THE UNIVERSITY OF MICHIGAN

COLLEGE OF ENGINEERING Department of Electrical Engineering Space Physics Research Laboratory

SOUNDING ROCKET FLIGHT REPORT

NASA 6.11 Thermosphere Probe Experiment

Prepared on behalf of the project by

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ORA Project 07065

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PROJECT PERSONNEL

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${\tt W}_{\circ}$	A.	Brence	Wallops Island Project	Wallops Island
			Engineer	
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\mathbb{D}_{\circ}	T.	Pelz	Calibration Physicist	GSFC
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I. INTRODUCTION

This report describes and discusses the results of the launching of NASA 6.11, an Aerobee 300 sounding rocket. The payload was the Thermosphere Probe (TP), described by Spencer, Brace, Carignan, Taeusch and Niemann (1965). The TP is an instrumented ejectable package developed by this laboratory in cooperation with the Goddard Space Flight Center, Laboratory for Atmospheric and Biological Sciences (GSFC) for the purpose of studying the variability of the earth's atmospheric parameters in the altitude region between 120 and 350 km. The NASA 6.11 payload included an omegatron mass analyzer (Niemann and Kennedy, 1966), an electron temperature probe (Spencer, Brace and Carignan, 1962), and a lunar aspect sensor. This complement of instruments permitted the determination of molecular nitrogen density and temperature, and electron density and temperature in the altitude range of approximately 140 to 280 km.

General description of the payload kinematics, orientation analysis, and data reduction techniques are given by Taeusch, Carignan, Niemann, and Nagy (1965). The orientation analysis and nitrogen data reduction were performed at this laboratory and the results are included in this report with a discussion of problem areas and probable errors. The electron temperature probe data were reduced at GSFC and are not discussed in this report.

The payload described in this report was launched 12 hours after a similar one (NASA 18.01) described in a separate report. The purpose of this dual launching was to establish the diurnal variation of the parameters measured to provide extra meaning to their use in studying the effect of the energy input to the atmosphere.

2. GENERAL FLIGHT INFORMATION

The general flight information for NASA 6.11 is tabulated below. The geophysical indices, the 10.7 cm solar radio flux, $F_{10.7}$; the five monthly averages of the solar 10.7 cm flux preceding the launch, $\overline{F}_{10.7}$; and the geomagnetic index, a_p , were obtained from the April, 1965, and May, 1965, issues of "Solar Geophysical Data" published by the U. S. Bureau of Standards. $F_{10.7}$ is given for the day preceding the launch and a_p is given for six hours previous to launch for convenient reference to the Harris and Priester (1964) model atmoshpere.

The Table of Events gives flight times and altitudes of significant events occurring during the flight. Some of these were estimated and are so marked. The others were obtained from the telemetry records and radar trajectory, where applicable.

Launch Date: March 20, 1965

Launch Time: 0042 EST; 05:42 GMT

Location: Wallops Island, Virginia

Longitude: 75.04°W Latitude: 37.54°N

Apogee Parameters:

Altitude: 326.9 km
Horizontal Velocity: 578 m/sec
Flight Time: 311.1 sec

Geophysical Indices:

 $\frac{F_{10.7}}{F_{10.7}} = 77.0$ $a_{p} = 75.3$

TP Motion:

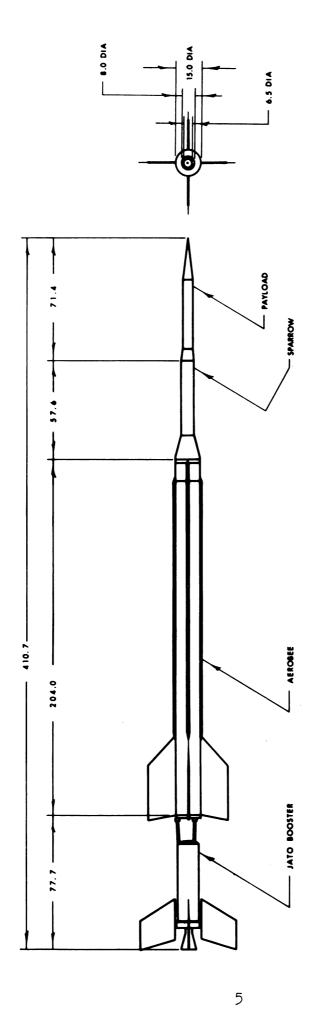
Tumble Period = 2.630 sec/tumble Roll Period = 1.406 sec/roll

TABLE OF EVENTS

Event	Flight Time (sec)	Altitude (km)	Remarks
Lift Off	0	Ó	
Aerobee Burn Out	53.9	34.4	(5,172 ft/sec)
Sparrow Ignition	54.6	35.4	
Sparrow Burn Out	57.6	41.4	(8,484 ft/sec)
TP Ejection	70.2	70.1	
Omegatron Breakoff	100.0(est)	130.0(es	t)
Omegatron Filaments On-to Mass 28	101.6	133.4	
Omegatron to Mass 32	155:00	220.0	
Omegatron to Mass 28	195.20	268.1	
Peak Altitude	311.1	326.9	
L.O.S.	575.0	30.0(est)

3. LAUNCH VEHICLE

The NASA 6.11 launch vehicle was an Aerobee 300, a three-stage JATO bottle-Aerobee-Sparrow combination. The JATO bottle weighed approximately 600 lb unburned, has a thrust of about 18,600 lb and burns for approximately 2.5 seconds. The Aerobee is a liquid propellent rocket weighing approximately 1350 lb unburned, has a thrust of 4,100 lb and burns for about 52 seconds. The Sparrow weighed 129 lb unburned and has a thrust of about 5500 lb for about 3 seconds. The total vehicle, including booster and payload, weighed about 2188 lb at lift off. The dimensions of the total vehicle and its components are shown in Figure 1. Photographs of the JATO bottle, Aerobee-Sparrow, and the Sparrow-payload combinations are shown in Figures 2, 3, and 4.



AEROBEE 300A DIMENSIONS - 6.5 IN. DIAMETER PAYLOAD

Figure 1. Aerobee 300A with thermosphere probe payload.

NASA W-65-262

Figure 2. JATO Bottle-Aerobee combination.

