoherent signal, which is equal to the singleelectron results multiplied by the total number of electrons which radiate). Second, it increases roughly as I^n , where *n* is the harmonic number, and gradually saturates when a_0 is of the order of unity⁶. Here $a_0 = eE/m_0\omega_0 c = 8.5 \times 10^{-10} \lambda I^{1/2}$ is the normalized vector potential, where *E* is the amplitude of laser electric field, $I = cE^2/8\pi$ (in W cm⁻²) is the laser intensity, λ is the laser wavelength (in μ m), e is the electron charge, m_0 is the electron mass, ω_0 is the laser frequency, and *c* is the speed of light. These are characteristically different from the behaviour of any other mechanisms. For instance, bremsstrahlung radiation should be proportional to the square of gas density $(N_e N_e \text{ or } N_e N_i)$ where N_e is the electron density and N_i is the ion density. In the experiment reported here, the intensity of the harmonic signal was determined from the peak intensity or the average intensity of the images of harmonics, when it was plotted as a function of the observing angle, gas density and laser power. Both showed the same variations. Figure 1 shows the variation of the second-harmonic signal as a function of laser power and plasma (electron) density. The experimental results show a reasonable match to the theoretical predictions. The first and third harmonics show the same match with the theory.

Although the above two observations are consistent with nonlinear Thomson scattering as the source of the harmonic signal, observations of the unique angular patterns are necessary to prove that the detailed dynamics of nonlinear Thomson scattering are indeed the same as the theoretical prediction. Figure 2a shows the ϕ dependence of the second-harmonic signal at $\theta = 90^\circ$. The experimental results match qualitatively the theoretical prediction, both having a quadrupole-type radiation pattern which is characteristically different from the dipole pattern for other mechanisms (1)– (4), and linear Thomson scattering. Other measurements, such as the ϕ -dependence of the second-harmonic light at $\theta = 51^\circ$ (an 'anti-dipole' pattern; Fig. 2b), and the ϕ -dependence of the thirdharmonic light at $\theta = 90^\circ$ (a 'butterfly' pattern; Fig. 3), were also made, all showing reasonable matches between the experimental





data and the theoretical predictions. Such angular radiation patterns directly prove the electrons do indeed oscillate with a figureof-eight trajectory in an intense (relativistic) laser field. The angular pattern of the first-harmonic light (linear component) of nonlinear Thomson scattering is also included in Fig. 2b for comparison.

The spectra of the second and the third harmonics each contain a peak at roughly the harmonic wavelength and a red-shifted broader peak, as shown in Fig. 4. The red-shifted broader peaks are believed to be part of the harmonic spectra generated by nonlinear Thomson scattering, because they vary in amplitude proportionally with the corresponding unshifted harmonic signals when the gas density and laser power are changed. It was expected that the spectra of the harmonics should be broadened tremendously for electrons in a high-fluid-velocity plasma wave5. A fast-phase-velocity electron plasma wave (with a maximum fluid velocity of $\sim 0.2c$) excited by stimulated Raman forward scattering¹⁹ was observed in this experiment at the highest laser power and gas density. But the spectral distribution of the harmonics was not observed to change significantly with variation of gas density and laser power, whereas the plasma wave amplitude did change with these variations. This indicates that such spectral structure has nothing to do with the collective drift motion of electrons in the plasma waves.

Although the angular radiation patterns of the harmonics could also be affected by such a 0.2c fluid-velocity oscillation, the changes are not significant enough (compare the solid and dashed lines in Fig. 2a) to be identified from the experimental data⁶. In other words, all measurements done in this experiment match qualitatively with the prediction of incoherent nonlinear Thomson scattering of electrons without drift motion; the results appear not to be affected by the existence of plasma waves, probably due to destructive coherent interference. The absolute scattering efficiency was measured to be 8×10^{-4} and 1×10^{-4} photons per electron per pulse for the second and third harmonics (including both the unshifted and red-shifted spectral components), respectively; these efficiencies were measured at $\theta = 90^{\circ}$ and $\phi = 50^{\circ}$, for an angle of collection of 7×10^{-3} sr. These numbers match reasonably well with the theoretical predictions for incoherent nonlinear Thomson scattering, which are respectively 8×10^{-4} and 5×10^{-4} photons per electron per pulse.

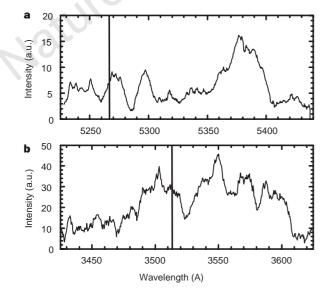


Figure 4 Spectra of harmonics. Shown are spectra of the second (**a**) and third (**b**) harmonics at $\theta = 90^\circ$, $\phi = 50^\circ$ for 0.8-J pulse energy and 6.2×10^{19} cm⁻³ electron density. Vertical lines indicate the wavelengths of the unshifted second and third harmonics. The intensities are plotted in arbitrary units. The spectra do not change with variation of ϕ at any specific θ , so the angular distributions measured are not affected by the choice of filter bandwidth.

Received 28 May; accepted 23 September 1998.