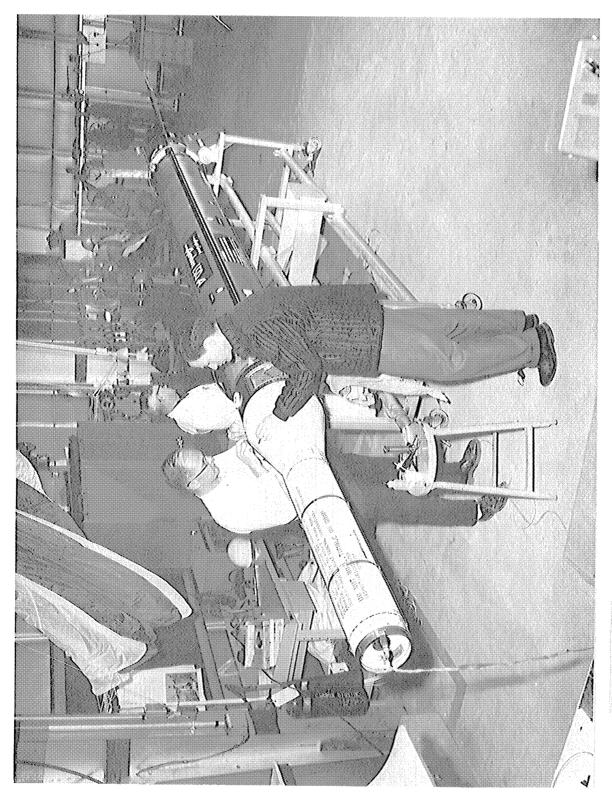
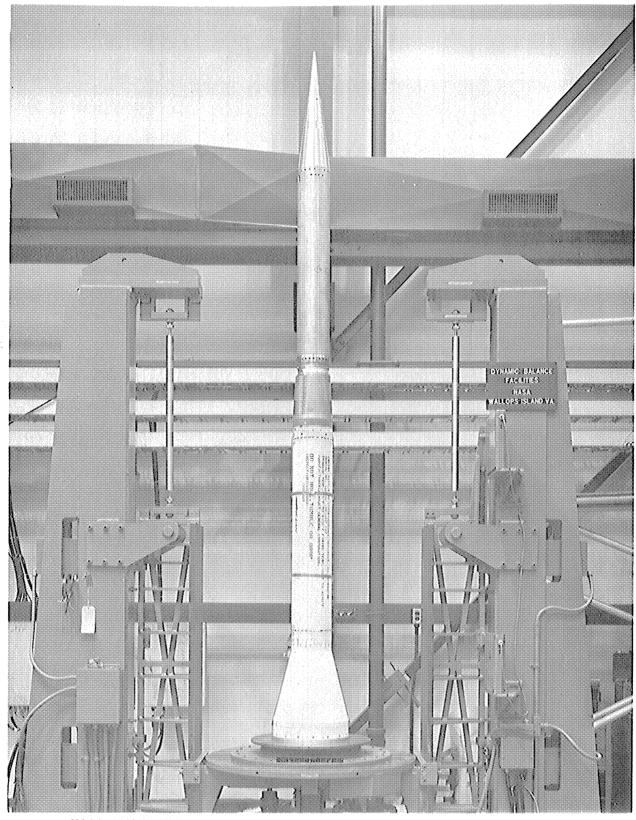


Figure 3. Aerobee-Sparrow combination.

NASA W-65-261



NASA W-65-158

Figure 4. Sparrow-payload on shake table.

4. NOSE CONE

A schematic of the total payload which includes the nose cone and adapter section is shown in Figure 5. Section A contains the batteries, timer, and pyrotechnics for the opening of the nose cone. Section B contains the plunger and the volume which holds the TP (E). Section C houses the ejection spring, plunger piston, and the lanyard negator motor. Dimensions and weights of the system are given on the schematic.

The nose cone is programmed to open at about 70 km altitude and the TP is ejected and tumbled. The breakoff device is removed at about 110 km and the omegatron turned on a few seconds later. The timing for this particular payload was described previously.

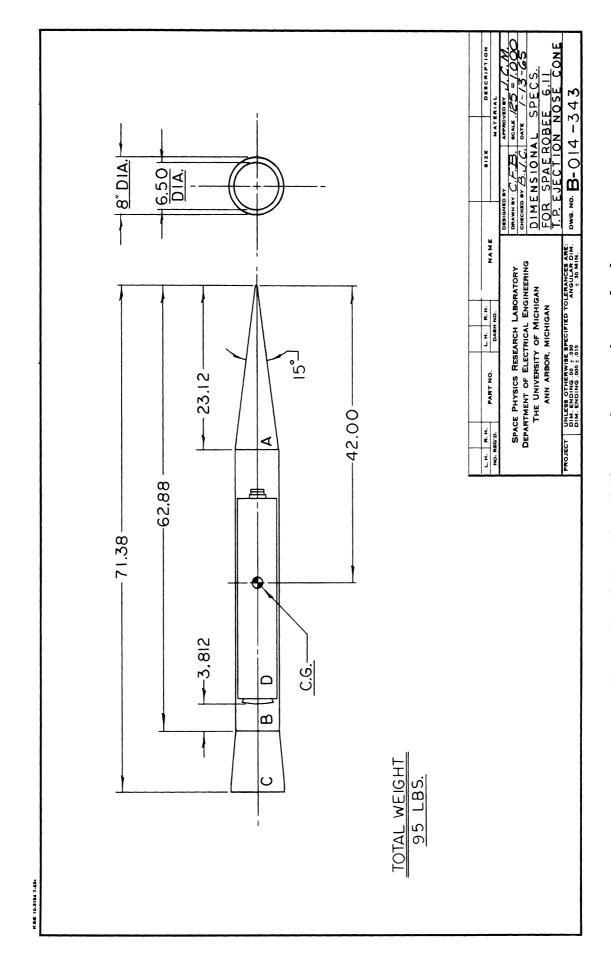


Figure 5. Schematic of thermosphere probe payland.

5. THE THERMOSPHERE PROBE (TP)

The TP used for the NASA 6.11 payload was a cylinder 32-7/64 in. long and 6 in. in diameter weighing 43 lb. The prime instruments for this payload were an omegatron mass analyzer (Niemann and Kennedy, 1966), and an electron temperature probe, (Brace, et al., 1963). Supporting instrumentation included a lunar aspect sensor for the determination of the TP aspect. The diagram shown in Figure 6 shows the instrumentation and supporting electronics location; and Figure 7 shows the block diagram. Figure 8 is a picture of the completely assembled TP.

5.1 OMEGATRON

The omegatron used in this payload was of the type described by Niemann and Kennedy (1966). An expanded view of the system is shown in Figure 9. Table I lists the operating parameters of the gauge and associated electronics. The characteristics of the linear electrometer amplifier current detector, used to monitor the omegatron output current, are also listed.

This omegatron was the second flight model to utilize a ceramic breakoff device which allowed vacuum sealing of the gauge. The breakoff configuration is shown in Figure 10. A new omegatron envelope was designed for this break-off device and is shown in Figure 11. The magnet used for this system is shown in Figure 12.

The calibration of this gauge, and that one used in NASA 18.01, was performed in February, 1965, at GSFC. The system used was an oil diffusion pump calibration system under the supervision of Mr. Carl Reber of GSFC. Other gauges used for reference were (1) a double focusing 180° magnetic deflection spectrometer, (2) a Westinghouse Bayard-Alpert (BA) gauge, (3) two Veeco BA gauges, calibrated by Ball Brothers, and (4) an AW5966 BA gauge. The two Veeco gauges were used as the standard for this calibration. They had been calibrated against a McLeod gauge to a stated absolute certainty of better than \pm 25%. As stated Previously, the NASA 18.01 omegatron was calibrated on this system at the same time; therefore, the relative accuracy between the two omegatrons is believed better than \pm 10%. A final relative calibration was performed at SPRL on February 26, 1965, at which time the NASA 18.01 omegatron was refocused. The NASA 6.11 calibration was used as the standard after the refocusing and a new sensitivity was determined for the NASA 18.01 gauge. The final calibration is shown in Figure 13.

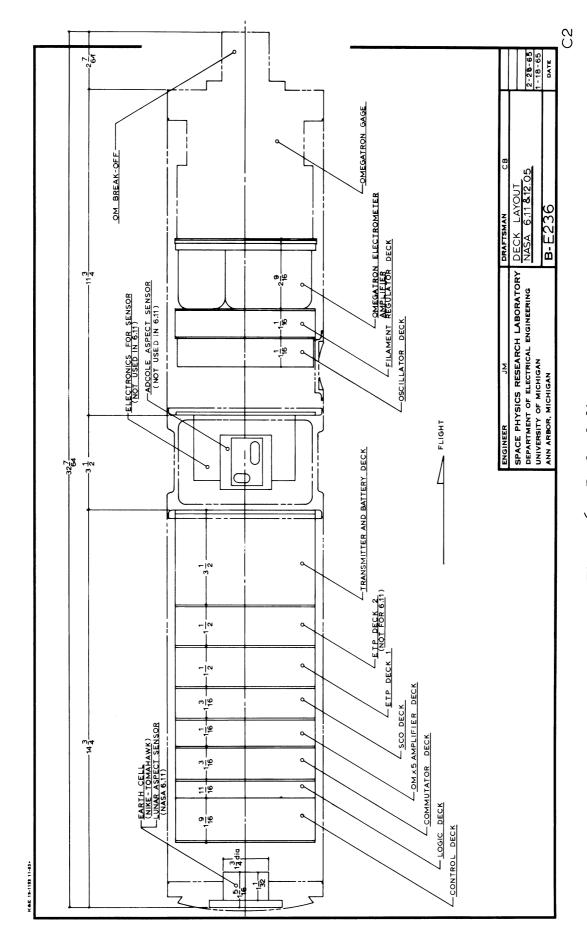


Figure 6. Payload diagram.

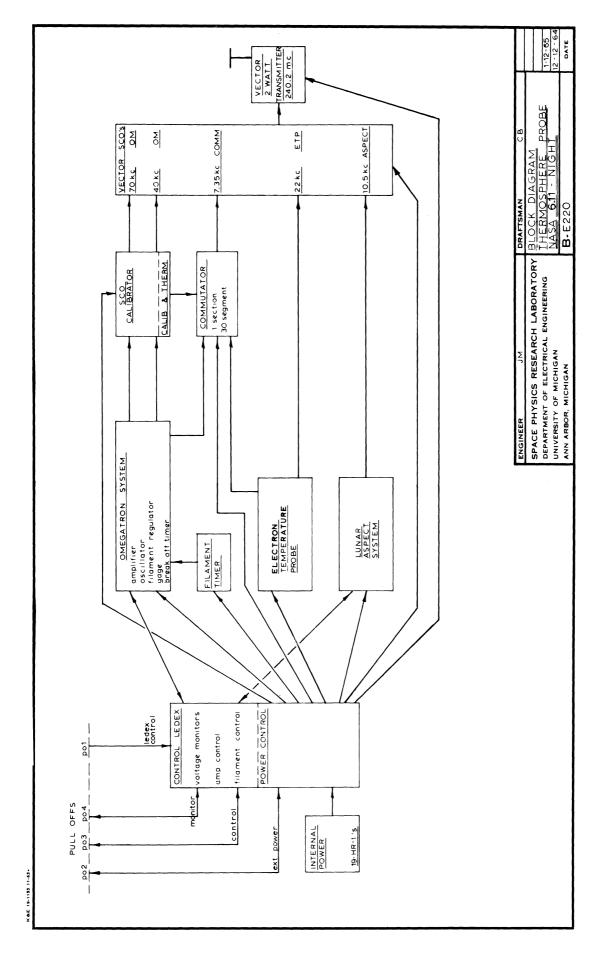


Figure 7. Block diagram.

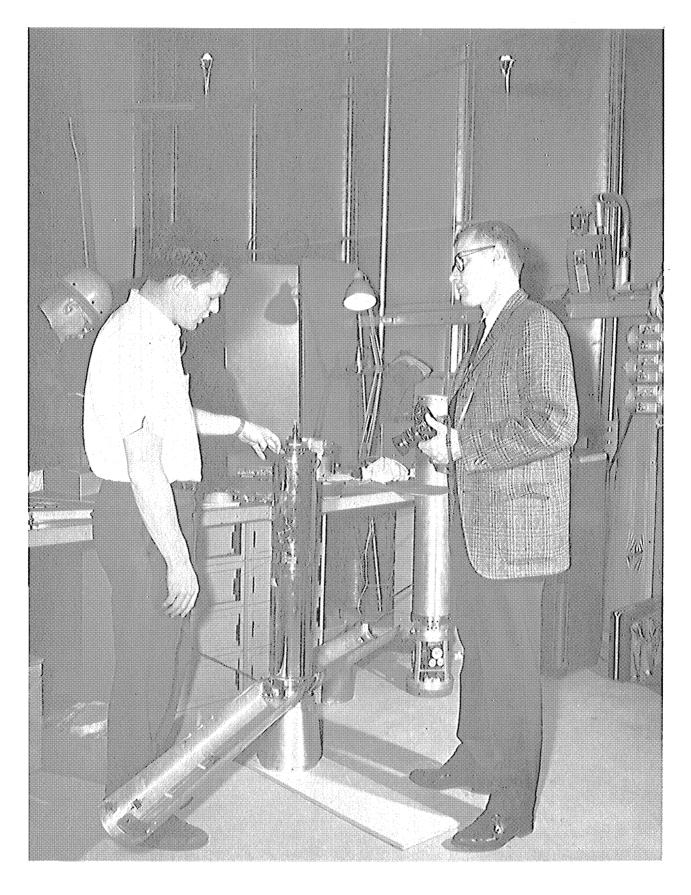


Figure 8. Thermosphere probe in nose cone.