- 1. Thomson, J. J. Conduction of Electricity through Gases (Cambridge Univ. Press, 1906).
- Vachaspati Harmonics in the scattering of light by free electrons. *Phys. Rev.* 128, 664–666 (1962).
 Brown, L. S. & Kibble, T. W. B. Interaction of intense laser beams with electrons. *Phys. Rev.* 133, A705-
- A719 (1964).
 Sarachik, E. S. & Schappert, G. T. Classical theory of the scattering of intense laser radiation by free
- electrons. *Phys. Rev. D* 1, 2738–2753 (1970).
 5. Castillo-Herrera, C. I. & Johnston, T. W. Incoherent harmonic emission from strong electromagnetic waves in plasmas. *IEEE Trans. Plasma Sci.* 21, 125–135 (1993).
- Esarey, E., Ride, S. K. & Sprangle, P. Nonlinear Thomson scattering of intense laser pulses from beams and plasmas. *Phys. Rev. E* 48, 3003–3021 (1993).
- Hartemann, F. V. & Luhmann, N. C. Jr Classical electrodynamical derivation of the radiation damping force. *Phys. Rev. Lett.* 74, 1107–1110 (1995).
- Hartemann, F. V. High-intensity scattering processes of relativistic electrons in vacuum. *Phys. Plasmas* 5, 2037–2047 (1998).
- Maine, P. et al. Generation of ultrahigh peak power pulses by chirped pulse amplification. IEEE J. Quantum Electron. 24, 398–403 (1988).
- Esarey, E. et al. Nonlinear analysis of relativistic harmonic generation by intense lasers in plasmas. IEEE Trans. Plasma Sci. 21, 95–104 (1993).
- Esarey, E. & Sprangle, P. Generation of stimulated backscattered harmonic generation from intenselaser interactions with beams and plasmas. *Phys. Rev. A* 45, 5872–5882 (1992).
- 12. Meyer, J. & Zhu, Y. Second harmonic emission from an underdense laser-produced plasma and filamentation. *Phys. Fluids* **30**, 890–895 (1987).
- Basov, N. G. et al. Investigation of 2ω₀-harmonic generation in a laser plasma. Sov. Phys. JETP 49, 1059–1067 (1979).
- Malka, V. *et al.* Second harmonic generation and its interaction with relativistic plasma waves driven by forward Raman instability in underdense plasmas. *Phys. Plasmas* 4, 1127–1131 (1997).
 Englert, T. J. & Rinehart, E. A. Second-harmonic photons from the interaction of free electrons with
- intense laser radiation. *Phys. Rev. A* 28, 1539–1545 (1983). 16. Bula, C. *et al.* Observation of nonlinear effects in Compton scattering. *Phys. Rev. Lett.* 76, 3116–3119
- (1996). 17. Brunel, F. Harmonic generation due to plasma effects in a gas undergoing multiphoton ionization in
- the high-intensity limit. J. Opt. Soc. Am. B 7, 521–526 (1990).
 Chen, S.-Y. et al. Evolution of a plasma waveguide created during relativistic-ponderomotive selfchanneling of an intense laser pulse. Phys. Rev. Lett. 80, 2610–2613 (1998).
- Le Blanc, S. P. et al. Temporal characterization of a self-modulated laser wakefield. *Phys. Rev. Lett.* 77, 5381–5384 (1996).

Acknowledgements. We thank G. Mourou, R. Wagner and X.-F. Wang for discussions. The work of S.C. and A.M. was supported by US NSF and that of D.U. by the Office of Energy Research, US DOE.

Correspondence and requests for materials should be addressed to D.U. (e-mail: dpu@umich.edu).

Electrostatically driven charge-ordering in Fe₂OBO₃

J. P. Attfield*†, A. M. T. Bell*, L. M. Rodriguez-Martinez*†, J. M. Greneche‡, R. J. Cernik§, J. F. Clarkel & D. A. Perkins

* Department of Chemistry, University of Cambridge, Lensfield Road, Cambridge CB2 1EW, UK † IRC in Superconductivity, University of Cambridge, Madingley Road, Cambridge CB3 0HE, UK

‡ Laboratoire de Physique de l'Etat Condensé, UPRESA CNRS 6087, Université du Maine, 72085 Le Mans, France

\$ Synchrotron Radiation Source, CLRC Daresbury Laboratory,

Warrington WA4 4AD, UK

Chemical Crystallography Laboratory, University of Oxford, Parks Road, Oxford, UK

Charge-ordering is an important phenomenon in conducting metal oxides: it leads to metal-insulator transitions¹ in manganite perovskites (which show 'colossal' magnetoresistances), and the Verwey² transition in magnetite (in which the material becomes insulating at low temperatures when the conduction electrons freeze into a regular array). Charge-ordered 'stripes' are found in some manganites^{3,4} and copper oxide superconductors⁵; in the latter case, dynamic fluctuations of the stripes have been proposed⁶ as a mechanism of high-temperature superconductivity. But an important unresolved issue is whether the chargeordering in oxides is driven by electrostatic repulsions between the charges (Wigner crystallization⁷), or by the strains arising from electron-lattice interactions (such as Jahn-Teller distortions) involving different localized electronic states. Here we report measurements on iron oxoborate, Fe₂OBO₃, that support the electrostatic repulsion charge-ordering mechanism: the system adopts a charge-ordered state below 317 K, in which Fe²⁺