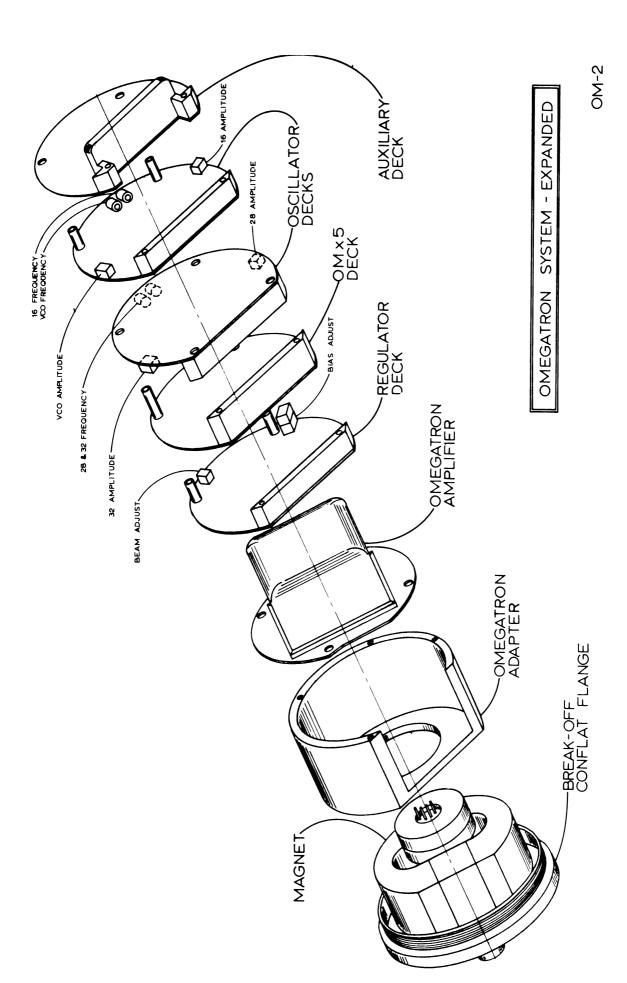


Figure 9. Expanded view of omegatron system.

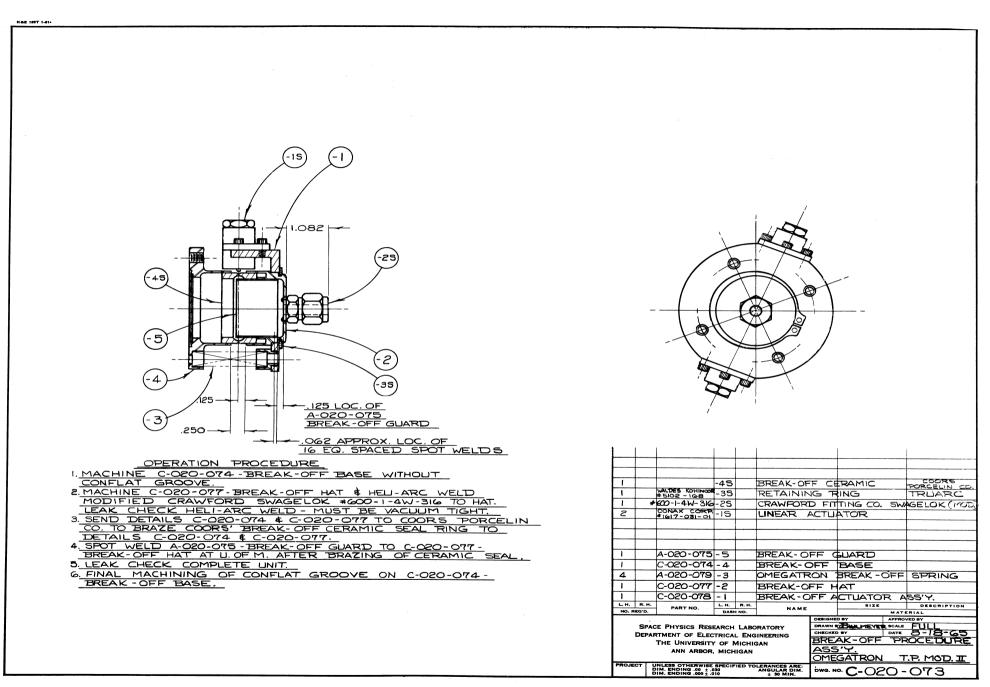


Figure 10. Breakoff configuration.

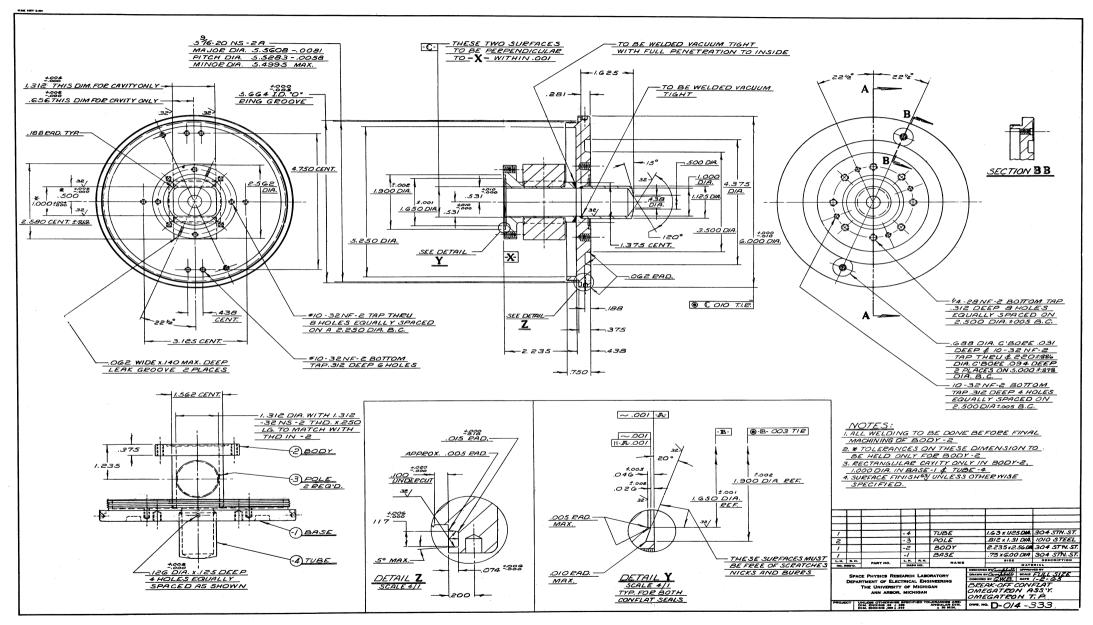


Figure 11. Omegatron envelope.

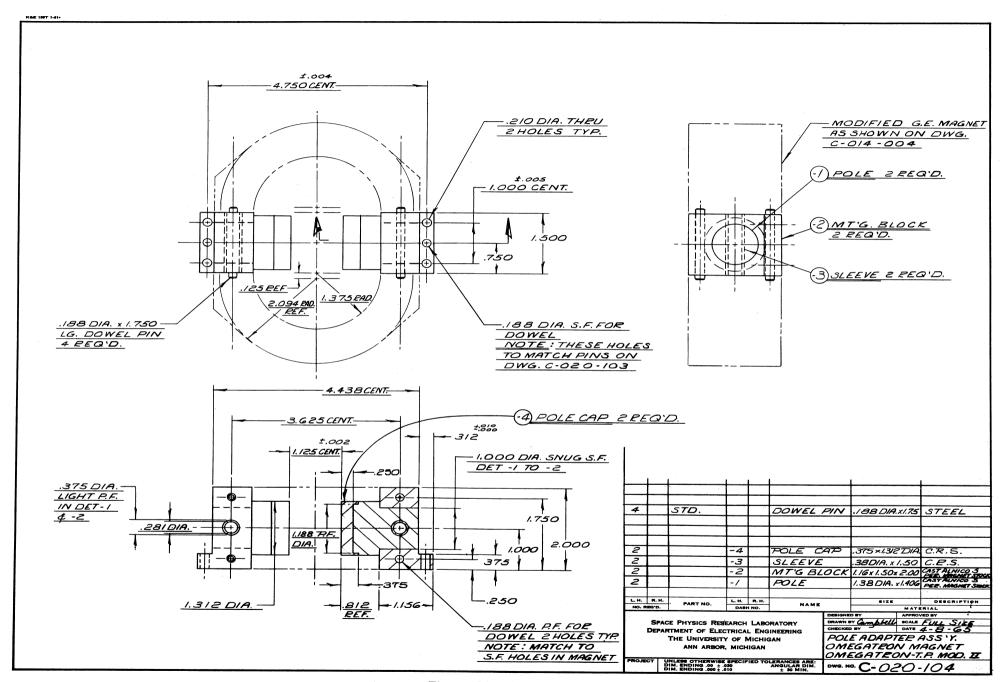


Figure 12. Magnet.

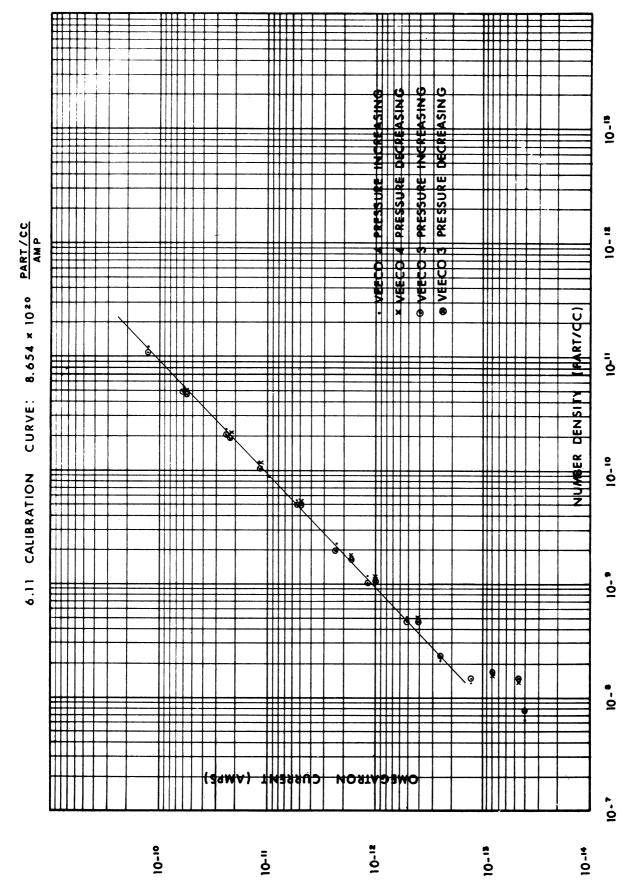


Figure 13. Omegatron calibration.

TABLE I

OMEGATRON DATA NASA 6.11

Omegatron Gauge Parameters

Beam current	3.96 µa
Electron collector bias	+77.0 v
Filament bias	- 91 . 6
Cage bias	-0.202 v
Top bias	-0.462 v
RF amplitude	
M 28	2.62 v P-P
$M_{\rm S2}$	2.57 v P-P

Monitor

Filament	
OFF	2.640 v
ON	4.901 v
Beam	
OFF	4.037 v
ON	2.740 v
Thermistor pressure*	1.86 v
Bias	4.010 v
RF	
M ₂₈	3.600 v
M ₃₂	1.766 v
<i>)</i> -	

Calibration

Sensitivity	3.75	х	10-5	amp/torr
Maximum linear pressure (5%)	~ 6	X	10-6	torr

Electrometer Amplifier

Range	Range Indicator	Range Resistor	ZPV
1	0.0 v	3.16 x 10 ¹⁰	5.01
2	~ 0.7 v	1.0 x 10 ¹¹	5.01
3	~ 1.4 v	3.16 x 10 ¹¹	5.01
4	~ 2.1 v	1.0 x 10 ¹²	5.01
5	~ 2.8 v	3.16 x 10 ¹²	5.01
6.	~ 3.5 v	1.0 x 10 ¹³	5.01
	Calibrate	0.855 v	

Miscellaneous

+ 28 power current all on	270 ma
Preflight gauge pressure	$\sim 10^{-3}$ torr
Magnetic field strength	~ 2200 gauss

^{*}Filament off, zero pressure

5.2 ELECTROSTATIC PROBE (ESP)

The ESP, described by Brace (1963), consists of a cylindrical probe, which is placed in the plasma, and an electronic unit, which measured the collected current. The electronic unit consists of a power convertor, ΔV generator, 3 range current detector, relays, and associated logic circuits.

The following are the specifications of the NASA 6.11 ESP system:

(a) Sensitivity:

Range No. 1 10.0 μ A full scale (4 v) Range No. 2 1.0 μ A full scale (4 v) Range No. 3 .1 μ A full scale (4 v)

(b) Input Power:

1.5 watts average at 28 v.

(c) ΔV Slope (dv/dt):

HI - 38 v/sec.LO - 13 v/sec.

(d) Output:

Voltage 0 ~ 6.9 v Resistance - < 2K

(e) System Bias Level:

l v

(f) System Calibration:

ON - 1.2 sec. Interval - 60 sec. Synchronized with ΔV .

(g) Sequence:

 ΔV - HI-LO alternated Range - Sequential and HI-LO ΔV per range.

5.3 SUPPORT MEASUREMENTS AND INSTRUMENTATION

5.3.1 Lunar Aspect Sensor

The NASA 6.11 TP utilized a lunar aspect sensor system for the determination of the angular momentum vector of the tumbling TP. This system, designed especially for the TP configuration, consists of a light sensitive element, a current detector and an aperture system of concentric rings which resembles a typical target used for firearms. The system viewed a cone of 60° half angle

and the apertures were 5° wide and separated by 5°. The system was mounted such that the axis of the cone of view was along the cylindrical axis of the TP. The output from this system yielded directly the angle between the plane of tumble and the earth-moon vector.

The particulars of the data reduction from this system are described by Taeusch, et al. (1965). The lunar aspect sensor worked as expected during the flight and the aspect of the TP was determined to an accuracy of approximately + 5°.

5.3.2. Telemetry

The payload data were transmitted in real time by a five-channel PAM/FM/FM telemetry system at 240.2 m hz with a nominal output of 2.5 watts. The telemetry system used five subcarrier channels, assigned as outlined below.

Transmitter: Driver TRPT 251 RBO-1 Serial Number 845
Power Amplifier TRFP-2V Serial Number 150
Mixer Amplifier-Type TA58 Serial Number 860

Subcarrier Channels (SCO-type TS54)

IRIG Band	Serial Number	Frequency	Nominal Freguency Response	Function
11	1681	7.35 K hz	110 hz	30 pps PAM Data
12	1697	10.5 K hz	160 hz	Lunar Aspect Data
.14	1712	22 K hz	330 hz	Electrostatic Probe Data
16 18	1736 1 0 85	40 K hz 70 K hz	600 hz 1050 hz	Omegatron Data Omegatron Data

Instrumentation power requirements totaled approximately 30 watts, which was supplied by a Yardney HR-l Silvercell battery pack of nominal 27.8 volts output.

5.3.3 Housekeeping Monitors

Outputs from various nomitors throughout the instrumentation provide information bearing on the operations of the electronics components during flight. These outputs were fed to a thirty-segment commutator which ran at one rps. The commutator assignments are as follows:

- 1. O v calibrate
- 2. l v calibrate
- 3. 2 v calibrate
- 4. 3 v calibrate
- 5. 4 v calibrate
- 6. 5 v calibrate
- 7. 5 v calibrate
- 8. ESP AV monitor
- 9. ESP AV monitor
- 10. Omegatron amplifier range
- 11. Omegatron amplifier output
- 12. Omegatron filament voltage monitor
- 13. Omegatron beam current monitor
- 14. Omegatron bias voltage converter monitor
- 15. Omegatron RF voltage and frequency monitor
- 16. ESP ΔV monitor
- 17. ESP ΔV monitor
- 18. Thermistor filament regulator temperature
- 19. Thermistor omegatron gauge temperature
- 20. Thermistor omegatron amplifier temperature
- 21. Thermistor omegatron internal pressure
- 22. Thermistor transmitter temperature
- 23. 4.5 v calibrate
- 24. 2.0 v calibrate
- 25. Battery voltage monitor (1 v out = 6.1 v battery)
- 26. Ground control position monitor
- 27. ESP ΔV monitor
- 28. ESP ΔV monitor
- 29. Omegatron amplifier range
- 30. Omegatron amplifier output

6. ENGINEERING RESULTS

The Thermosphere Probe instrumentation operated normally throughout the flight. The probe was identical to that described in the report on NASA 18.01 (Taeusch and Carignan, 1966) except that a lunar aspect sensor especially developed for the TP application was used.

The use of the newly developed breakoff device which permits more accurate measurements to higher altitudes also greatly simplifies the field support requirements, an important factor in permitting more wide spread application of the technique.

The launch of 6.11 will probably be the last use of the Aerobee 300A (Spaerobee) for the TP experiment. The new two-stage solid propellant systems, which have essentially the same performance, reduce the cost and operational complexity of carrying out the TP experiment. The Aerobee 300A has been used by the experimenters for a total of thirteen launches of an ejectable experiment package dating back to November 1958 when the Dumbbell experiment was successfully flown from Ft. Churchill on rocket ABM 10.200. None of the thirteen rockets failed to perform properly although one was mistakenly cut-down by range safety before third-stage ignition.

7. DATA ANALYSIS

The telemetered data were recorded on magnetic tapes at two stations, Wallops Island Main Base Telemetry Station and Goddard Space Flight Center Station "A." One set of real time "paper" records for quick look evaluation of the results were also obtained. The omegatron, housekeeping and aspect data were reduced to engineering parameters from paper records, run at 10 in/sec, using a Gerber GDDRS data reader and scanner. The paper records used for data reduction were recorded from the magnetic tape masters.

Tracking data for trajectory determination were obtained from FPS-16 and FPQ-6 radars. Continuous data were obtained from +7 to +264 sec by FPS-16 radar and from +65 to +662 sec by FPQ-6 radar.

7.1 Trajectory

The trajectory and velocity information used to determine the aspect, density, and temperature data as a function of altitude were obtained by fitting a smooth theoretical trajectory to the measured radar data. The theoretical trajectory is programmed for computer solution similar to that described by Parker (1962). The output format is shown in Figure 14. The trajectory is shown in Figure 15. The analysis of minimum angle of attack (α) is described by Taeusch et al. (1965) is also incorporated in the program and the output of the computer furnishes α and $\cos \alpha$ versus time, altitude, etc. A plot of the NASA 6.11 α versus altitude is given in Figure 16.

7.2 Ambient N_2 Density

The neutral molecular nitrogen density was determined from the measured gauge partial pressure as described by Spencer, $\underline{\text{et}}$ $\underline{\text{al}}$. (1965), using the basic relationship:

$$\left(n_{\mathbf{a}} = \frac{\Delta n_{\mathbf{i}} U_{\mathbf{i}}}{2 \sqrt{\pi} V \cos \alpha}\right) N_{2}$$

where

 n_a = Ambient N2 number density

 Δn_1 = Maximum minus minimum gauge number density during one tumble

N A S A 6 . 1 1 T P 7

LAUNCH TIME (GMT)

YEAR 1965 DAY 79 HGUR 5 MINUTE 42 SECOND .000

INITIAL CONDITIONS

PEAK PARAMETERS

38.825 .678 .7791 DEGREES RADIANS COSINE ALPHA ELEVATION DEGREES INERTIAL WRT EARTH 88 AZIMUTH DEGREES INERTIAL WRT EARTH 115.04 131.63 1440 -1234 -48 × > VELOCITY WRT EARTH FIXED VEL 578 578 -0 LONGITUDE HORIZONTAL HORIZCNTAL HORIZONTAL HORIZONTAL HORIZONTAL VERTICAL VERTICAL 1897 1897 -0 908 908 -0 INERTIAL VELOCITY FEET METERS 2978 2978 -0 POSITION 36.96 FEET METERS 509605 155328 RANGE FEET METERS GEOPOT METE 1072359 326855 310903 ALTITUDE 311.10 TIME SEC

Figure 14. Trajectory program output format.

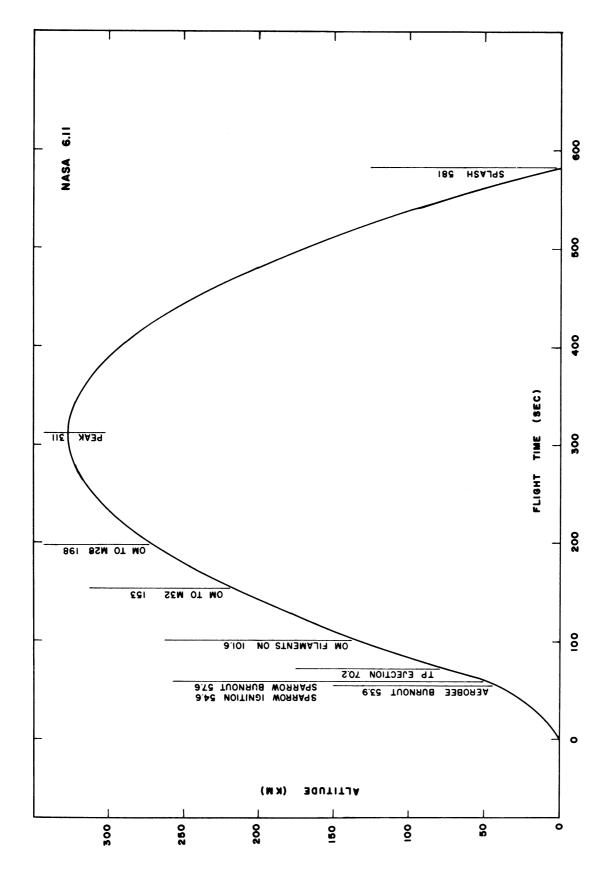


Figure 15. Trajectory with timing.

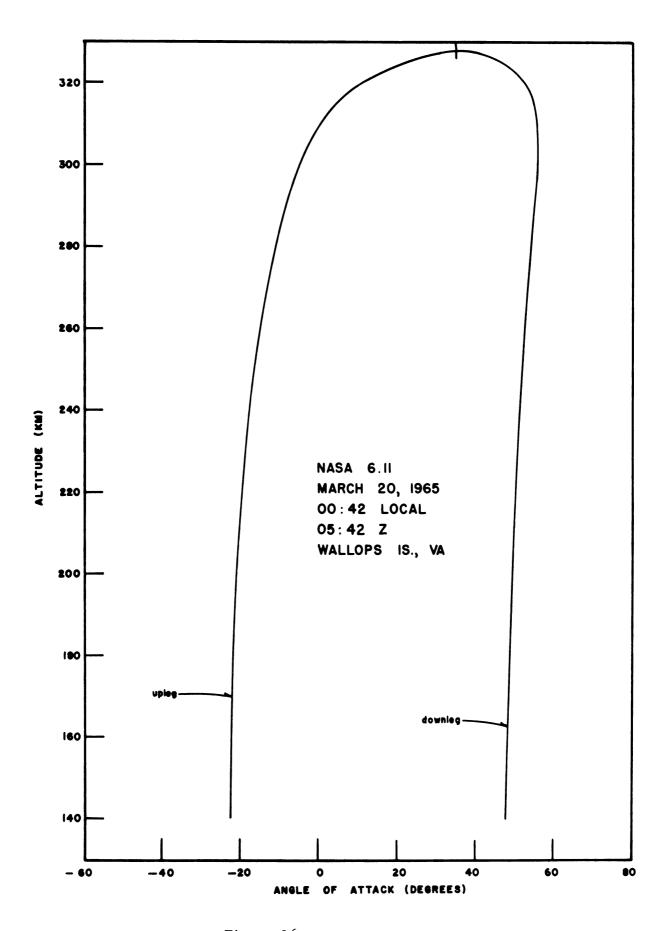


Figure 16. α vs. altitude.

 $U_i = \sqrt{2kTi/m}$ most probable velocity of particle inside gauge

T; = Gauge wall temperature

V = Vehicle velocity with respect to the earth

 α = Minimum angle of attack for one tumble

 $\Delta I_{\rm i}$, the difference between the maximum (peak) omegatron gauge current and the minimum (background) gauge current is shown versus flight time in Figure 17. The background current is also shown in the figure. The background current is the result of the outgassing of the gauge walls and the inside density due to atmospheric particles which have high enough energy to overtake the TP and enter the gauge. The outgassing component is assumed constant for one tumble and effects both the peak reading and the background reading; and, therefore, does not effect the difference. From calibration data obtained by standard techniques, the inside number density, $\Delta n_{\rm i}$, is computed for the measured current. As described by Spencer, Taeusch, and Carignan (1965), the uncertainty in these data is believed to be \pm 10.2% rms relative to other measurements using the same calibration system and \pm 25.1% rms absolute.

 $U_{\rm i}$, the most probable thermal speed of the particles inside the gauge, is computed using the measured gauge wall temperature shown in Figure 18. The uncertainty in this measurement is believed to be + 2.2% rms absolute.

V, the vehicle velocity with respect to the earth; is believed known to better than \pm 1% absolute. It is obtained from the trajectory curve fitting described previously and is the most accurately known quantity obtained from the analysis.

Cos α is obtained from the aspect analysis described by Taeusch, et al. (1965). Since the uncertainty in cos α depends upon α , for any given error in α , each particular case and altitude range must be considered separately. As can be seen in Figure 16, the upleg data were obtained for angles of attack less than 30°, which results in an uncertainty in cos α approximately \pm 3% for an uncertainty in α of approximately \pm 5°.

The resulting ambient N_2 number density, obtained from the measured quantities described above, is shown in Figure 19. The uncertainty in the ambient density due to the combined uncertainties in the measured quantities is \pm 10.9% rms relative and \pm 25.4% rms absolute.

7.3 Temperature

The ambient N_2 temperature profile, shown in Figure 20, was obtained by integrating the density profile to obtain the pressure and then relating the known density and pressure to the temperature through the ideal gas law. The

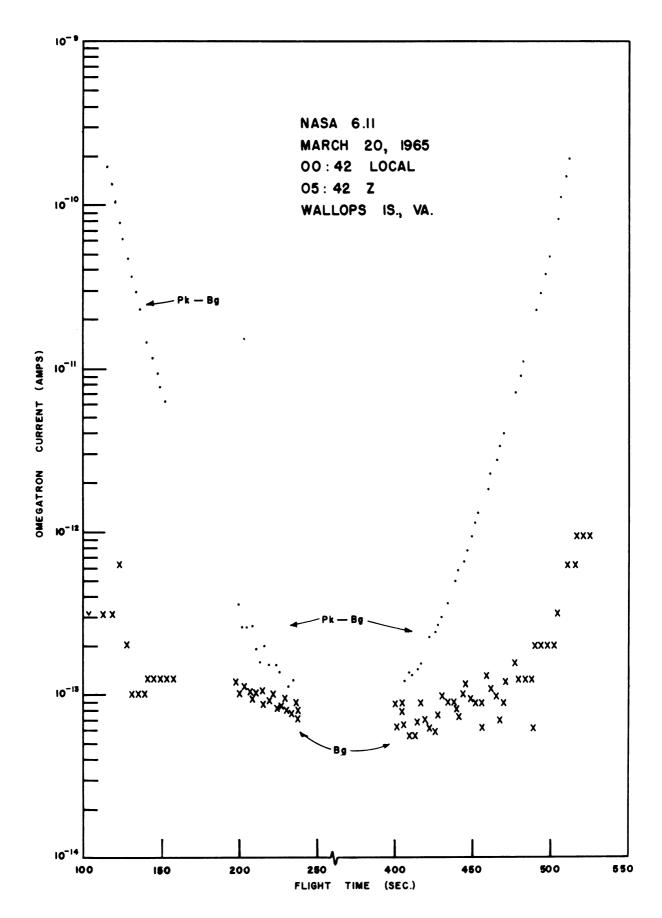


Figure 17. Peak minus background current vs. time.

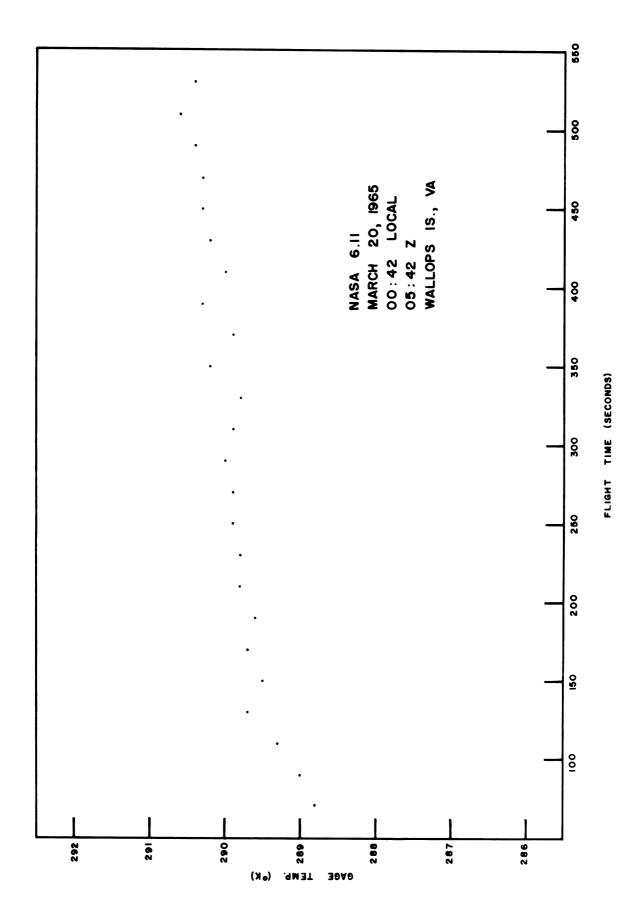


Figure 18. Gauge temperature vs. flight time.

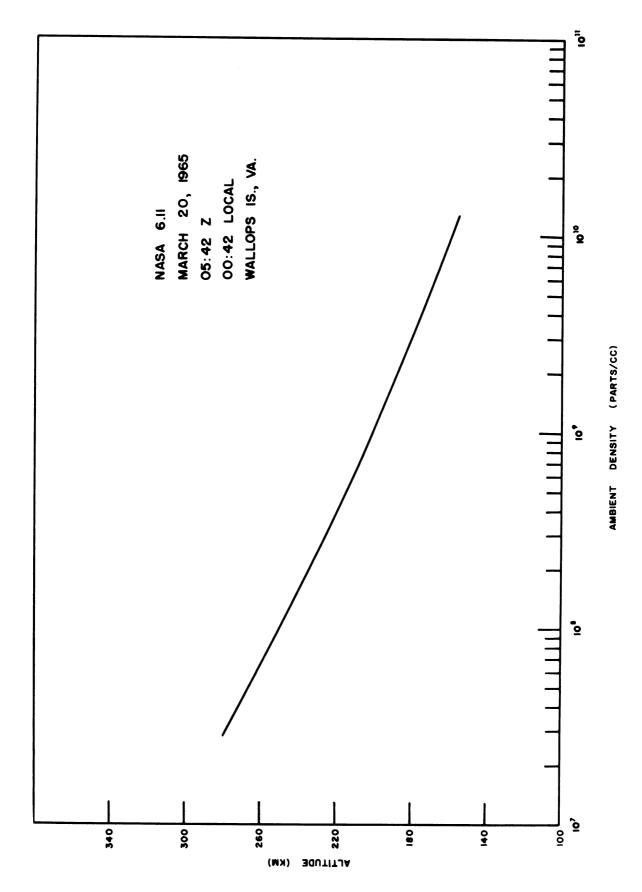


Figure 19. Ambient N_2 number density vs. altitude.

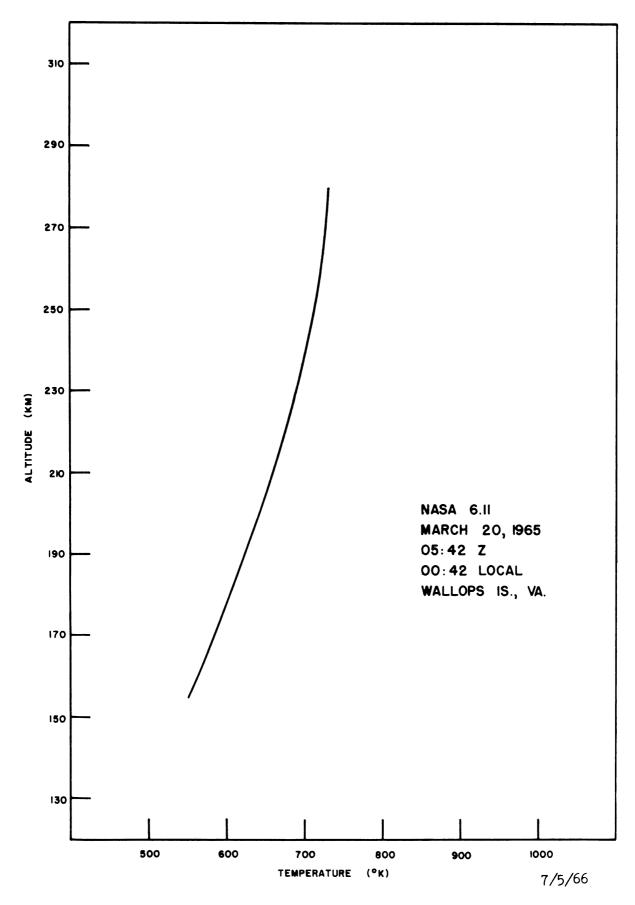


Figure 20. Ambient N_2 temperature vs. altitude.

assumption that the gas is in hydrostatic equilibrium and behaves as an ideal gas is implicit. Since the temperature depends only upon the shape of the density profile and not its magnitude, it is believed that the uncertainty in its magnitude is \pm 5% absolute.

FINAL N₂ DENSITY AND TEMPERATURE (after application of η , geometry correction factor)

NASA 6.11

March 20, 1965

05:42 Z

00:42 LOCAL

Wallops Is., Va.

ALTITUDE	DENSITY (Part/cc)	TEMPERATURE (°Kelvin)
155	1.28 x 10 ¹⁰	551
160	9.58 x 10 ⁹	562
165	7.18	572
170	5•39	583
175	4.67	594
180	3.04	604
185	2.28	614
190	1.76	624
195	1.35	633
200	1.04 x 10 ⁹	642
2 0 5	8.03 x 16 ⁸	651
210	6, 28	660
215	4.90	668
220	3.85	676
225	3.0 4	683
230	2.40	69 0
235	1.91	6 96
240	1.52	7 0 2
245	1.22 x 10 ⁸	7 0 8
250	9.79 x 10 ⁷	713
255	7.89	718
260	6. 38	722
265	5.14	725
270	4.13	728
275	3.34	730
279	2.83×10^{7}	731

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